

CREAM:

AT THE HIGH ENERGY COSMIC RAY FRONTIER

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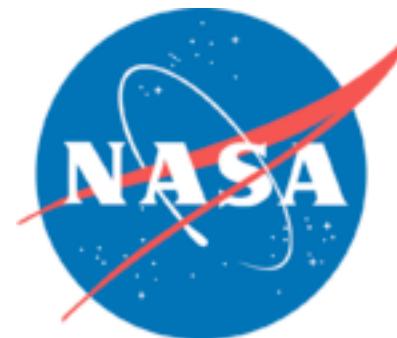
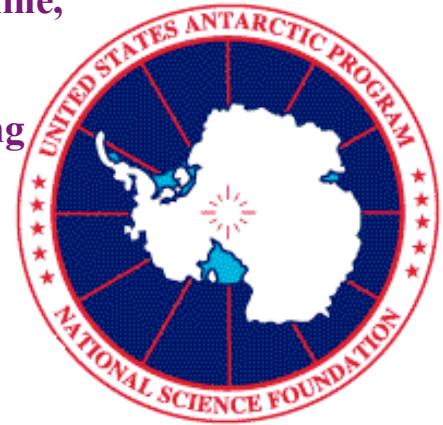
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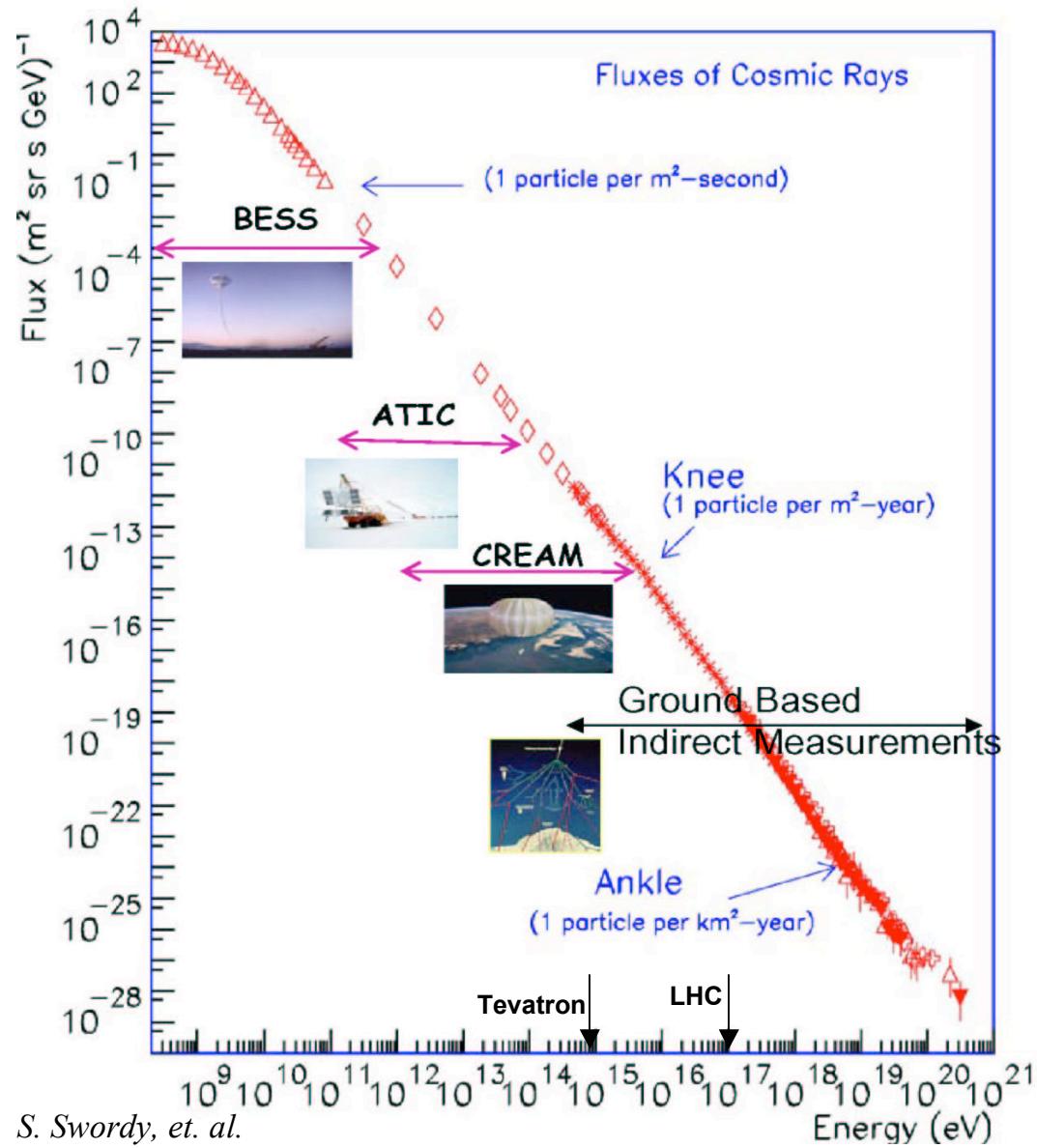
and WFF



Also thanks to:

NSF, INFN, KICOS, and MOST

All-particle CR Spectrum



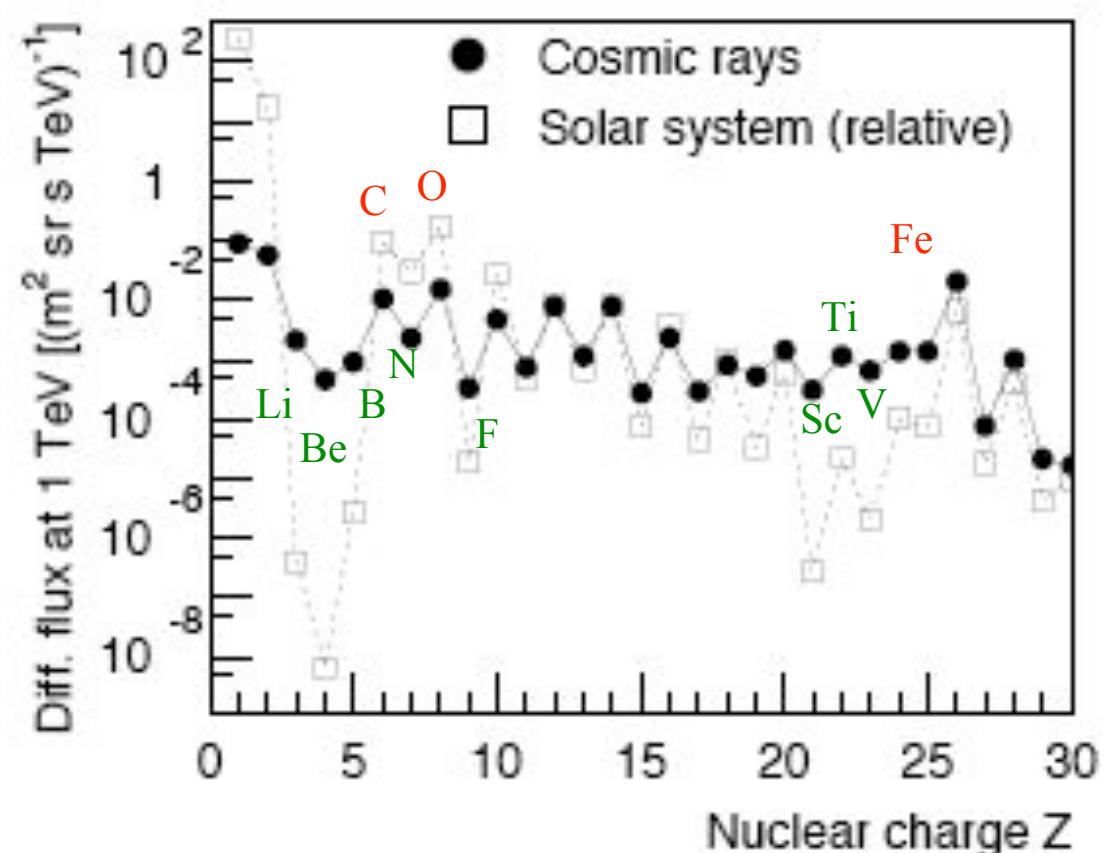
Abundance

Cosmic ray abundances

- trace solar system values
- implying stellar origin.

Even-odd effect:

- Even-charge elements more abundant due to stability in stellar nucleosynthesis.
- Over-abundance of unstable elements in CRs implies spallation of primaries during propagation.



Hörandel

Cosmic Ray Energetics and Mass Particle Detector

Charge and Energy:

- $+1 \leq Z \leq 26$
- $10^{12} \text{ eV} < \sim E < \sim 10^{15} \text{ eV}$

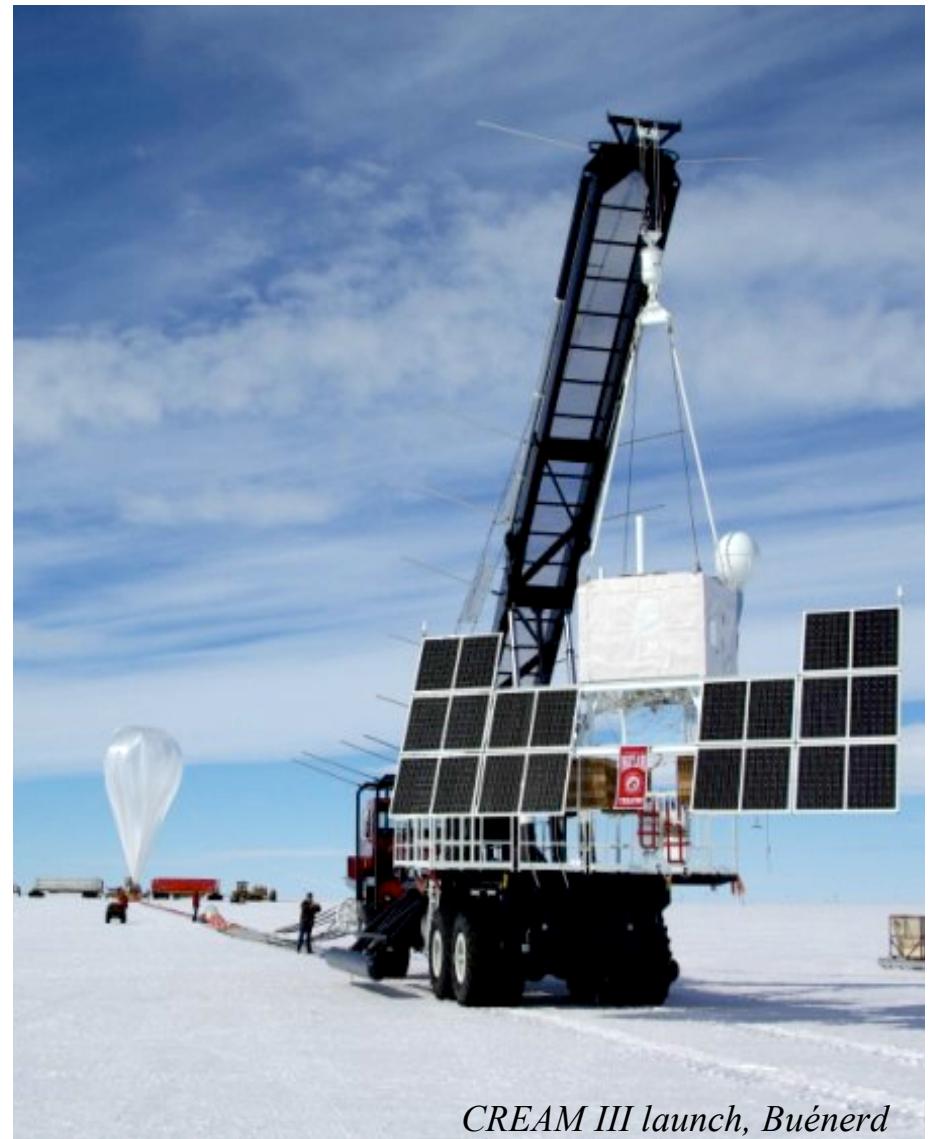
from 4 main subsystems:

- TCD (Z), TRD (E), SCD (Z), Cal (E)

Timing Charge Detector, Transition Radiation Detector,
Silicon Charge Detector, and Calorimeter

Flown on a Long-Duration Balloon over
Antarctica

- CREAM I: 42 days
(16 Dec 04 - 27 Jan 05)
- CREAM II: 28 days
(16 Dec 05 - 13 Jan 06)
- CREAM III: 28 days
(19 Dec 07 - 16 Jan 08)

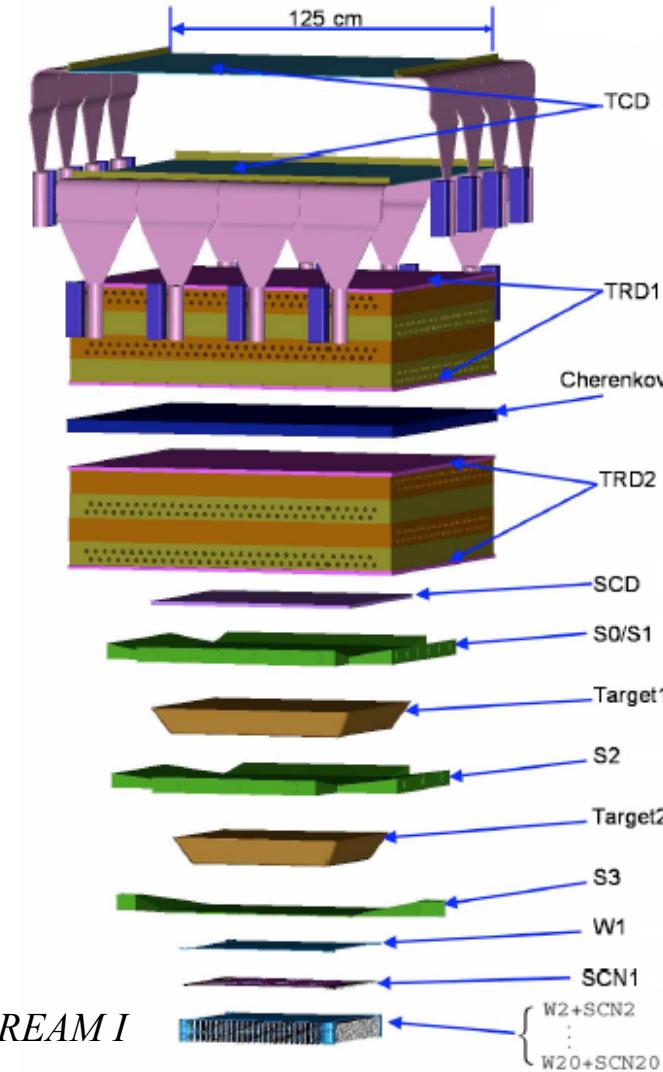


CREAM III launch, Buénerd

How CREAM (I) does all that.

Timing Charge Det.:

- Scintillation $\Rightarrow Z, \beta$
- $\sigma_Z \sim 0.2\text{-}0.35e$
- High Z trigger
- PMTs read out with v. fast electronics



Silicon Charge Det.:

- Si $\Rightarrow Z, E$
- $\sigma_Z \sim 0.2e$
- smaller (area \sim half) acceptance than TCD

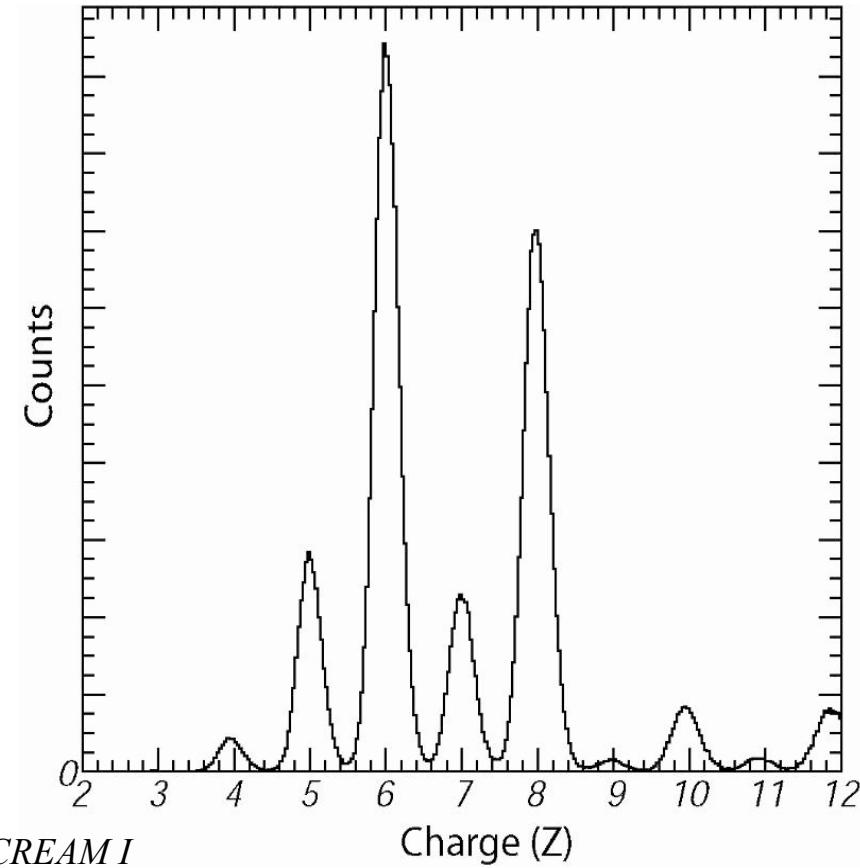
Transition Radiation Det. (+Cherenkov):

- Ionization & TR $\Rightarrow \gamma, Z$
- Tracking (mm) and E
- Threshold: $Z > 3$
- Cher $\rightarrow \gamma \rightarrow E$

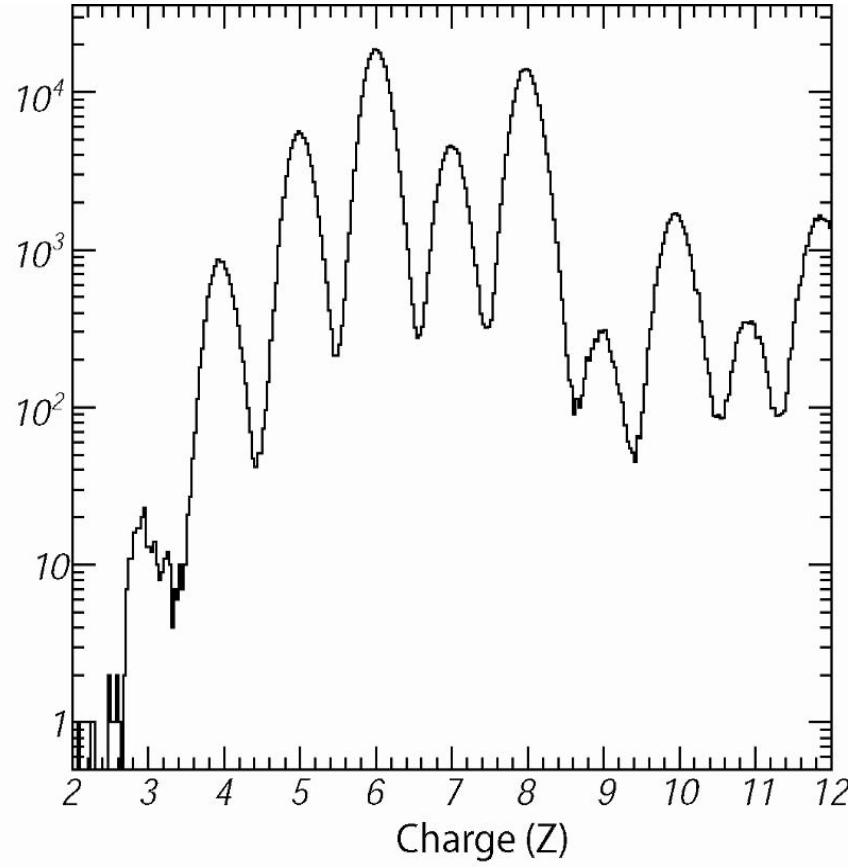
Calorimeter:

- Inelastic interactions in Carbon target
- Track secondaries
- W & Scin. fibers measure E_{dep}
- Hi E trigger

Charge



CREAM I



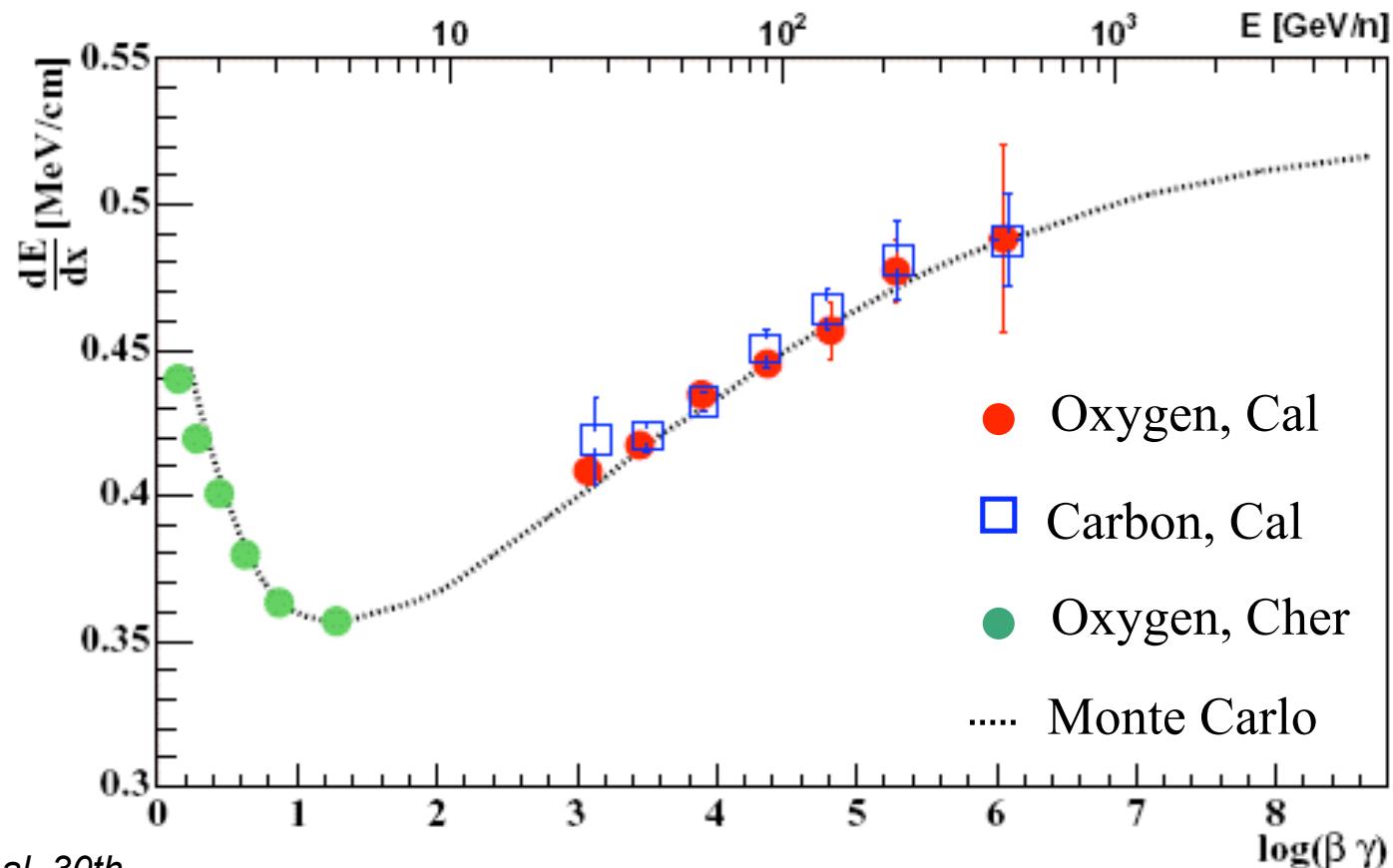
Carbon: $\sigma = 0.16$ e

Oxygen: $\sigma = 0.17$ e

from TCD and SCD

Energy

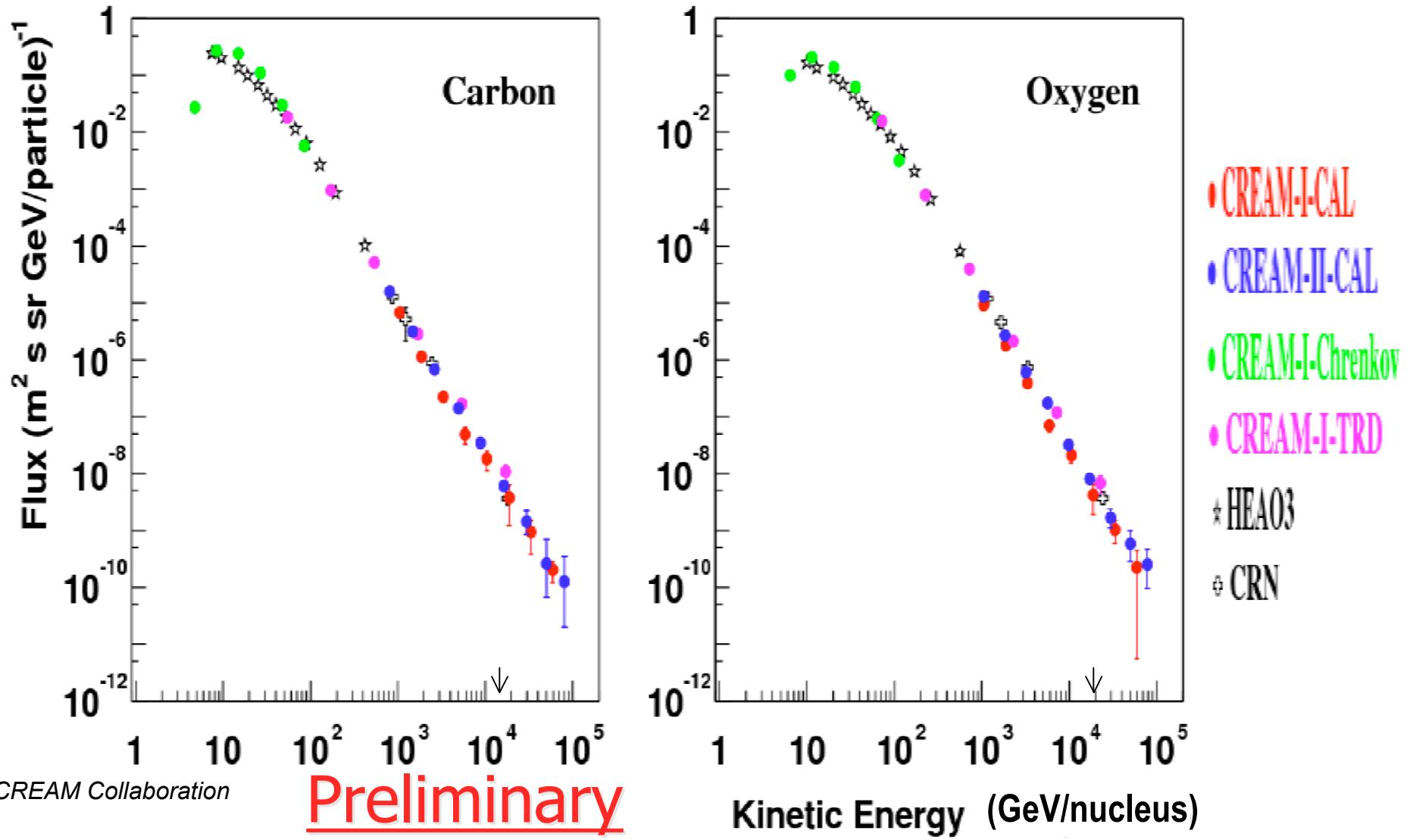
Cross-calibration of TRD and Calorimeter energy measurements



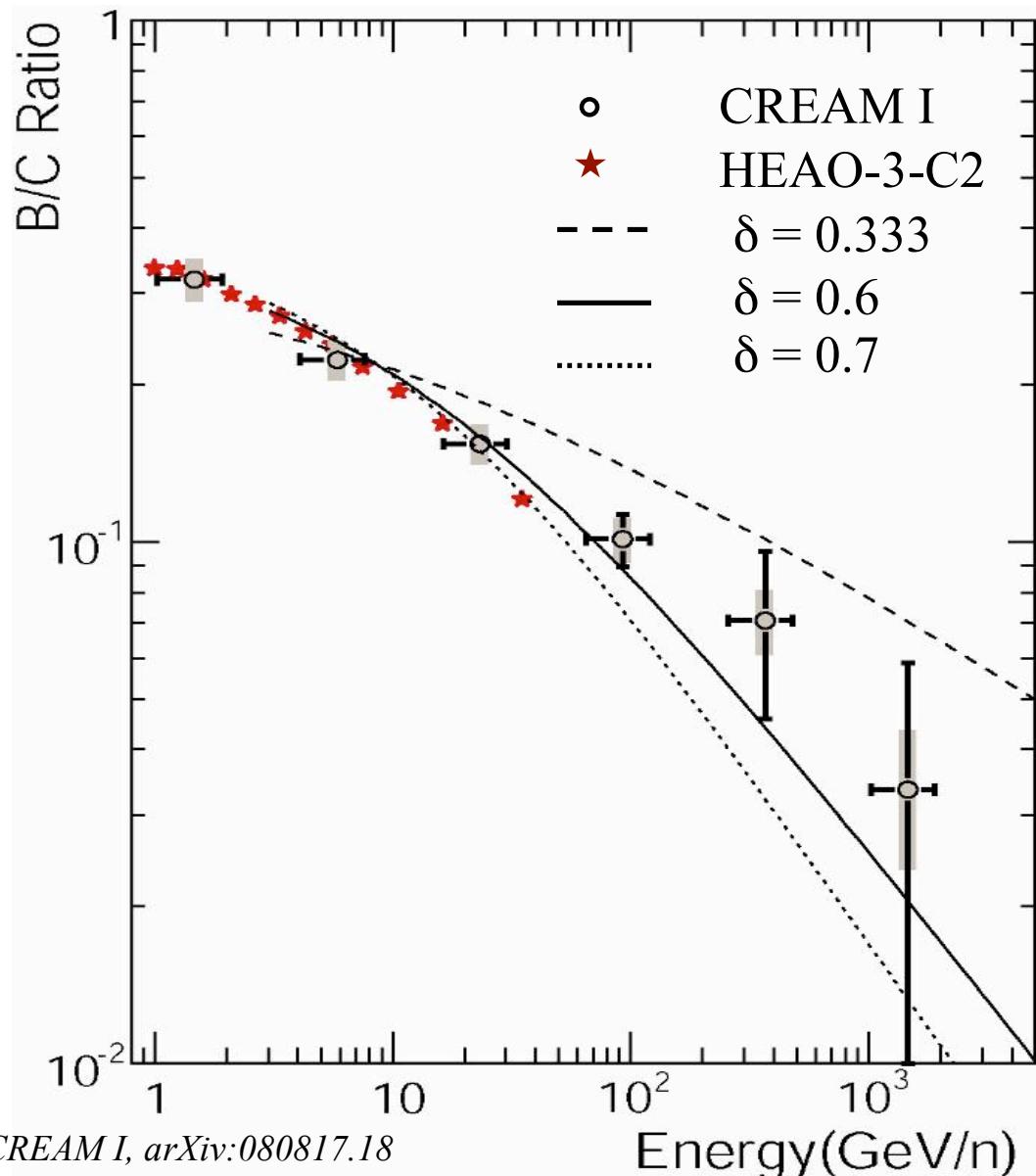
Maestro et al. 30th
ICRC, Merida, 2007

Preliminary Results

Carbon and Oxygen Spectra



Boron:Carbon



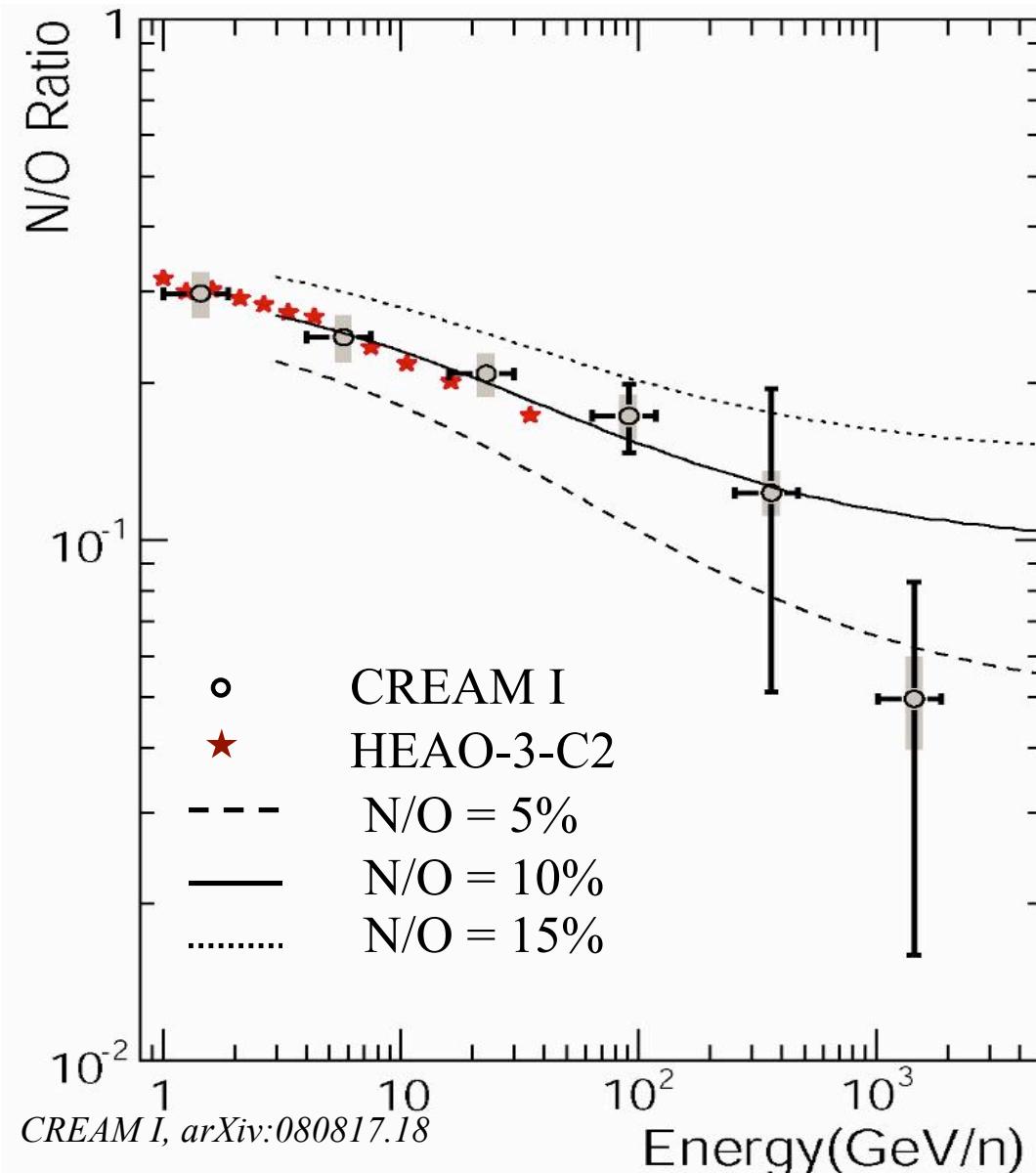
Lines:

- Simple Leaky-box model with range of propagation index δ

Error Bars: (Vertical)

- Line: Statistical
- Shaded: Systematic, dominated by uncertainties in atmospheric secondary production at $E > 300\text{GeV}/n$

Nitrogen:Oxygen



Lines:

- Simple Leaky-box model assuming $\delta = 0.6$ with range of N/O in source

Error Bars: (Vertical)

- Line: Statistical
- Shaded: Systematic, dominated by uncertainties in atmospheric secondary production at $E > 300\text{GeV}/n$

Conclusions

- CREAM data agrees with previous measurements and has extended the Carbon and Oxygen spectra in energy.
- CREAM has extended the Boron to Carbon and Nitrogen to Oxygen ratios with good statistics 2 orders of magnitude higher in energy.
- The galactic propagation index from B:C corresponds to a high energy propagation pathlength of $\sim 1 \text{ g/cm}^2$ and, in conjunction with the N:O ratio, suggests a source N abundance at high energies of $\sim 10\%$.
- CREAM II and III have increased our statistics; further analysis is forthcoming. Systematic errors can be reduced by improving our understanding of nuclei interactions.
- CREAM IV is being prepared to fly again this winter.