## Superconducting Nanowire Single Photon Detectors Opportunities for HEP



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### **SNSPD** Device Concept

- Heat from a single photon triggers detector out of superconducting state
- Resistance jumps to few  $k\Omega$  in picoseconds, shunting detector current into readout
- Highest performance single-photon detector available, from UV to mid-infrared
- 1 4 Kelvin operating temperature



An SNSPD is simply a current-biased superconducting wire in parallel with a readout circuit.

With the current through the nanowire reduced, the hotspot cools off, returning the wire to its original state.

The current density surrounding the hotspot exceeds the critical curent, and the entire wire width goes normal. The current is redirected through the measurement circuit, creating a detectable voltage pulse. When a photon hits the wire, it creates a hotspot, where a small region of the wire goes normal.

The current diverts around the hotspot.

after [1] Gol'tsman et al. (2001)



### WSi SNSPD Architecture





# **Established SNSPD Applications**

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#### Quantum Information Science

- Quantum Optics
- Trapped Ion Quantum Computing
- Linear Optical Quantum Computing

#### Free-Space Laser Communication

- Lunar Laser Comm Demo
- Deep Space Optical Comm Demo (Psyche)
- Space-to-Ground
  Quantum Communication

#### Physics & Astronomy

- Dark Matter searches
- Exoplanet Transit Spectroscopy
- Ultrafast transients







With its mid-infrared transit spectrometer, OST will search for **bio-indicators** (H<sub>2</sub>O and CO<sub>2</sub>) and biosignatures (O<sub>2</sub> and CH<sub>2</sub>) in nearby explanets to determine if we are **alone in the Universe**. OST can measure water's **DH Ingerprint in over 500 comets** to provide the leap needed to understand the delivery of water to our own inhabited planet. OST places our **solar system in context** by characterizing Kuiger belt objects and imaging Kuiger belt analogs in other solar systems.



NIST

Parameter	SOA 2019	SOA 2016	Group
Efficiency	98% @ 1550 nm	93% @ 1550 nm	NIST
Dark Counts	< 1e-5 cps	< 0.1 cps	MIT / NIST
Energy Threshold	0.125 eV	0.250 eV	NIST / JPL / MIT
Timing Jitter	2.7 ps	16 ps	MIT / JPL / NIST
Active Area	0.92 mm <sup>2</sup>	0.058 mm <sup>2</sup>	NIST / JPL
Max Count Rate	1.2 Gcps	0.3 Gcps	JPL
Pixel Count	1024 (32x32)	64 (8x8)	NIST / JPL
Photon Number Resolution	1 from 2 or more w/ 95% confidence	None	MIT / JPL
Operating Temp	4.3 K @ 1550 nm	2.5 K @ 1550nm	Single Quantum



NASA Deep Space Optical Communications (DSOC) Technical Demonstration Mission

Longest range demonstration of free-space laser communication by ~1000x

Demonstrating optical communications from deep space (0.1 - 2.7 AU) at rates up to 267 Mbps to validate:

- Link acquisition laser pointing control
- High photon efficiency signaling

Ground Laser Transmitter Table Mtn, CA 1 m OCTL telescope 5 kW laser power

**Ground Laser Receiver** Palomar Mtn, CA 5 m Hale telescope

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Pre-Decisional Information -For Planning and Discussion Purposes Only

Psyche spacecraft

Optical Platform Assembly 22 cm mirror 4 W laser power







1550 nm downlink

jpl.nasa.gov

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## 64-Pixel SNSPD Array for DSOC

- SNSPD planned for DSOC Ground Laser Receiver at 200 inch Palomar telescope (5.1 m)
- 64-element WSi SNSPD array with >79,000 μm<sup>2</sup> area (equiv. to 318.5 μm diameter)
- Divided into four spatial quadrants for fast beam centroiding
- 160 nm WSi nanowires on 1200 nm pitch each wire ~1 mm in length (~7000 squares)
- Free-space coupling to 1 Kelvin cryostat, with cryogenic filters and lens
- 78% system detection efficiency at 1550 nm
- < 80 ps FWHM timing jitter</li>
- ~1.2 Gcps maximum count rate







Electron Microscope Image of Nanowire Structure

CAD Design of SNSPD focal plane array
 CAD Design showing one of 16 individual sensor elements per guadrant



### **DSOC** Readout Electronics

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- Worked with industry on 64-channel TDC capable of streaming 900 Mtags / sec over PCIe
- Each nanowire sensor element has its own dedicated readout channel
- DC-coupled cryogenic amplifiers used at 40 K stage of cryostat

### **MUQDevices**



#### 40 Kelvin Cryogenic Amplifier Board





#### Streaming 64-channel time-to-digital converter

# **Ultra-high time resolution in SNSPDs**

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- Reduced timing jitter in SNSPDs from ~15 ps (2016 record) to 2.7 ps FWHM
- Advances in superconducting materials and low-noise readout technology
- Devices fabricated at MIT, measured with low-noise readouts at JPL and NIST
- Short nanobridges were used to avoid geometric jitter
- Proves that low jitter is achievable, but short nanobridges have negligible efficiency



Specialized low-jitter NbN SNSPD

Korzh et al, arXiv 1804.06839 (2018)



Dependence of timing jitter on photon energy



Ultra-low jitter measurement setup at JPL

# **Differential SNSPDs with Cavity**



- **Differential readout** cancels geometric jitter
- 61% SDE @ 1550 nm
- 12 ps jitter (300K LNA)
- Photon number • resolution



# **Differential SNSPDs with Cavity**

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# **Photon Counting Lidar With SNSPDs**

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- Differential SNSPDs infused into photon counting lidar field trials in Edinburgh, UK
- 12 ps FWHM timing jitter enables sub-cm depth resolution for every received photon



Depth image through 325 m turbulent air

# Acquired photon number 10-15 photons per pixel





Scanner



#### 325 m distance

jpl.nasa.gov



## **Mid-Infrared Operation**

- SNSPDs are near-ideal detectors at 1550 nm but can be modified to work well in mid-infrared
- NIST/JPL/MIT team recently demonstrated single photon sensitivity out to 9.92 μm (0.125 eV)
- Capability exists to develop time-resolved single photon mid-IR focal plane arrays with >60% efficiency and kilopixel form factors
- Advance enabled by WSi with reduced superconducting gap energy





# **Scaling to Kilopixel Arrays**

Demonstration of 1024-pixel imaging array with >99% pixel yield

50µm

- First ever demonstration of SNSPD array larger than 64 pixels
- 32x32 imager read out using only 64 readout lines
- 30x30 µm active area on 50 µm pitch total area 0.92 mm<sup>2</sup>



c30

c25

14

a)

c20 c10 c15 c5

 $\geq \mathbf{R}_{1}$ 



30µm







**FABRICATED AT NIST** 

**TESTED AT JPL** 

# **Thermal Row-Column Arrays**

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- Enables arrays with near-unity fill factor
- More scalable than conventional row-column
- Technique can be used at mid-IR wavelengths
- Working toward commercialization with NSF STTR
- 32x32 TRC array has 59 out of 64 channels working and 67% system detection efficiency at 1550 nm



Cross-polarized device design



# Dark Matter Detection with SNSPDs

- SNSPDs are ultra-low noise detectors with 1 0.1 eV energy thresholds
- Multiple proposed applications in searches for low-mass dark matter
  - Dark photon "haloscope"
  - Electron recoil in cryogenic semiconductors
  - Resonant absorption/scattering in molecular gases
  - Electron recoil in superconducting nanowire
- Technical challenges in scaling active area and pixel count to increase experimental reach





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- SNSPDs are a powerful technology for time-resolved single photon counting
- Progress in performance enhancement has been extremely rapid
- SNSPDs have much to offer for dark matter detection experiments
- For more information, contact mattshaw@jpl.nasa.gov

