

Structure Function Uncertainties and Neutrino Oscillation Experiments

Outline:

1. Neutrino oscillation experiments
2. Characteristic systematics worries
 - low Q^2 suppression in nuclei
 - E_ν from few-GeV calorimetry
 - ν_e appearance backgrounds
3. MINOS cross section modeling and plans

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April 29, 2005

ν Oscillation Experiments



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-Nakagawa-Sakata matrix

The Pontecorvo-Maki

$$= \begin{pmatrix} v_e \\ v_\mu \\ v_\tau \end{pmatrix} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$$

$$P_{V_\mu \rightarrow V_\tau} = \cos^4 \theta_{13} \sin^2 2\theta_{23} \sin^2 1.27 \Delta m_{32}^2 L / E_\nu$$

$$P_{V_\mu \rightarrow V_e} = \cos^2 \theta_{13} \sin^2 2\theta_{13} \sin^2 1.27 \Delta m_{32}^2 L / E_\nu$$

$$P_{V_e \rightarrow V_\tau} = \cos^2 \theta_{13} \sin^2 2\theta_{13} \sin^2 1.27 \Delta m_{32}^2 L / E_\nu$$

Oscillations between all 3 flavors
at each Δm^2 —
3 sets of 3 equations like these

The diagram illustrates the Earth's magnetic field lines, divided into two main regions: NORMAL and INVERTED.

- NORMAL Region:** On the right, labeled "NORMAL", there are three vertical bars representing the field direction. The top bar (3) has a small hatched section at the top. The middle bar (2) has a larger hatched section in the middle. The bottom bar (1) has a small hatched section at the bottom. Red arrows point from the labels v_τ , v_μ , and v_e towards these hatched sections, indicating the direction of the field lines.
- INVERTED Region:** On the left, labeled "INVERTED", there are three vertical bars representing the field direction. The top bar (3) has a large hatched section at the top. The middle bar (2) has a small hatched section in the middle. The bottom bar (1) has a large hatched section at the bottom.
- Solar Influence:** A bracket labeled "Solar { " is positioned above the first two bars (3 and 2). Another bracket labeled " } Solar" is positioned below the last two bars (1 and 2).
- Atmospheric Layer:** A bracket at the bottom is labeled "<--Atmospheric-->".

Goals of next generation of expts:

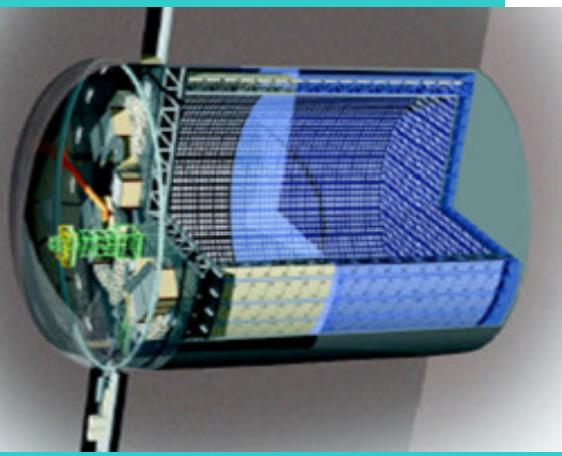
- 1) Better precision on parameters
 - 2) Observing **Oscillations**
 - 3) θ_{13} and δ_{CP} in the PMNS matrix
 - 4) “Normal” or “Inverted” mass hierarchy

ν Oscillation Experiments

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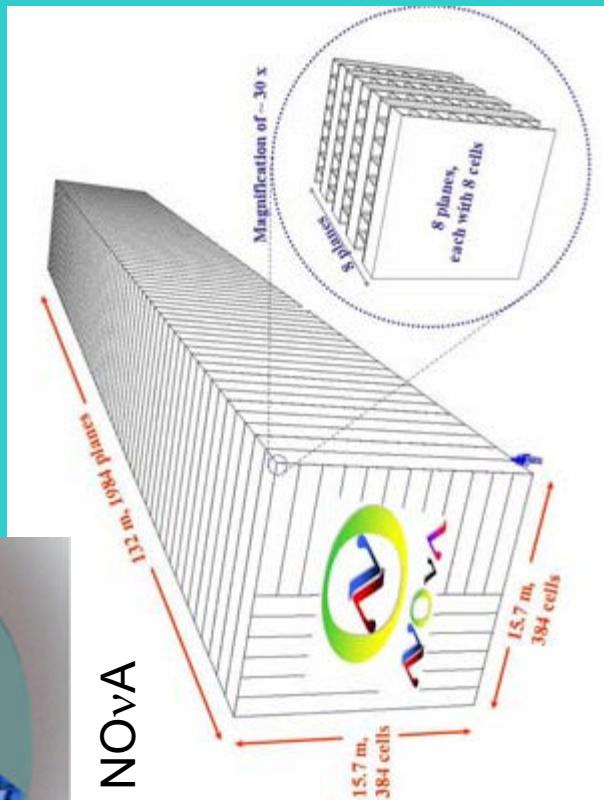
ν Oscillation Experiments

Super-Kamiokande

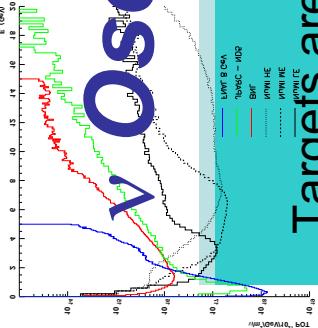


Current/Future long baseline experiments;

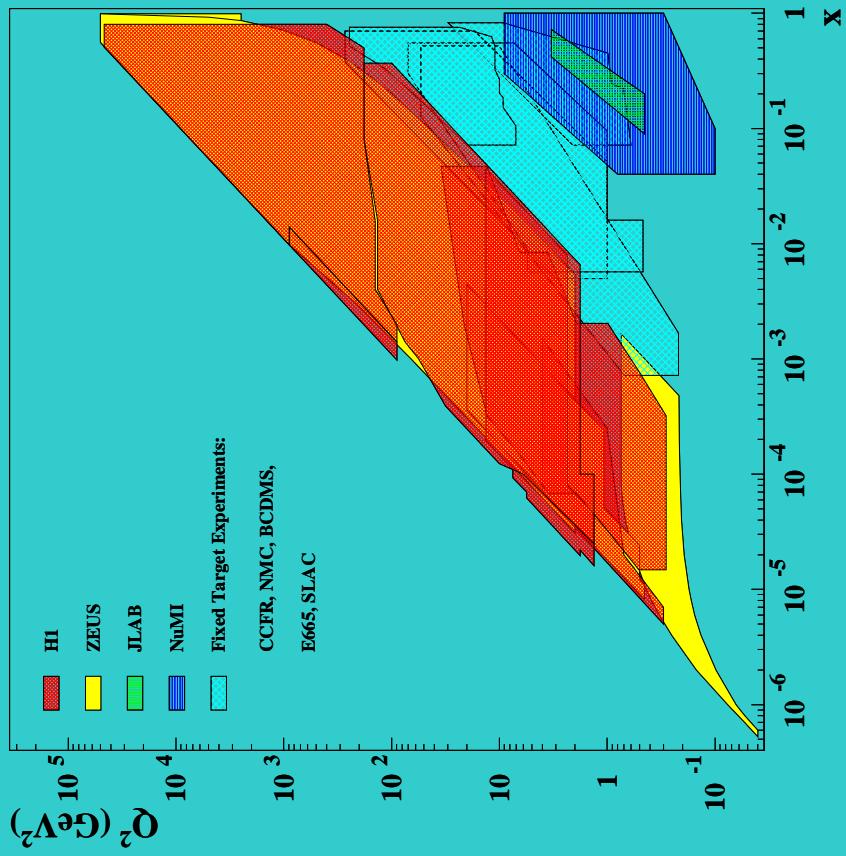
Expt.	Dates	E_ν (GeV)	Mass (kt)
K2K	1998-2005	1	50
MINOS	2005+	~3-10	5.4
T2K	2008+	0.7	50
NO _ν A	?	2.2	30



Oscillation Experiments

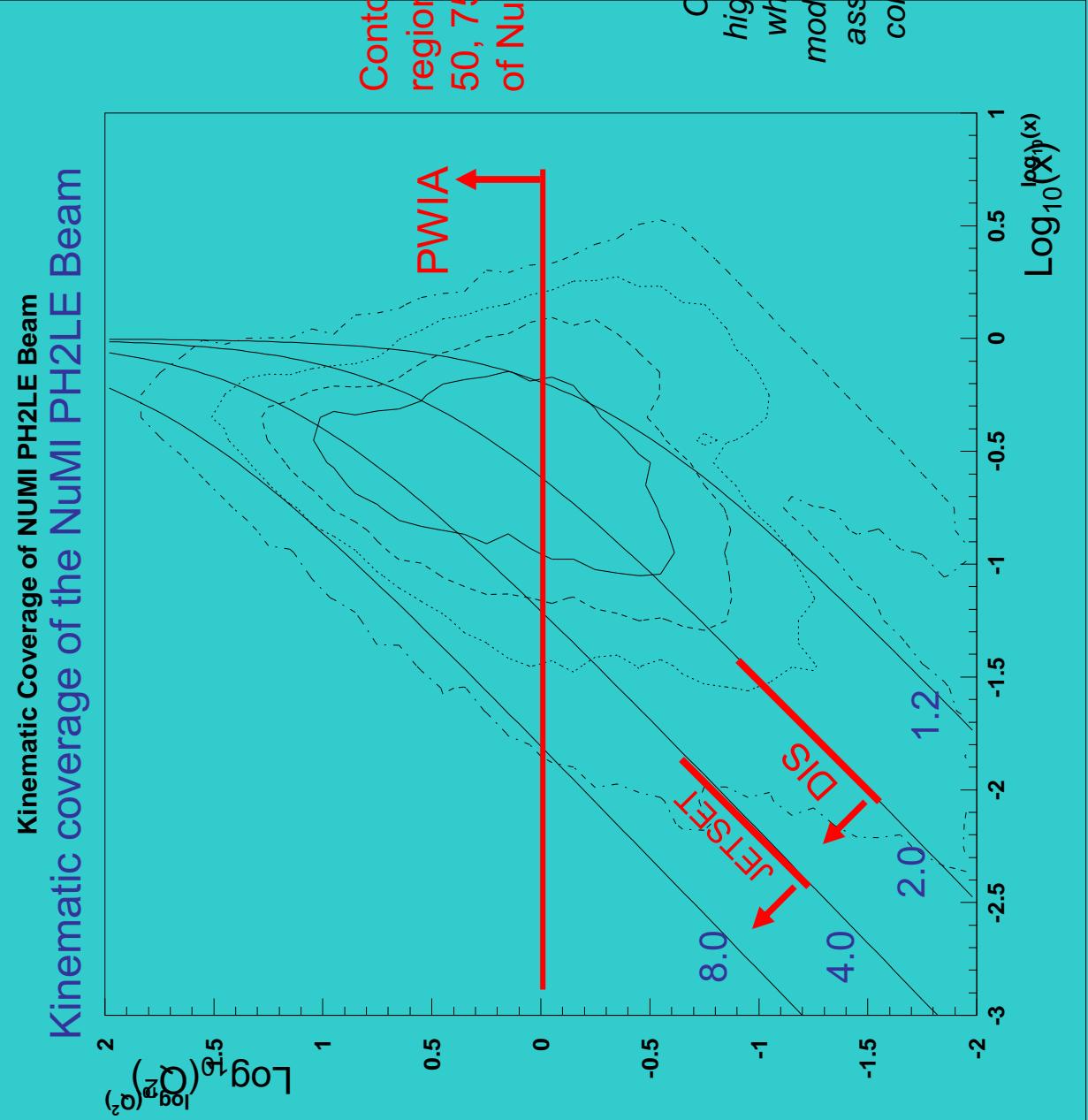


Targets are mainly CH_2 , H_2O , or Fe.
Variety of options for beam energies:
On/off-axis conventional beams,
“SuperBeams”



Large degree of overlap between the kinematics of the NuMI beam and those under exploration at JLab and elsewhere.

Kinematic Coverage



Near Detectors

Structure functions, cross sections, nuclear effects... that stuff is messy!
That's why we built a near detector...



“Near/Far comparison” does cancel many systematics, but:

1. Near and far detectors slightly different (2% relative uncertainty for MINOS)
2. Alternate ways of describing ND data might extrapolate to FD differently
3. Near and far beams different – decay kinematics and beam optics (smaller effect), oscillations (bigger effect)
4. Not everything cancels: Δm^2 is coupled to E_ν , knowledge of the absolute neutrino energy scale is required.

Oscillation Experiments

Most important uncertainties related to cross sections and modeling of neutrino-nuclear interactions in nuclei depend on beam energy and analysis.

θ_{23} measurements:

Δm^2 measurement coupled to E_ν
Understanding neutrino energy scale

K2K – E_ν based on QEL reconstruction

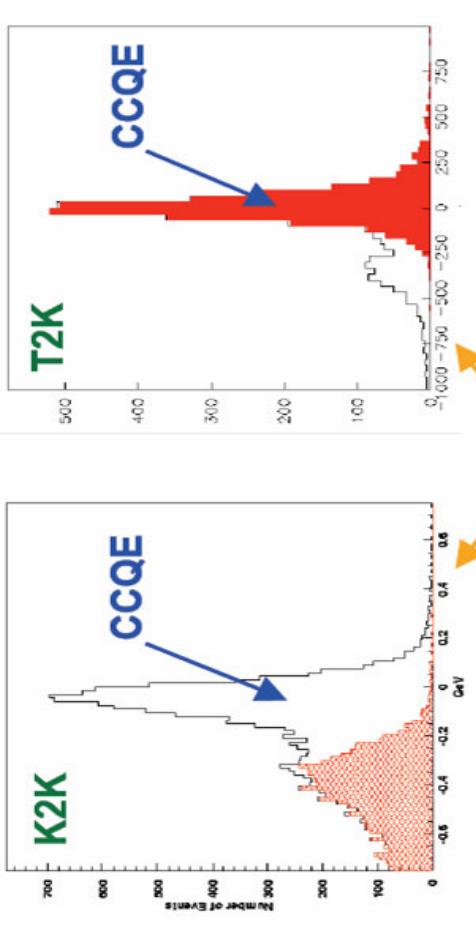
MINOS – Calorimetry /

NC mis-id at low E_ν

θ_{13} measurements:
understanding NC backgrounds

π° production

$E_\nu(\text{reconstruct}) - E_\nu(\text{True})$ (MeV)



T. Nakaya, NuINT 04

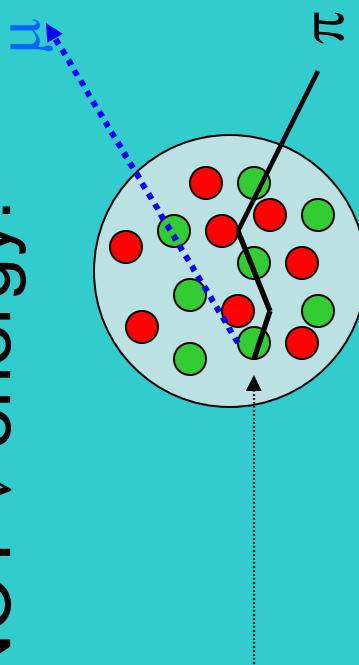
Complete understanding of systematic uncertainties and sensitivities involves many sources:
near – far extrapolations
cross sections – inclusive and exclusive (topological identification)

hadronization models – affect NC/CC mis-id, ν_e background
nuclear physics – on cross section and re-scattering of produced particles

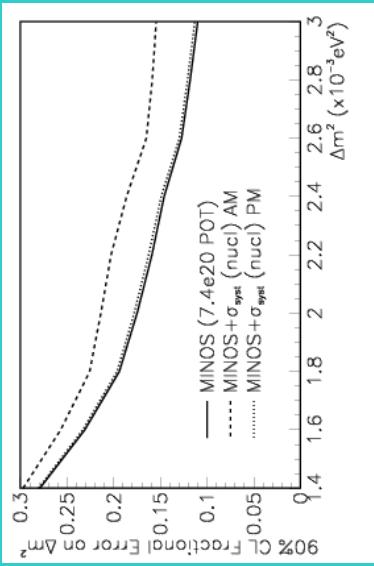
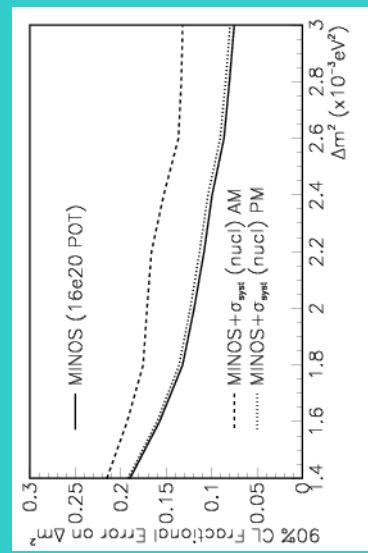
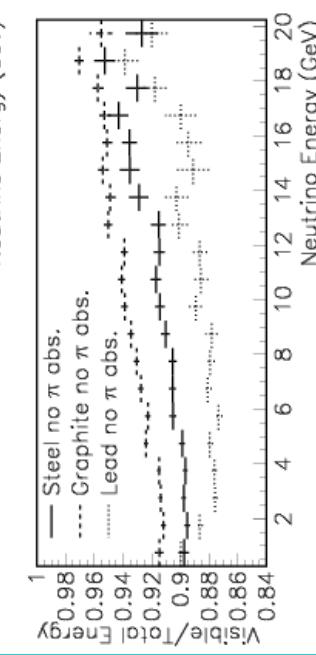
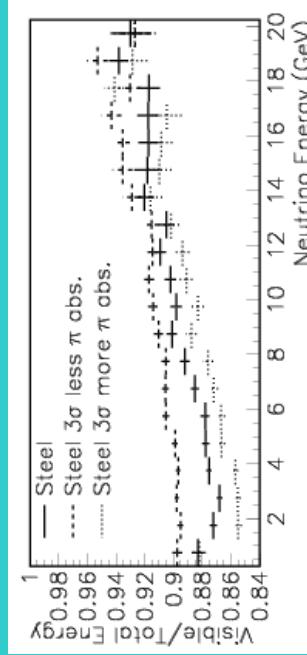
Example: E_ν from Calorimetry

Visible Energy in Calorimeter is NOT ν energy!
 π absorption, rescattering
final state rest mass

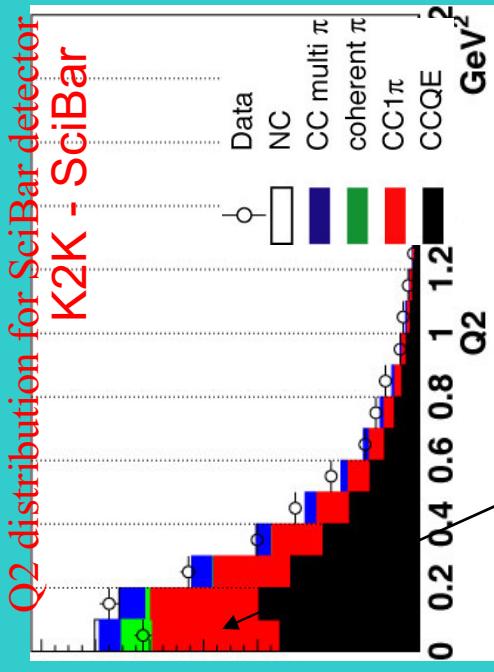
(D. Harris et al., hep-ex/0410005)



Nuclear effects studied in charged lepton scattering, from deuterium to lead, at high energies, but nuclear corrections may be different between e/μ and ν scattering



Example: Low Q^2 suppression



One such effect is the larger than expected suppression of events at low Q^2 from K2K and miniBoone.

All “known” nuclear effects taken into account: Pauli suppression, Fermi Motion, Final State Interactions.

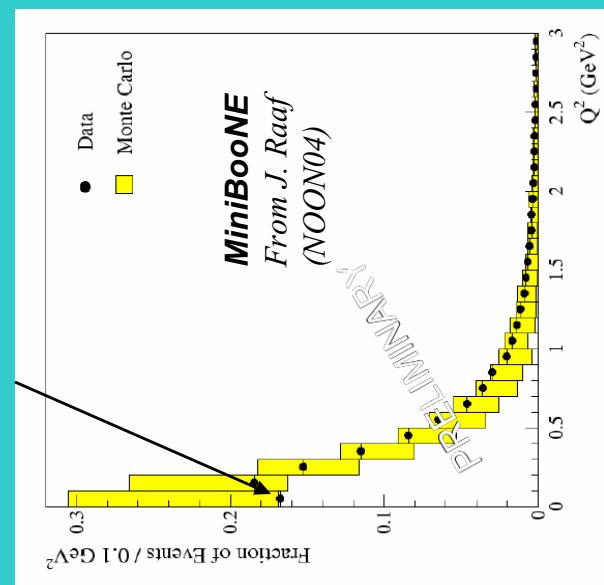
Variety of explanations examined:

• Pauli blocking of Δ states?

- Smaller than expected CC coherent contribution?
- Missing lepton mass terms in resonance production cross sections.

(Lalakulich-Paschos, Phys. Rev. D71:074003, 2005)
• Nuclear shadowing (Kopeliovich, hep-ph/0409079)

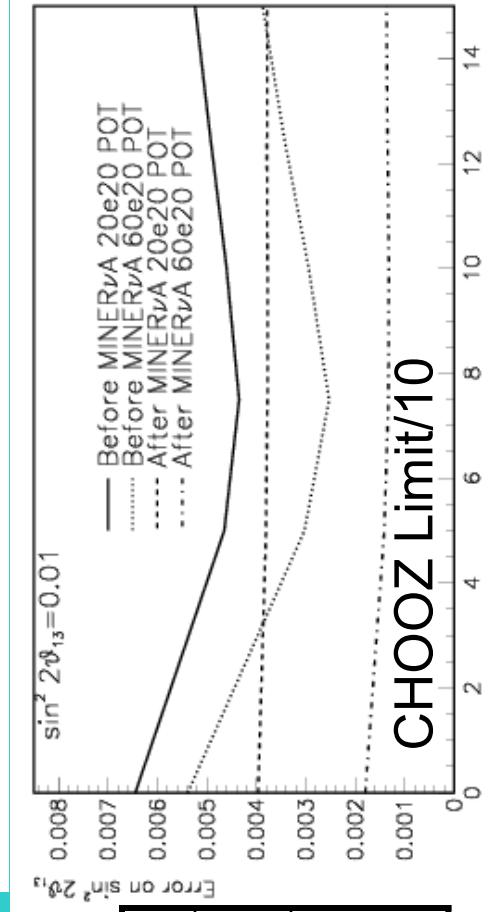
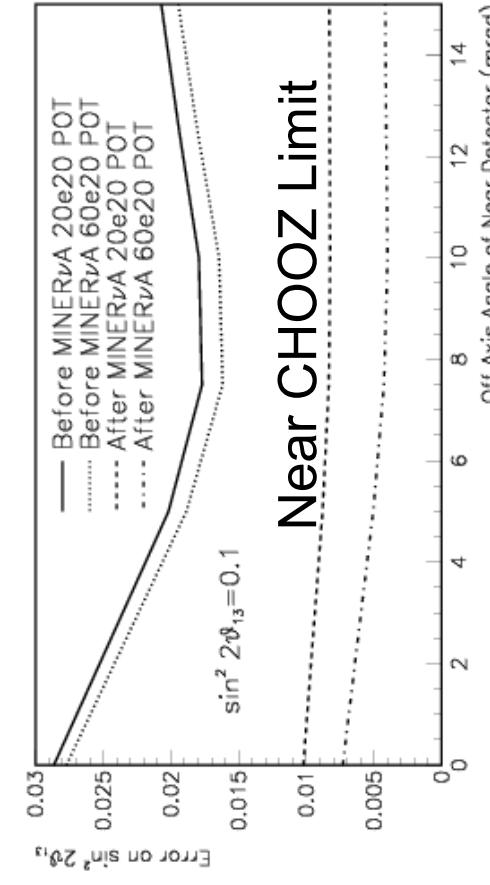
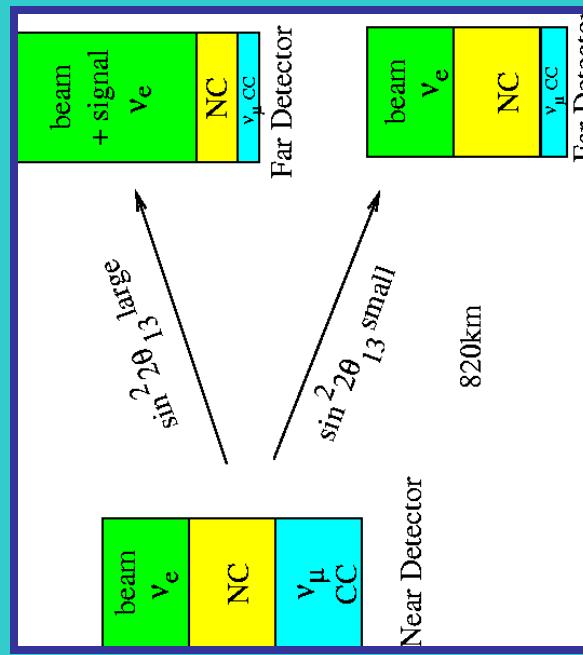
Importance of using correct form factors (non-dipole) for M_A extraction \rightarrow affects Q^2 distribution shape (Budd, Bodek, Arrington, hep-ex/0410055).



After considerable study, K2K parametrized the deficit, folded back into their MC for the far detector. No effect on the oscillation fits (but low statistics).

Example: Future θ_{13} Expts

(D. Harris et al., hep-ex/0410005)



Process	QE	RES	COH	DIS
δσ/σ NOW (CC,NC)	20%	40%	100%	20%
δσ/σ after MINERvA (CC/NC)	5%/na	5%/10%	5%/20%	5%/10%

Without MINERvA, NOvA may be limited by cross section systematics



Model Tuning (MINOS)

Modeling of ν -nucleus interactions had been based primarily on old models and tuned primarily to neutrino bubble chamber data.

LO DIS calculations

Rein-Seghal model for $W < 2$ GeV

(Annals of Physics 133, 79, 1981)

FGM for nuclear effects

There has been a renewed interest in the subject, and many papers in the last 4 years detailing improvements to the modeling

1) JLAB data / improving resonance models

Lalakulich-Paschos, Phys. Rev. D71:074003, 2005

T. Sato et al. Phys. Rev. C67, 065201 (2003)

2) Nuclear modeling / spectral functions

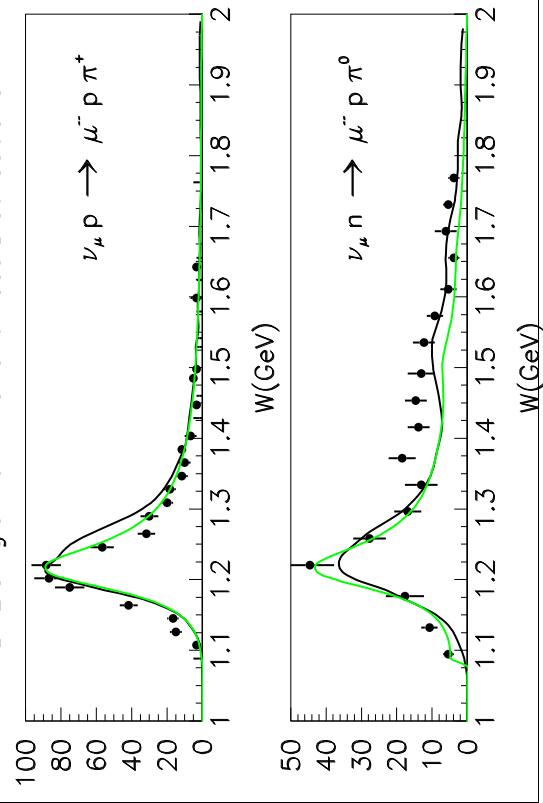
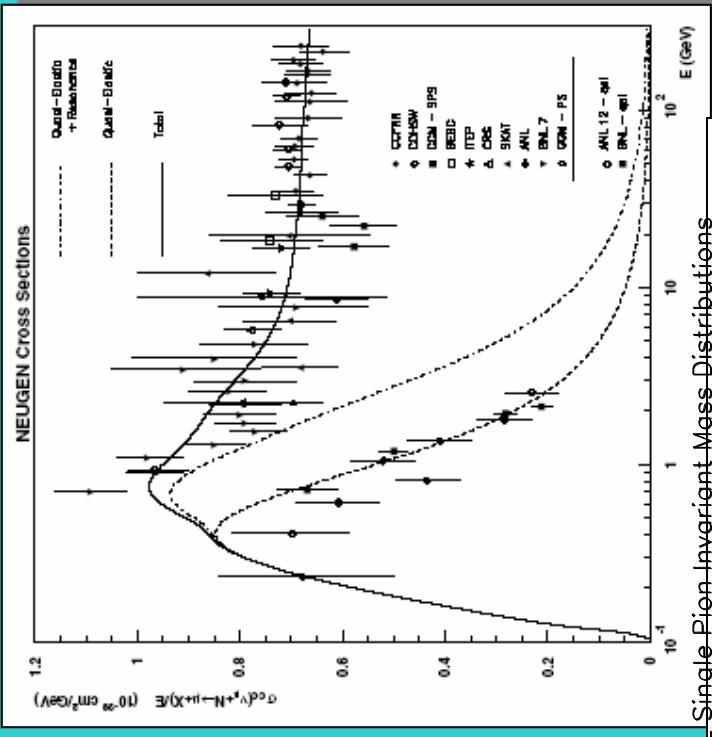
O. Benhar, Nucl. Phys. B620, 302

Nieves et al. nucl-th/0408008

J. E. Amaro et al. nucl-th/0409078

3) Bridging the pQCD/non-pQCD gap

Bodek et al. hep-ph/0411202

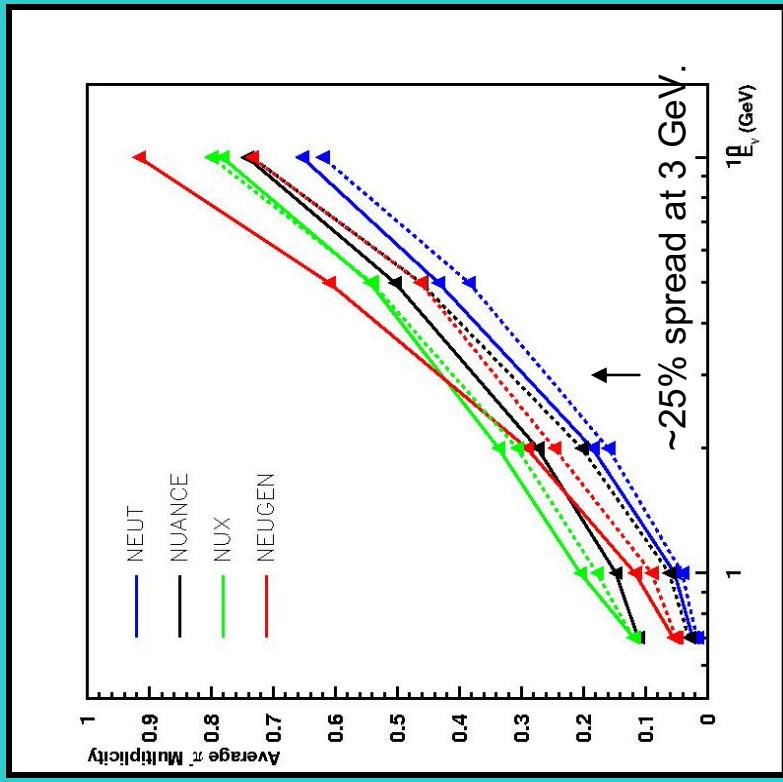
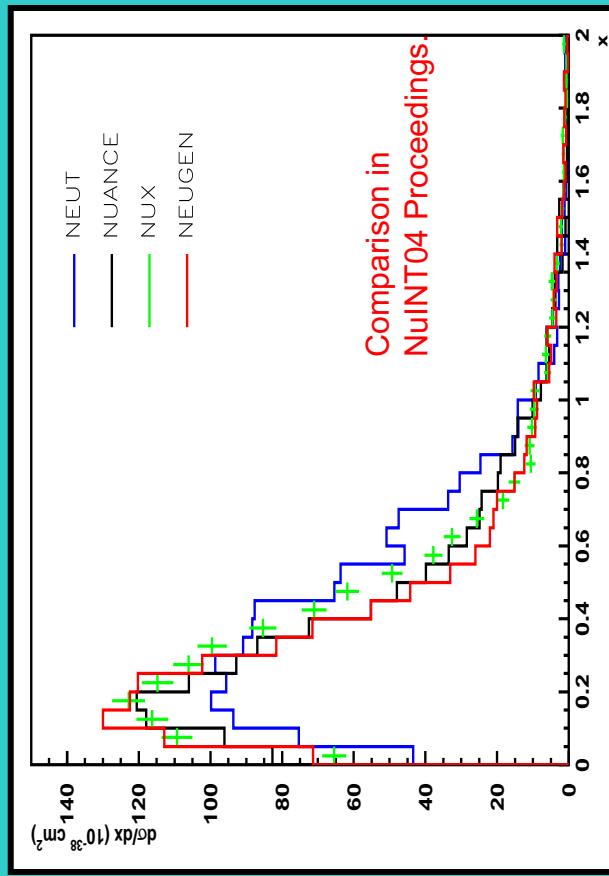
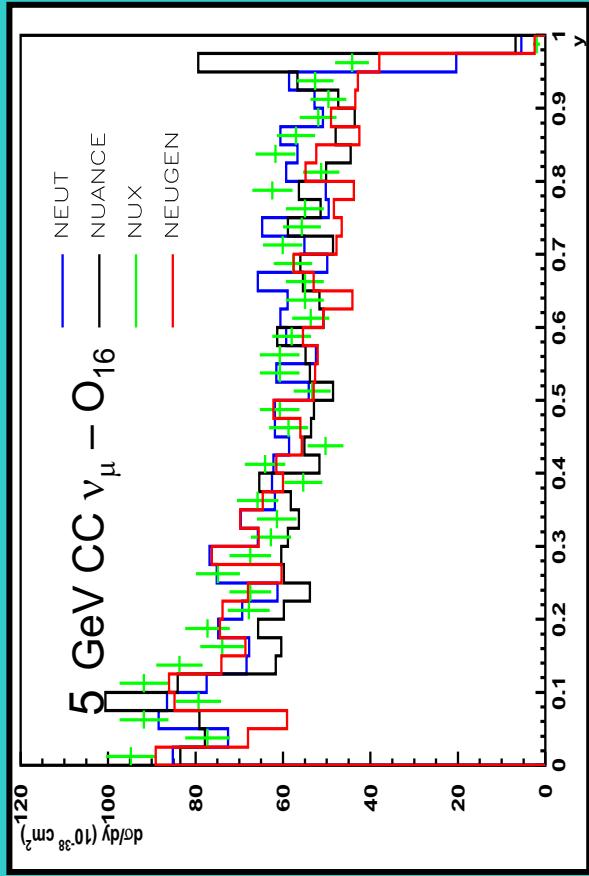


$\nu_\mu n \rightarrow \mu^- p \pi^+$

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Model Tuning (MINOS)

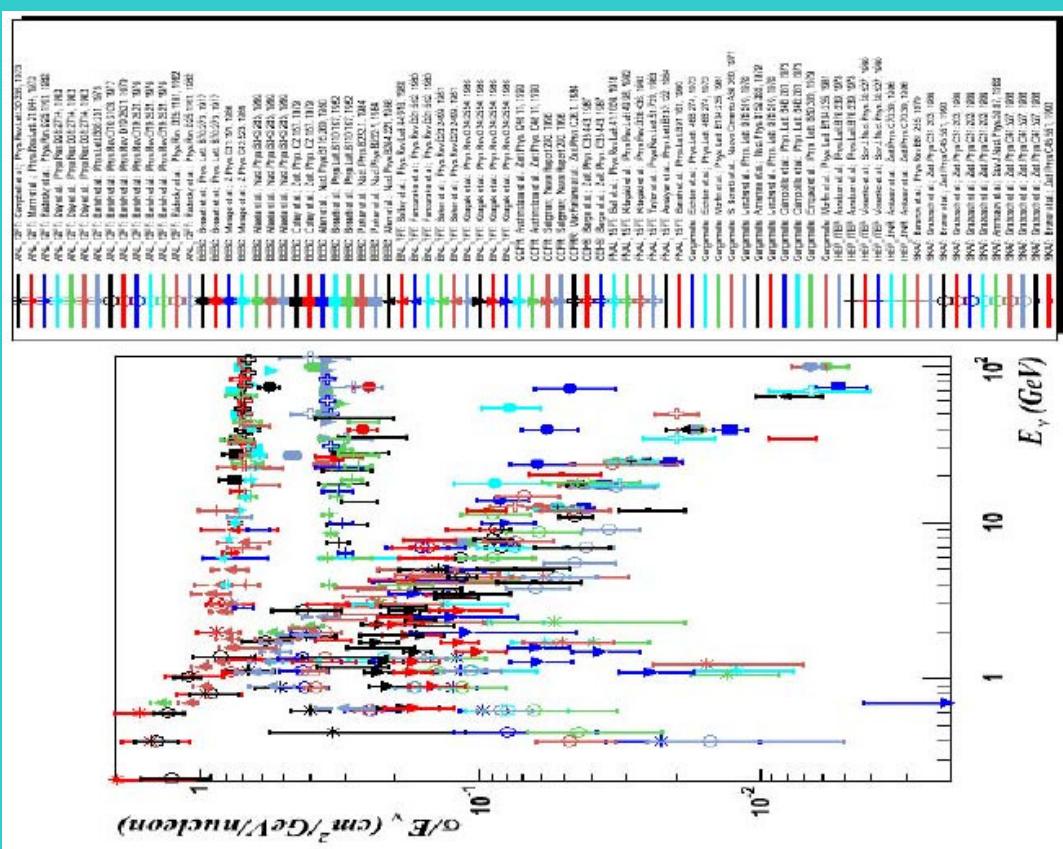
NUANCE (SuperK/K2K/miniBoone):
D. Casper, hep-ph/0208030
NUX-FLUKA (NOMAD/ICARUS):
F. Battistoni *et al.* NuInt 02 Proceedings
NEUT (SuperK/K2K):
Y. Hayato, Nucl. Phys. Proc. Suppl. 112, 171 (2002)
NEUGEN (MINOS, Sudan 2):
H. Gallagher, Nucl. Phys. Proc. Suppl. 112 188 (2002)



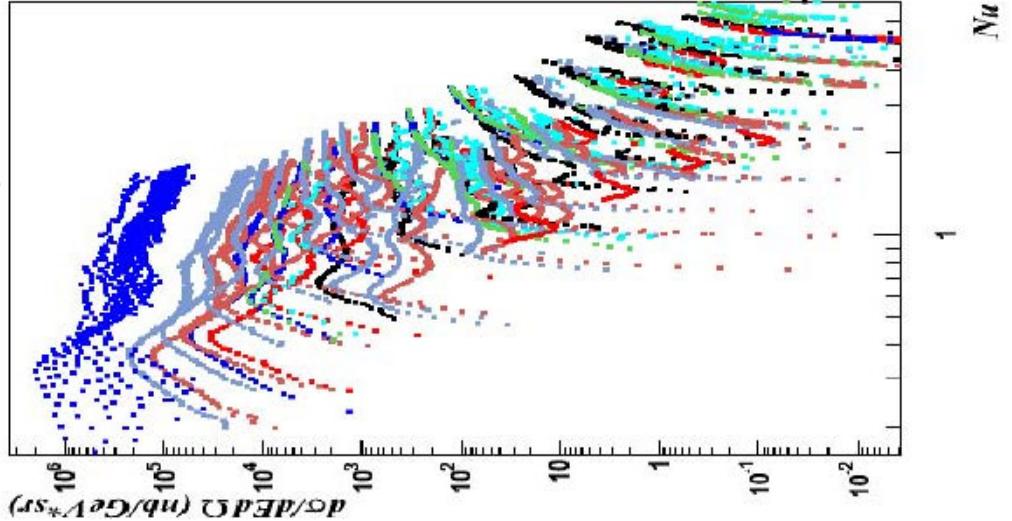
Comparison in
NuInt04 Proceedings.

Model Tuning (MNOS)

Past tunings have used this data – v cross sections



*But not this data, nor SF
data explicitly*



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E133: S. Wood et al.; e133.xl; 2004

E133: S. Wood et al.; e133.xl; 2004

E140: S. Wood et al.; e140.xl; 2004

E140: S. Wood et al.; e140.xl; 2004

E140: S. Wood et al.; e140.xl; 2004

E140C: S. Wood et al.; e140c.xl; 2004

E140D: S. Wood et al.; e140d.xl; 2004

E140D: S. Wood et al.; e140d.xl; 2004

E140E: S. Wood et al.; e140e.xl; 2004

E140F: S. Wood et al.; e140f.xl; 2004

E140G: S. Wood et al.; e140g.xl; 2004

E140H: S. Wood et al.; e140h.xl; 2004

E140I: S. Wood et al.; e140i.xl; 2004

E140J: S. Wood et al.; e140j.xl; 2004

E140K: S. Wood et al.; e140k.xl; 2004

E140L: S. Wood et al.; e140l.xl; 2004

E140M: S. Wood et al.; e140m.xl; 2004

E140N: S. Wood et al.; e140n.xl; 2004

E140O: S. Wood et al.; e140o.xl; 2004

E140P: S. Wood et al.; e140p.xl; 2004

E140Q: S. Wood et al.; e140q.xl; 2004

E140R: S. Wood et al.; e140r.xl; 2004

E140S: S. Wood et al.; e140s.xl; 2004

E140T: S. Wood et al.; e140t.xl; 2004

E140U: S. Wood et al.; e140u.xl; 2004

E140V: S. Wood et al.; e140v.xl; 2004

E140W: S. Wood et al.; e140w.xl; 2004

E140X: S. Wood et al.; e140x.xl; 2004

E140Y: S. Wood et al.; e140y.xl; 2004

E140Z: S. Wood et al.; e140z.xl; 2004

JAB: S. Wood et al.; jabc1.xl; 2004

JAB: S. Wood et al.; jabc2.xl; 2004

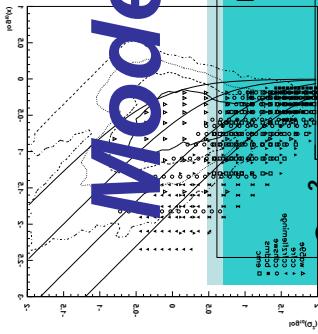
JAB: S. Wood et al.; jabc3.xl; 2004

NE11: S. Wood et al.; ne11.xl; 2004

NE11: S. Wood et al.; ne11.xl; 2004

OPEN/H/F: S. Wood et al.; open-h/f.xls; 2004

Model Tuning (MINOS)



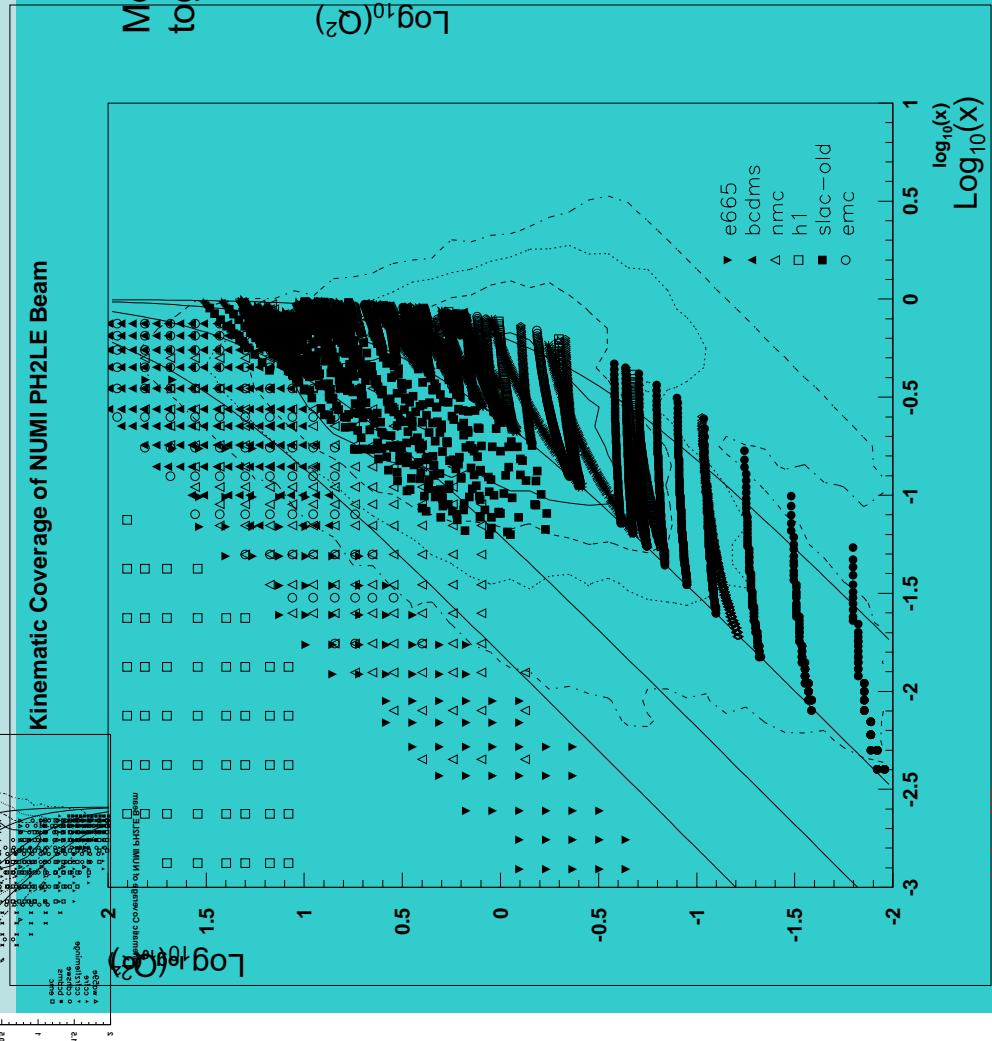
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Mostly structure function data,
together with some cross section data.

Nuclear targets

$\log_{10}(Q^2)$



H_2/D_2 targets

$\log_{10}(x)$

JUPITER data will help fill in here

Model Tuning (*MINOS*)

Well understood: (better than 10%)

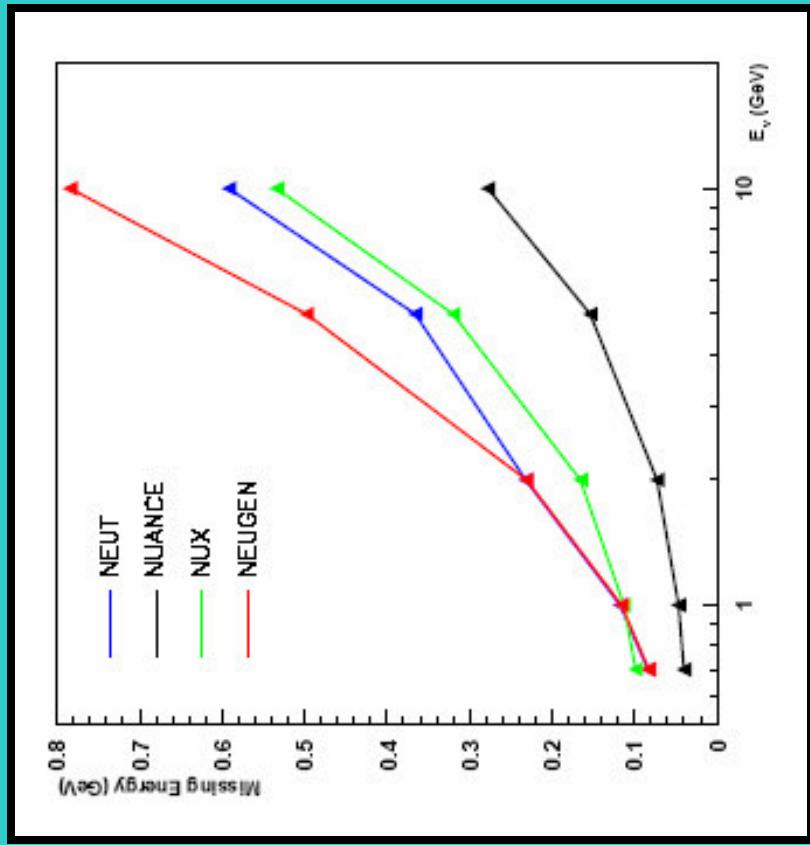
- High energy ($E > 10$ GeV) CC cross section (few %)
- location of peak of the beam spectrum
- Quasi-elastic / Δ cross section (> 0.1 GeV^2 or so)
- y-distributions at higher energies ($E > 10$ GeV)

Somewhat well known: (~ 10-20%)

- flux normalization and shape
- low energy CC total cross sections
- x-distributions at medium energy

Not very well known:

- non-scaling contribution (QEL/RES) at low Q^2
- hadronic system multiplicities, e/h,
- and leading pion distributions
- intranuclear rescattering – “missing energy”



Model Tuning (MINOS)

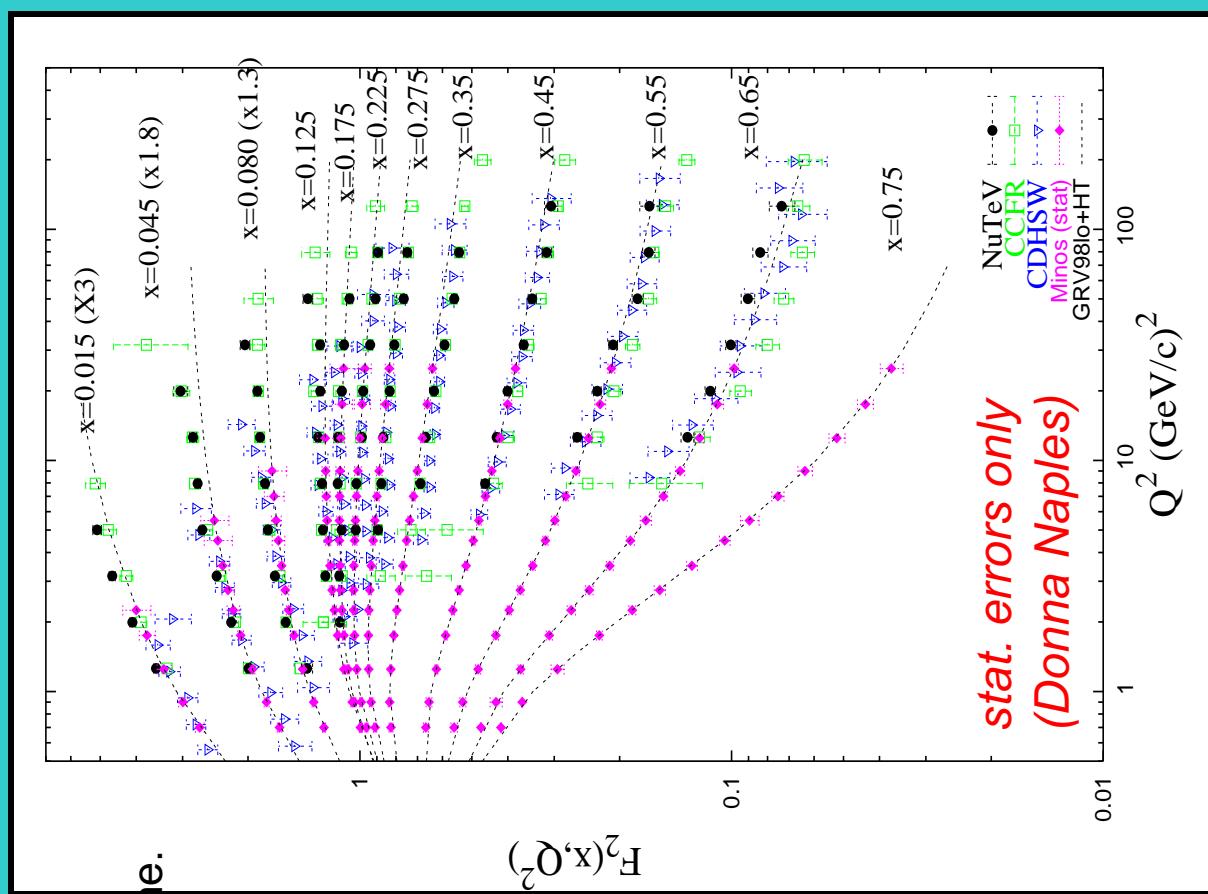
The MINOS near detector will have very high statistics: 78k-257k CC ν_μ events/10²⁰ POT/ton depending on the beam tune.

Challenge comes in having to simultaneously tune beam and interaction physics models. By the end of MINOS we expect improved knowledge on the beam from MIAPP and on cross sections from Minerva (see J. Morfin talk tomorrow).

General approaches:
“bottoms-up” – identify areas of disagreement and try to identify and improve models

“top-down” – simply parametrize differences in the measured distributions.

In the middle are “physics-based” parametrizations, for example the Bureas-Gaemers iron PDF fits used by NuTeV. (Donna Naples, U. Pittsburgh).



Conclusions

1. Correct modeling of ν -A cross sections over a broad range in kinematics and A important for current and future experiments.
2. Additional contributions related to hadronization and modeling of intranuclear cascade.
3. Building a decent simulation and tuning to experimental data requires close communication between experimentalists and theorists, as well as experimentalists and experimentalists.