

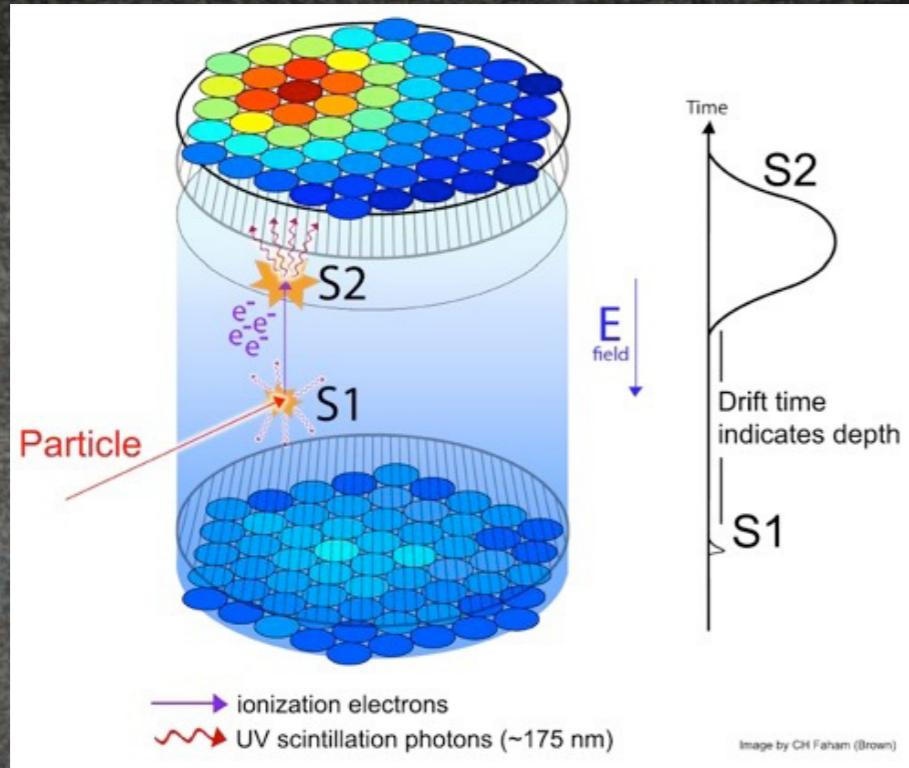
# Some slides on MiniCLEAN

KJ Palladino

6/7/16

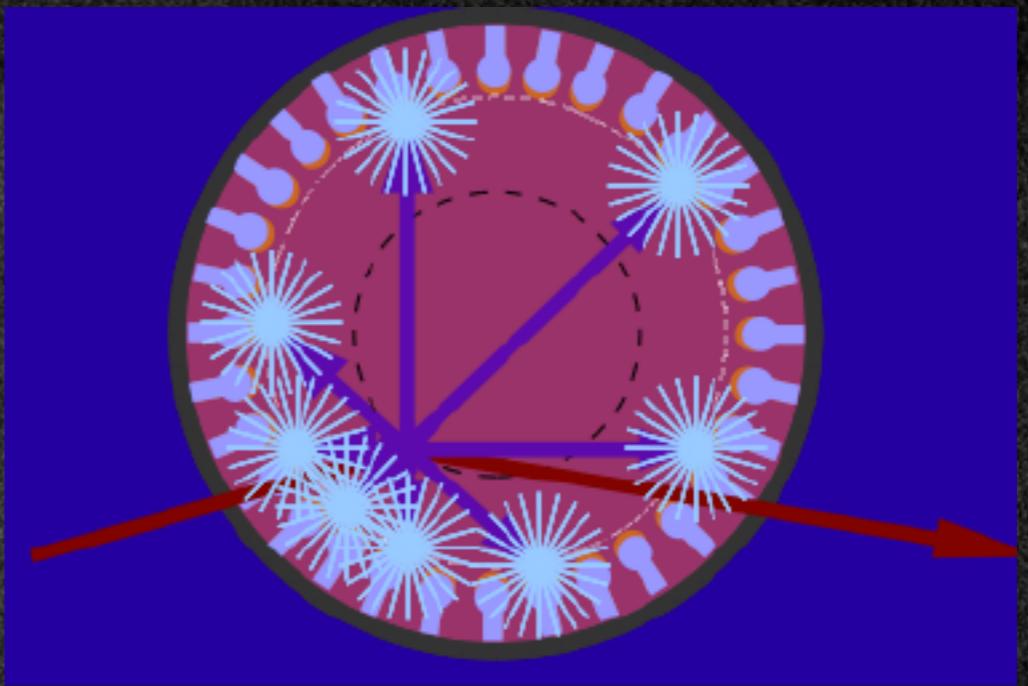
## Double

# Double vs Single Phase



- Ionized and excited states
- Primary Scintillation (S1) with some recombination and de-excitation in the liquid
- Ions drift in TPC electric field
- Amplification region in gas creates proportional light (S2)
- S2/S1 provides particle ID
- Events are hundreds of microseconds (set by electron drift velocity)
- Strong position reconstruction

## Single



- Ionized and excited states
- Scintillation with recombination and de-excitation in the liquid
- Pulse shape discrimination provides particle ID
- Events are tens of microseconds (set by triplet lifetime)
- Position reconstruction with charge distribution, t.o.f. with larger detectors

# MiniCLEAN



**CLEAN**  
Cryogenic  
Low Energy  
Astrophysics with  
Noble liquids

**CLEAN**

**AREA OF RESEARCH:**  
Direct detection of dark matter and low-energy neutrinos

**TECHNOLOGY:**  
Single-phase liquid argon detector

**FUN FACT:**  
CLEAN can operate with interchangeable targets of liquid argon and liquid neon

**INSTITUTIONS:**  
Boston University; Los Alamos National Laboratory; MIT; National Institute of Standards and Technology; Royal Holloway University of London; SNOLAB; Syracuse University; University of New Mexico; University of North Carolina, Chapel Hill; University of Pennsylvania; University of South Dakota; Yale University

**CLEAN**

**TARGET:**  
LAr, LNe

**Mass:**  
500 kg target, 150 kg fiducial

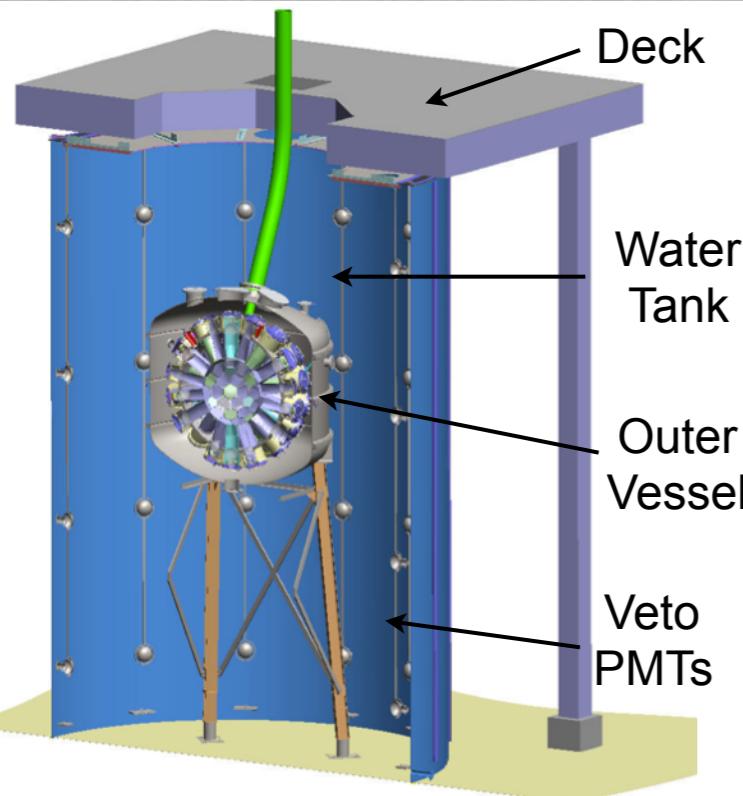
**Light collection:**  
92 8" Hamamatsu R5912-02 MOD PMTs

**Vessel:**  
stainless steel, with modular optical cassettes inserted

**Shielding:**  
10 cm acrylic & 20 cm Ar, in ~8m water shield

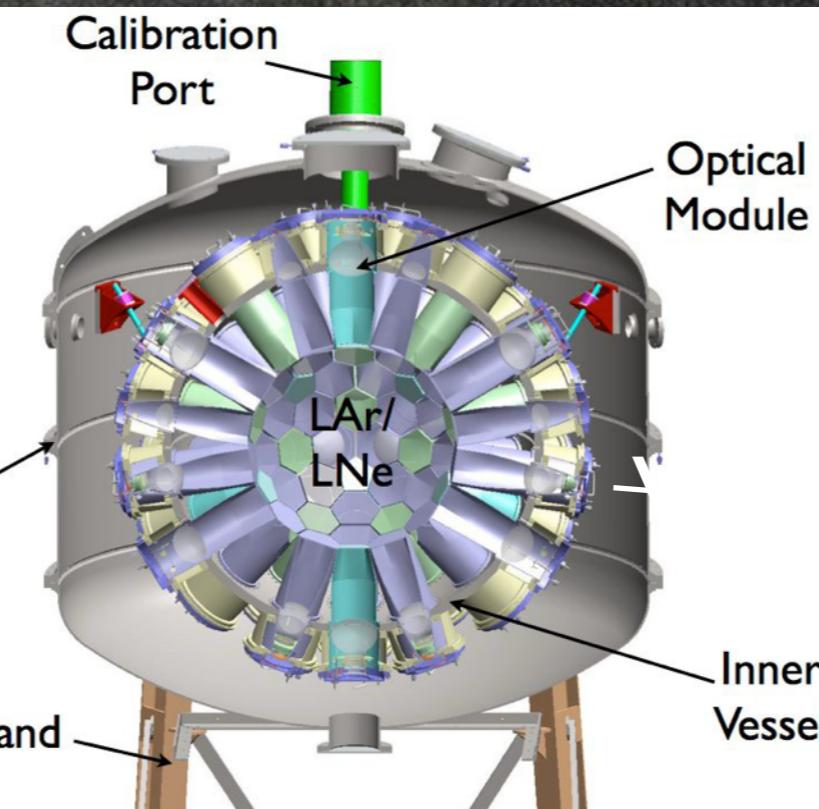
**Sensitivity:**  
SI cross section  $2 \times 10^{-45} \text{ cm}^2$

# MiniCLEAN Design

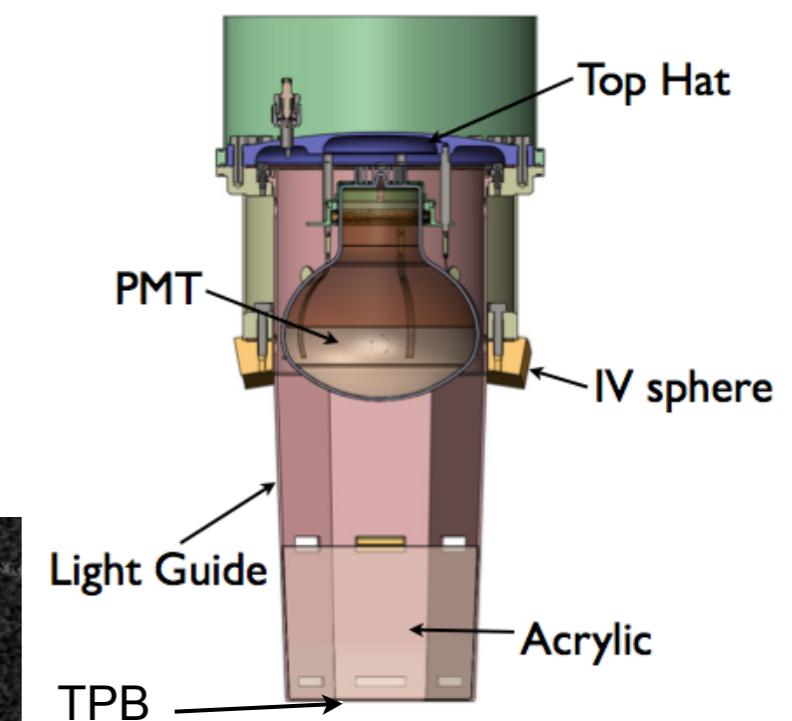


Water veto tank:  
 25 ft. tall  
 18 ft. diameter  
 48 veto PMTs  
 ~1.5 m water on sides  
 ~3.5 m water from  
 bottom

Outer Vessel: ~1.25 m radius  
 Inner Vessel: ~ .7 m diameter,  
 ~500 kg LAr  
 Target radius: ~.4 m  
 Fiducial radius: ~.3 m, 150 kg LAr



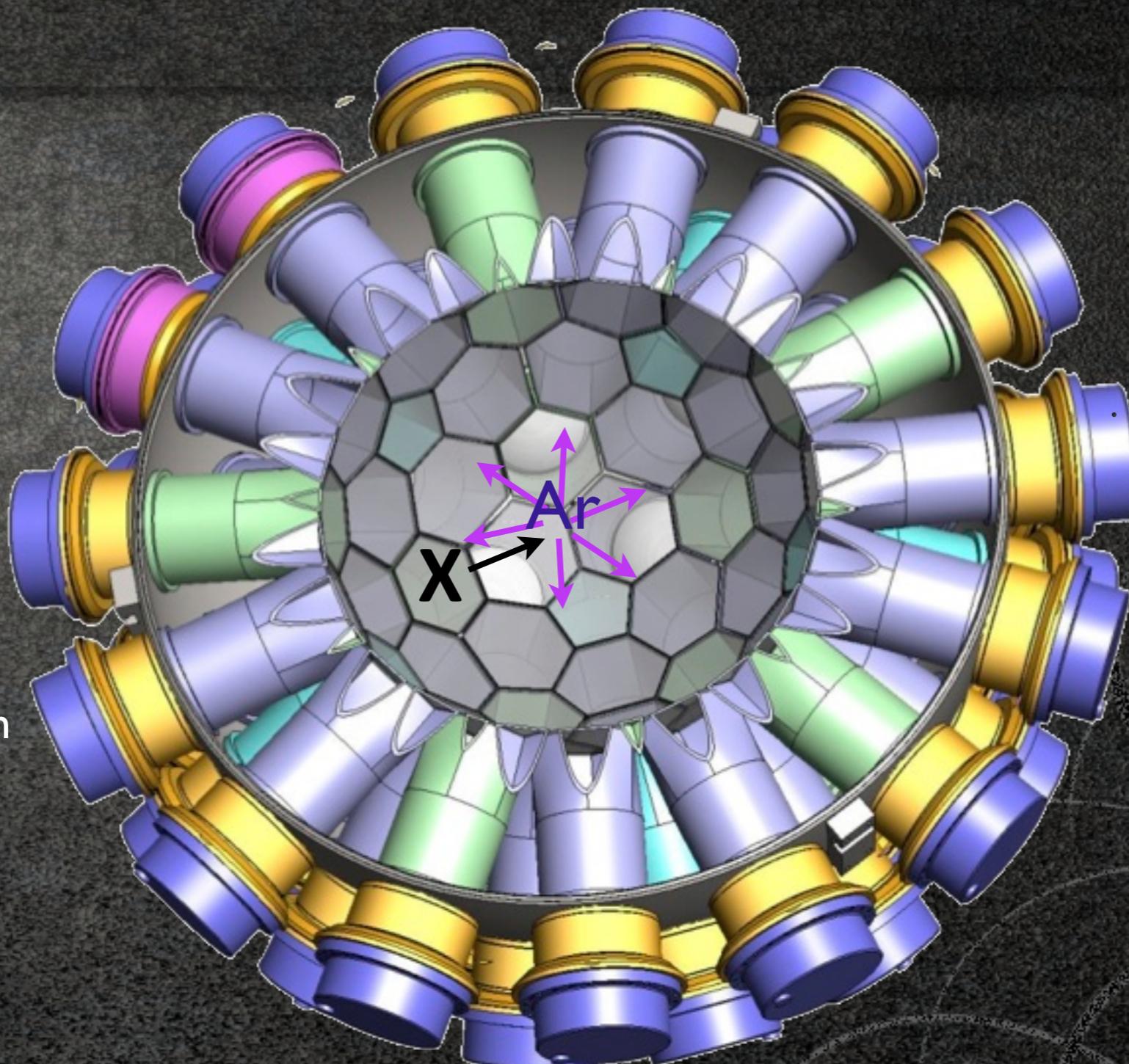
4 $\pi$  coverage  
 92 modular optical cassettes  
 PMTs are cold



# A WIMP Event

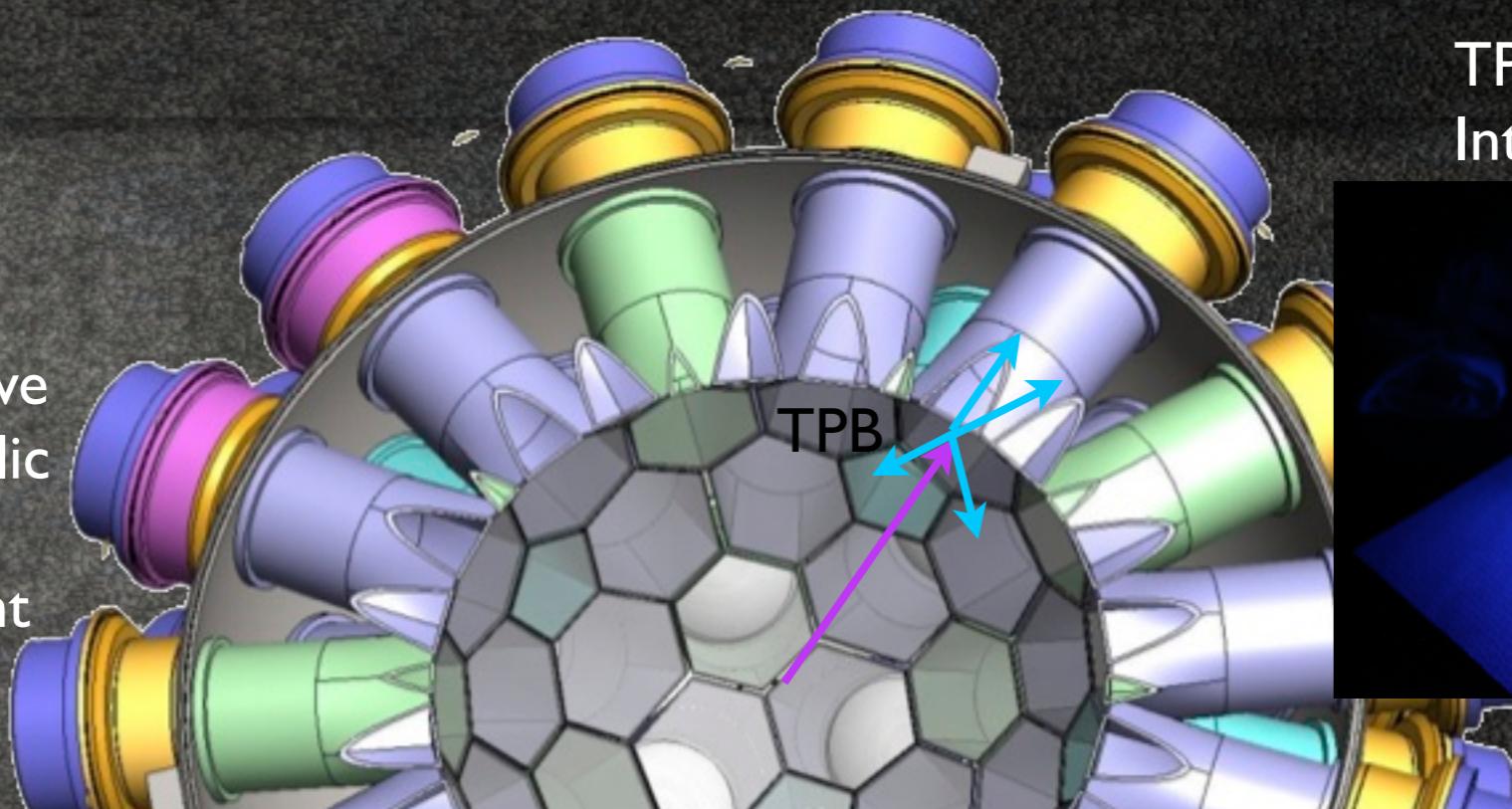
A WIMP, **X**,  
has a coherent  
elastic scatter  
with an argon  
nucleus

The nuclear recoil  
then causes EUV  
scintillation, 128 nm

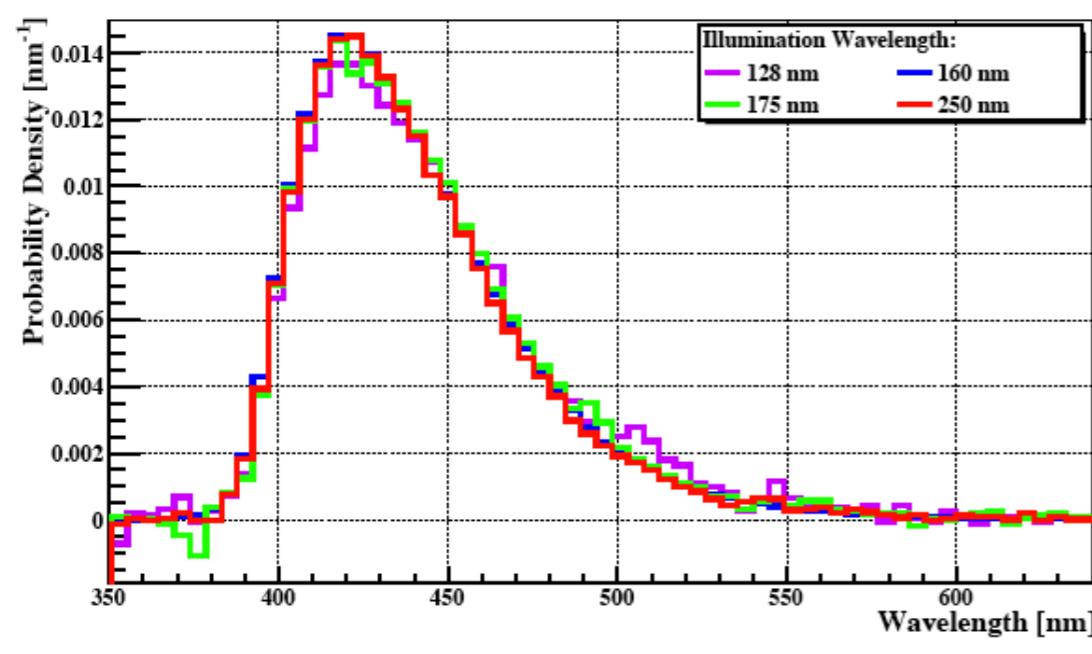


# TPB Re-emission

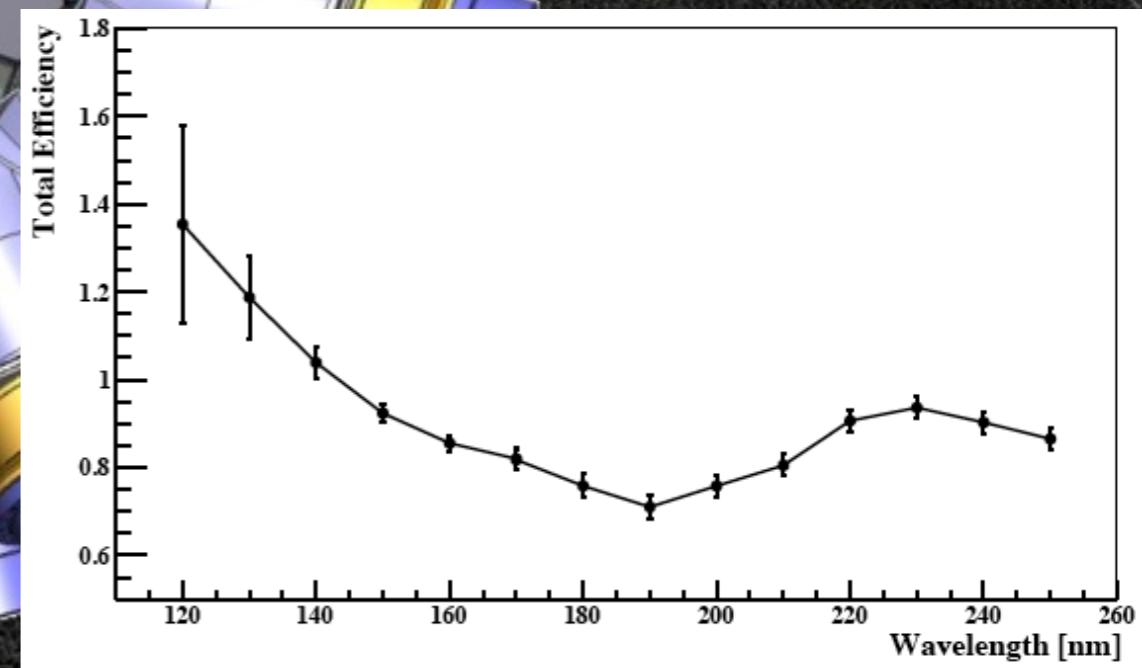
The UV light is absorbed by the TPB wavelength shifter, an evaporative coating on the acrylic face, which re-emits visible light



TPB Coating done at International Vacuum



TPB Re-emission spectrum  
Gehman et al. NIM A654 (2011) 116-121



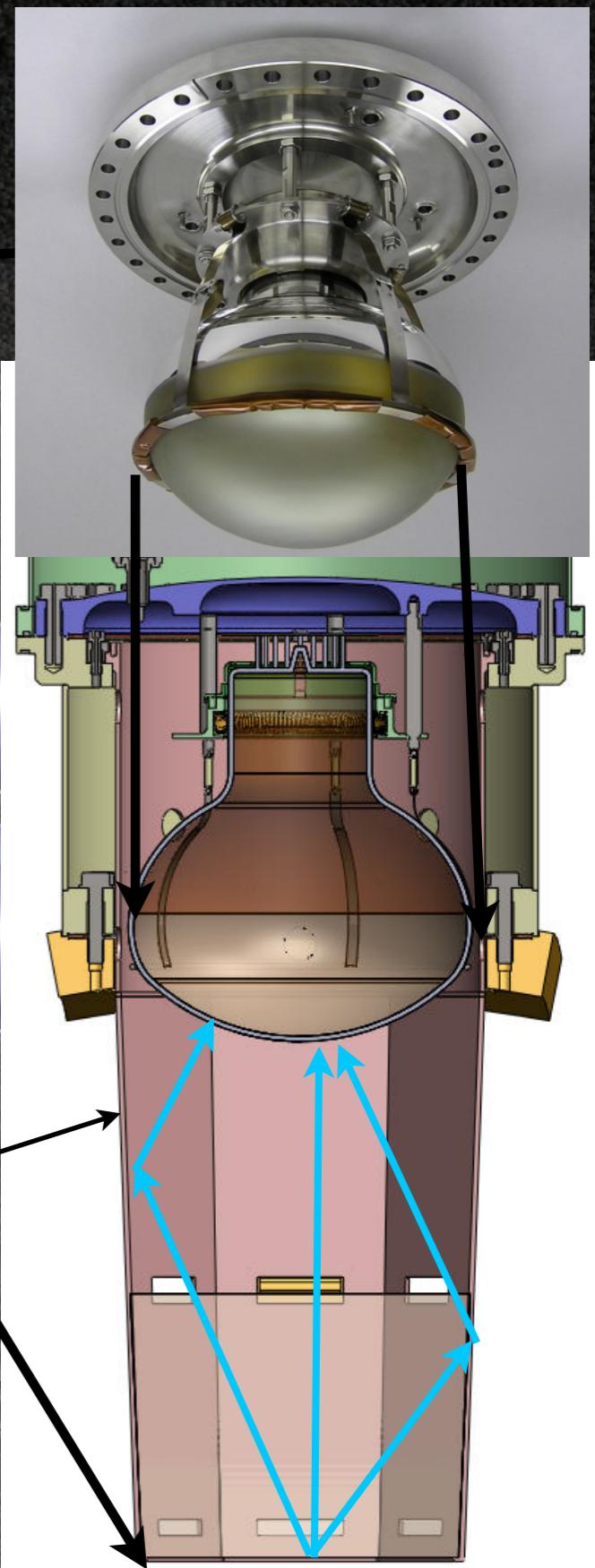
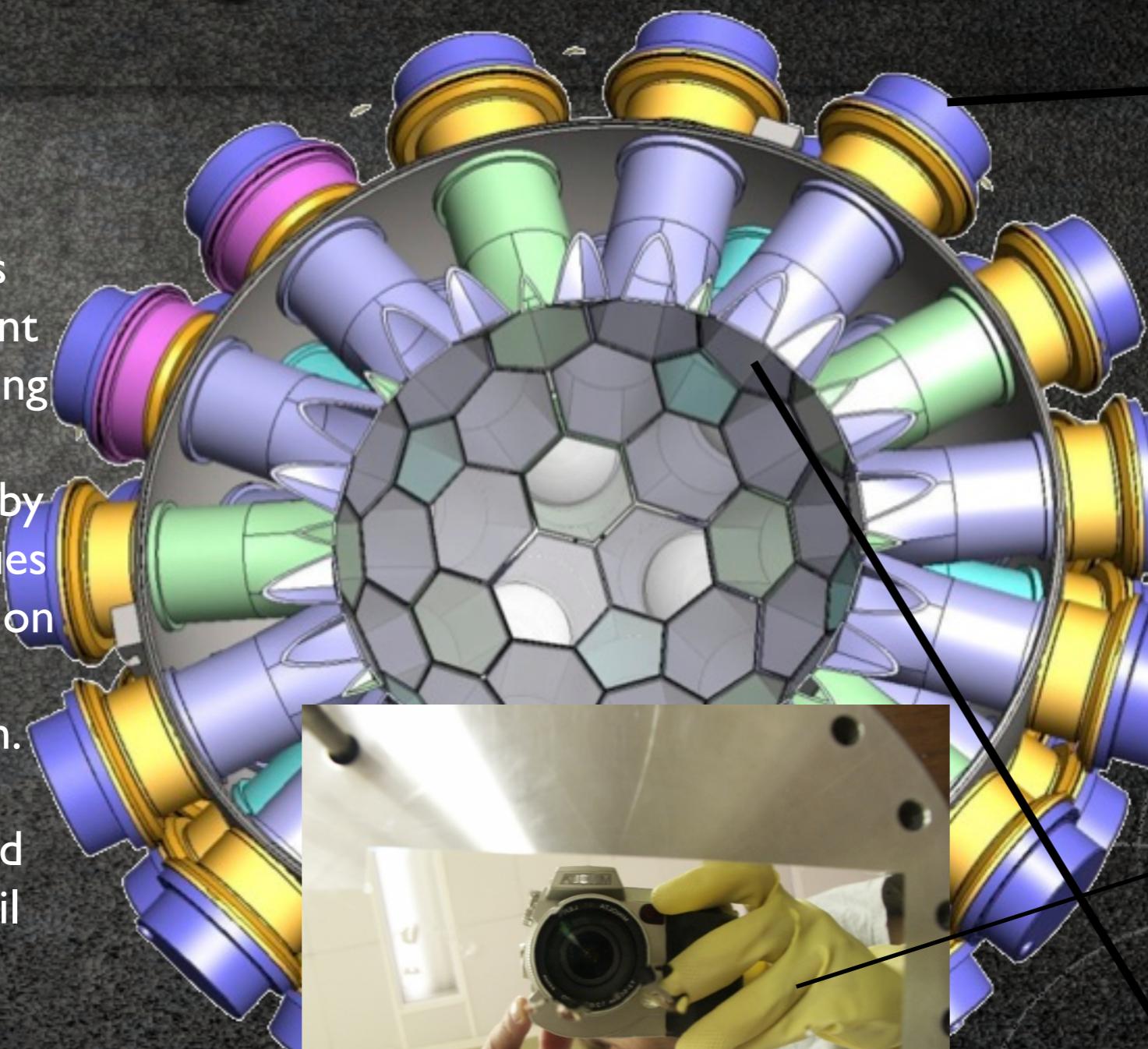
TPB Emission Efficiency

# Optical Cassette

Light then travels through transparent acrylic. UV absorbing

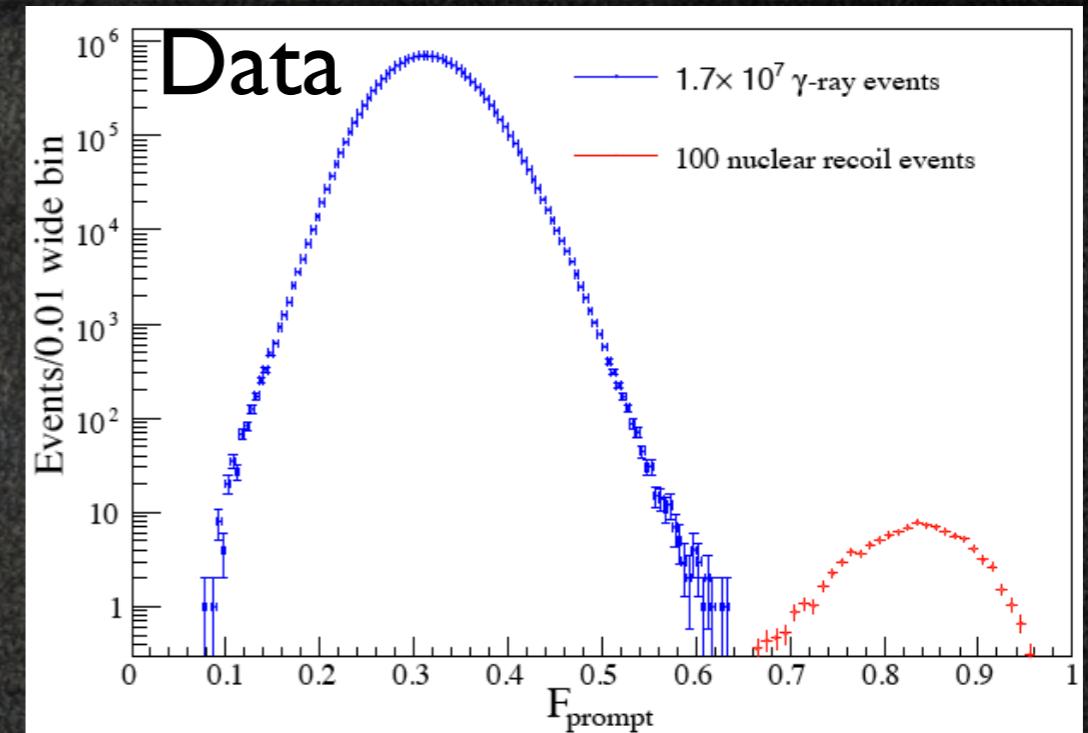
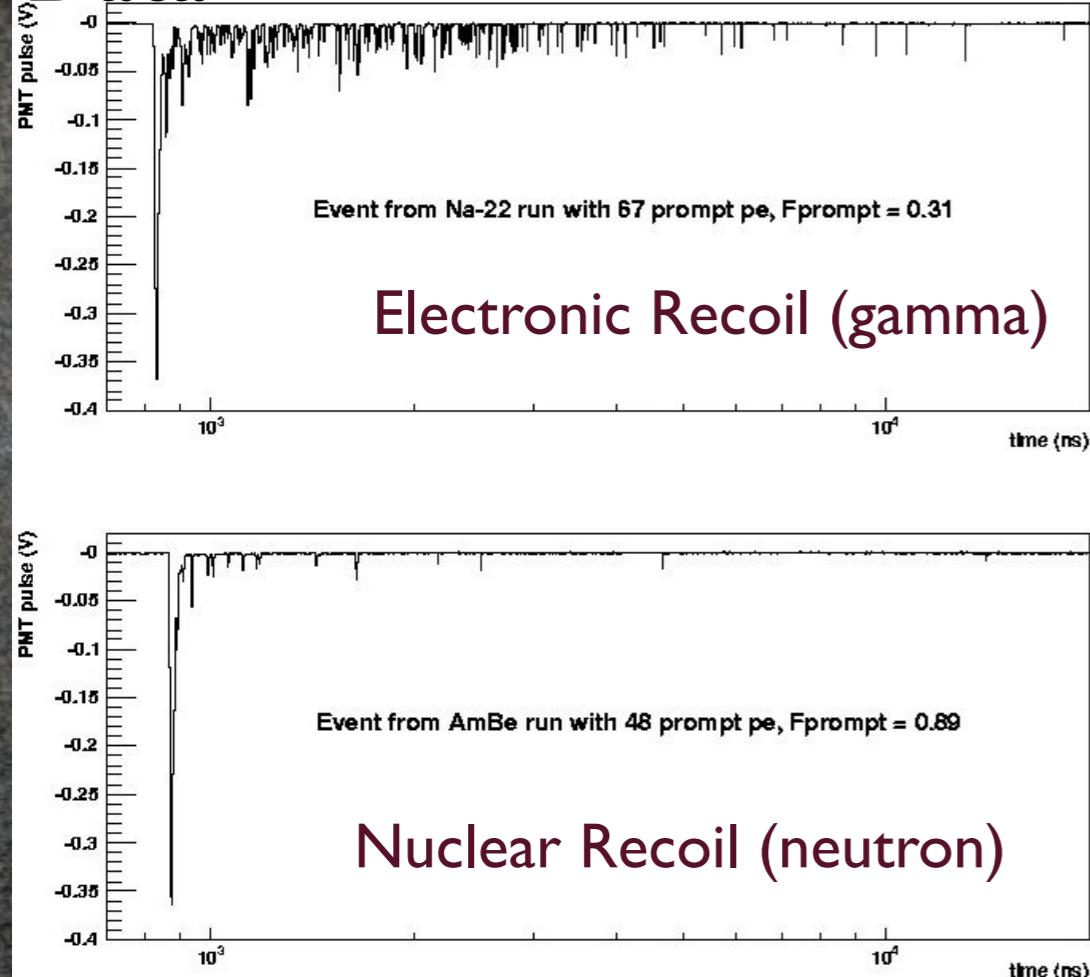
acrylic was surprisingly found by our DEAP colleagues to be best, absorption lengths of ~few meters at 420 nm.

Lightguide covered with 3M DESR foil reflector.

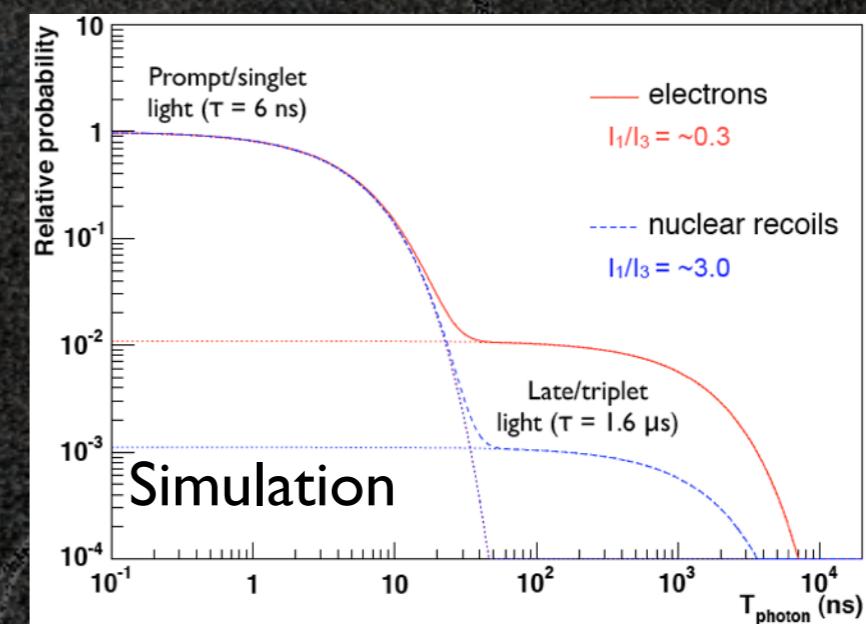


# Pulse Shape Discrimination

## Data



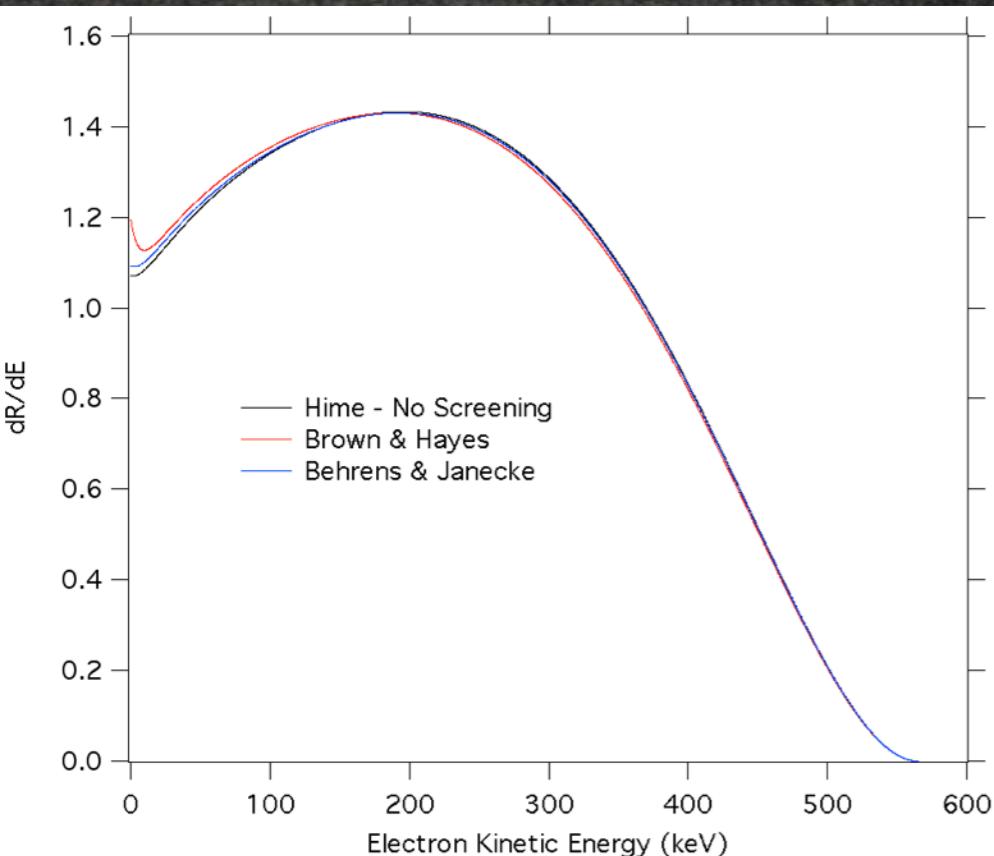
Boulay et al. arXiv:0904.2930



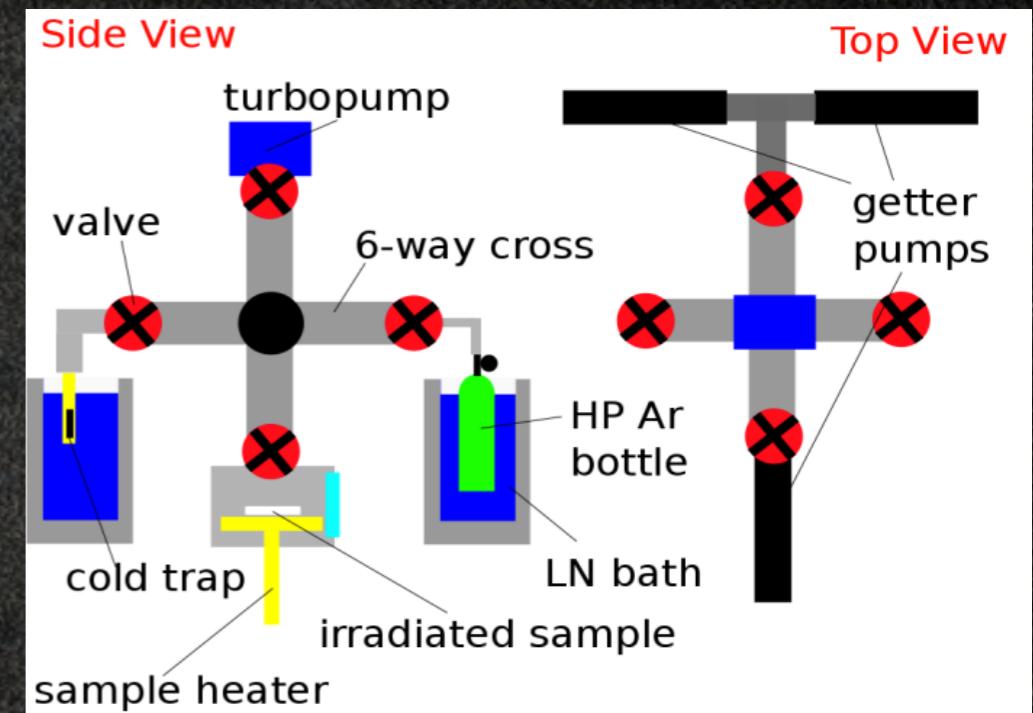
Boulay and Hime, Astropart. Phys. **25**, 179 (2006)

In LAr,  $t_{\text{singlet}} = 6$  ns,  $t_{\text{triplet}} = 1.6$   $\mu\text{s}$   
 $F_{\text{prompt}} \sim .3$  for electron recoils,  $.7\text{--}.8$  for nuclear  
How well can discrimination work?  
DEAP-1 has demonstrated (stat. Limited)  
 $< 6 \times 10^{-8}$   $43 < E < 86$  keVee  
Necessary in LAr because of  $^{39}\text{Ar}$

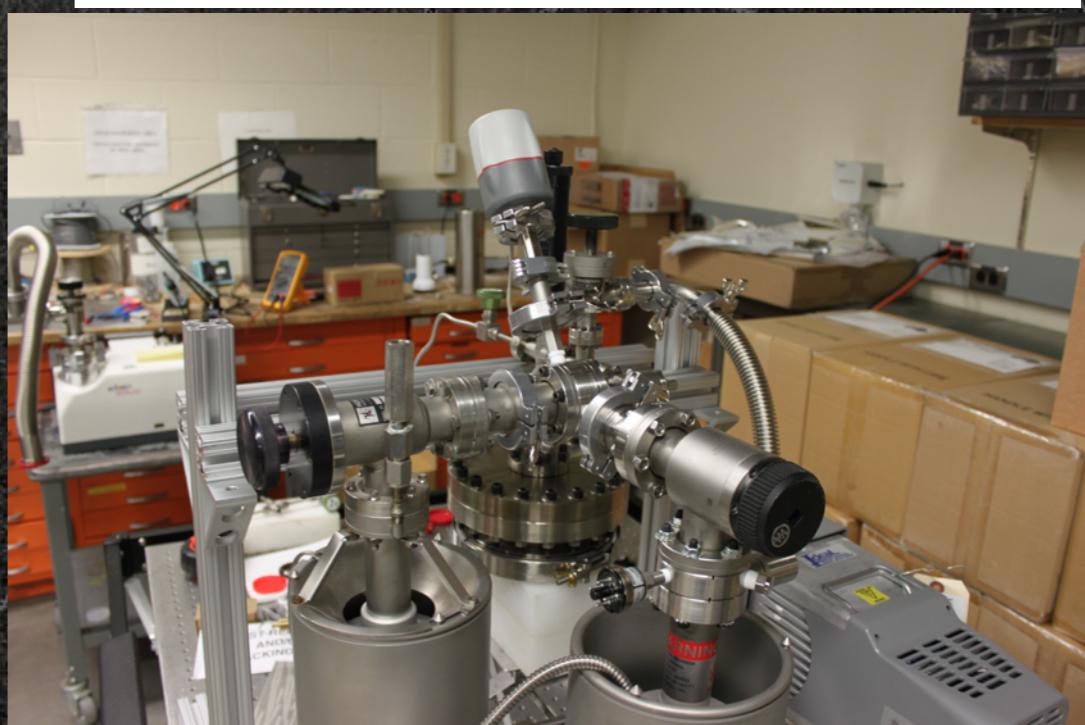
# $^{39}\text{Ar}$ Spike



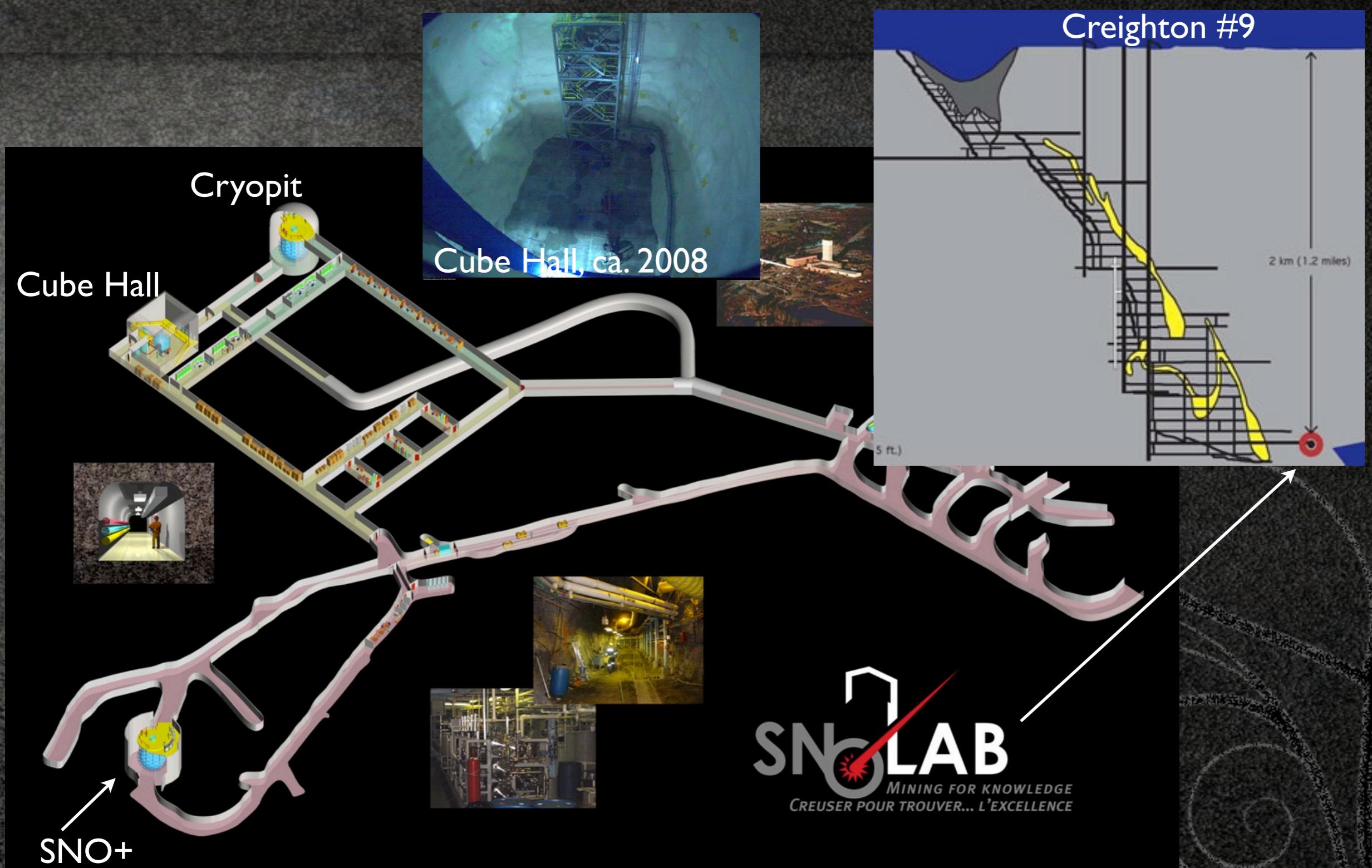
MiniCLEAN will dope with  $^{39}\text{Ar}$  to establish PSD for larger detectors (rejection of  $10^{-10}$ ) plan to increase event rate 3-10 x



KCl target from TRIUMF proton beam, other LANL groups want Si and Al isotopes, we have obtained  $\sim 1.7 \mu\text{Ci}$  of  $^{39}\text{Ar}$  to spike our natural argon



# SNOLab Cube Hall



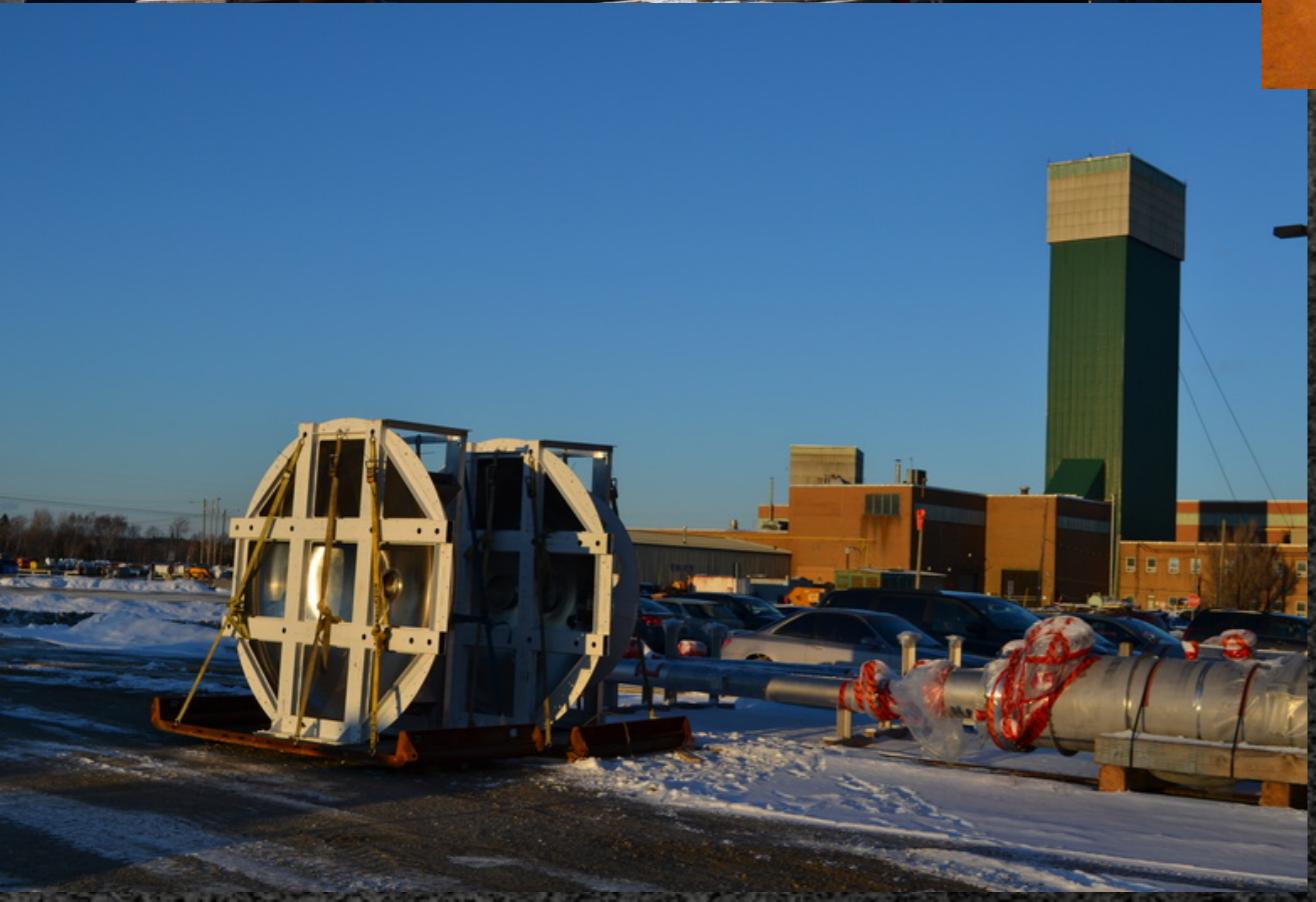
# Construction: Veto Tank



Veto String Test Assembly at Bates Lab, MIT



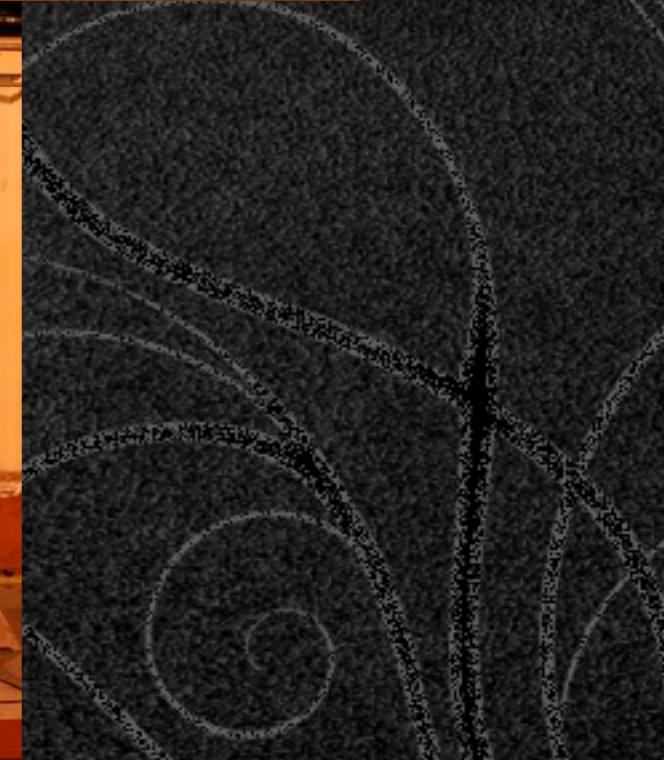
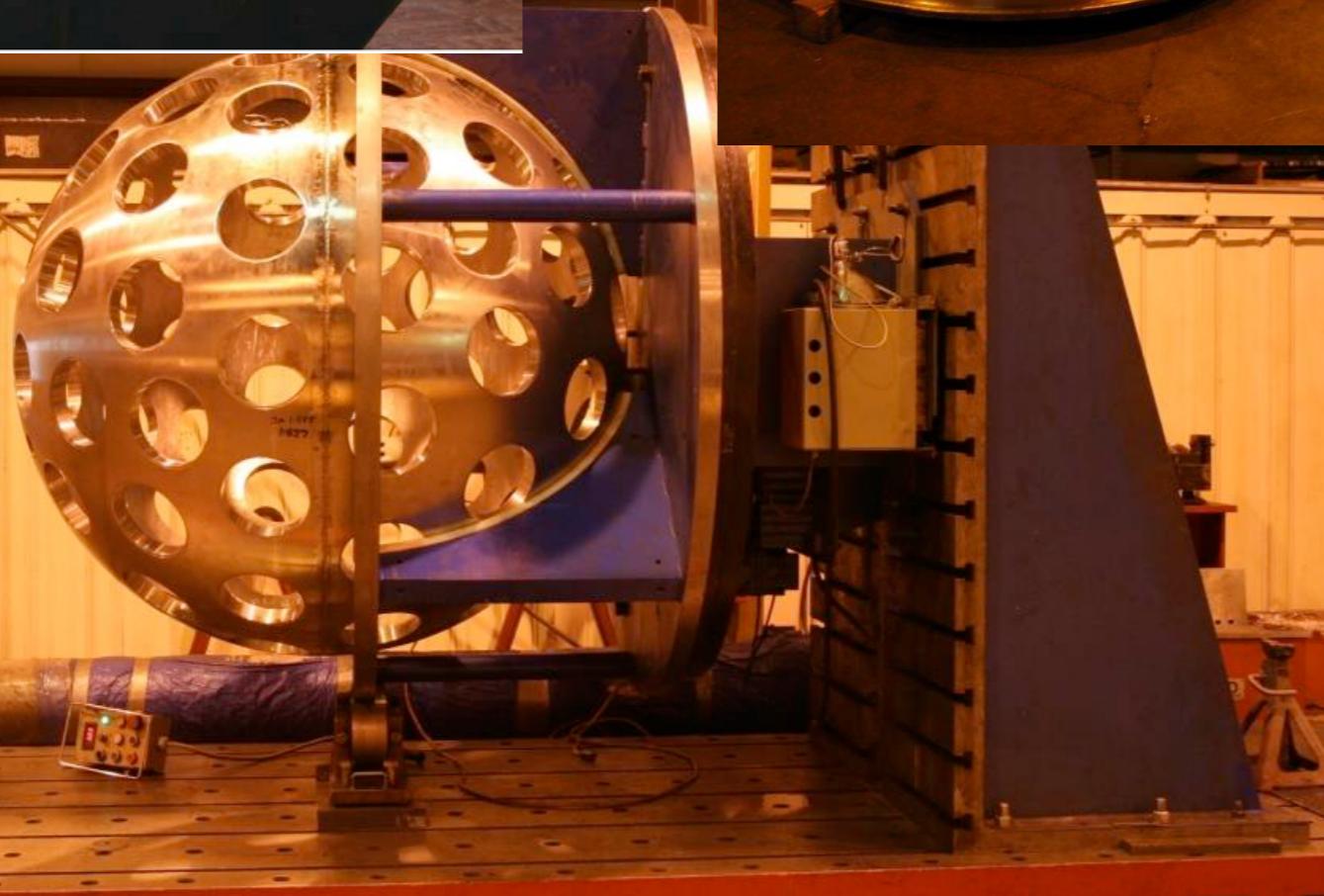
# Outer Vessel Transport



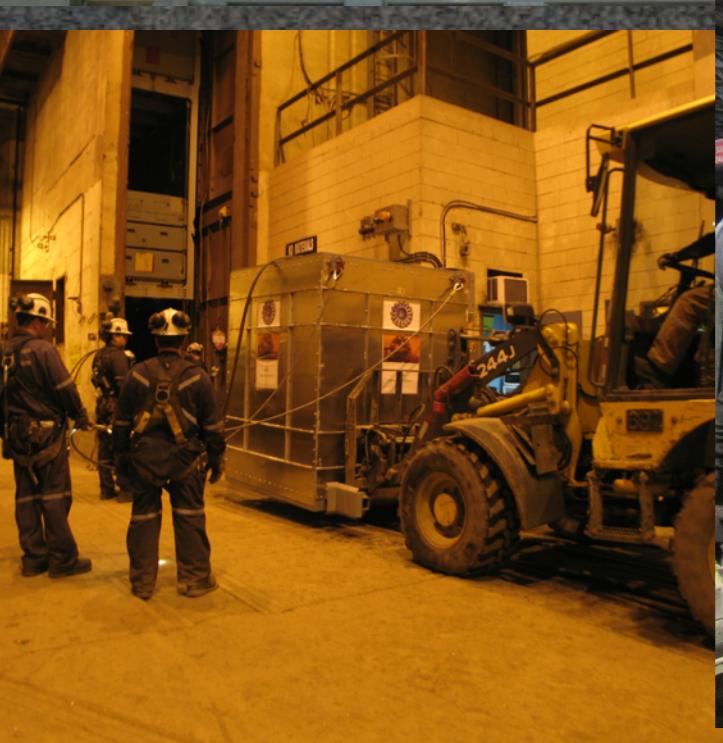
# Outer Vessel Assembly



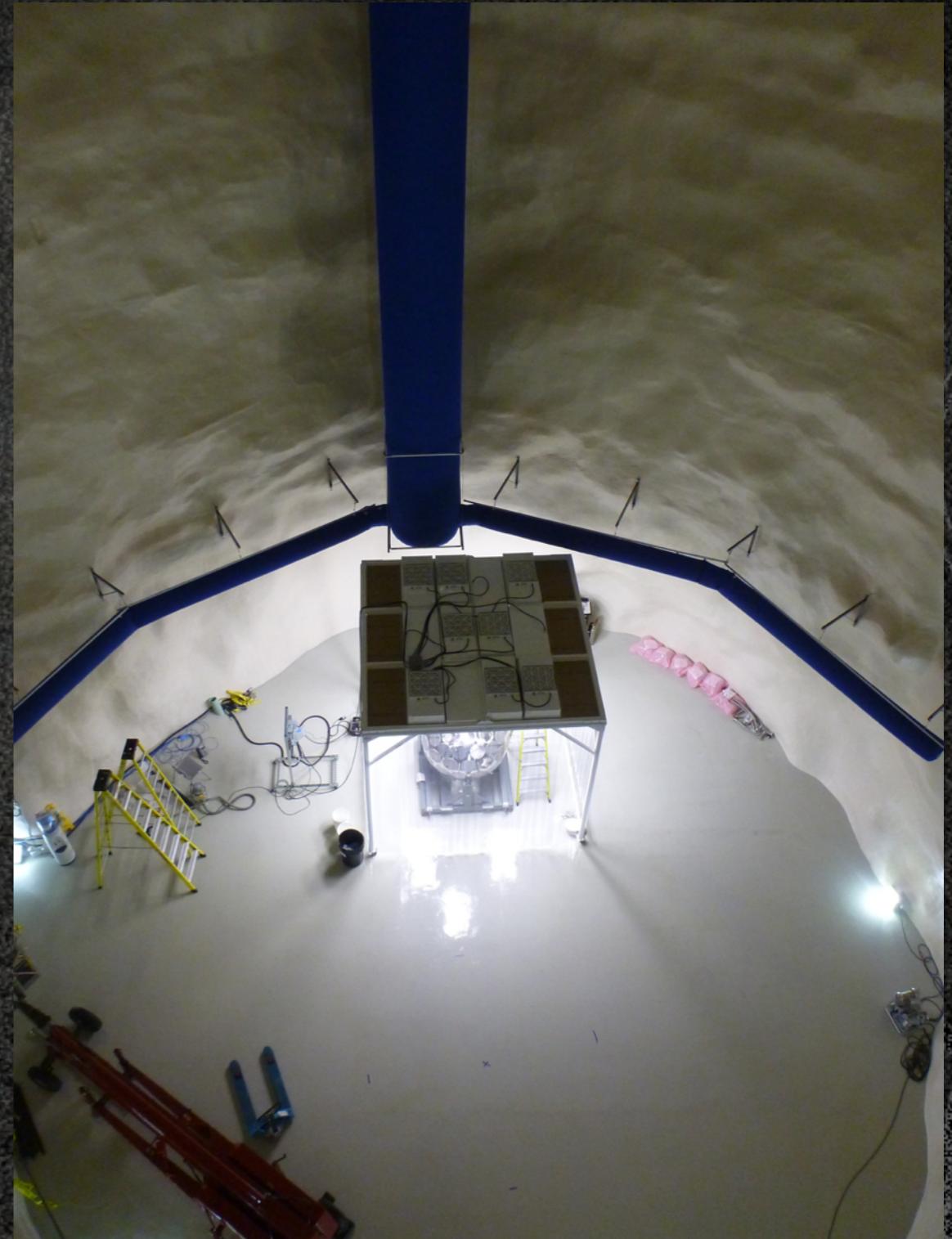
# Construction: Inner Vessel



# Inner Vessel Transport



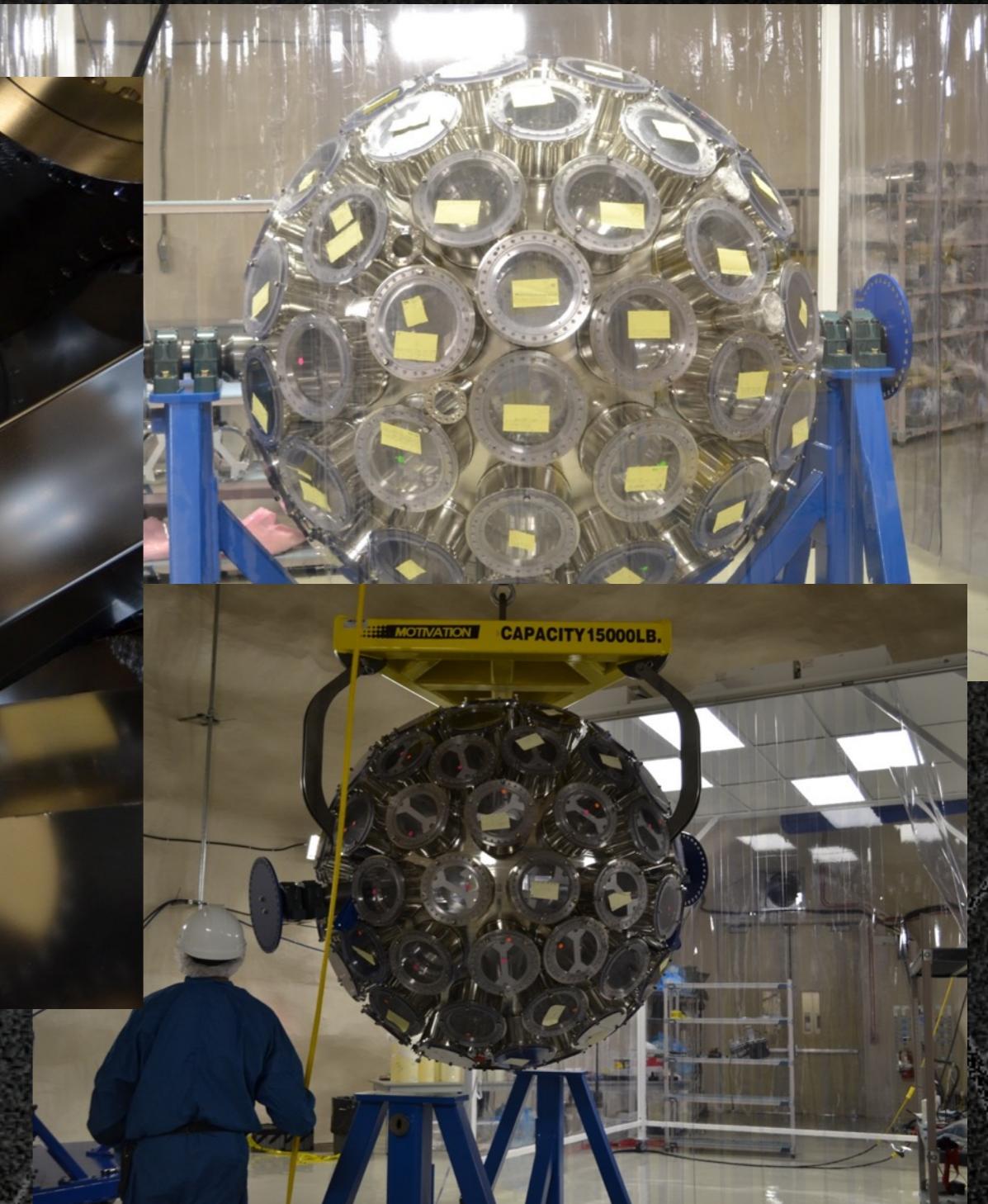
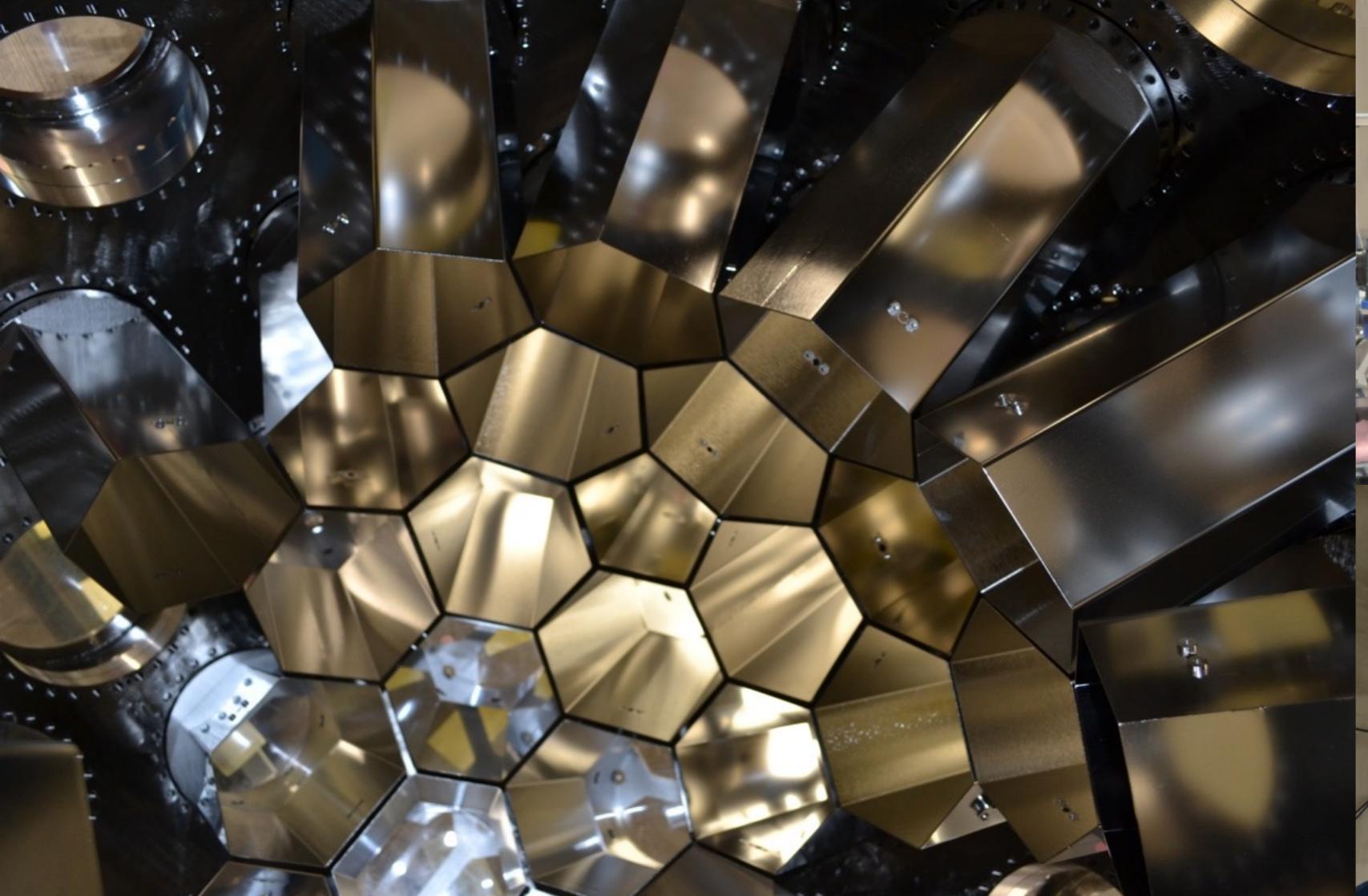
# IV Underground



# IV Preparation



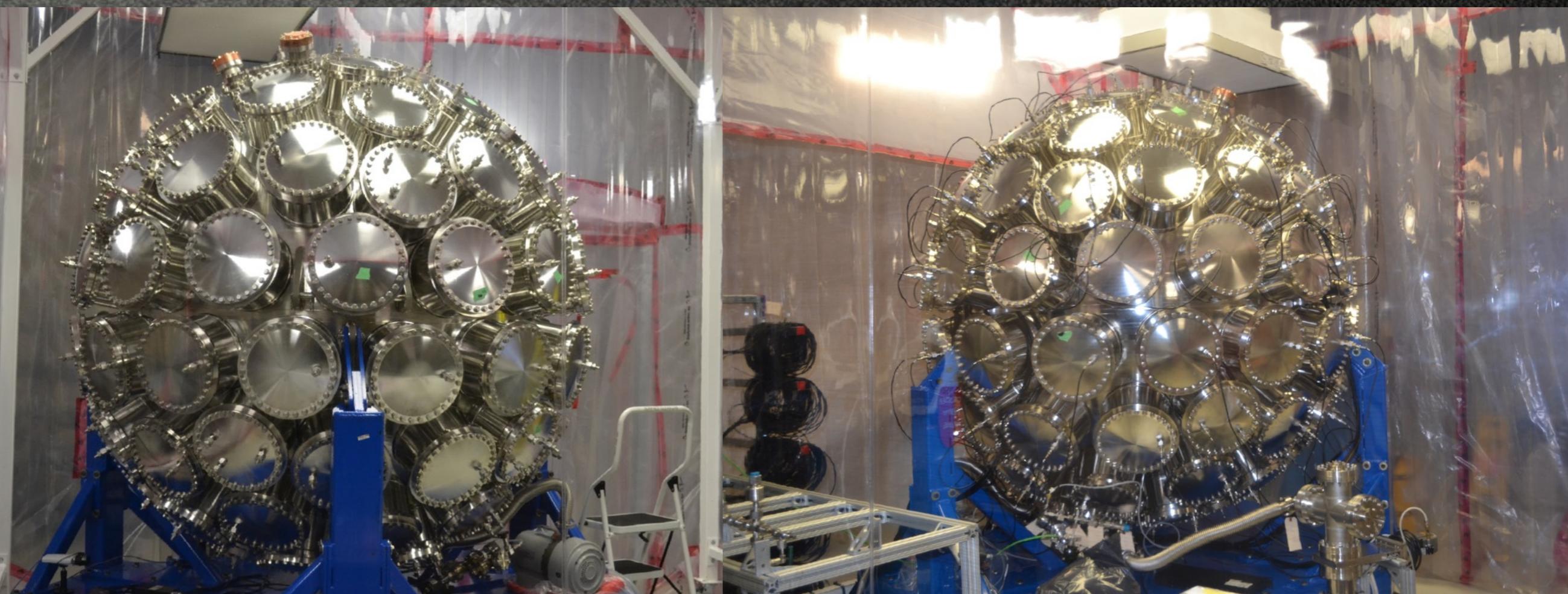
# IV Assembly



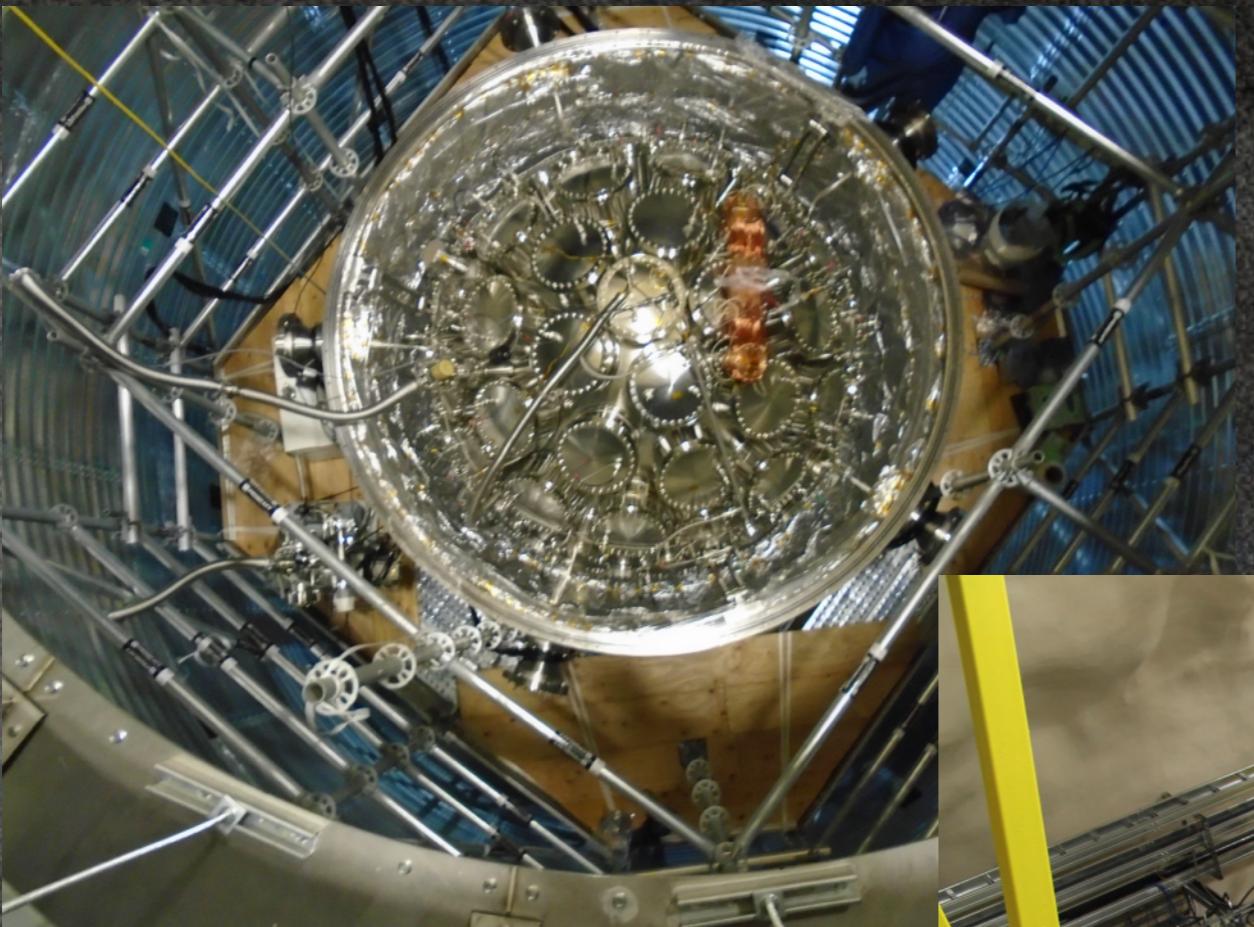
# Cassette Insertion



# IV Complete



# Installed in Cubehall



# Backgrounds

Event Selection	$^{39}\text{Ar}$	PMT (a,n)	Surface as	$\gamma\text{s}$
<b>Raw Rate</b>	1 Bq/kg	42000 n/yr	10000 a/yr	$1.4 \times 10^{10} \gamma/\text{yr}$
<b>Energy Between 12.5-25 keV<sub>ee</sub></b>	$4.2 \times 10^8$	$352.2 \pm 2.1$	3360	$6.0 \times 10^6$
<b>Fiducial Cut</b>	$1.2 \times 10^8$	$91.6 \pm 1.1$	$0.82 \pm 0.09$	$3 \times 10^5$
<b>Fraction of Prompt Light</b>	$75 \pm 1.1$	$7.1 \pm 0.3$	$0.24 \pm 0.05$	<0.36
<b>Bayesian Discriminant</b>	$0.3 \pm 0.2$	$3.8 \pm 0.2$	$0.14 \pm 0.04$	-
<b>Tagging</b>	-	$0.9 \pm 0.1$	-	-
	PSD	Fiducial	Fiducial	PSD

# Schedule

- **Technology Demonstration**
  - Energy scale, position resolution, pulse shape discrimination, surface alpha rate
- **Dark Matter Search**
  - expected  $\sim 2 \times 10^{-45} \text{ cm}^2$
- **$^{39}\text{Ar}$  Spike**
  - Test pulse shape discrimination at level required for 150 tonne Generation 3 detector (assuming underground Argon depleted of  $^{39}\text{Ar}$ )
  - Inform DEAP-3600 in case of signal observation
- **Liquid Neon Run**
  - Demonstrate target exchange capability for Generation 3 dark matter detector
  - Prototype for precision pp-Solar neutrino measurement