





### **Novel Small-Gap Materials as Photodetectors**

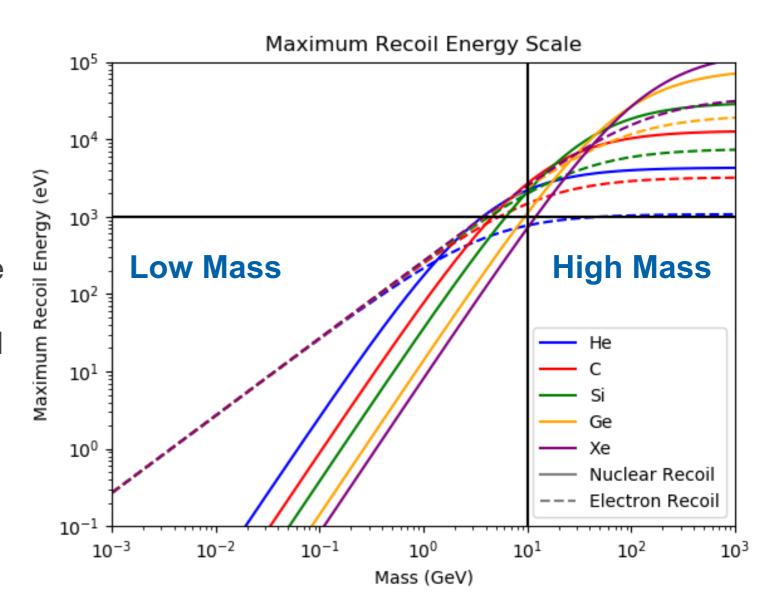
Noah Kurinsky, Yoni Kahn (UIUC) **CPAD 2019** December 10, 2019

#### **DM Collision Kinematics**

- Recoil energy for a typical WIMP velocity depends on target mass and recoil type
- Electron and nuclear recoils have different kinematics; nuclear recoils are simple elastic collisions, electron recoils are largely inelastic and depend on electron orbital and kinematics within the bound electron-atom system
- In addition to momentum transfer for a fixed velocity, using a velocity and angular distribution yields an expected energy spectrum

$$\Delta E_{NR} \le \frac{1}{2m_N} q_{max}^2 = \frac{m_N v^2}{2} \left( \frac{2m_\chi}{m_\chi + m_N} \right)^2$$

$$\Delta E_{ER} \le \frac{1}{2} \mu_{N\chi} v^2 = \frac{m_N v^2}{2} \left( \frac{m_\chi}{m_\chi + m_N} \right)$$



$$m_{\chi,NR} \ge \frac{\sqrt{2m_T\sigma_E}}{v}$$

$$m_{\chi,ER} \ge \frac{2\sigma_E}{v^2}$$



#### **General Idea**

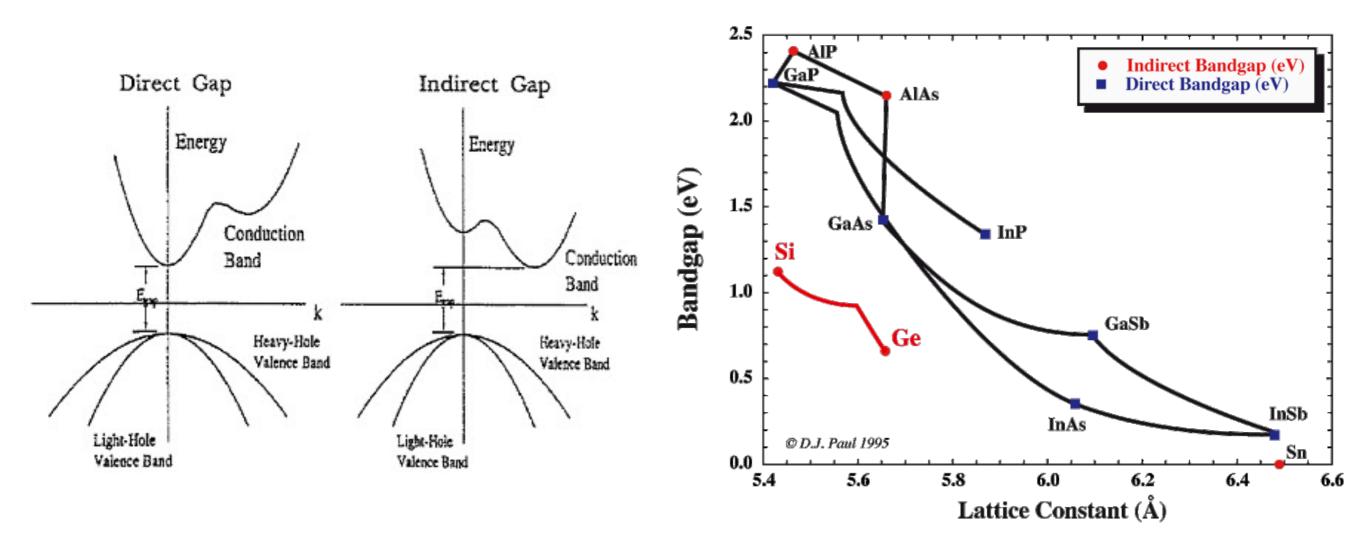
- Low-mass dark matter searches would greatly benefit from new detector materials with small gaps or long-lived phonon modes
- In the case of materials with small energy gaps, the technology challenges are identical to the early days of Si, Ge, etc photodiodes, with some exceptions
  - We have the advantage of years of R&D into single-charge resolving Si detectors
  - We can reliably make APDs from many materials down to the single electron level, and SiPMs are a mature technology
- Technology challenge: develop single charge resolved materials with small gaps at temperatures compatible with low dark rates
  - Driven by DM science
  - Inherently useful for other sub-eV processes
- Motivates developing technology driven by *intrinsic* microphysics rather than heterostructures

#### https://arxiv.org/pdf/1910.10716.pdf

Target	$E_g$ [eV]	$\overline{arepsilon}_{\infty}$	$\overline{\omega}_{ m O}  [{ m meV}]$
Si	1.11	_	62.3
Ge	0.67	_	34.8
NaI	5.8	3.27	12.4 - 20.0
CsI	6.14	2.70	6.9 - 10.0
CaWO <sub>4</sub>	5.2	3.84	8.48 - 106
GaAs	1.42	10.9	31.8 - 34.9
Al <sub>2</sub> O <sub>3</sub>	8.8	3.26	35.6 - 104
Diamond	5.47*	-	161
$SiO_2$	9.2	2.41	13.7 - 149
PbTe	0.19*	26.3	3.91 - 13.5
InSb	0.24*	23.7	20.5 - 21.5
AlN	6.20	4.54	29.4 - 109
$CaF_2$	11.81	2.26	28.4 - 55.6
GaN	3.43*	6.10	16.7 - 88.9
GaSb	0.720	21.6	26.4 - 27.3
LiF	14.2	2.02	33.5 - 77.2
$MgF_2$	12.4	1.97	12.1 - 73.7
MgO	7.83	3.38	46.3 - 82.6
NaCl	8.75	2.44	19.1 - 30.6
NaF	11.5	1.78	29.6 - 49.9
PbS	0.29*	15.0	7.27 - 26.9
PbSe	0.17*	19.5	4.86 - 17.1
ZnO	3.3	6.13	11.1 - 63.4
ZnS	3.80*	5.91	32.8 - 41.0



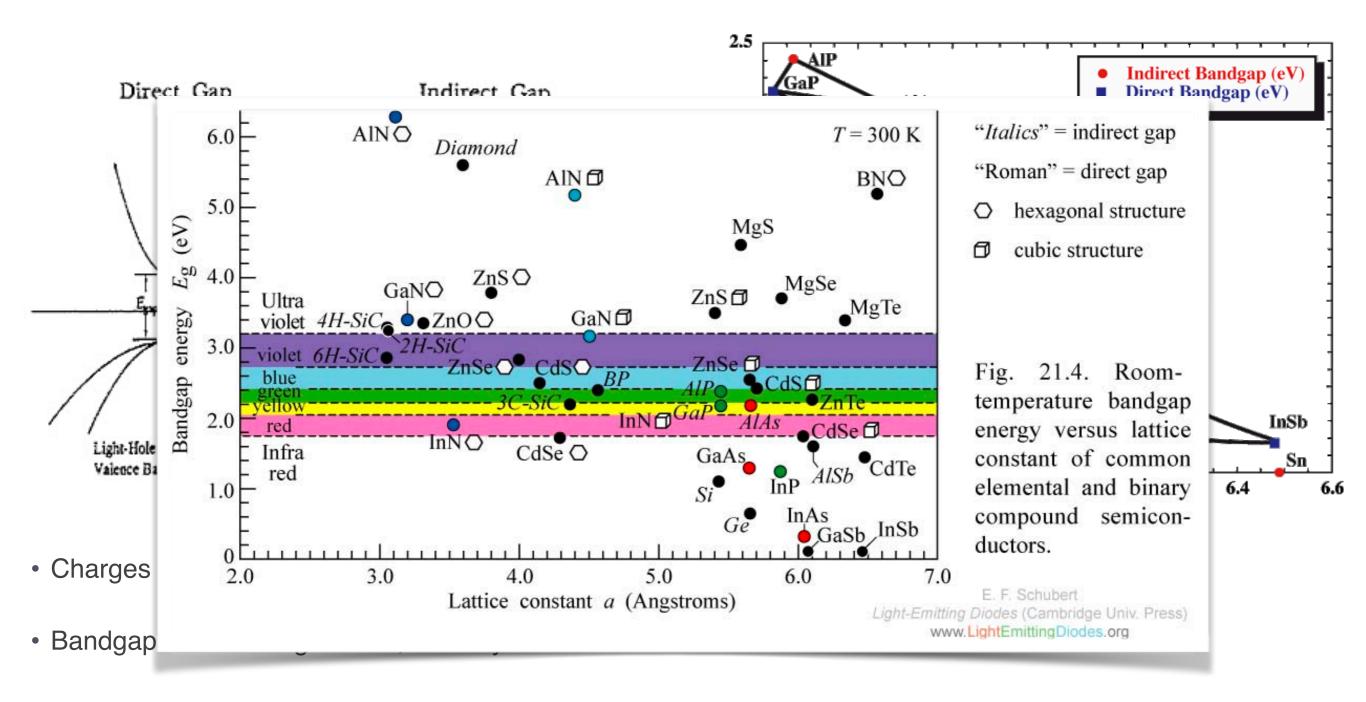
#### **Classical Picture of Semiconductors**



- Charges produced, and minimum photon energy, determined by material bandgap
- Bandgaps can be engineered, but only to some extent
- Indirect bandgaps require more energy to liberate electrons thermally, but are still sensitive radiation down to bandgap energies (though the efficiency is reduced)



### Classical Picture of Semiconductors Updated



 Indirect bandgaps require more energy to liberate electrons thermally, but are still sensitive radiation down to bandgap energies (though the efficiency is reduced)



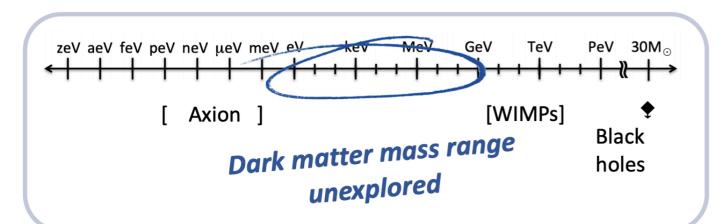
### From Sub-GeV DM Workshop (L. Wagner)

https://astro.fnal.gov/ldm/

### Spin-orbit semiconductors for dark matter detection

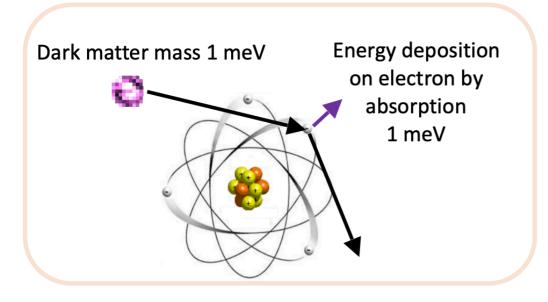
Inzani, Griffin

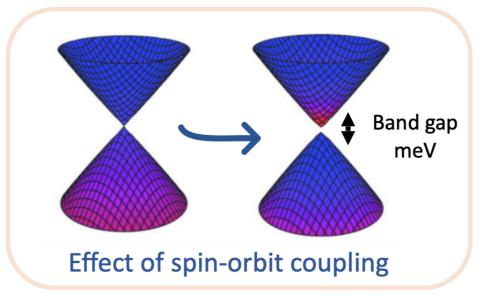




- Direct detection of sub-GeV masses is within the reach of short, small-scale experiments
- Small band gap semiconductors could be used to observe absorption or scattering events

**Aim:** To identify semiconductors with millielectronvolt band gaps





Strategy: Search for materials with band gaps opened by spinorbit coupling

https://indico.fnal.gov/event/20385/session/55/contribution/39/material/slides/0.pdf



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### Spin-orbit semiconductors for dark matter detection

Inzani, Griffin



Method I: High-throughput computational screening

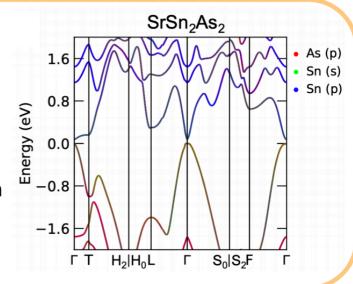


**86,412** inorganic materials

Spin-orbit interactions in 4,357 compounds

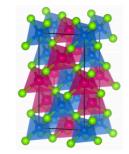
1 Family of materials with meV-scale band gaps

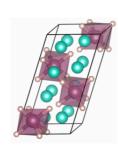
Tin pnictides  $ASn_2Pn_2$ Band gap variable by composition 0 – 200 meV

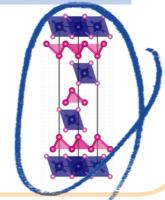


Method II: Refined electronic structures calculated by density functional theory

3 "would-be metals" with band gaps opened up through spin orbit coupling







Candidate materials for dark matter detection identified

Synthesis pending...

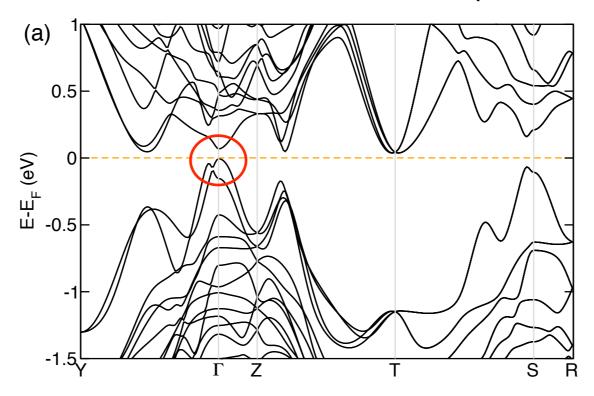
https://indico.fnal.gov/event/20385/session/55/contribution/39/material/slides/0.pdf



# Dirac materials for DM

[Hochberg, **YK**, Lisanti, Zurek, Grushin, Ilan, Liu, Weber, Griffin, Neaton, Phys. Rev. D 2018, 1708.08929]

### 3D Dirac semimetal (ZrTe<sub>5</sub>)

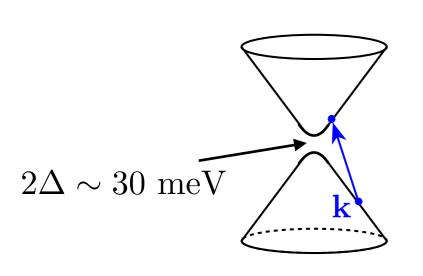


- meV excitation energies
- Anisotropic (bands and crystal)
- (Theoretically) insulating at zero temperature

Potential new class of materials for DM detection!

### What is a Dirac material?

ZrTe<sub>5</sub> is a "Dirac material" with highly anisotropic band structure



DM has a quadratic dispersion, so solving energy-momentum conservation with a linear dispersion leads to curious effects:

$$v_{\min}(|\mathbf{q}|,\omega_{\boldsymbol{\ell},\boldsymbol{\ell}+\mathbf{q}}) = \frac{\sqrt{v_{F,\perp}^2(\boldsymbol{\ell}+\mathbf{q})_{\perp}^2 + v_{F,z}^2(\ell_z + q_z)^2} + \sqrt{v_{F,\perp}^2\boldsymbol{\ell}_{\perp}^2 + v_{F,z}^2\ell_z^2}}{|\mathbf{q}|} + \frac{|\mathbf{q}|}{2m_{\chi}}$$

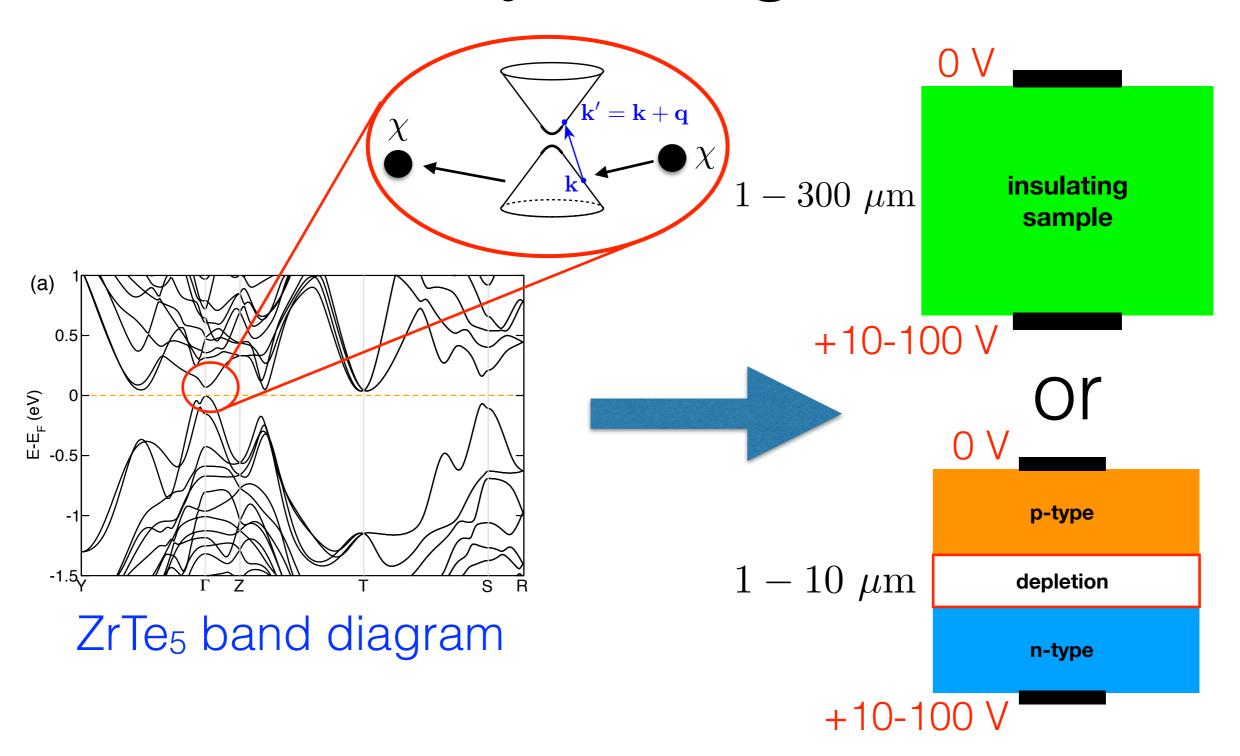
$$E(\mathbf{k}) = \pm \sqrt{v_{F,x}^2 k_x^2 + v_{F,y}^2 k_y^2 + v_{F,z}^2 k_z^2 + \Delta^2}$$
$$v_{F,x} \sim v_{F,y} \sim 1.6 \times 10^{-3} c$$
$$v_{F,z} \sim 6.5 \times 10^{-4} c \ll v_{F,x}, v_{F,y}$$

Like a rescaled relativistic ("Dirac") electron if space were anisotropic

No scattering if DM is slower than v<sub>F</sub>, but this depends on direction of q!

Strong directional dependence

### Detection by charge avalanche



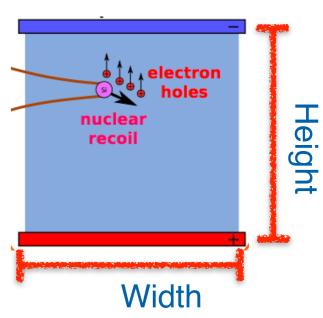
Would operate like an APD or a SiPM:

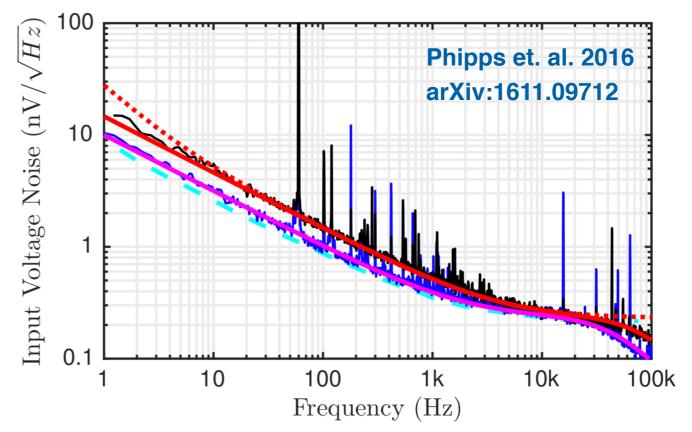
e-h pair created in insulating layer is accelerated, **impact ionization leads to charge cascade**: mA for exponential gain, 10-100 e for linear gain

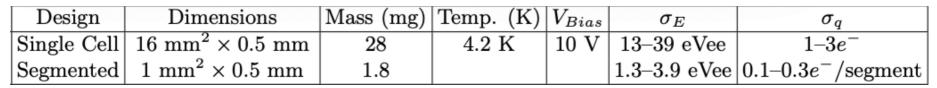
### **Cryogenic Single Charge Detector**

- Charge resolution is determined by detector+readout capacitance, voltage noise, and 1/f cutoff
- Capacitance and volume directly trade off, allowing for larger pixels to have the same charge resolution in a material with lower permittivity, or lower resolution in pixels of the same size
- Recent HEMT amplifiers have achieved performance at 4.2K sufficient for singlecharge ionization-chamber style detectors for low-rate signals at ~mg masses
- For diamond (epsilon~4) examples explored in table below

$$\sigma_q \ge \frac{N_v(C_{\rm det} + C_{\rm in})}{\epsilon_q \sqrt{\tau}}$$
 
$$\sigma_q \approx 35 \frac{(C_{\rm det} + C_{\rm in})/(250 \text{ pF})}{(C_{\rm in}/100 \text{ pF})^{1/4}}$$





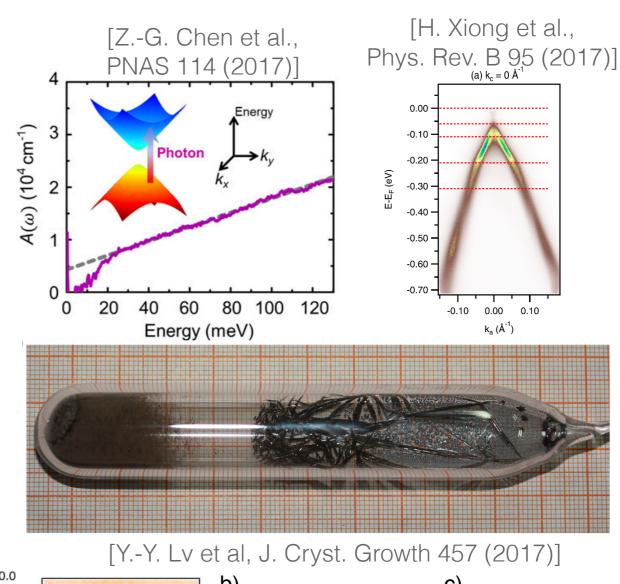


Kurinsky, Yu, Hochberg, Cabrera (1901.07569)



# ZrTe5 state of the art

- Gap is ~20 meV at 4.5 K, band structure is very close to linear near BZ center
- ~50 mg single crystals have been grown at Brookhaven and in China
- Fermi level can be manipulated with epitaxial strain and/or Te concentration
- Need to measure dielectric function and Fermi velocities in an insulating sample: UIUC has unique expertise and equipment to do this



[Y.-Y. Lv et al, J. Cryst. Growth 457 (2017)]

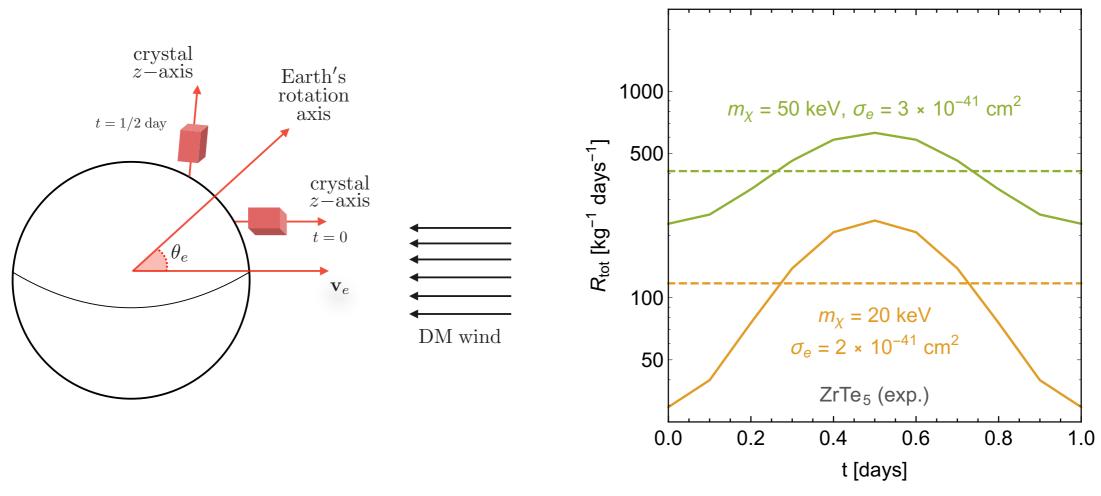
b)

c)

column and the second se

### Directional detection

With a charge detector, there is always a "dark rate" (impurities, thermal, stray light, etc), but this is isotropic

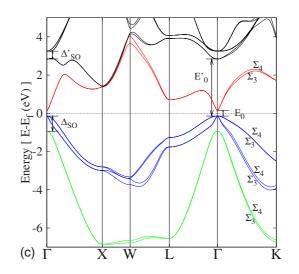


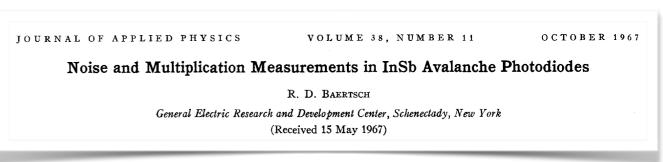
[Coskuner, Mitridate, Olivares, Zurek, arXiv:1909.09170; Geilhufe, Kahlhoefer, Winkler, arXiv:1910.02091]

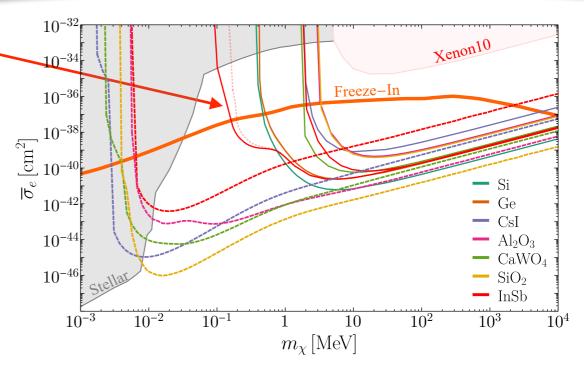
Daily modulation of event rate is a smoking gun for DM

### Alternative: InSb APD

- Making an avalanche device out of a new material is hard. Practice on an easier material
- InSb has a gap of ~235 meV and was first made into an APD in 1967 (!!)
- Large gap isn't great for lightest DM, but still sensitive to DM down to 200 keV
- To do: make a device, test at mK temperatures, calculate DM sensitivity
- Novel elements: cryogenic temperatures, SQUID current sensor readout in linear gain mode







[Griffin et al., arXiv:1910.10716]

### **Summary of Work Needed**

- Identify candidate materials
  - So far InSb, ZrTe5, new materials identified by materials project
- Establish testing program to validate theoretical properties
  - Resistance vs. temperature (RRR) and current-mode dark count measurements
  - Room temperature/cryogenic permittivity
  - Charge lifetimes and diffusion lengths
  - Characterize impact ionization in pure samples; determines monolithic or avalanche readout mode
- Establish cryogenic readout
  - ASICs being developed by CNRS for use with low capacitance HEMT amplifiers
  - Work starting at FNAL to develop integrated charge amplifiers at 4K
- Dark matter search is the 'black box' test that established technological viability; dark rates approaching single events/day are the long-term goal, but even Hz/mm is an excellent start
  - High dark rates in low-gap materials still translate to low dark rates at higher energies with adequate timing resolution



### Current state of the collaboration



P. Abbamonte (UIUC)



F. Mahmood (UIUC)





Y. Kahn (UIUC)



L. Wagner (UIUC)

calculation of event rate, materials search

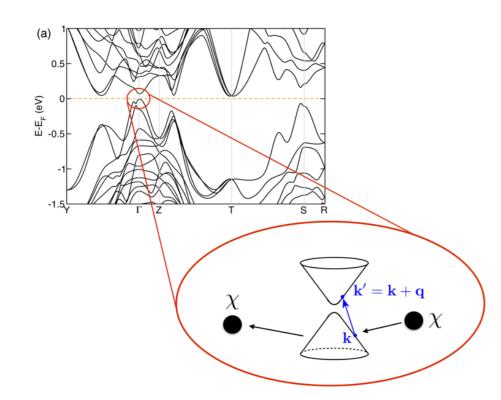


J. Eckstein (UIUC)



D. Van Harlingen (UIUC) detector prototype

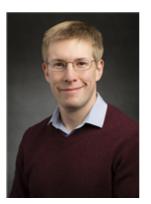
fabrication (MBE, SQUID current sensor)



measurement of  $S(\mathbf{q},\omega)$  with

THz/IR spectroscopy

and M-EELS



J. Filippini (UIUC)



N. Kurinsky (FNAL)

prototype testing at low-background cryogenic facility