

SCOTT KRAVITZ, PETER SORENSEN, SCOTT HASELSCHWARDT, MATTHEW FORMAN LAWRENCE BERKELEY NATIONAL LAB

> CPAD WORKSHOP, DEC 8 2019

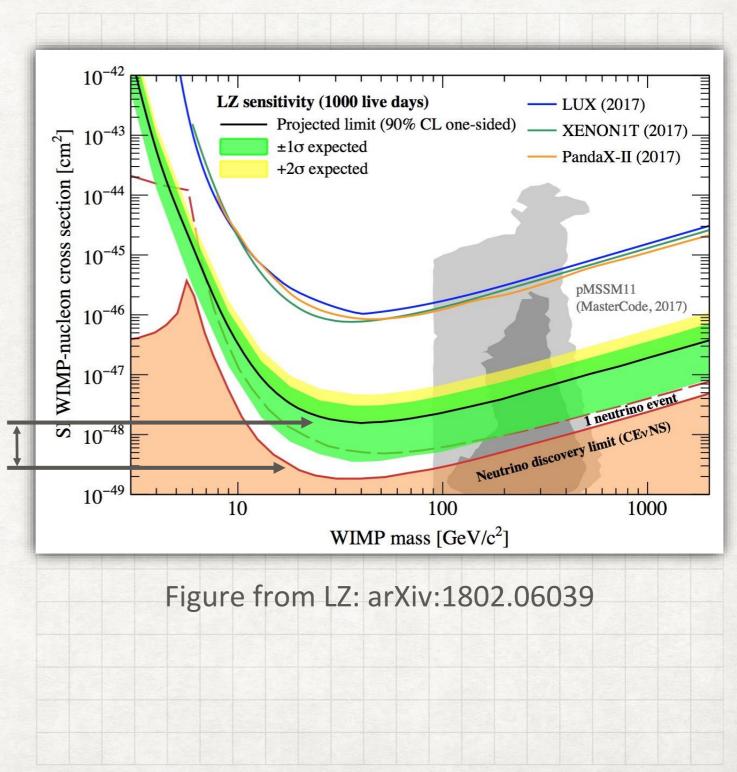
Where will dark matter direct detection be in 2025?

• Either:

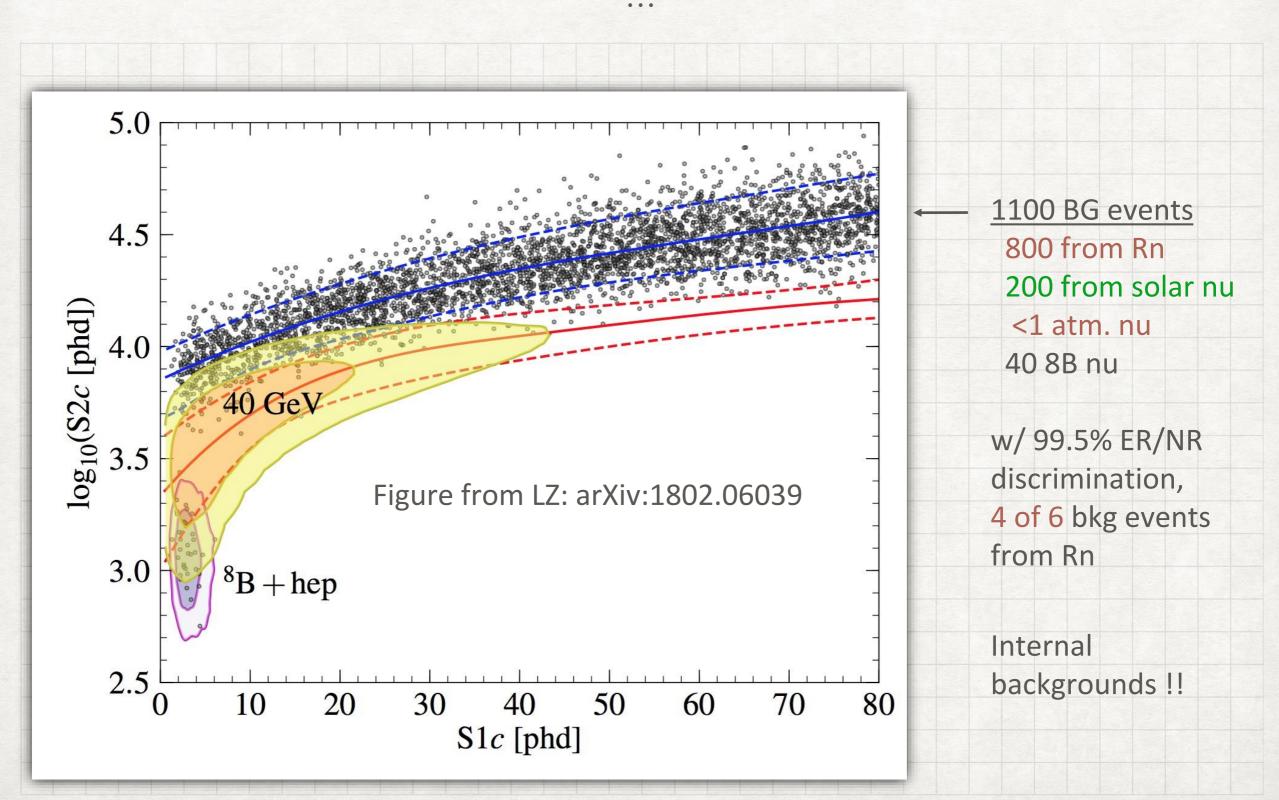
- no evidence for WIMPs
- a hint of WIMPs
- a handful of WIMPs

But must reach the neutrino floor: 1

- A G3 (i.e. 20+ tonne) xenon TPC would require a lot of:
 - time (10+ years)
 - money (\$200+ M)
- Not obvious this is the best path

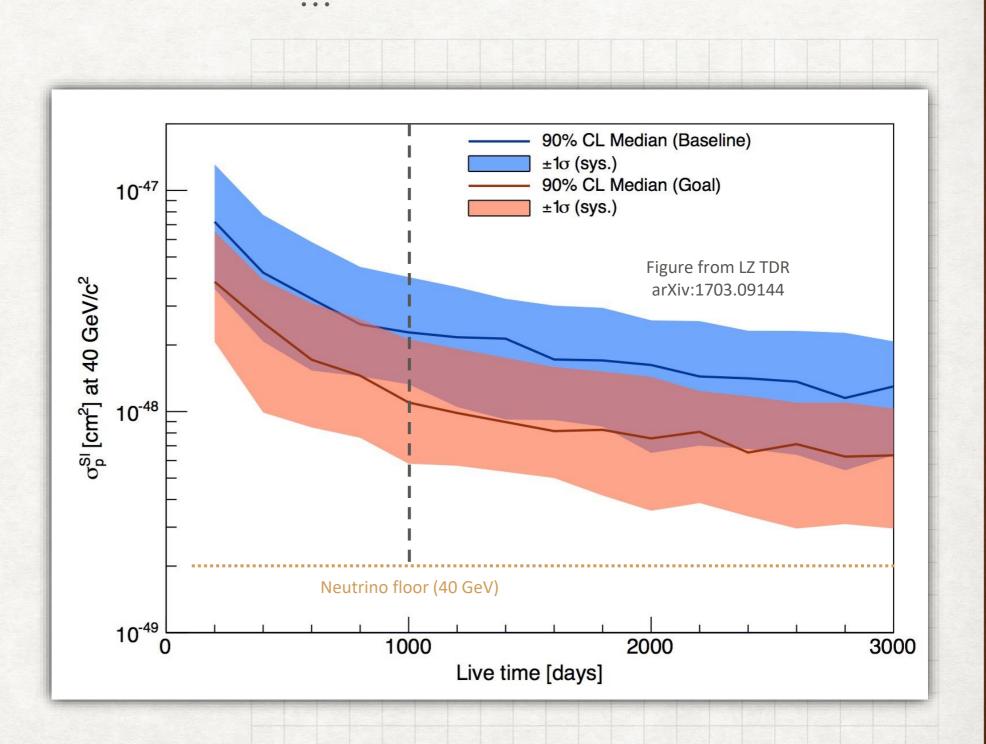


Limitations from LZ backgrounds



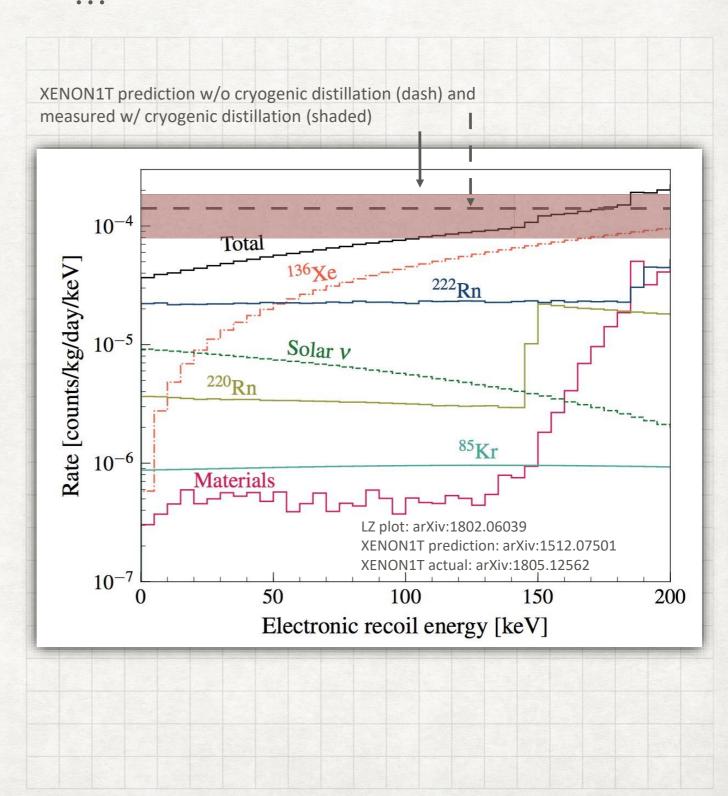
Run LZ for longer?

Doesn't work.
Backgrounds win,
mostly radon



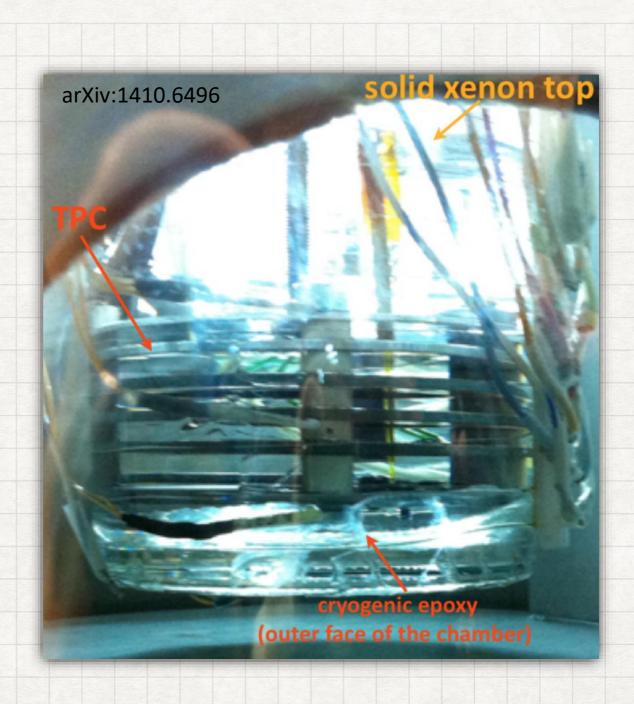
Get better at radon reduction?

Active area of R&D. HARD.



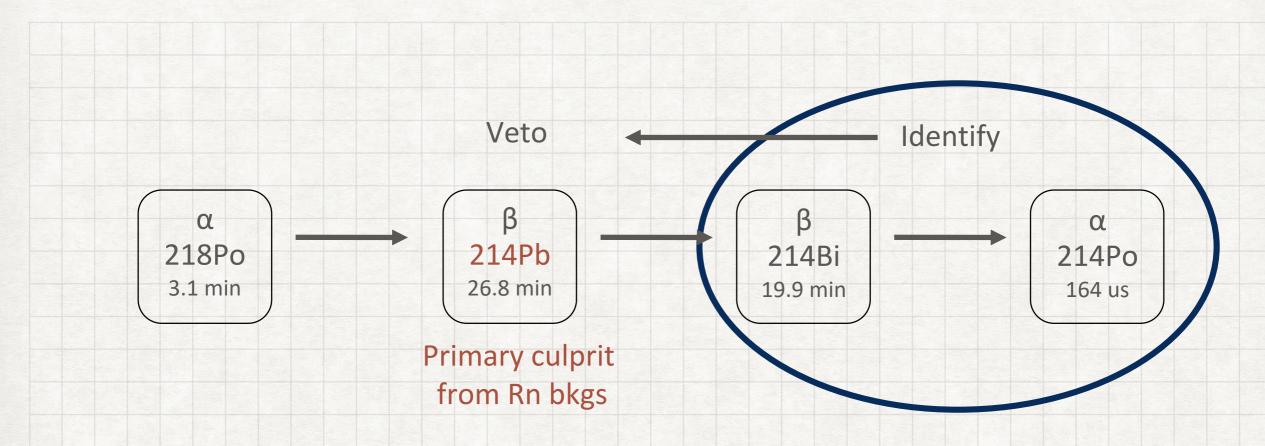
Solution: CrystaLiZe

- Freeze LZ:
 Radon emanated from surfaces
 now excluded from solid bulk
- In a crystal, radon decay daughters would stay at the same (x,y,z) as the parent, enabling tagging/veto
- DAMIC has used this trick in Si arXiv:1506.02562



Rn daughter tagging

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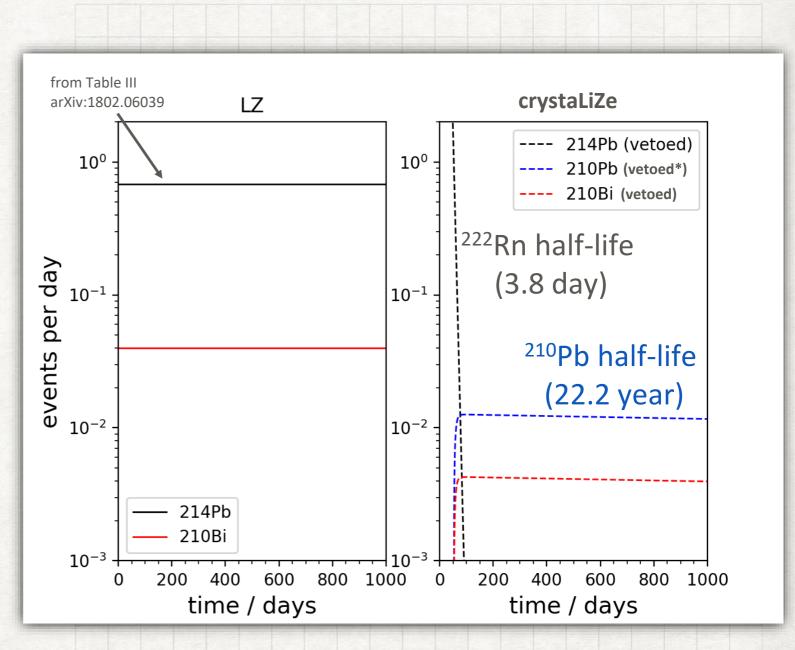


This is an easy sequence to tag if all the decays happen in exactly the same spot in a low-background instrument

(similar situation for 220Rn chain)

Radon tagging by the numbers

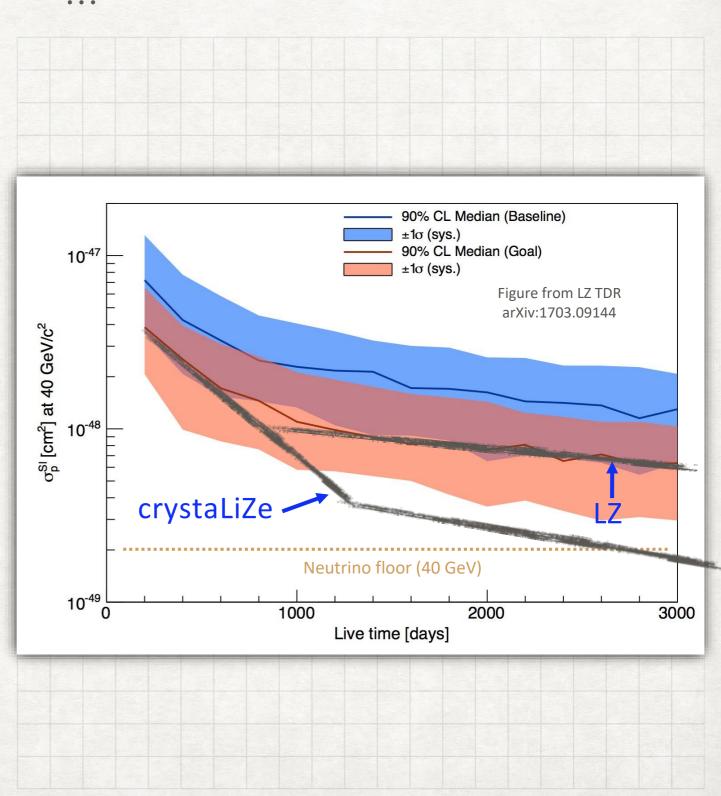
- In LZ, Rn emanates into the liquid and flows in to the bulk
- In crystaLiZe, Rn would emanate into the bulk for O(100) days while the crystal is growing; then, Rn in bulk target would be fixed and decay away



same LZ emanation and dust assumptions

Two slopes towards discovery

a crystalline xenon TPC
 ("crystaLiZe") would progress
 along the first slope for longer



crystaLiZe looks technically feasible

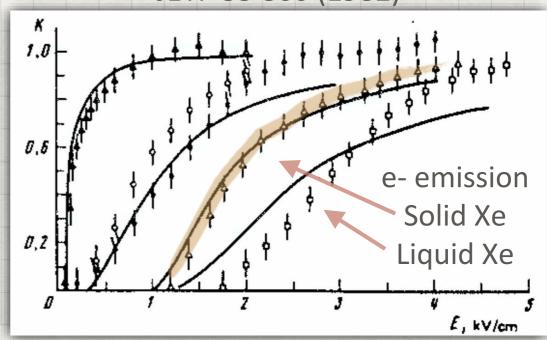
- Solid and liquid xenon have similar physical properties
- Solid xenon emission TPC is expected to perform as well as a liquid xenon emission TPC
 - band gap (hence W-value)
 - electron mobility (doubled)
 - electron emission
 - density (20% bonus!)
 - high voltage **
- Similar scintillation signal observed in solid and liquid
- cf. arXiv:1410.6496 and arXiv:1508.05903

Phys Rev B 10 4464 (1974)

TABLE II. Comparison of transport parameters in solid and liquid xenon. Values of other data used in the calculations are also quoted.

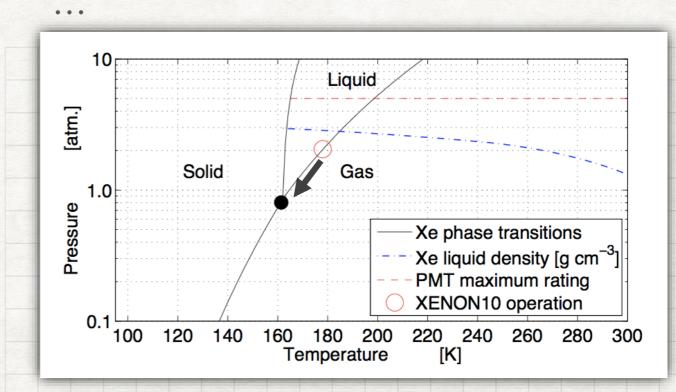
	Solid T = 161.2 °K	Liquid T = 163 °K	Unit		
E_G	9,272	9,22	eV		
E_G	1.063	1.084	eV		
ϵ_{∞}	2.00 a	1.85 b	• • •		
m* μ	0.31^{c} $4.5 \times 10^{3} d$	0.27 $2.2 \times 10^{3} e$	electron mass cm ² V ⁻¹ sec ⁻¹		
$\stackrel{ au_p}{L}$	8.0×10^{-13} 7.1×10^{-6}	3.4×10^{-13} 3.3×10^{-6}	sec cm		
β	1.36×10 ^{10 f}	0.58×10^{10} g	dyn/cm ²		
a	3.8×10^{-9}	4.2×10^{-9}	cm		
$ E_{1CB} $	0.93	1.01	eV		

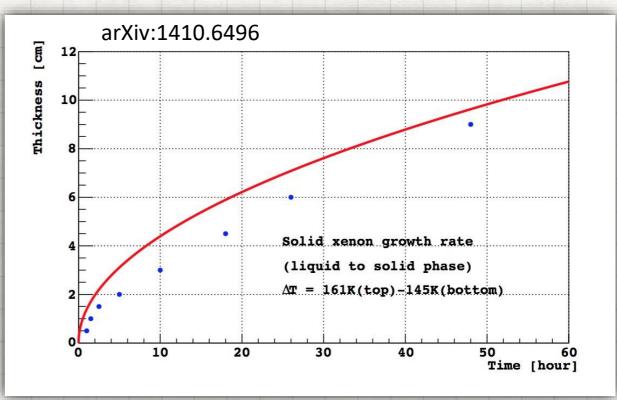
JETP 55 860 (1982)



Technical Challenges

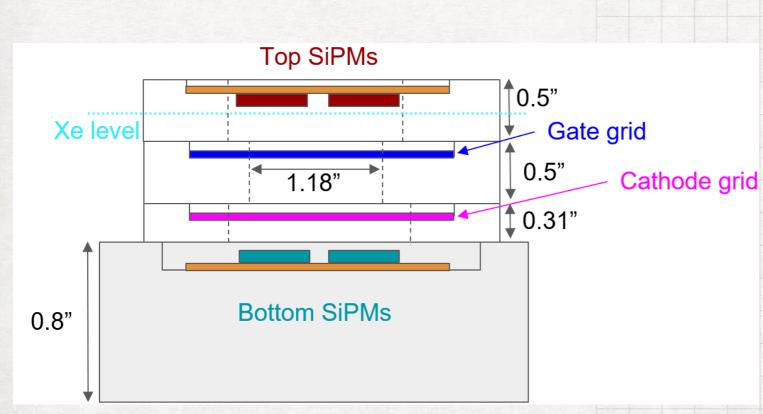
- Single e- sensitivity?
- Retaining high purity while crystallizing
 - likely requiring elevated temperature bakeout
 - would take multiple months to crystallize LZ w/o defects
- Precise temperature gradients require more elaborate control/measurement of T
- R&D: use small scale crystalline Xe TPC test bed to gauge performance





Test Bed Design

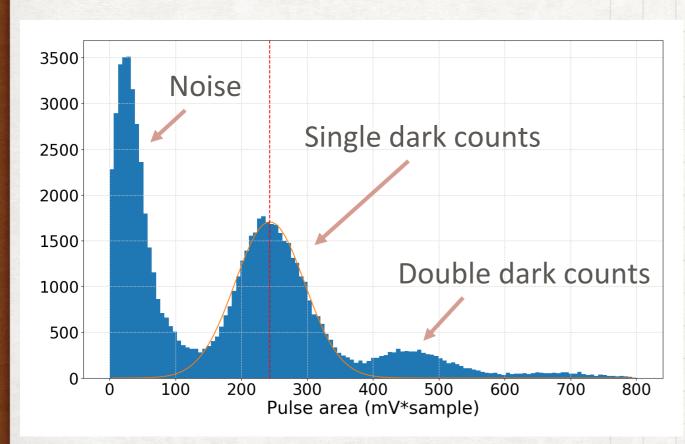
- Two phase Xe mini-TPC
- ~500 g Xe when full
- S1 and S2 readout:
 SiPM arrays (4 top, 4 bottom)





Why SiPMs?

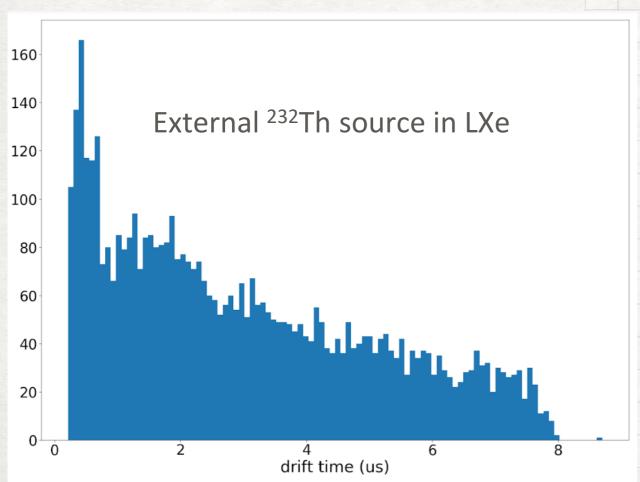
- Compact less Xe needed
- No vacuum space no structural concerns, esp. during freezing
- QE extends to deeper-UV light that may be present in ice

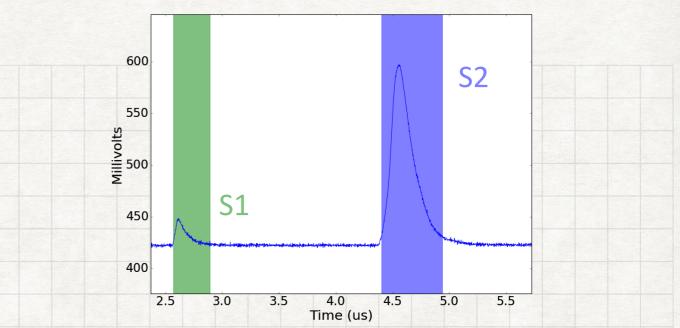


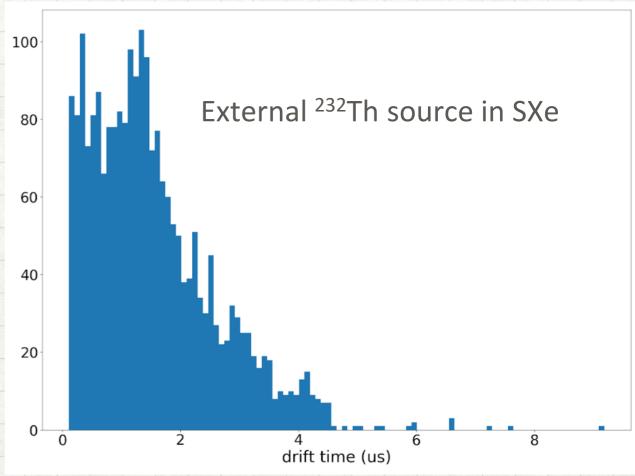


TPC Tests

- Observe S1s and S2s in Xe
- Clear indications of freezing:
 - Vapor pressure below triple point
 - Drift time halves







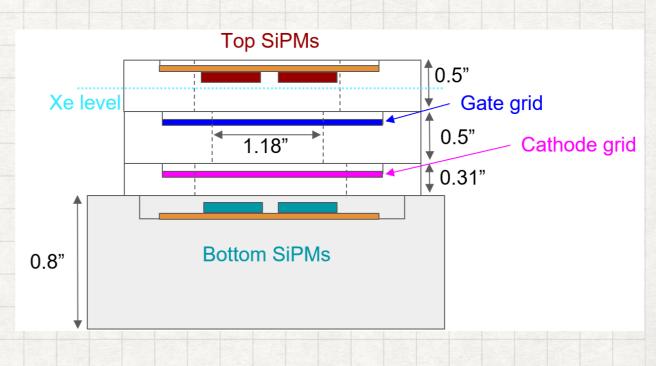
CRYSTA LIZE: SUMMARY



- Performance of large Xe detectors expected to be limited by Rn before reaching the neutrino floor
- CrystaLiZe is a logical upgrade path for LZ to reach the neutrino floor (+solar neutrino science), if it works
- Near future: establish if crystalline xenon emission TPC maintains all the benefits of the liquid xenon emission TPC

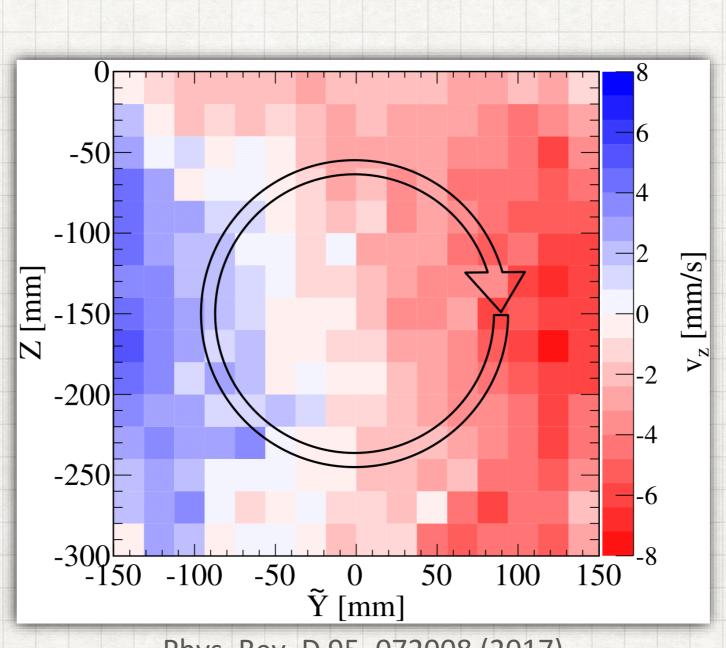
EXTRA SLIDES FOLLOW

Test Bed Design



LXe convection ~6 cm/min i.e. too fast





Phys. Rev. D 95, 072008 (2017)

Operation of a crystalline Xe TPC

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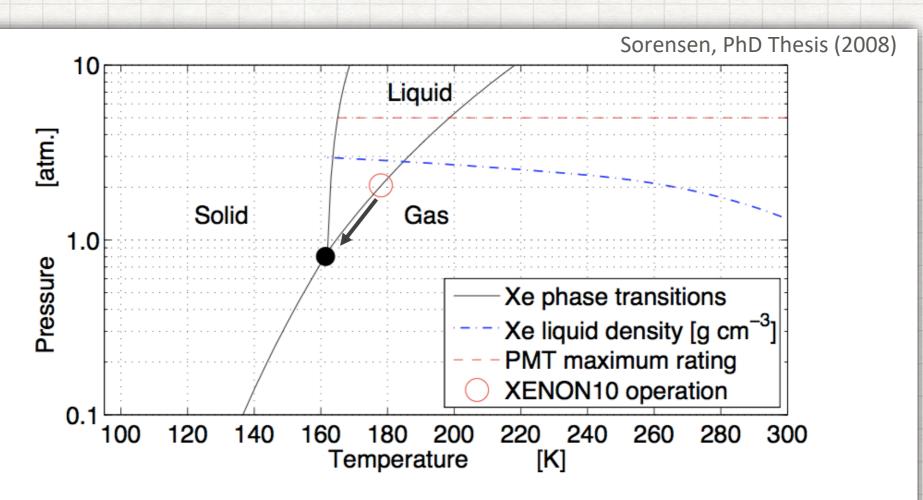


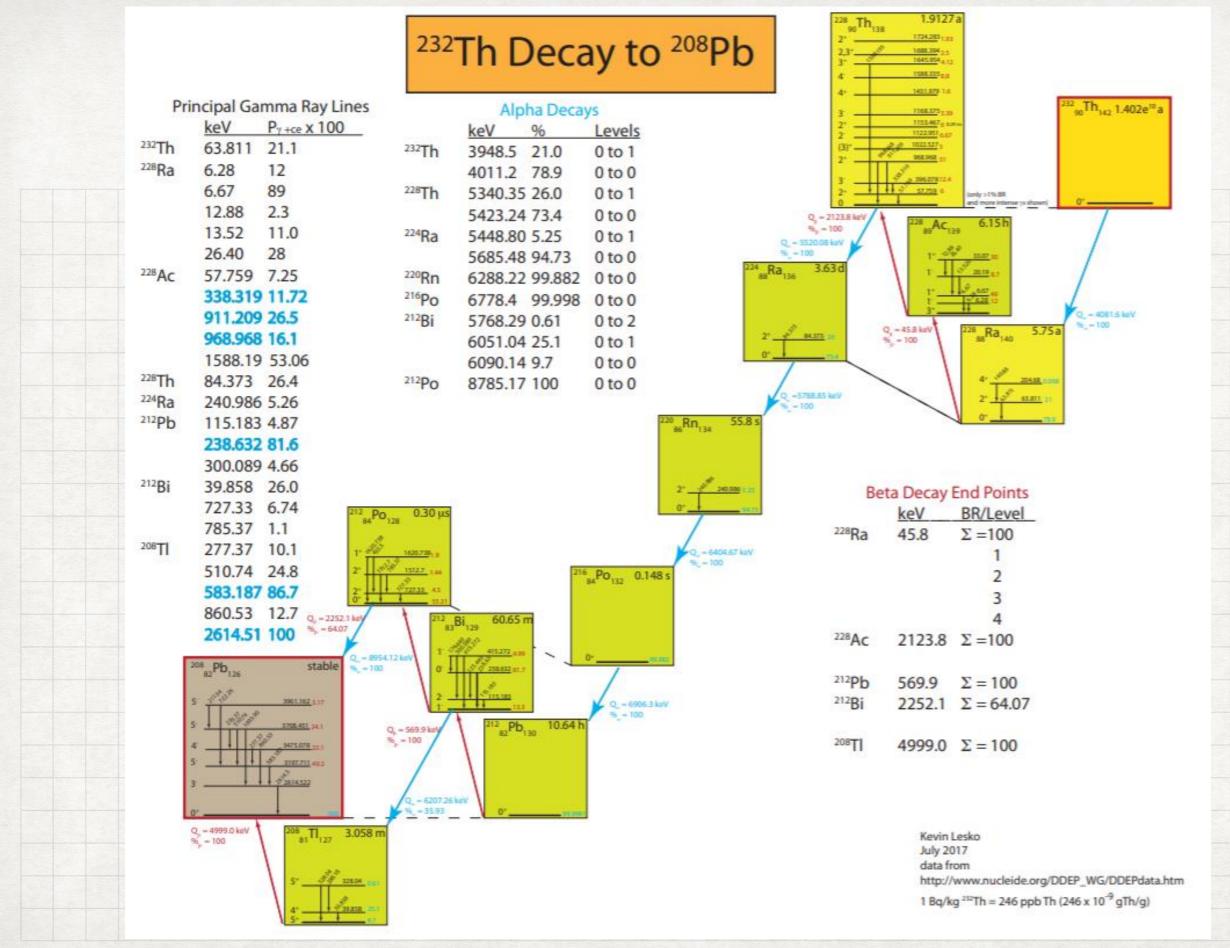
Figure 3.22: Phase Diagram for Xe (data from [127]) and liquid Xe density (data from [123]). For dual-phase operation as in XENON10, the comfortable operating window is indicated by the red circle. Higher pressures are excluded out of consideration for the PMTs (which are rated to a maximum of 5 atm, above which they will crush). Lower temperatures dictate a lower vapor pressure, which reduces the efficiency of the S2 proportional scintillation (Sec. 2.2.4). It is also necessary to stay safely above the 161 K, to exclude the possibility of sublimation.

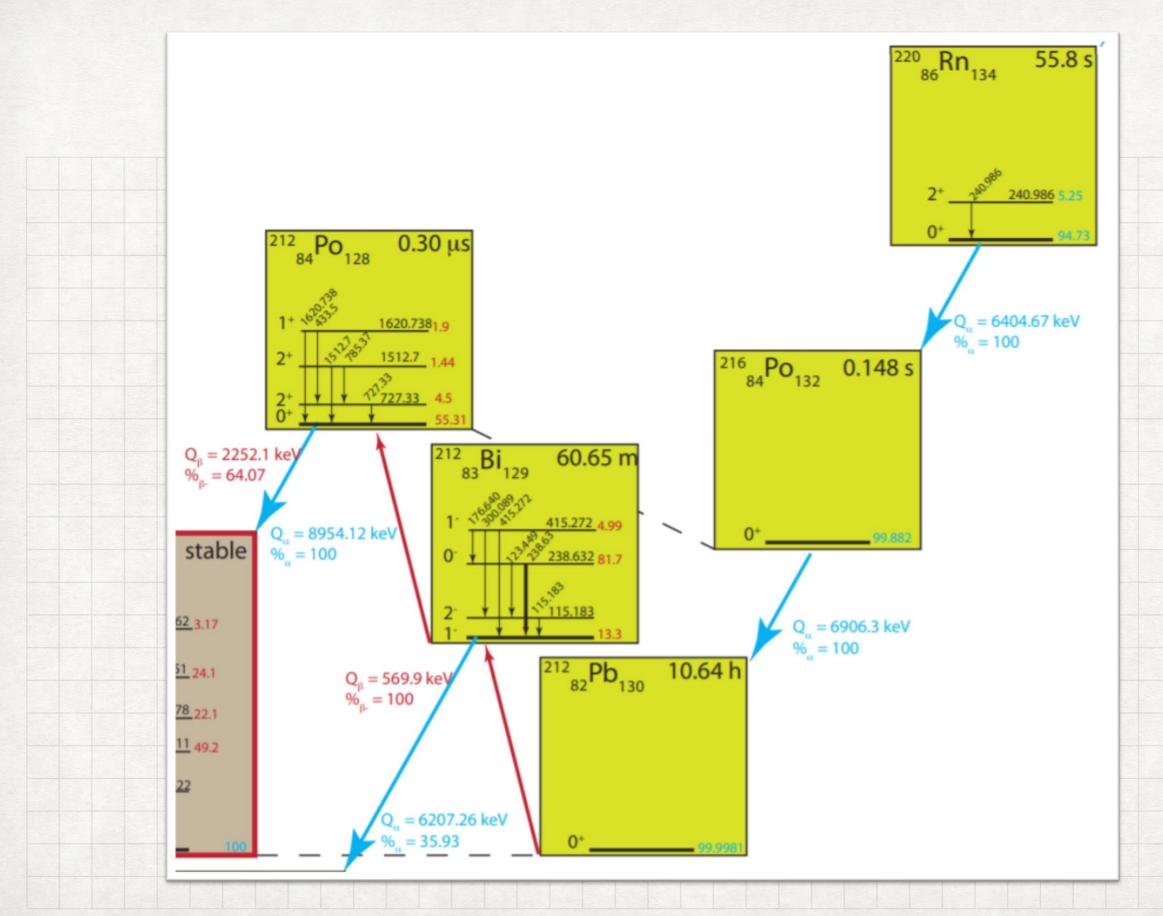
LZ Backgrounds

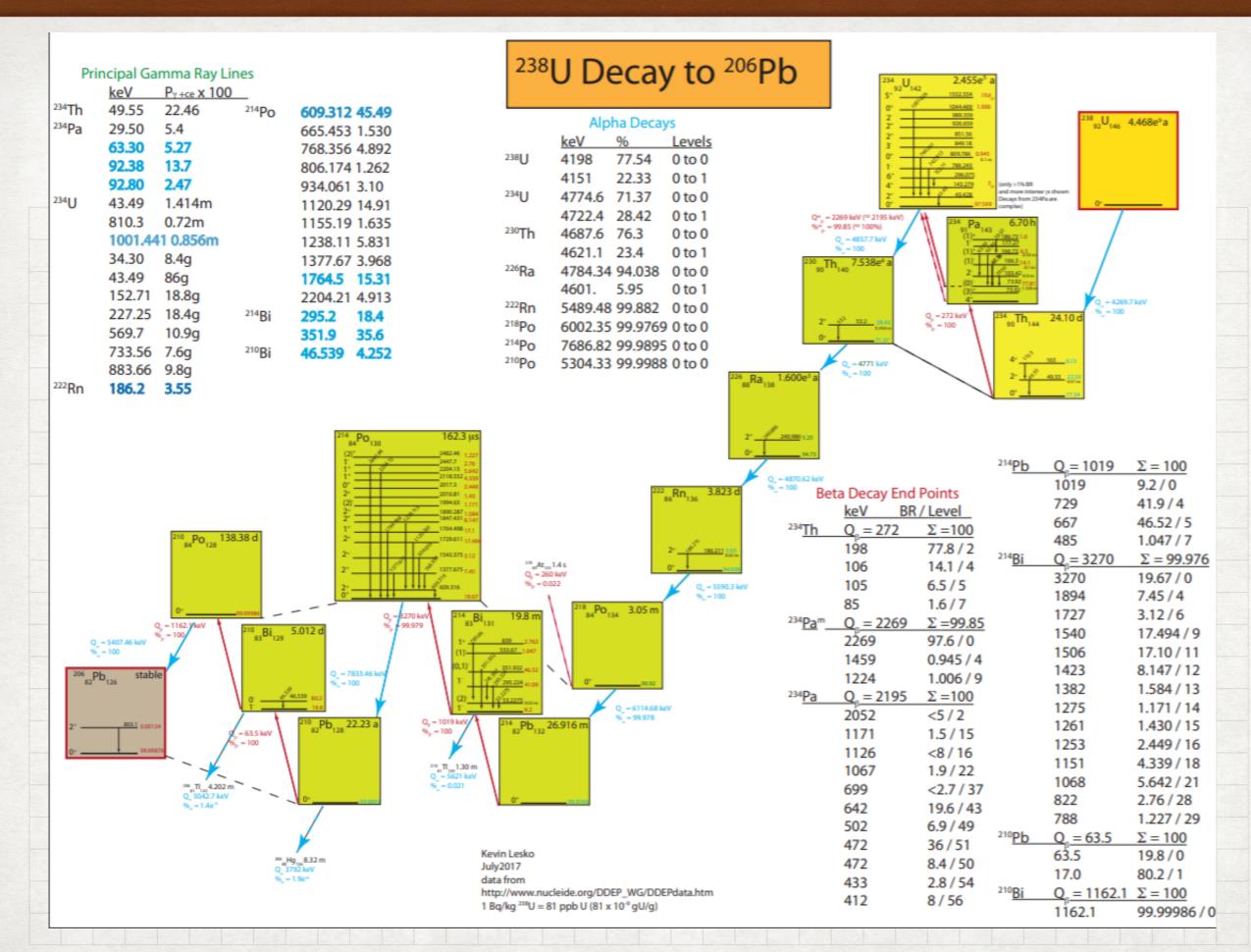
TABLE III. Estimated backgrounds from all significant sources in the LZ 1000 day WIMP search exposure. Counts are for a region of interest relevant to a $40~{\rm GeV/c^2}$ WIMP: approximately $1.5-6.5~{\rm keV}$ for ERs and $6-30~{\rm keV}$ for NRs; and after application of the single scatter, skin and OD veto, and $5.6~{\rm tonne}$ fiducial volume cuts. Mass-weighted average activities are shown for composite materials and the $^{238}{\rm U}$ and $^{232}{\rm Th}$ chains are split into contributions from early- and late-chain, with the latter defined as those coming from isotopes below and including $^{226}{\rm Ra}$ and $^{224}{\rm Ra}$, respectively.

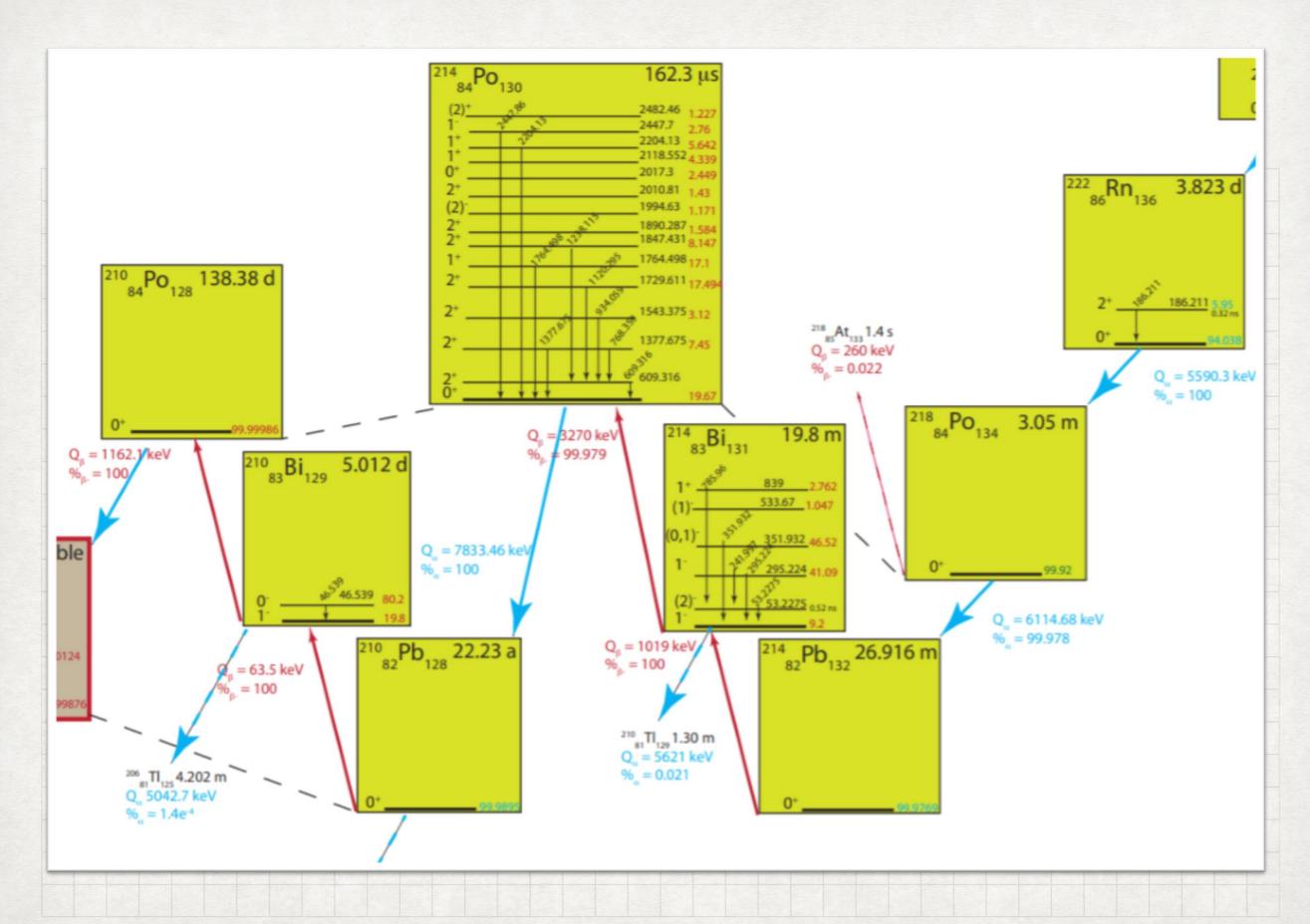
Background Source	Mass	$^{238}U_e$	$^{238}\mathrm{U}_{l}$	$^{232}\mathrm{Th}_e$	$^{232}\mathrm{Th}_{l}$	⁶⁰ Co	$^{40}\mathrm{K}$	n/yr	ER	NR
	(kg)	mBq/kg							(cts)	(cts)
Detector Components										
PMT systems	308	31.2	5.20	2.32	2.29	1.46	18.6	248	2.82	0.027
TPC systems	373	3.28	1.01	0.84	0.76	2.58	7.80	79.9	4.33	0.022
Cryostat	2778	2.88	0.63	0.48	0.51	0.31	2.62	323	1.27	0.018
Outer detector (OD)	22950	6.13	4.74	3.78	3.71	0.33	13.8	8061	0.62	0.001
All else	358	3.61	1.25	0.55	0.65	1.31	2.64	39.1	0.11	0.003
	subtotal								9	0.07
Surface Contamination	1									
Dust (intrinsic activity, 500 ng/cm ²)										0.05
Plate-out (PTFE panels, 50 nBq/cm ²)										0.05
²¹⁰ Bi mobility (0.1 μBq/kg LXe)										-
Ion misreconstruction (50 nBq/cm ²)										0.16
²¹⁰ Pb (in bulk PTFE, 10 mBq/kg PTFE)										0.12
	1/ 0	, ,					sı	ıbtotal	40	0.39
Xenon contaminants										
222 Rn (1.8 µBq/kg)										_
220 Rn (0.09 µBq/kg)										_
nat Kr (0.015 ppt g/g)										_
nat Ar (0.45 ppb g/g)										
111 (0.10 PPO 8/8/							SI	ıbtotal	2.5 819	0
Laboratory and Cosmo	ogenics									
Laboratory and Cosmogenics Laboratory rock walls										0.00
Muon induced neutrons									4.6	0.06
Cosmogenic activation										_
o consequent description							sı	ıbtotal	5	0.06
Physics										
136 Xe $2\nu\beta\beta$									67	_
Solar neutrinos: $pp+^{7}Be+^{13}N$, $^{8}B+hep$										0*
Diffuse supernova neutrinos (DSN)										0.05
Atmospheric neutrinos (Atm)										
•							sı	ıbtotal	258	0.51
Total									1131	1.03
Total (with 99.5% ER discrimination, 50% NR efficiency)									5.66	0.52
Sum of ER and NR in						ith all a	analysi	s cuts	6.	18
* Below the 6 keV NR thres	hold used	l here.								

arxiv:1802.06039









PREVIOUS WORK

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- 1. arXiv:1410.6496 and arXiv:1508.05903
 - optical photon collection comparison to liquid mild decrease
 - charge drift comparison to liquid factor x2 faster
- 2. solid xenon bolometers (2013, Drexel)

Old Paradigm



Phases of the XENON program

LUX



XENON10

XENON100

XENON1T / XENONnT



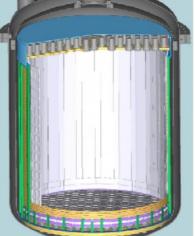
2005-2007 15 cm drift TPC – 25 kg

2008-2016 30 cm drift TPC - 161 kg

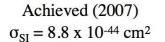




2013-2018 / 2019-2023

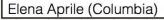


100 cm / 144 cm drift TPC - 3200 kg / ~8000 kg



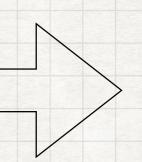
Achieved (2016) $\sigma_{SI} = 1.1 \times 10^{-45} \text{ cm}^2$

Projected (2018) / Projected (2023) $\sigma_{SI} = 1.6 \text{ x } 10^{-47} \text{ cm}^2 / \sigma_{SI} = 1.6 \text{ x } 10^{-48} \text{ cm}^2$



XENON1T: First Results @ Andes 2017, June 30, 2017





20-50 tonne liquid xenon TPC