



Managed by Fermi Research Alliance, LLC for the U.S. Department of Energy Office of Science

Readout and control of superconducting detectors

Gustavo Cancelo

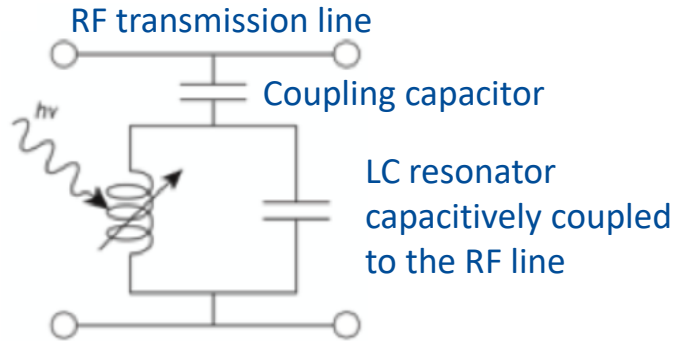
CPAD 2019, Madison, WI
December 8th, 2019

Some examples of superconducting detectors in astrophysics and astronomy

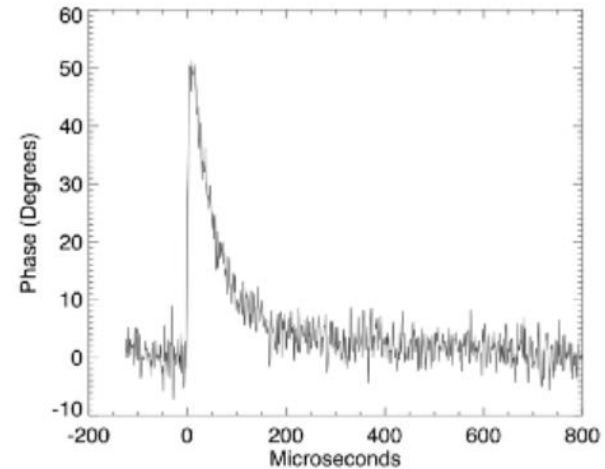
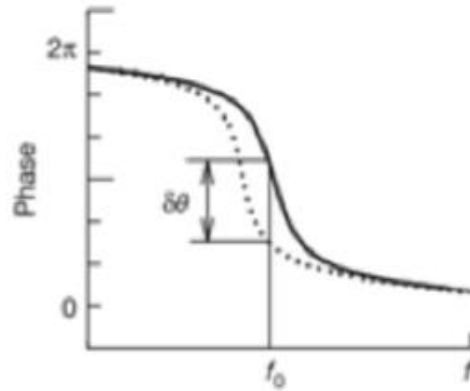
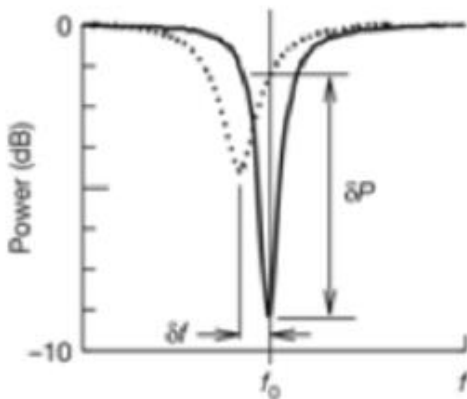
- TES-SQUIDs: B. Cabrera (Stanford), K. Irwin and NIST, Caltech, JPL, ANL, LBNL, etc.
 - Dark matter: CDMS, SuperCDMS.
 - CMB 1st, 2nd, 3rd generation of experiments.
 - Now applications from IR-Optical-Xray.
- MKIDs: Zmudzinas (Caltech), B. Mazin (UCSB), E. Shirokoff (UC).
 - Astronomy: Exoplanets, NIKA2.
 - Dark Energy: Large survey low resolution spectroscopy (future) .
 - Applications from IR-Optical-Xray
- SNSPD: D. Semenov (Moscow), R. Hadfield (Glasgow), U Rochester.
 - Quantum photonics. High resolution photon timing ~10ps. Q. Communications.
- 3D Superconducting qubits (Yale, UCSB, Google, IBM, UC, etc)
 - Dark matter: Axion detection, ADMX?
- **So far, all superconducting detectors are controlled and readout by non superconducting electronics between 4K and 300K.**
 - The field of superconducting readout electronics is progressing but won't be able to replace warm electronics in the next decade. The consumer market of electronics is still based on 300K devices.
- **This talk addresses the similarities (and differences) of highly multiplexed readouts for systems that use, typically, a portion of the RF spectrum (e.g. 4GHz to 8GHz).**

MKIDs are RF resonators

A microwave or optical photon deposits energy on the inductor breaking Cooper pairs into electrons

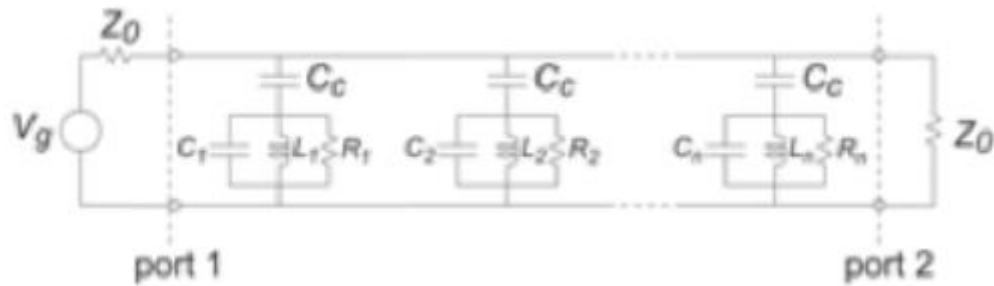


The LC resonator has a large $Q \sim 30K$.
 The photon generates electrons and add a resistor in parallel to LC.
 The Q lowers, the resonance shifts to the left (amplitude and phase).
 The electrons go back to Cooper pairs in 100 usec time.



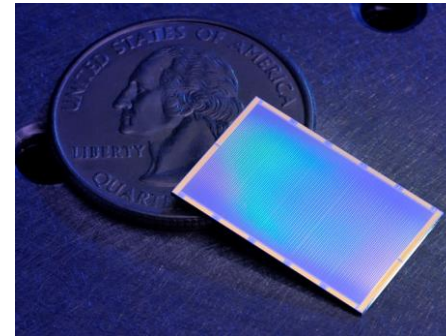
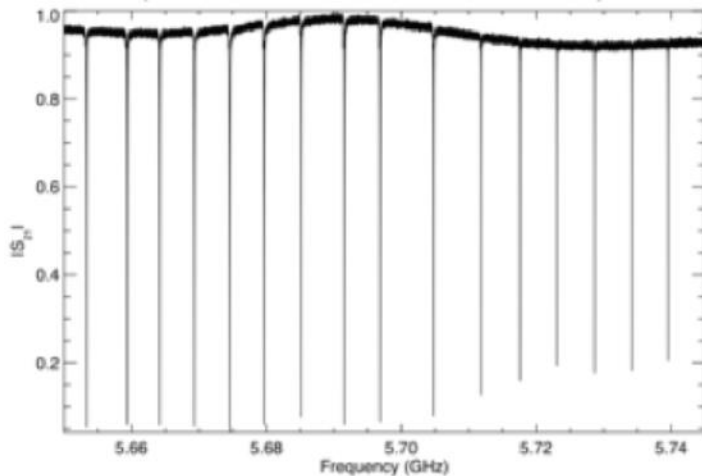
- The superconducting part of the inductance is the kinetic inductance and used for detection. The capacitors can be used for frequency placement and coupling (loaded Q).
- Typical bandwidth $\sim 250KHz$

MKIDs can be frequency multiplexed

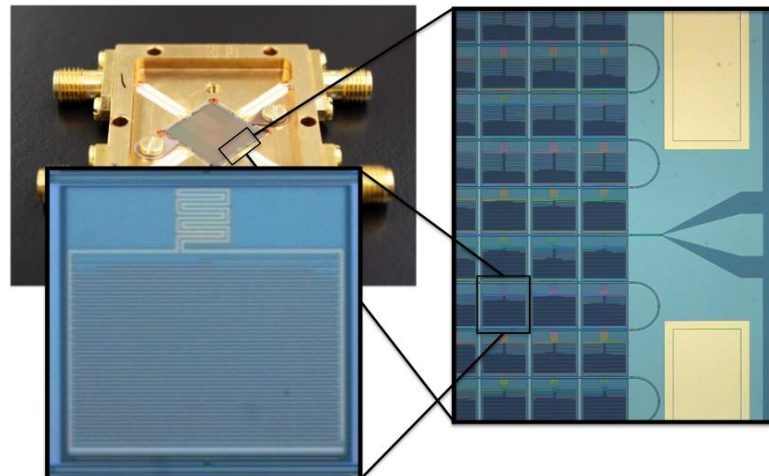


- A typical detector may have 2K MKIDs separated by 2MHz on a single RF line.

A frequency scan made by a VNA instrument of the S_{21} amplitude

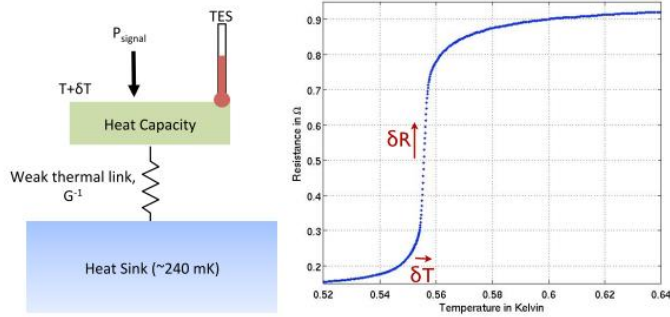


10,000 pixels
Mazin et al.
(UCSB)



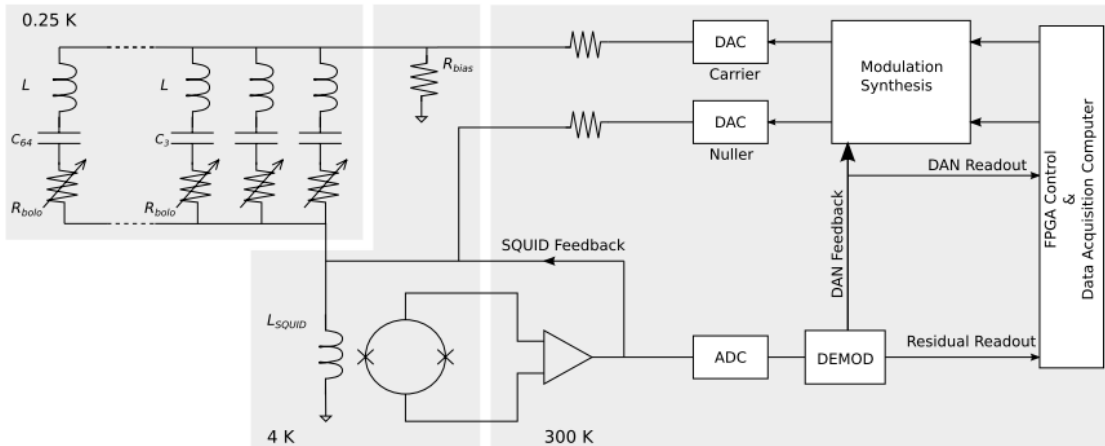
Different size capacitors define the resonant frequency

TES-SQUIDs



- TES is a bolometer coupled to a SQUID with feedback to keep it operating at the middle of the high sloped temperature transition.
 - They are low noise.
 - Used on the sky for many years and many telescopes.
 - Technology of choice for CMB-S4 (main frequencies).

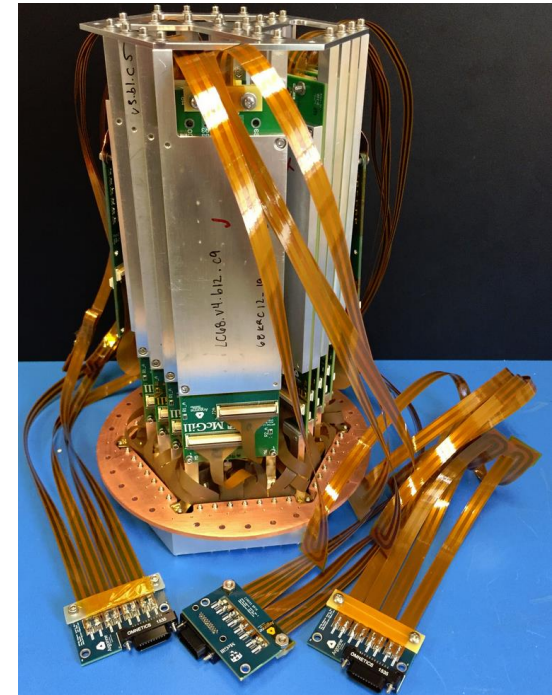
FDM architecture of SPT3



Issues:

TES couple to a very nonlinear device, the SQUID.

Very large arrays (>100K channels) pose a readout interconnection challenge.



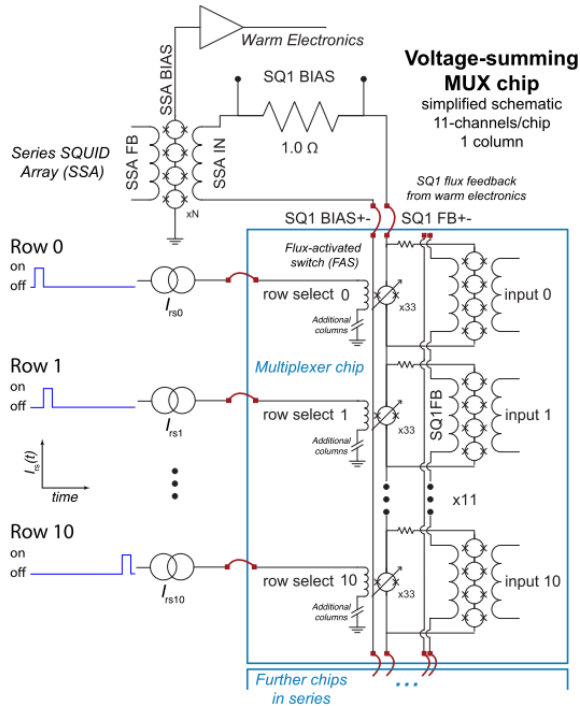
Experiments that need >100K superconducting detectors



- 3 Large Aperture Telescopes
- 18 Small Aperture Telescopes
- 2 sites: South pole, Atacama.
- 8 frequencies of observation.
- ~500,000 channels.
- Amazing science: Inflation, light relics, galaxy clusters, gamma-ray burst, neutrinos, &more.
- Down selection of detectors, readout, to be done in 2021.

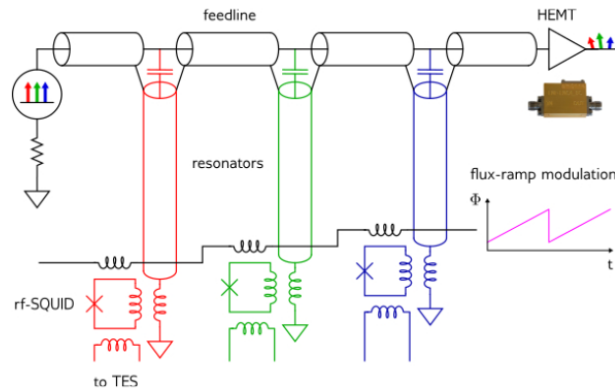
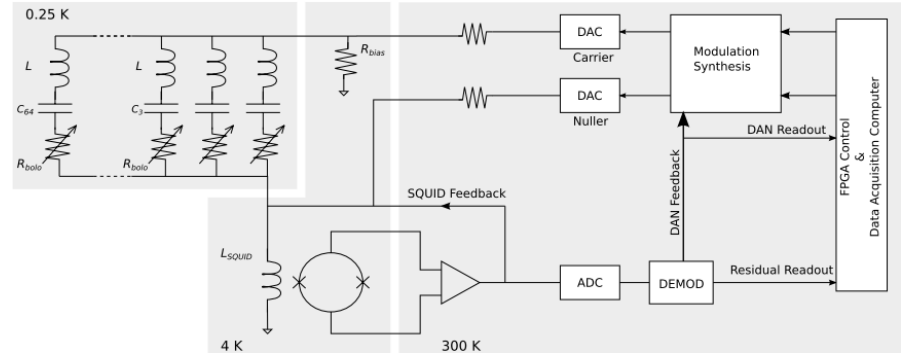
- Possible readout methods:
 - Time domain multiplexing.
 - Frequency domain multiplexing.
 - RF Frequency domain multiplexing.

TDM architecture of Biceps



- Readout alternatives:
 - FDM and TDM (time domain multiplexing): require a large number of interconnections with the warm electronics. Dedicated cold electronics and ASICs.
 - uMUX: solves the interconnection problem with the help of a dedicated RF MUX device (K. Irwin, B. Mates).

FDM architecture of SPT3

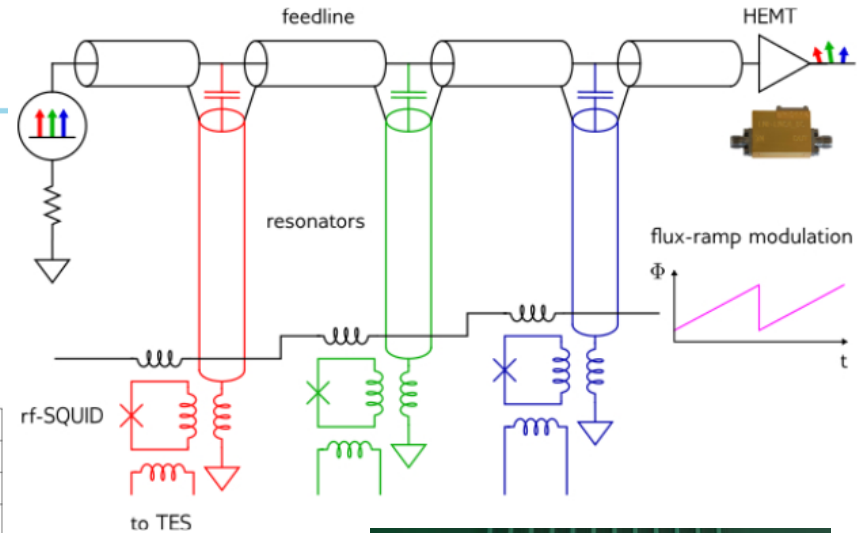
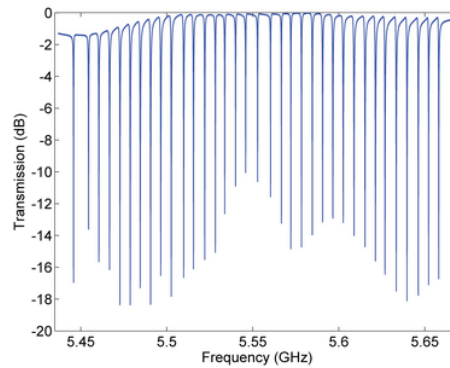
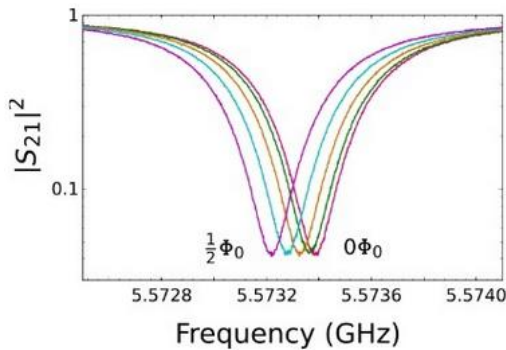


uMUX: RF multiplexed architecture: Simons Obs.

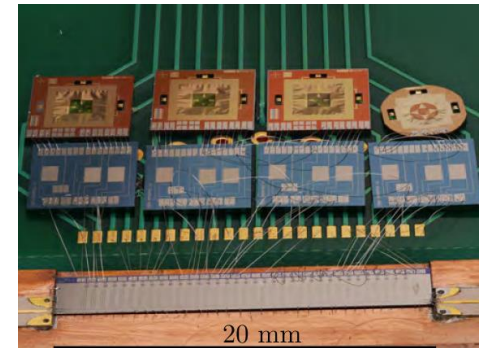
uMUX electronics

- The resonators remain at fixed frequency and can reach high channel multiplexing > 2000.
- 4GHz – 8GHz RF band OK.

Very similar to MKIDs S_{21}



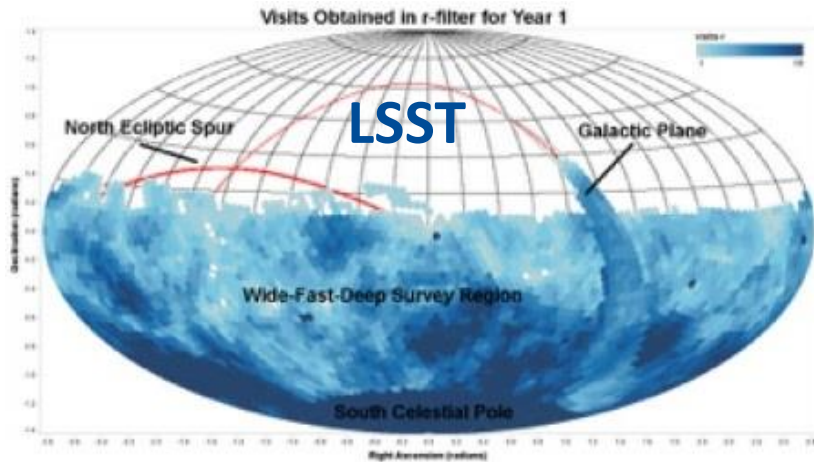
B. Mates, 2011 (NIST)



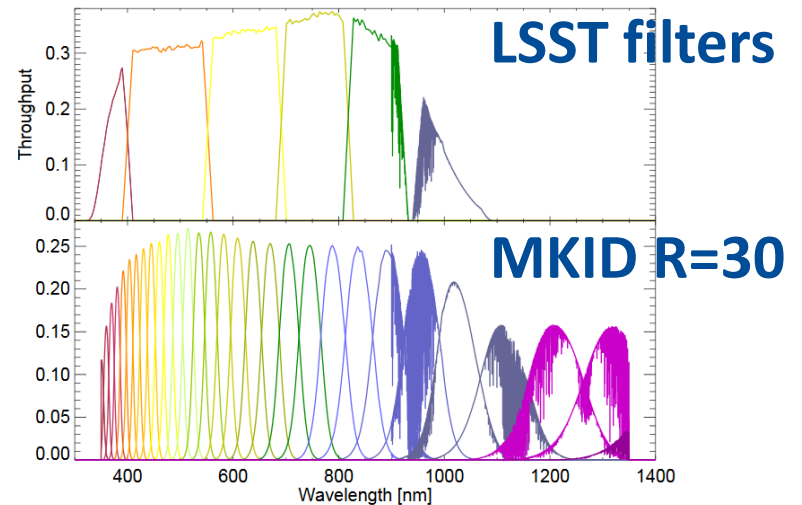
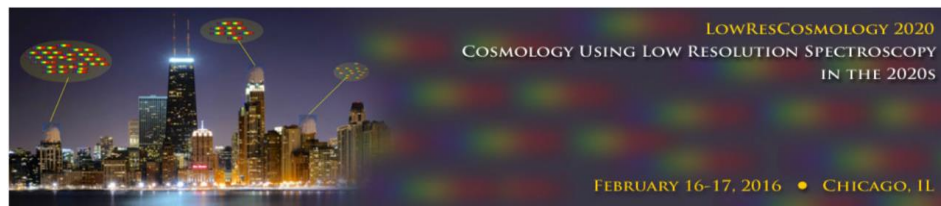
Still some problems remain:

- Channel frequency separation is non uniform. Some channels will overlap with neighbor channels and will be non usable.
- SQUID nonlinearity: requires a ramped flux bias to extract a “linearized” signal.
- Multiplexing 2000 channels (or even less) on a single RF line makes the LNA (HEMT) work near saturation and in the nonlinear region: Problem is solved by “tone tracking” where the S_{21} power of each tone is minimized to avoid overpowering the HEMT.

MKID as a tool for Dark Energy



- The DES and LSST will produce unprecedented imaging data. An MKIDs based low spectrograph surveying >1B galaxies reaching to NIR would be of great help.
- MKIDs have been recognized by P5 as a technology that could dramatically leverage investments.



Luis R. Abramo,¹ Rafael Arcos-Olalla,² Begoña Ascaso,³ Narciso Benitez,⁴ Francisco J Castander,⁵ Eduardo Cypriano,⁶ Ross Cawthon,^{6,7} Scott Dodelson,^{8,6,7} Olivier Dore,⁹ Renato A Dupke,^{10,11} Tim F Eifer,^{12,13} Martin Eriksen,¹⁴ Juan Estrada,⁸ Samuel Flender,¹⁵ Tommaso Giannantonio,¹⁶ Gaston Gutierrez,⁸ Gourav Khullar,⁶ James Lasker,^{6,7} Jeffrey Newman,¹⁷ Dan Scolnic,⁷ Marcelle Soares-Santos,⁸ Bjoern Soergel,¹⁶ Albert J. Stebbins,⁸ Chris Stoughton,⁸ and Guangtun Zhu¹⁸

¹Universidade de Sao Paulo

²Universidad de Guanajuato

³Universite Paris

⁴IAA-CSIC, Granada

⁵ICE, IEEC-CSIC, Barcelona

⁶Department of Astronomy & Astrophysics, University of Chicago, Chicago IL 60637

⁷Kavli Institute for Cosmological Physics, Enrico Fermi Institute, University of Chicago, Chicago, IL 60637

⁸Fermi National Accelerator Laboratory, Batavia, IL 60510-0500

⁹Jet Propulsion Laboratory, California Institute of Technology

¹⁰National Observatory, Rio de Janeiro

¹¹Univ. of Michigan, Ann Arbor

¹²Jet Propulsion Laboratory, Pasadena, CA

¹³California Institute of Technology, Pasadena, CA

¹⁴Leiden University

¹⁵Argonne National Laboratory

¹⁶Institute of Astronomy & Kavli Institute for Cosmology, University of Cambridge, UK

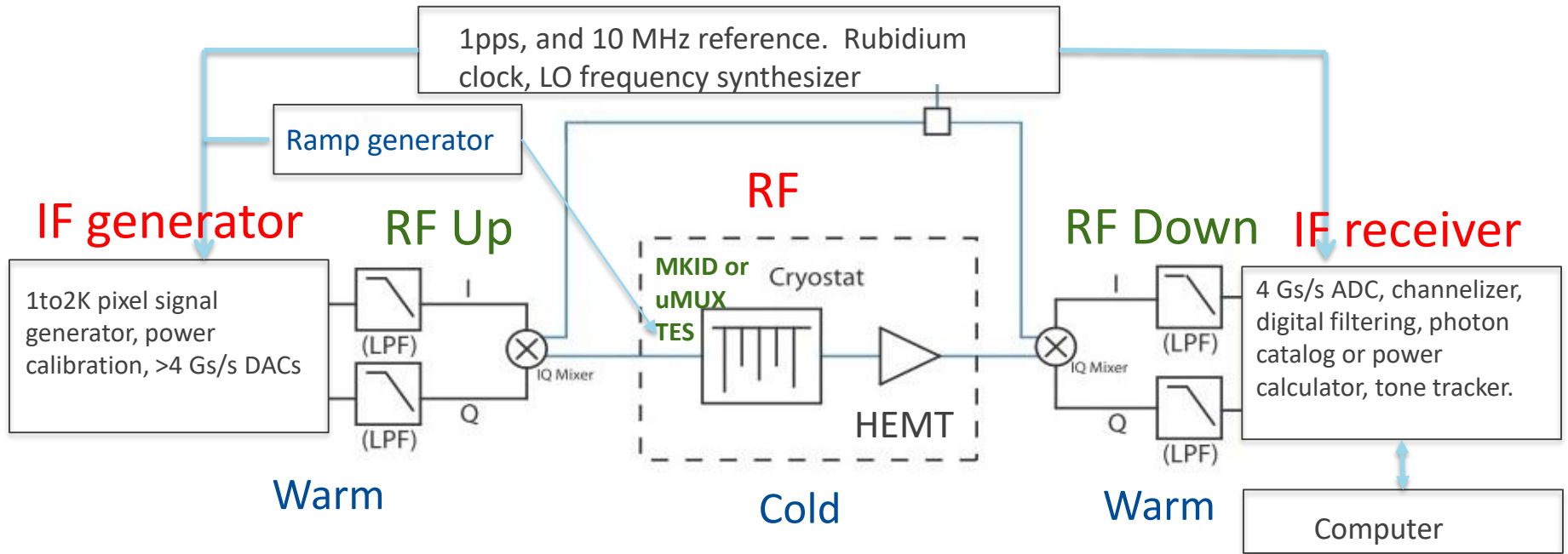
¹⁷University of Pittsburgh / PITT PACC

¹⁸Johns Hopkins University

Improvements in BAO FOM.

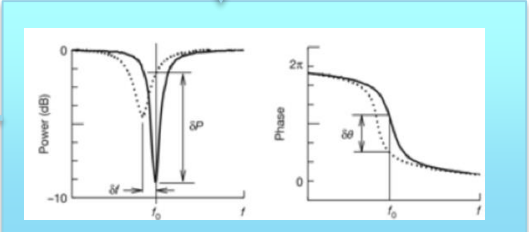
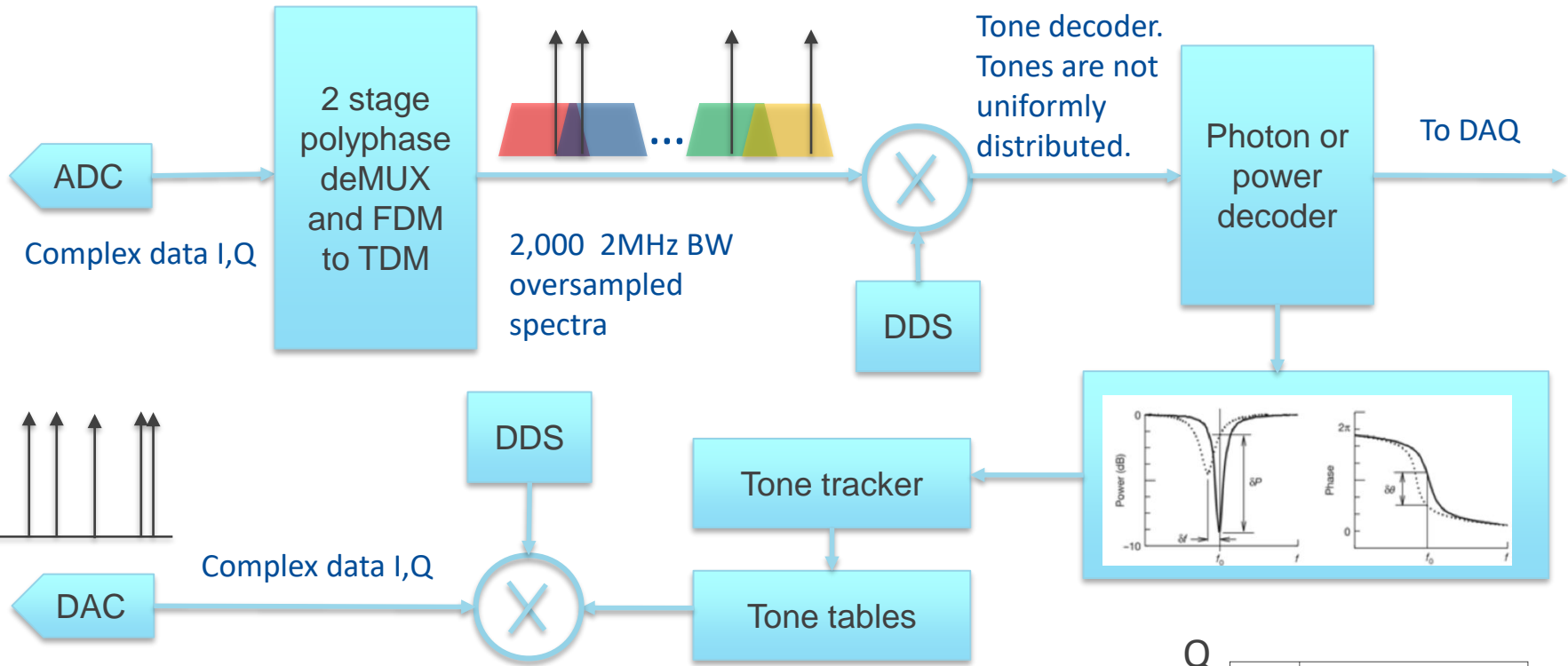
MKIDs detectors need to improve in spectral resolution and uniformity.

fMESSI DAQ block diagram



- The IF generator and receiver and reference timing are in the same box.
 - Sophisticated firmware in FPGA, plus full DAQ.
- RF Up/down converters and LO generator share same pc board.
- The RF electronics need to have flexible input and output powers ($\pm 60\text{dB}$) to accommodate 1 to 2000 tones.
 - The mixers should work at \sim constant power to avoid nonlinearities and bad mixing products.
 - The RF amplifiers should provide close to constant power across the 4-8GHz spectrum.
 - Tones should allow for individual power and phase rotation calibration.
 - Channel crosstalk should be minimized.

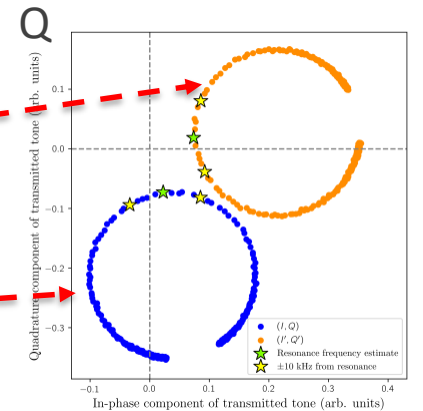
Firmware and software (RF part not shown)



- Software must find resonant frequencies, calibrate powers, rotate I, Q loops, calculate resonator parameters, program Tone tables or DDS.
- Take data and make images.

Rotated/calibrated resonator

Unrotated/uncalibrated resonator

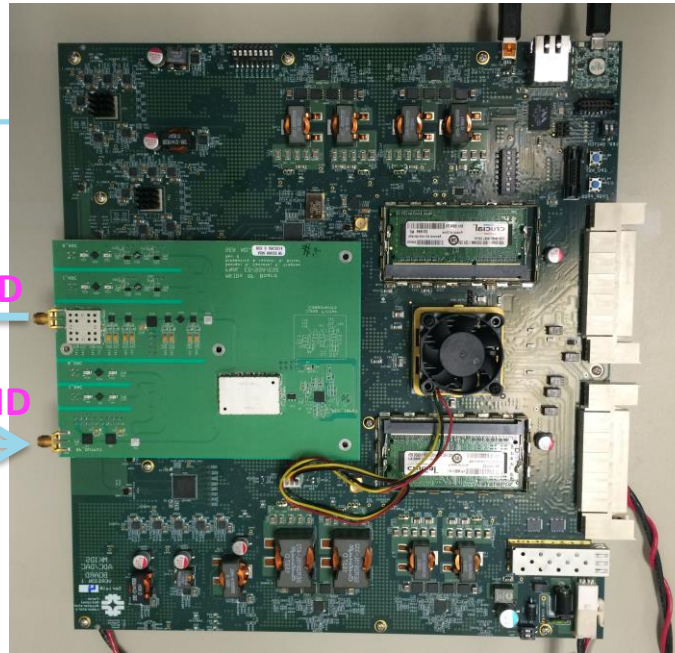


FMESSI DAQ

Designed by
Fermilab in
2016.

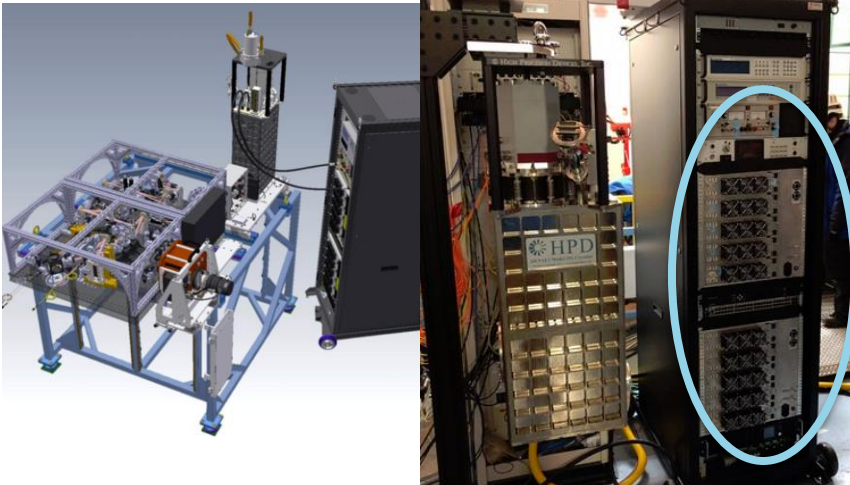
Production ~100
Still in use

To MKID
from MKID



- fMESSI main features:
- 1M photons/s per board.
- Noise: 12dB below HEMT.
- $<0.1\text{Hz}$ 1/f knee w/o compensation.
- 40 watts of power for the entire board running all ADCs and DACs at max sampling rate.
- Low cost: \$6/channel.

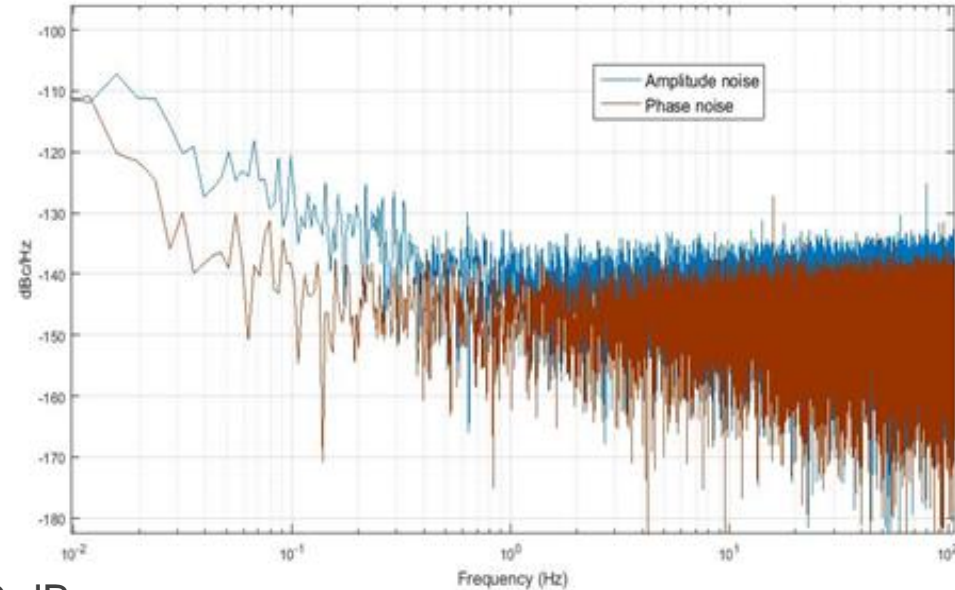
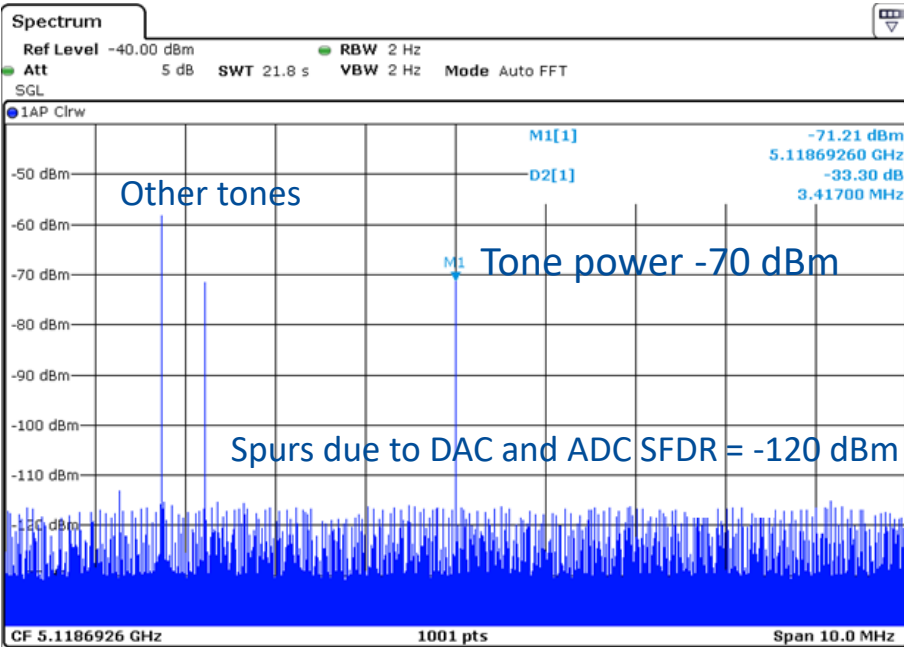
MEC: 20K MKID
system operational
at 8m telescope at
Subaru



DARKNESS 10Kpixels



fMESSI performance 1000 tones



- Noise floor from spurs of 1000 tones at -120 dBm.
- WG noise floor -145 dBm
- That is 2016, things get better now (next slides).

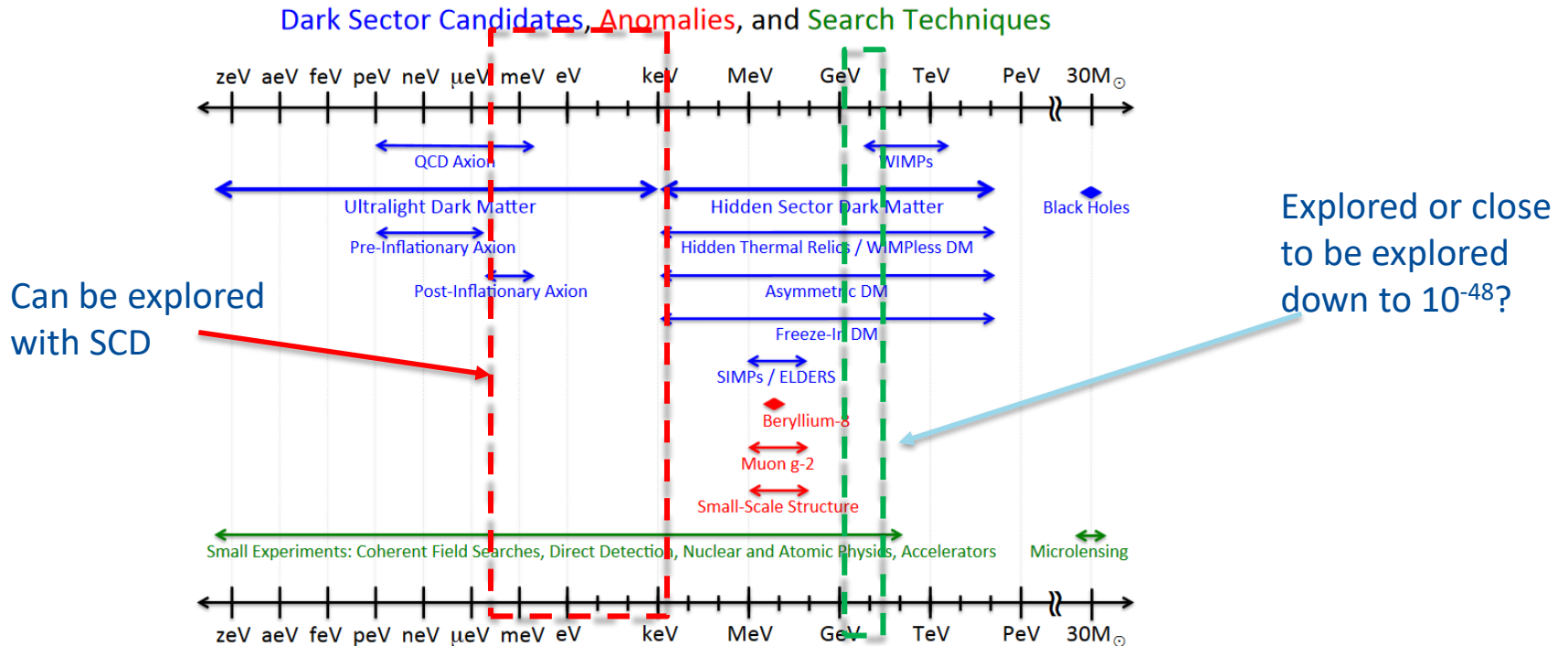
DARKNESS: A Microwave Kinetic Inductance Detector Integral Field Spectrograph for High-Contrast Astronomy

arxiv.org/abs/1803.10420 , DOI: [10.1088/1538-3873/aab5e7](https://doi.org/10.1088/1538-3873/aab5e7)

Also SMURF paper (SLAC) [arXiv:1809.03689](https://arxiv.org/abs/1809.03689)

Superconducting detector for light dark matter search

- Light Dark Matter: searching for low energy WIMPs and Axions.



Axion detectors can make use of superconducting technology developed for QIS. There is natural synergy between the readout and control electronics for MKIDs, uMUX TES and QIS with superconducting elements.

Superconducting devices for Quantum Computing

- Quantum computing uses quantum-mechanical phenomena such as superposition and entanglement to perform computation.
 - Promise to solve BQP (bounded error quantum probabilistic polynomial) problems exponentially faster than traditional computers.
 - Physical Quantum Computers have not yet achieved Quantum Supremacy.

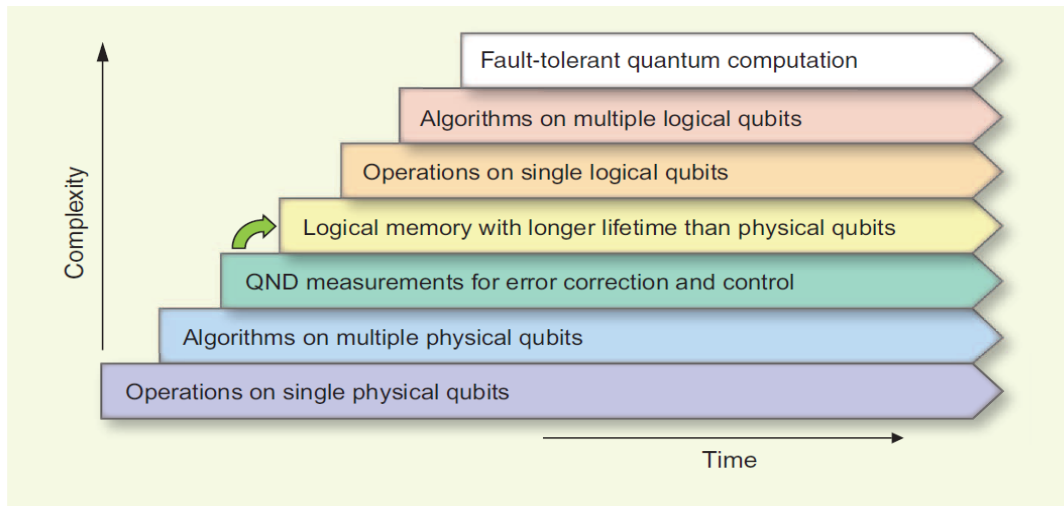


Fig. 1. Seven stages in the development of quantum information processing. Each advancement requires mastery of the preceding stages, but each also represents a continuing task that must be perfected in parallel with the others. Superconducting qubits are the only solid-state implementation at the third stage, and they now aim at reaching the fourth stage (green arrow). In the domain of atomic physics and quantum optics, the third stage had been previously attained by trapped ions and by Rydberg atoms. No implementation has yet reached the fourth stage, where a logical qubit can be stored, via error correction, for a time substantially longer than the decoherence time of its physical qubit components.

What is a qubit?

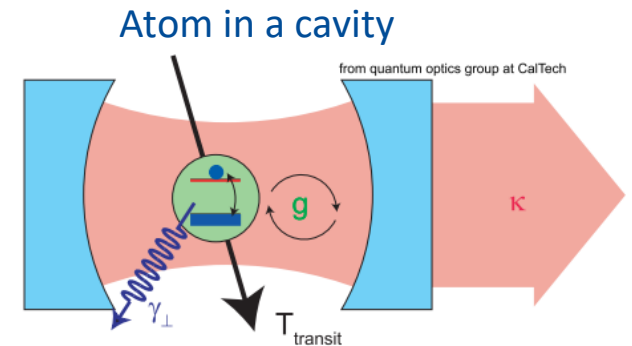
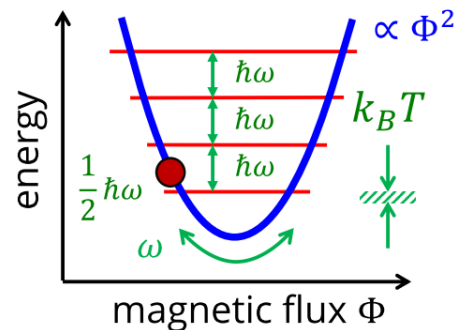
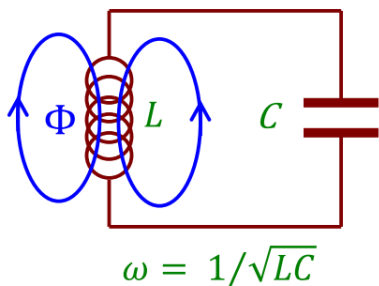
- A quantum mechanical quantum bit is the basic unit of quantum information.
- It is a quantum mechanical two-level system defined by the Jaynes-Cummings Hamiltonian in a 2-dimensional complex Hilbert space.

$$H_{JC} = \hbar\omega_r \left(a^\dagger a + \frac{1}{2} \right) + \hbar\omega_a \sigma_z + \hbar g_c (a^\dagger \sigma^- + a \sigma^+)$$

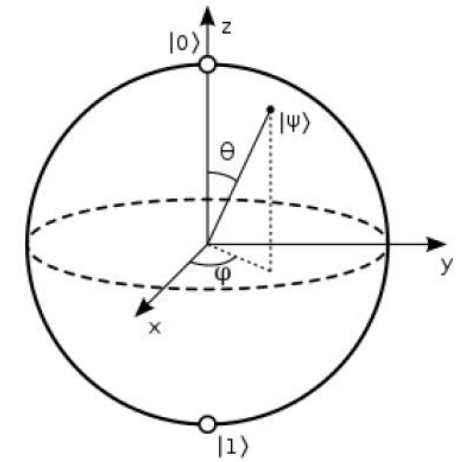
$$\varphi(x, t) = \cos\left(\frac{\theta}{2}\right) \varphi_0(x) + \sin\left(\frac{\theta}{2}\right) \varphi_1(x) e^{i\phi}$$

A general wave function state $|\varphi\rangle$ is in a superposition of the base states and collapses to $|0\rangle$ or $|1\rangle$ when measured.

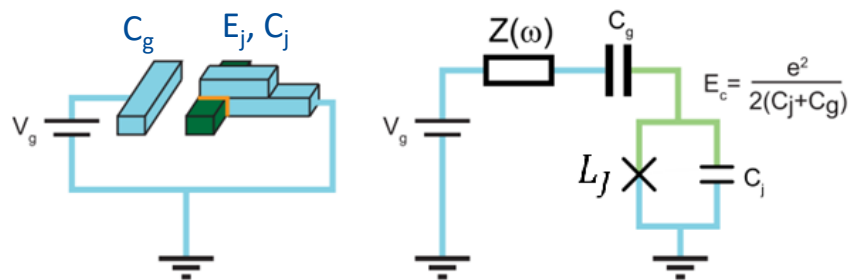
The harmonic oscillator does NOT work, we need an harmonic eigenenergies



The Bloch sphere

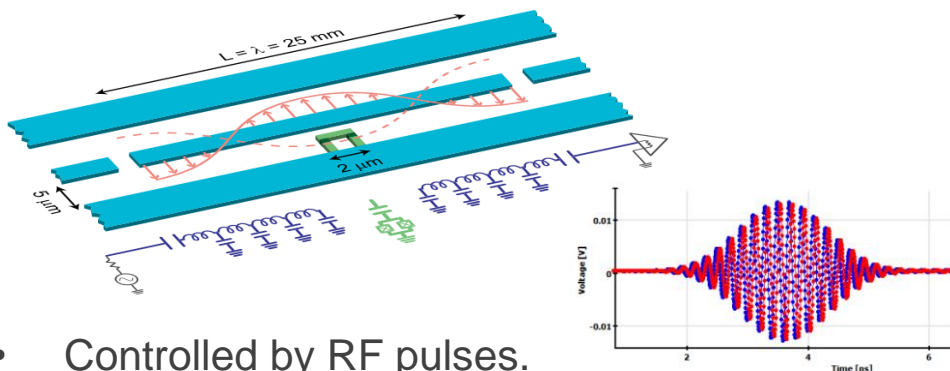
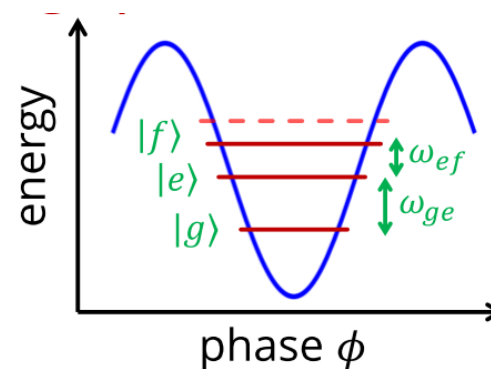


The Cooper pair box, 2D and 3D xmons: RF cavities and TJS

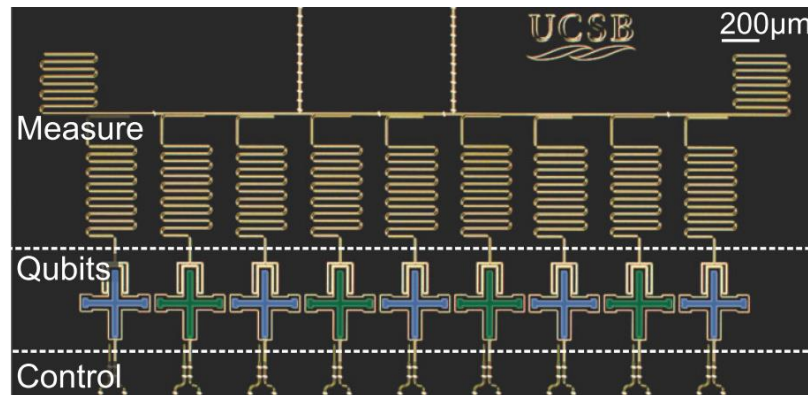


The oscillator's inductance has been replaced by a nonlinear inductance, a Josephson junction.

Now the energy levels are not equally spaced. We can control that device using different RF frequencies.



- Controlled by RF pulses.
- Multiplexed readout in 4-8 GHz like MKIDs.



J. Martinis (Google, UCSB), A. Cleland (UC)

3D qubits couple xmons with 3D RF cavities

Observation of High Coherence in Josephson Junction Qubits Measured in a Three-Dimensional Circuit QED Architecture

Hanhee Paik,¹ D. I. Schuster,^{1,2} Lev S. Bishop,^{1,3} G. Kirchmair,¹ G. Catelani,¹ A. P. Sears,¹ B. R. Johnson,^{1,4} M. J. Reagor,¹ L. Frunzio,¹ L. I. Glazman,¹ S. M. Girvin,¹ M. H. Devoret,¹ and R. J. Schoelkopf¹

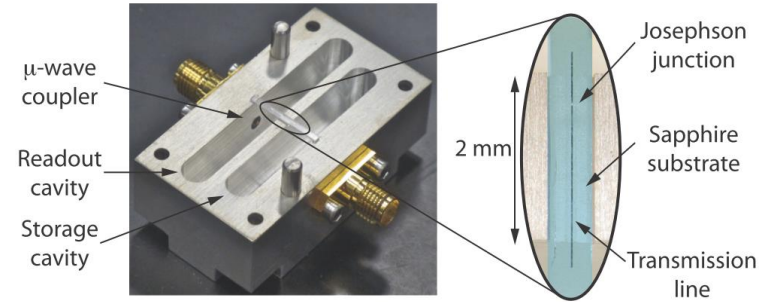
¹Department of Physics and Applied Physics, Yale University, New Haven, Connecticut 06520, USA

²Department of Physics and James Franck Institute, University of Chicago, Chicago, Illinois 60637, USA

³Joint Quantum Institute and Condensed Matter Theory Center, Department of Physics, University of Maryland, College Park, Maryland 20742, USA

⁴Raytheon BBN Technologies, Cambridge, Massachusetts 02138, USA

(Received 3 July 2011; revised manuscript received 15 September 2011; published 5 December 2011)

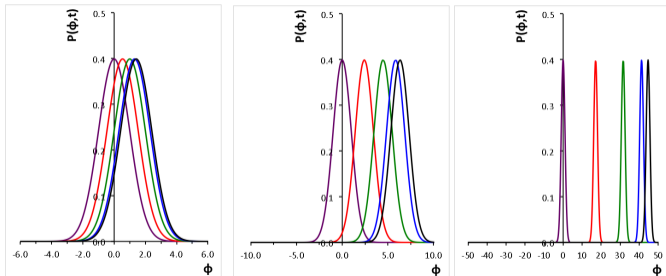


Three-dimensional superconducting resonators at $T < 20$ mK with the photon lifetime up to $\tau = 2$ seconds*

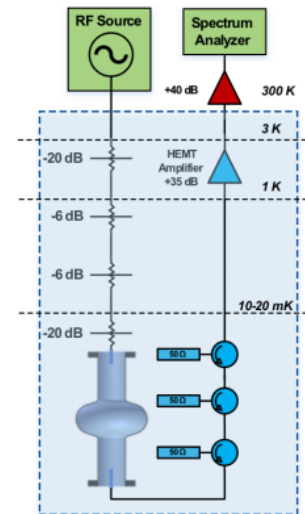
A. Romanenko,¹ R. Pilipenko, S. Zorzetti, D. Frolov, M. Awida, S. Posen, and A. Grassellino
 Fermi National Accelerator Laboratory, Batavia, IL 60510, USA
 (Dated: October 10, 2018)

Cavity State Manipulation Using Photon-Number Selective Phase Gates

Reinier W. Heeres, Brian Vlastakis, Eric Holland, Stefan Krastanov,
 Victor V. Albert, Luigi Frunzio, Liang Jiang, and Robert J. Schoelkopf
 Departments of Physics and Applied Physics, Yale University, New Haven, Connecticut 06520, USA
 (Dated: March 4, 2015)



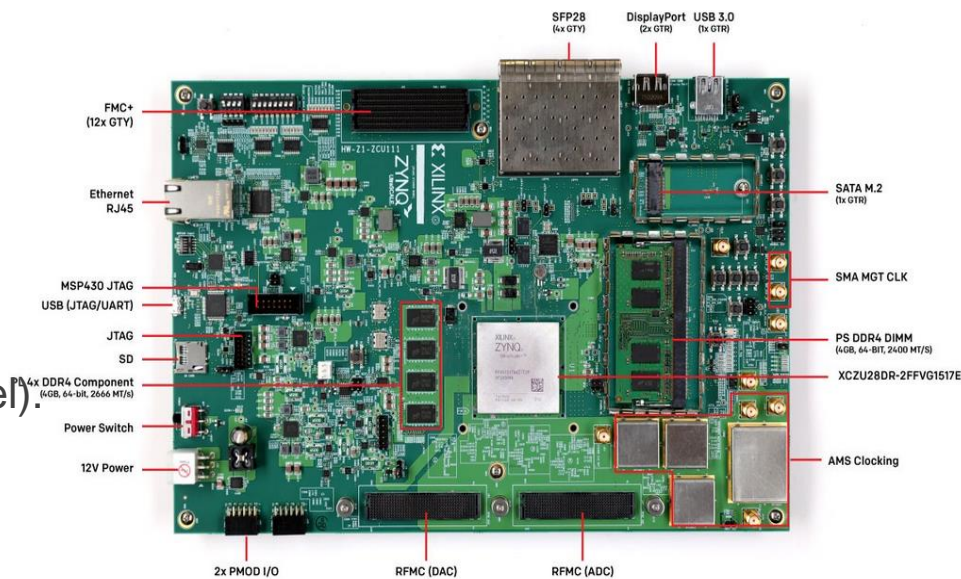
HEP is already using all this technology available. Readout and controls is key



How are we moving forward with multiplexed readout and controls?

- Commercial technology has made a quantum leap in FPGA and analog ADC/DAC integration. Xilinx ZU series integrates:
 - 8 ADCs at 5Gs/s
 - 8 DACs at 10 Gs/s
 - Huge FPGA logic
 - 4 core ARM processor (ready for Linux)
 - lots of peripherals.
- Almost all we need is in a single chip!
- A single FPGA can control and readout up to 8000 channels (2MHz separation/channel).
- Advantages:
 - Cost: \$1/channel.
 - Power: <10mW/channel.
 - Multiplexing factor and DAQ size reduction.

Xilinx ZCU111 evaluation board



How are we moving forward with multiplexed readout and controls?

- The priority is the firmware and the software.
 - Complex FPGAs require sophisticated firmware techniques from firmware experts.
 - Mathematical functions from theory need to be translated into efficient algorithms for FPGAs.
 - Algorithms must use FPGA resources efficiently at >500Mhz of clock.
 - Latency is important in qubit control.
- Hardware developments:
 - For now we use the ZCU111. Projects such as CMB-S4 will require new hardware designs.
 - A multichannel IF/RF hardware is needed.

Fermilab has started a by weekly collaboration meeting with US labs and Univ. You are welcome to join: Every other Tuesday 11am Central time (please email me at cancelo@fnal.gov to be added to the list) <https://fnal.zoom.us/j/470735194>

Summary

- Superconducting detectors are being chosen for large experiments approaching millions of channels.
- Although some electronics is moving to <1K and 4K, it is expected that the rapid development of commercial electronics for 300K will lead the way to readout and Control implementations.
- New experiments, in particular for cosmology, will continue to request new hardware firmware and software developments for Readout and Controls.
- Multiplexed Readout and Controls will lead the way for these large size experiments.