

# First MINOS Results from the NuMI Beam

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for the MINOS Collaboration

Argonne • Athens • Benedictine • Brookhaven • Caltech • Cambridge • Campinas • Fermilab  
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# Oscillation Physics

- There is experimental evidence that neutrino flavor is not conserved and thus **neutrinos have non-zero masses**.
- A neutrino of flavor  $\alpha = e, \mu, \tau$  is a superposition of mass eigenstates described by the lepton mixing matrix  $U$ :

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \quad s_{ij} = \sin\theta_{ij}, \quad c_{ij} = \cos\theta_{ij}$$

$$|v_\alpha\rangle = \sum_i U_{\alpha i} |v_i\rangle; \alpha = e, \mu, \tau, \dots; i = 1, 2, 3, \dots$$

$|U_{\alpha i}|^2$ : probability that  $v_i$  interaction will produce  $v$  of flavor  $\alpha$

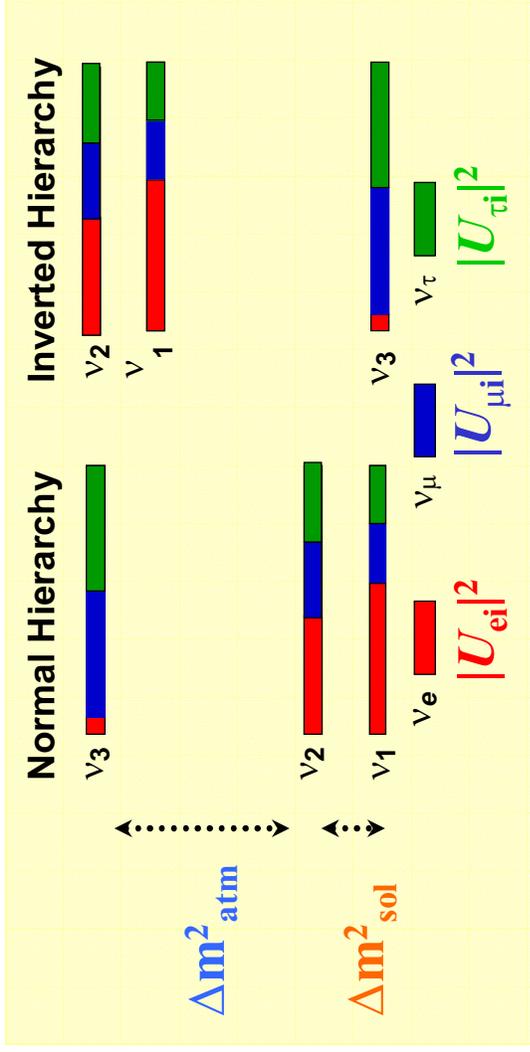
- In the two-flavor approximation:

$$P(v_\alpha \rightarrow v_\beta) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 (eV^2) L (km)}{E_\nu (GeV)} \right), \alpha \neq \beta$$

→ **L/E determines what  $\Delta m^2$  is probed**

# Neutrino Mass Spectrum

(3 mass eigenstates - assume no  $\nu_{\text{sterile}}$ )



**Two mass splittings:**

**Large ( $\Delta m^2_{\text{atm}}$ ):**

Atmospheric (up-down asymmetry)  
Long baseline

**Small ( $\Delta m^2_{\text{sol}}$ ):**

Solar + reactor

- $|U_{e3}|^2$  is not known (bounded by reactor expts) and neither is the mass hierarchy.
- Solar data suggests  $|U_{e2}|^2 \sim 1/3$ , and thus  $|U_{e1}|^2 \sim 2/3$  (unitarity and small  $|U_{e3}|^2$ ).
- Also, data suggests  $\nu_\mu \rightarrow \nu_\tau$  dominant at  $\Delta m^2_{\text{atm}}$  and maximal mixing angle,

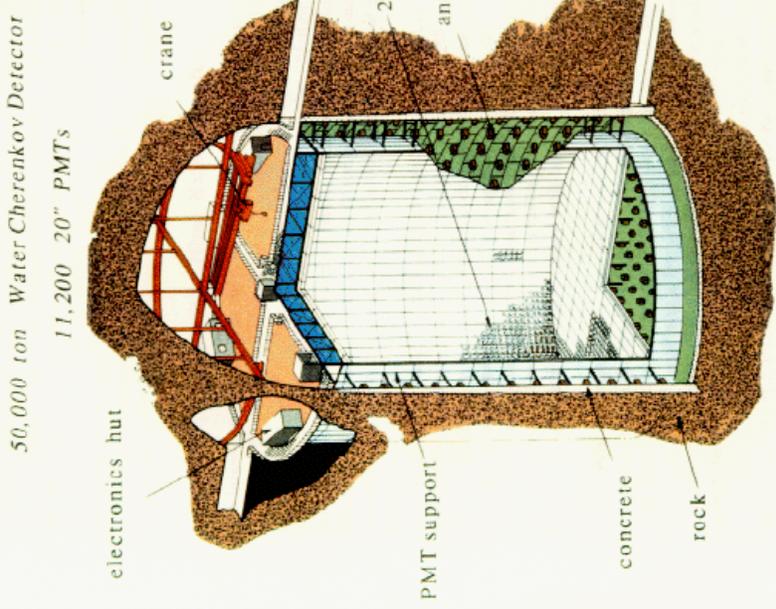
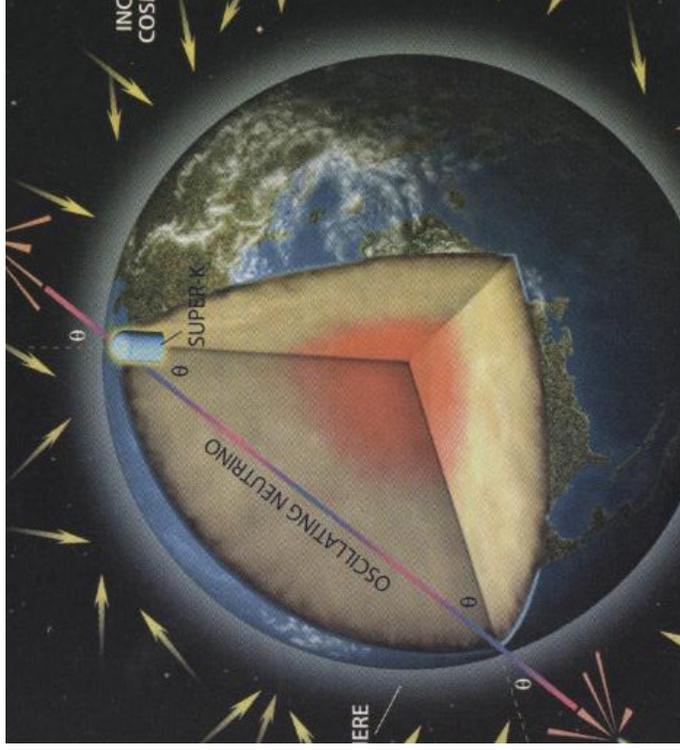
$$|U_{\mu i}|^2 \sim |U_{\tau i}|^2$$

With  $L/E \sim 10^3$  MINOS is sensitive to  $\Delta m^2_{\text{atm}}$   
This scale will be the focus of this talk

# Super-Kamiokande

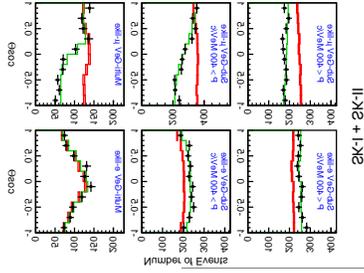
Ref. hep-ex/0501064, T. Kajita WHEPP9

- Neutrinos produced by the interaction of primary cosmic rays in the Earth's atmosphere. Wide range of energies and travel distances.
- Measurement relies on knowledge of flux w/o oscillations

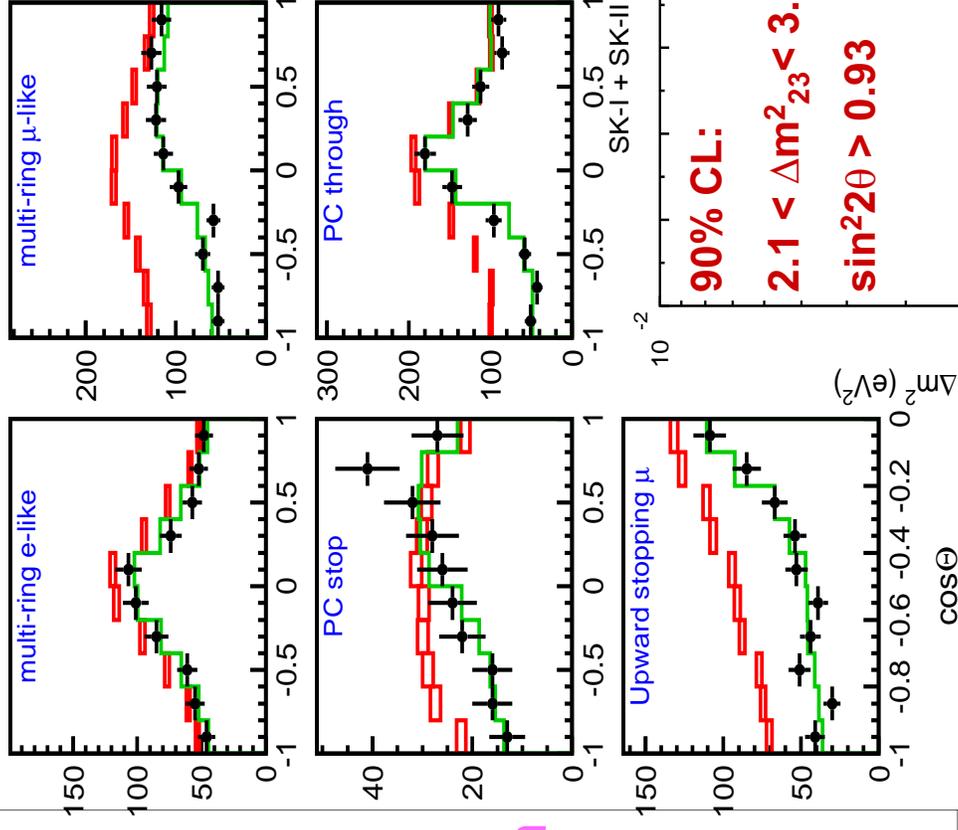


**2,700 m.w.e. overburden**  
**50 kT water Cherenkov detector**

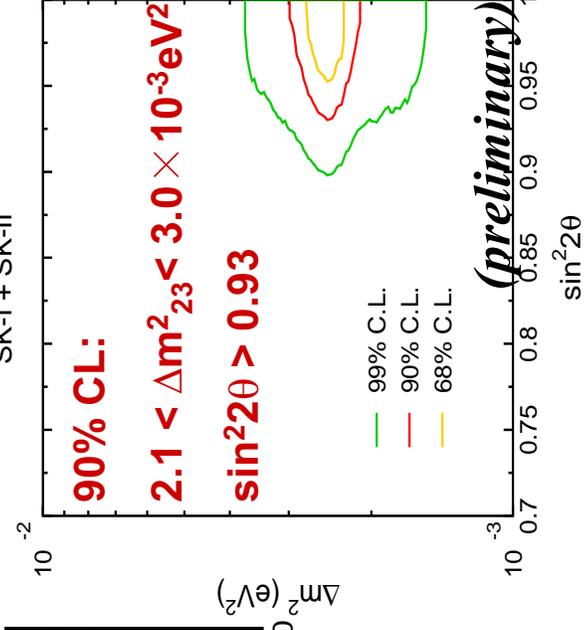




# Super-K (I+II) - Zenith



Up ... Down

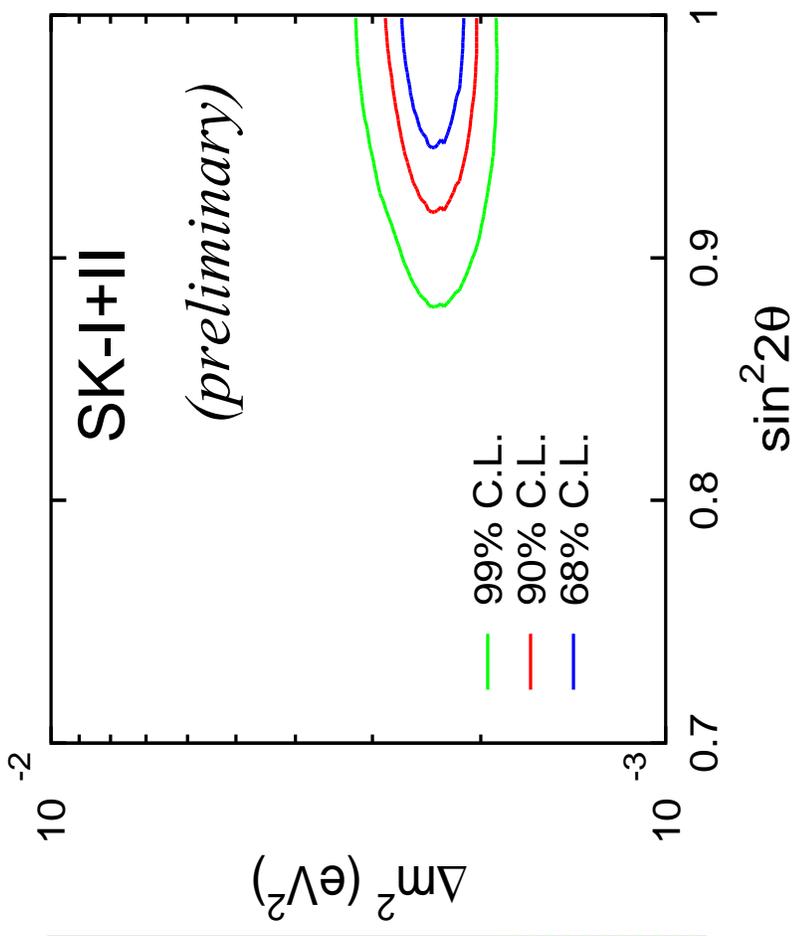
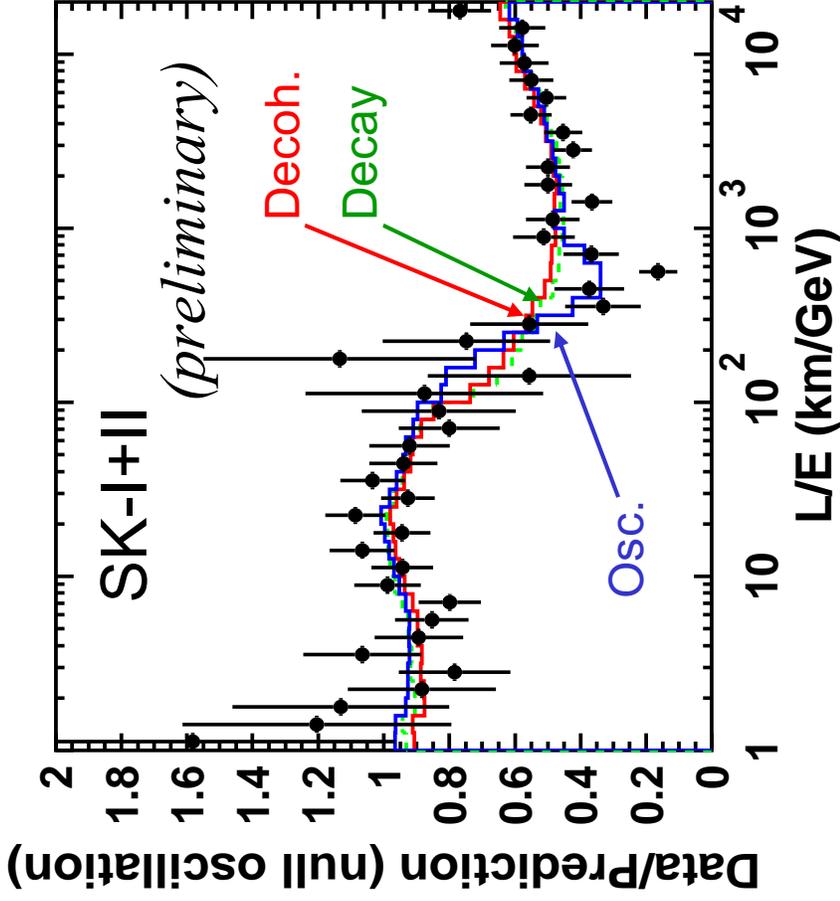


**Unoscillated MC**  
**Best fit:  $\Delta m^2_{23} = 2.5 \times 10^{-3} \text{eV}^2$ ,  $\sin^2 2\theta_{23} = 1$**   
**Data**



# Super-K (I+II) – L/E

- Consistent with Zenith angle analysis results.
- At 90% CL:  $2.0 < \Delta m^2_{23} < 2.9 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2\theta_{23} > 0.92$

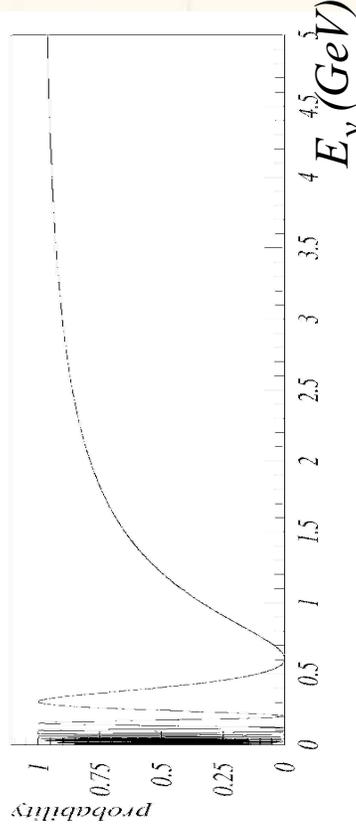


# K2K

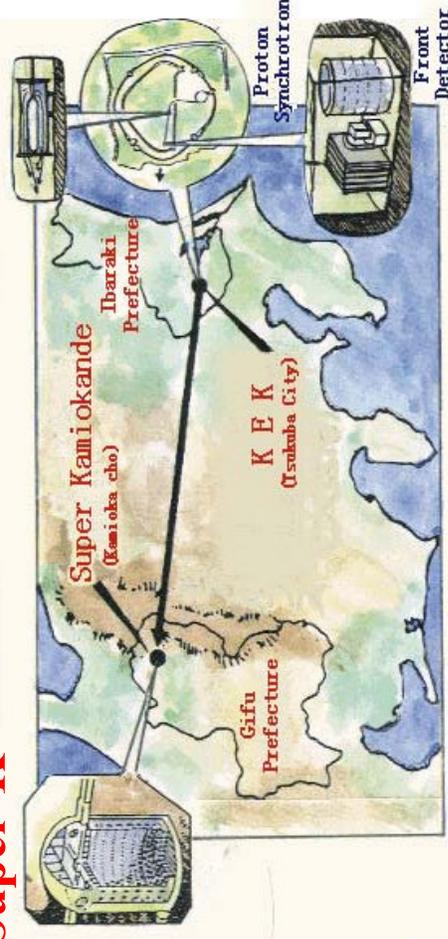
Ref. hep-ex/0411038, hep-ex/0512041

- First **long baseline** accelerator neutrino experiment.
- Accelerator produced  $\nu_\mu$  beam (**1.3 GeV mean energy**) travels **250 km** toward Super-K. The near detector is located 300 m from the production target.
- **Observe energy dependent  $\nu_\mu$  disappearance.**

$$\Delta m^2 = 0.003 \text{eV}^2, L = 250 \text{km}$$



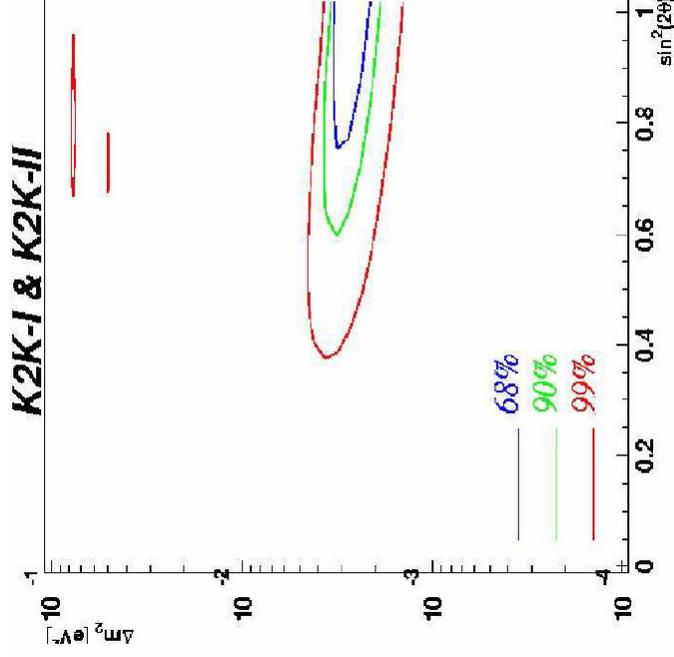
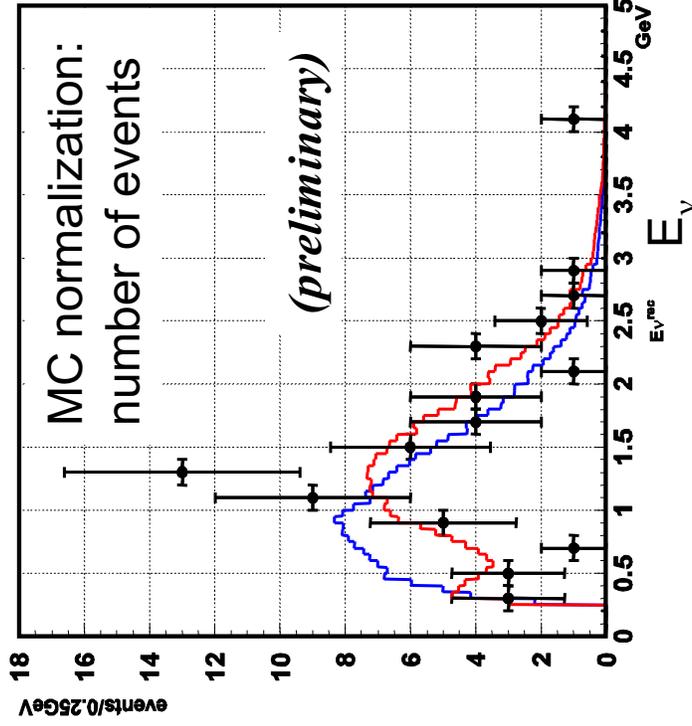
## Super-K



**Near Detectors (SciBar,  
1 kton water Cherenkov)**

# K2K - Results

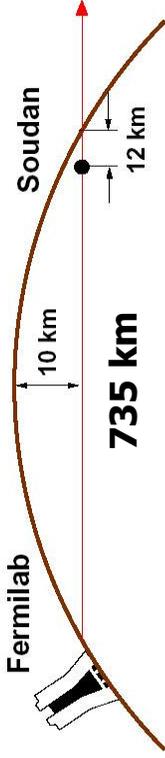
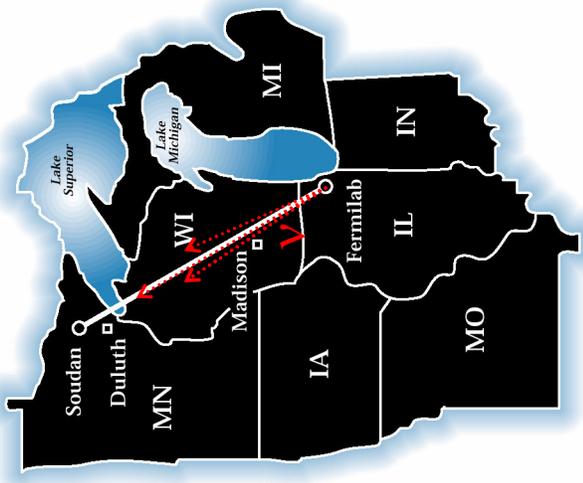
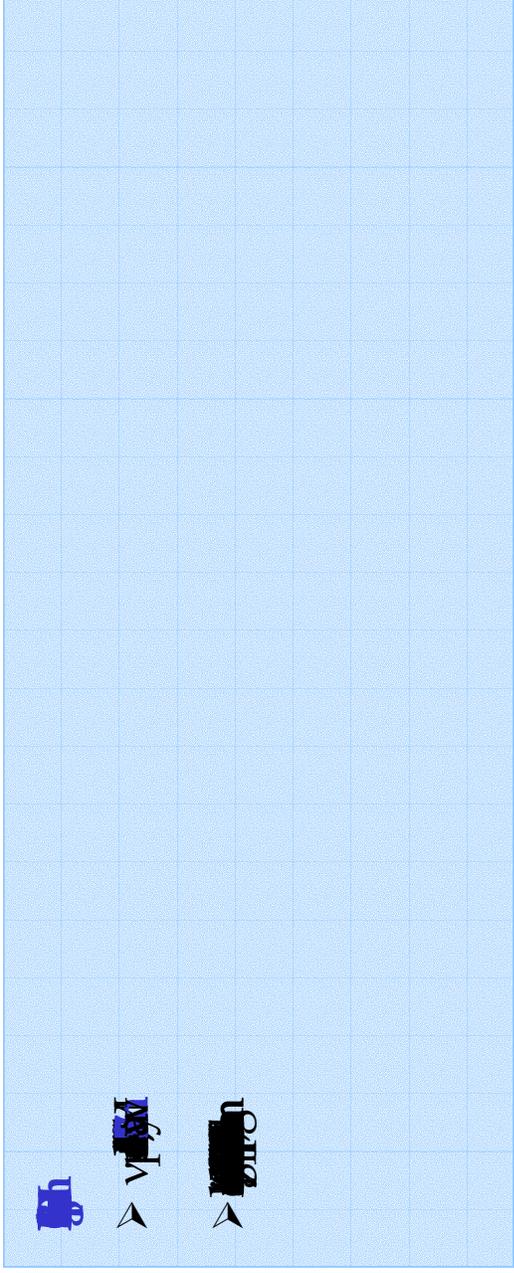
- Observe deficit of  $\nu_\mu$  at SK and distortion in the energy spectrum:  
**112 events at SK, expect  $155.9^{+13.6}_{-15.6}$**
- Well described by  $\nu_\mu \rightarrow \nu_\tau$  oscillation



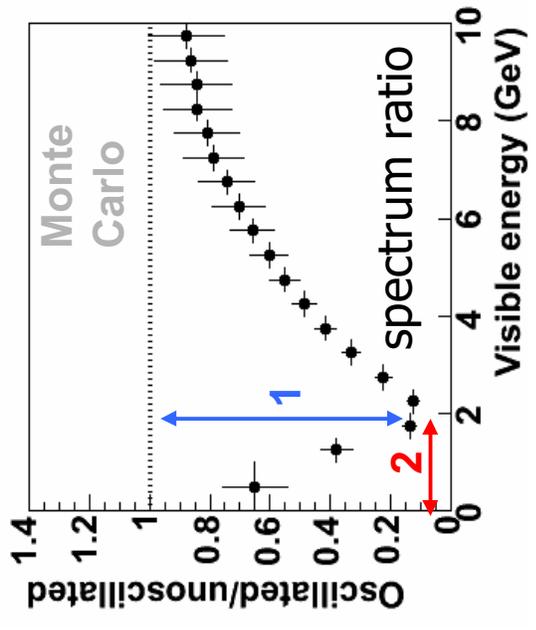
Best fit:  $\Delta m^2_{23} = 2.76 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2\theta_{23} = 1$   
 For  $\sin^2 2\theta_{23} = 1$ : 90% CL  $1.9 < \Delta m^2_{23} < 3.5 \times 10^{-3} \text{ eV}^2$

# MINOS

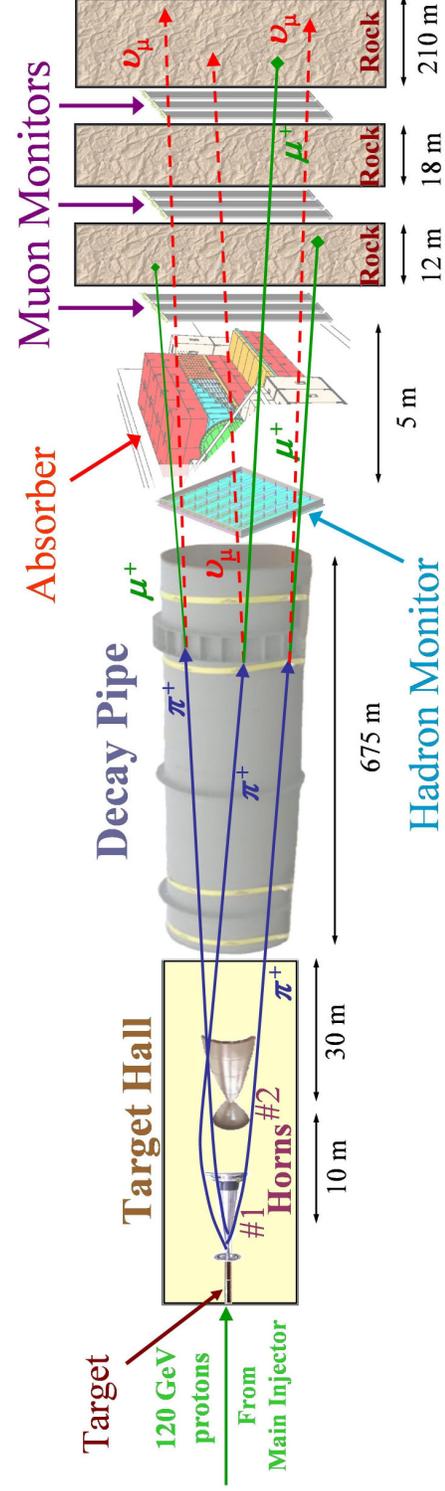
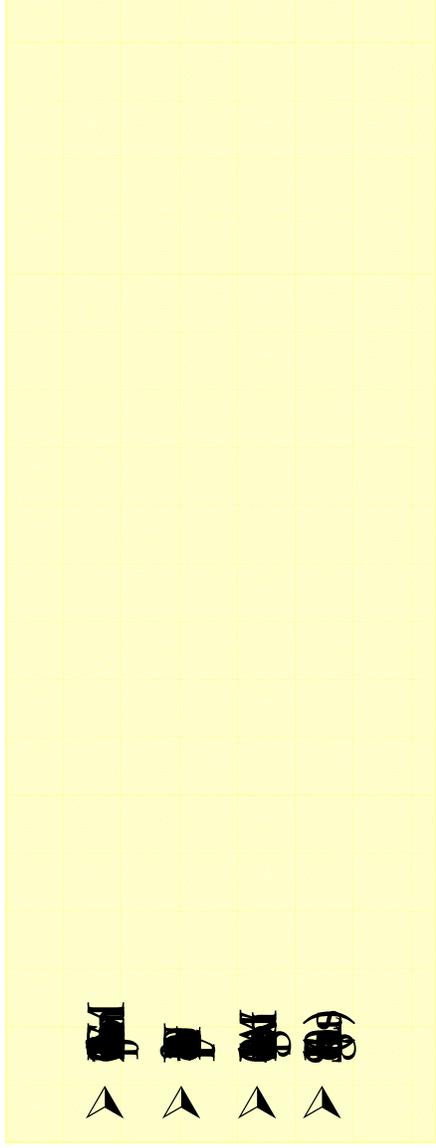
## Main Injector Neutrino Oscillation Search



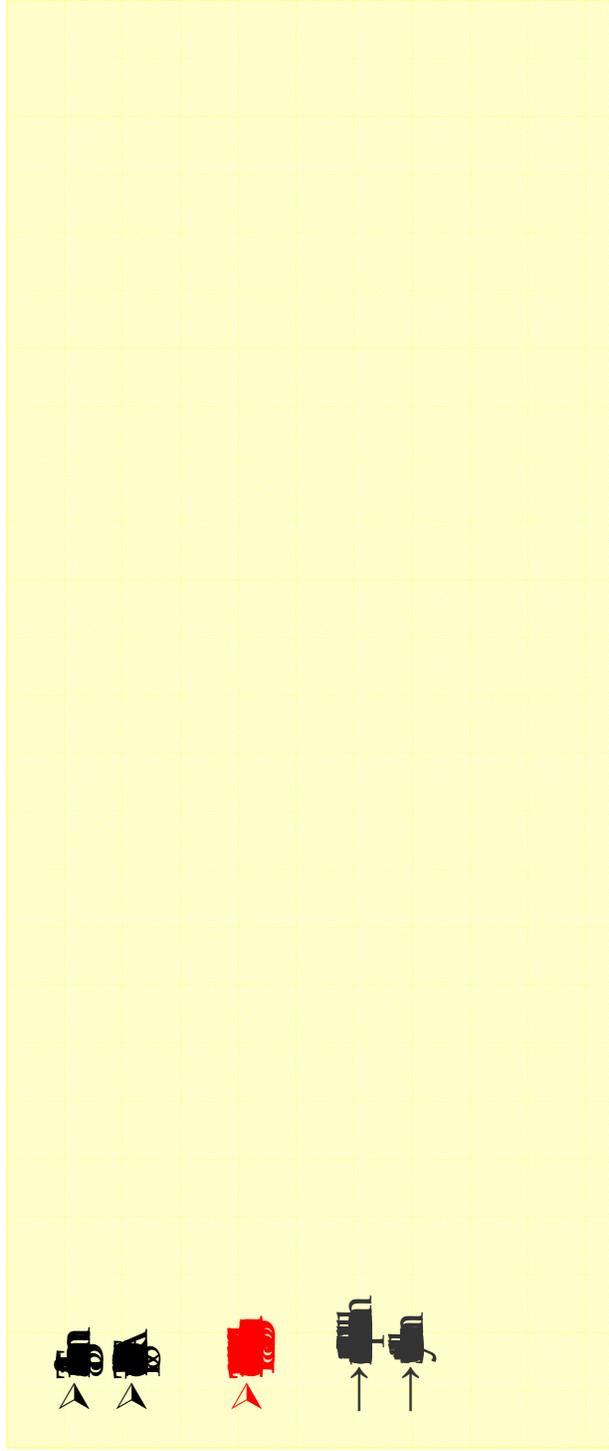
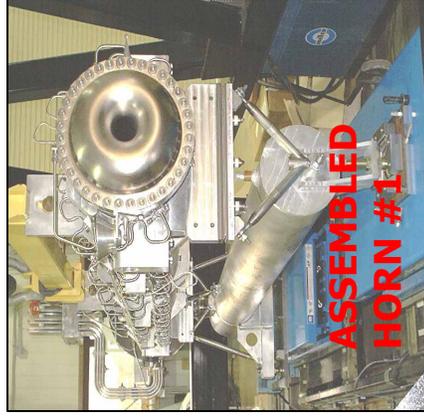
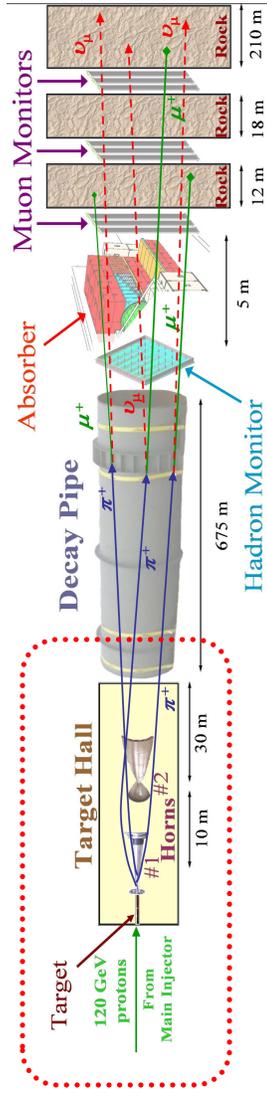
$$P(\nu_{\mu} \rightarrow \nu_{\mu}) = 1 - \sin^2 2\theta_{23} \sin^2 \left( 1.27 \frac{\Delta m_{23}^2 (eV^2) L(km)}{E_{\nu} (GeV)} \right)$$



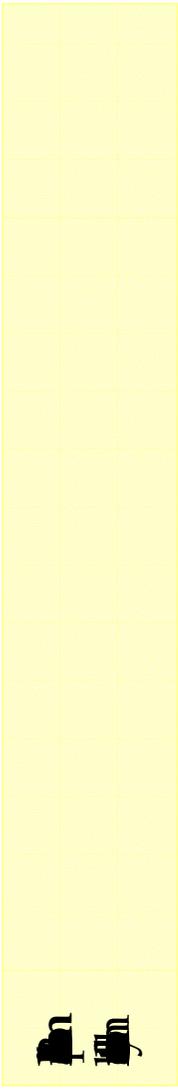
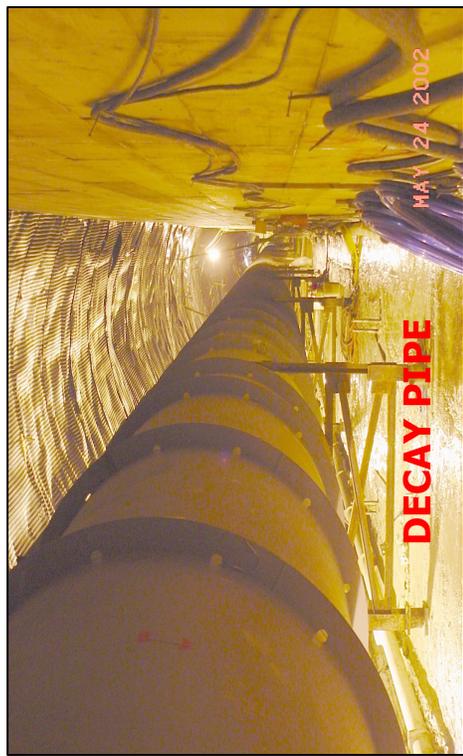
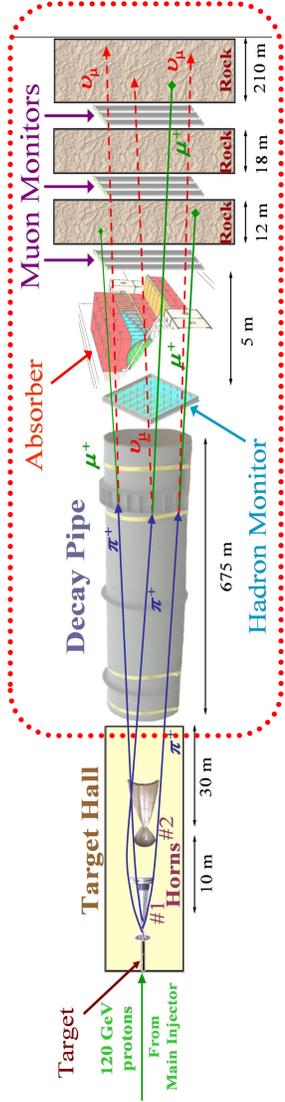
# NuMI Beamline



# NuMI Target and Horns



# Hadron and Muon Monitors

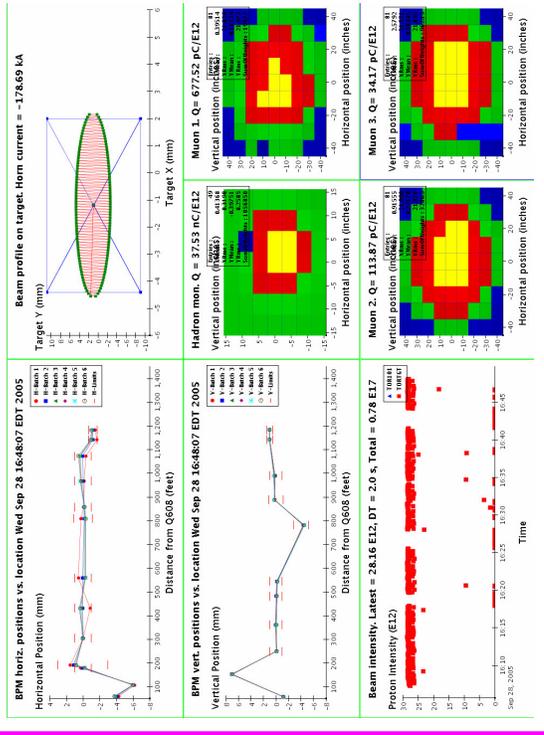


**HADRON MONITOR**



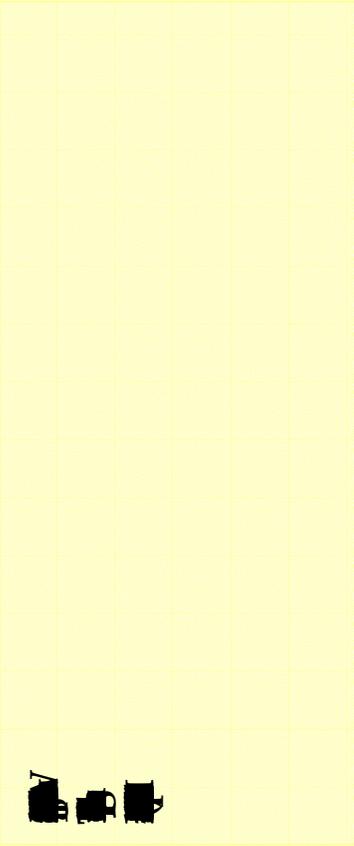
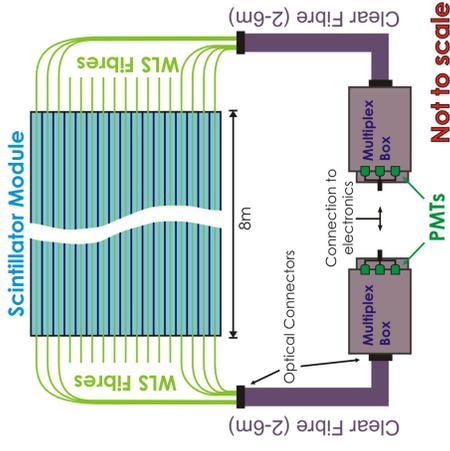
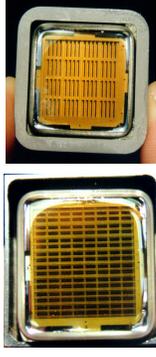
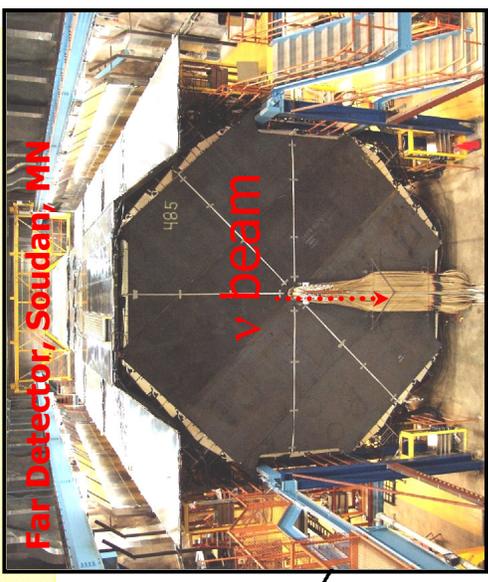
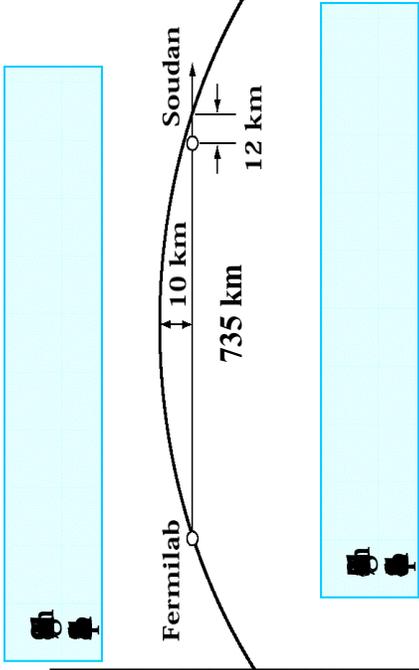
**MUON MONITOR**

## BEAM MONITORING

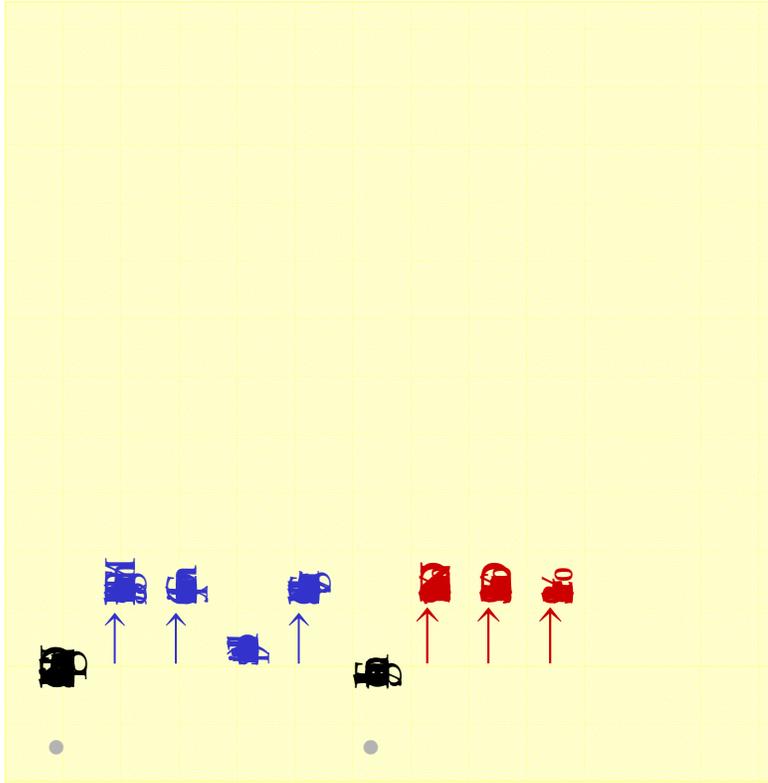
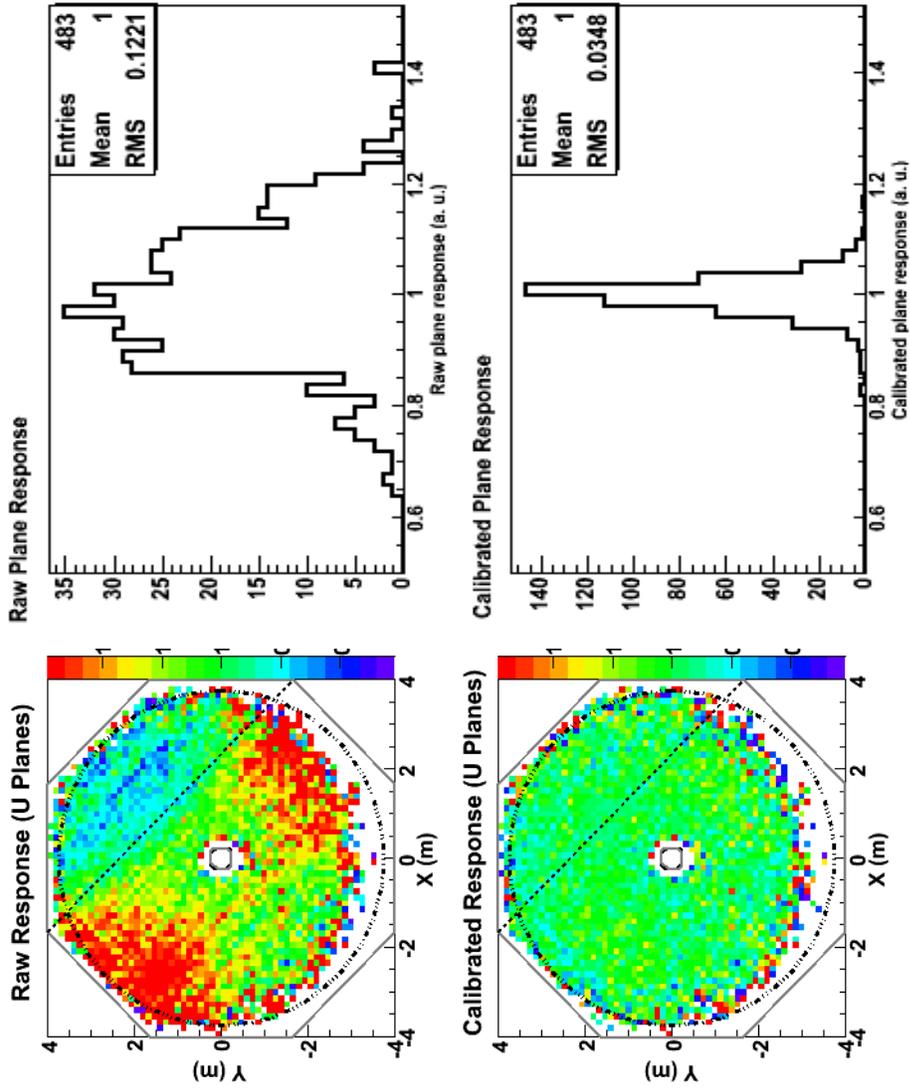


# The MINOS Detectors

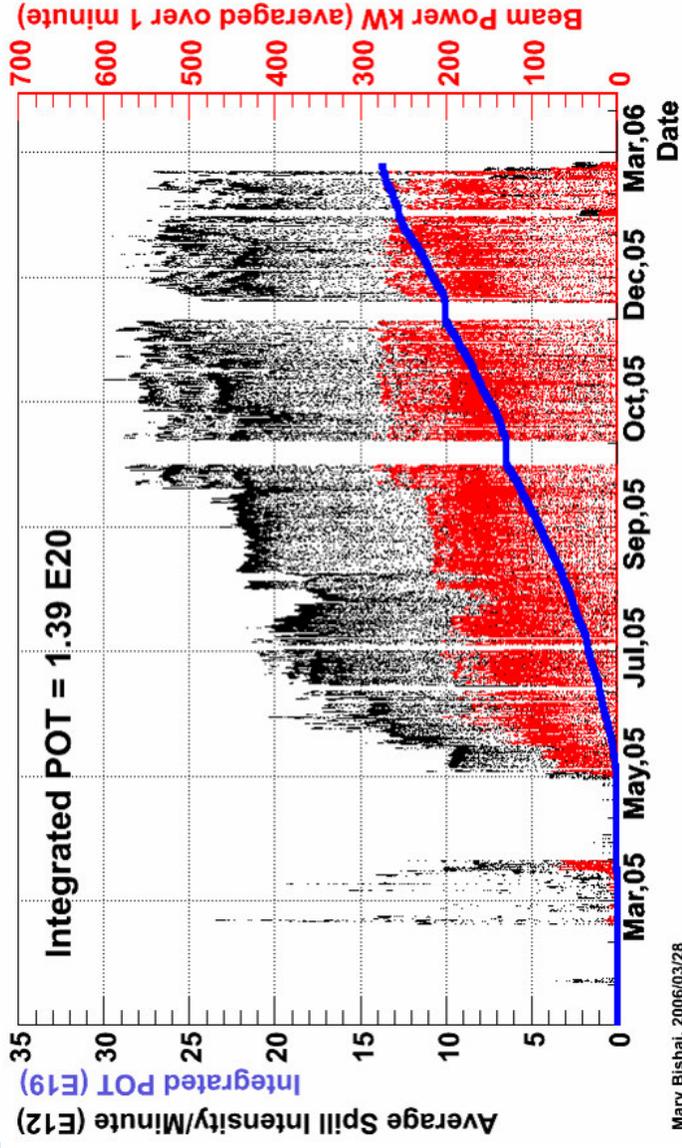
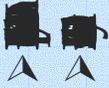
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# Calibration



# Start-Up and Running



Mary Bishai, 2006/03/28

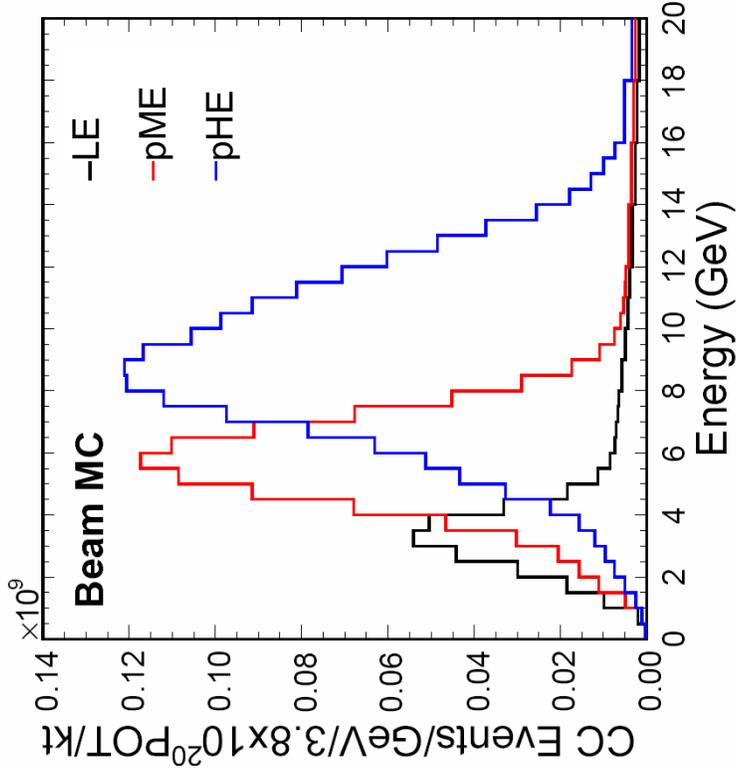


# Neutrino Energy Spectrum



$\delta m^2_{23}$

8.5%  $\nu_\mu + \nu_\mu$  (6.5%  $\nu_\mu$ ), 1.5%  $\nu_e + \nu_e$

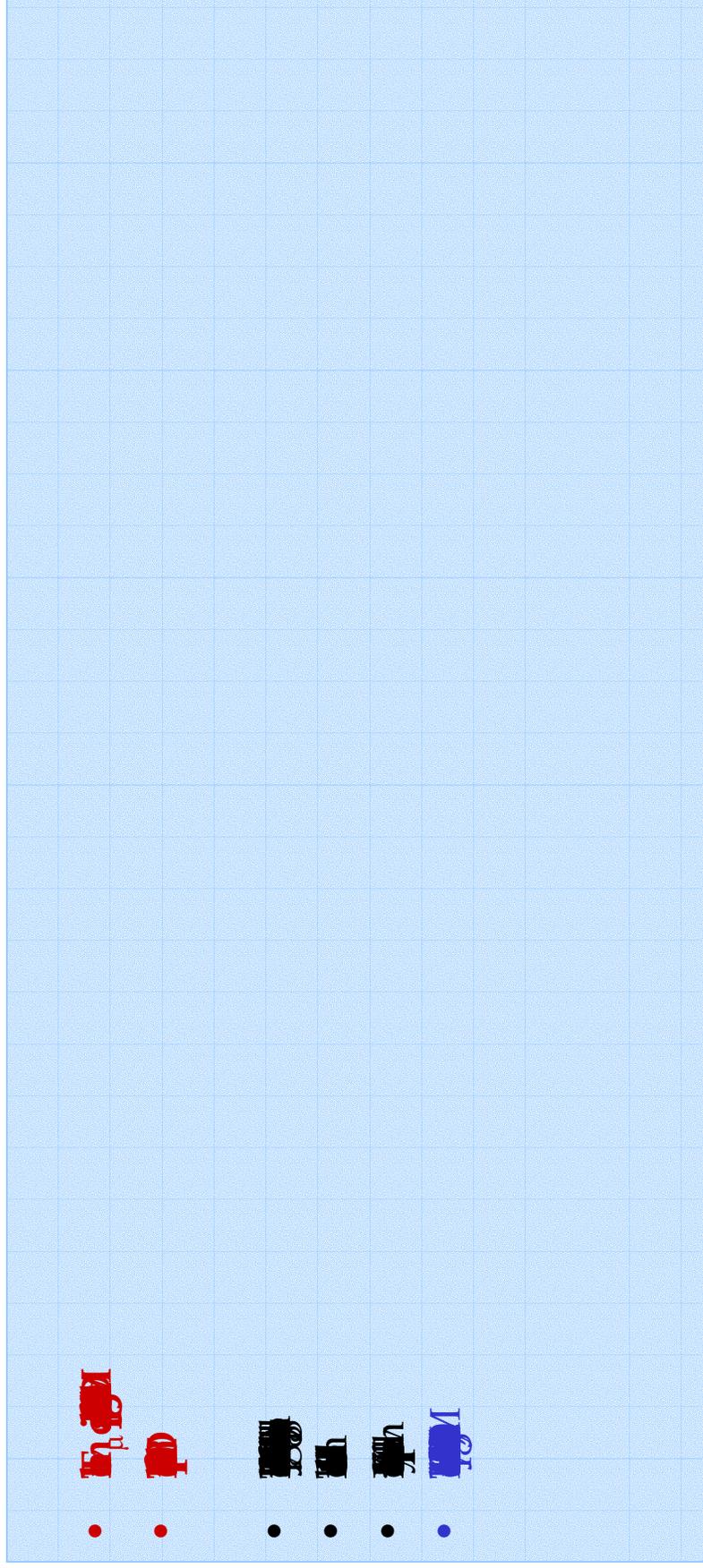


Beam	Target z position (cm)	FD Events (no osc) per 1 x 10 <sup>20</sup> POT
LE-10	-10	390
pME	-100	970
pHE	-250	1340

Events in fiducial volume



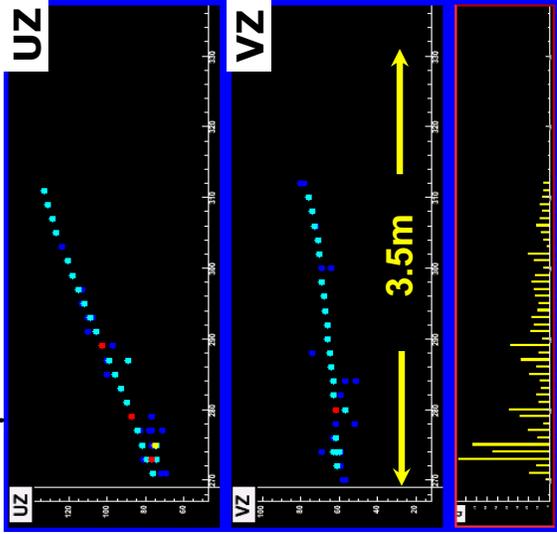
# Oscillation Analysis



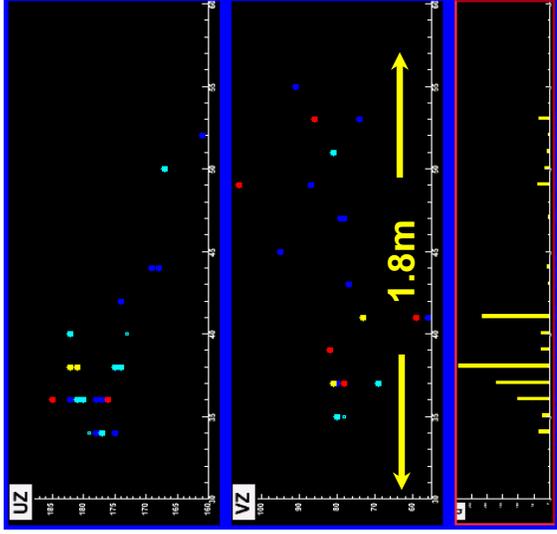
# Event Topologies

Monte Carlo

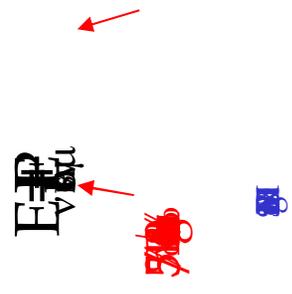
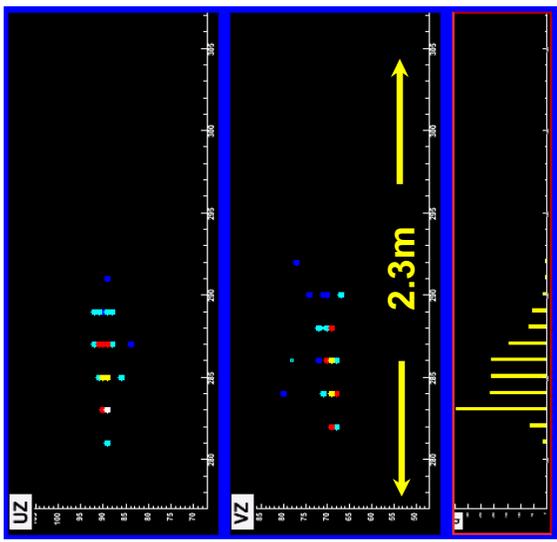
$\nu_{\mu}$  CC Event



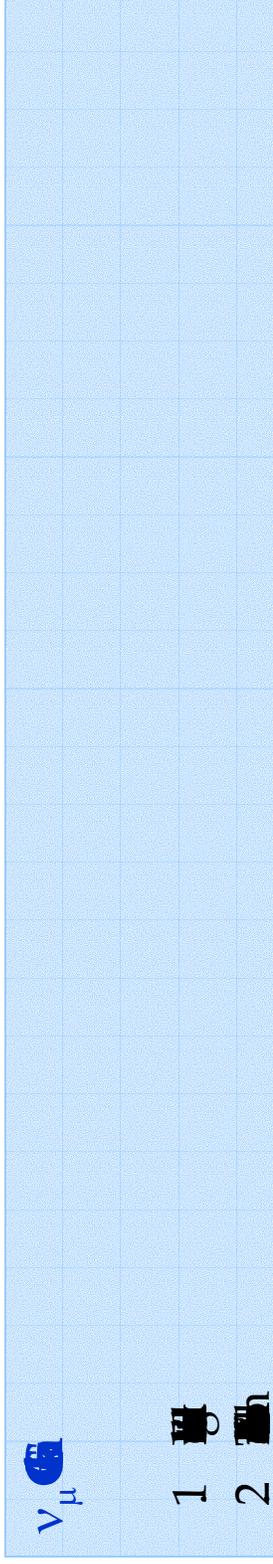
NC Event



$\nu_e$  CC Event



# Data Selection



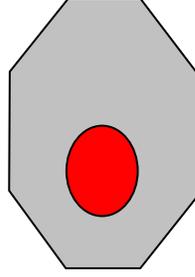
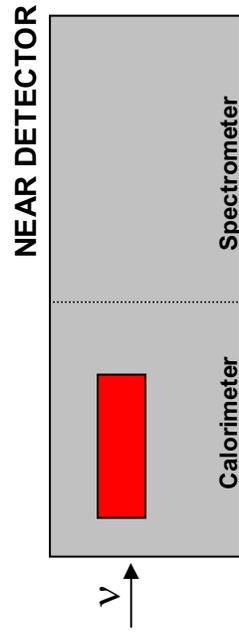
**NEAR:**  $1\text{m} < z < 5\text{m}$  ( $z$  from front face)

$R < 1\text{m}$  from beam center

**FAR:**  $z > 50\text{cm}$  from front face

$z > 2\text{m}$  from rear face

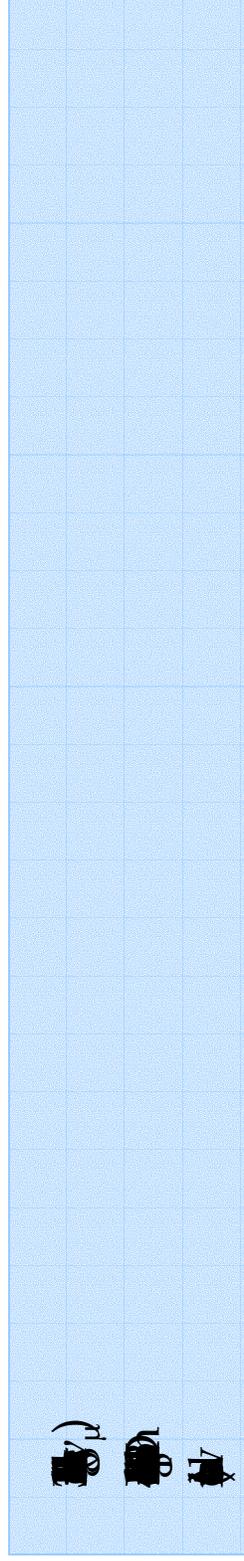
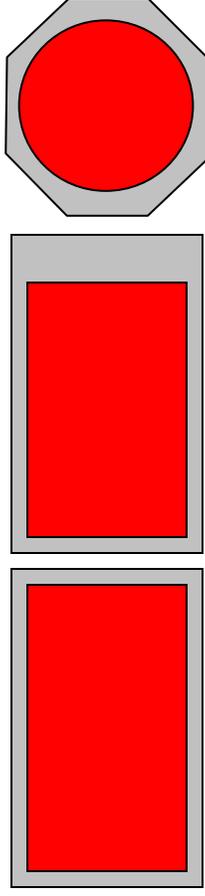
$R < 3.7\text{m}$  from center of detector



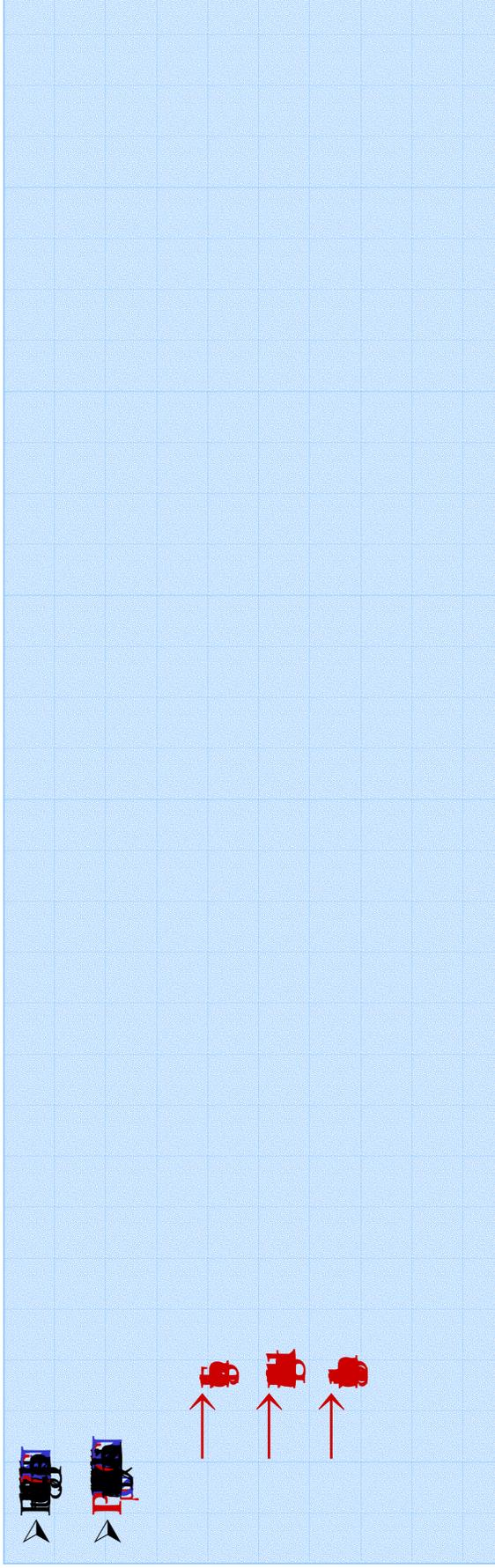
Fiducial Volume



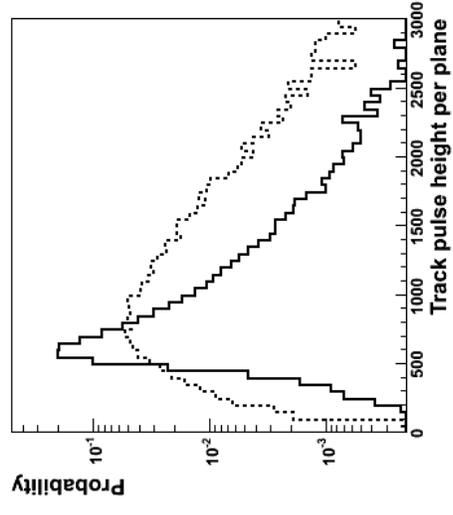
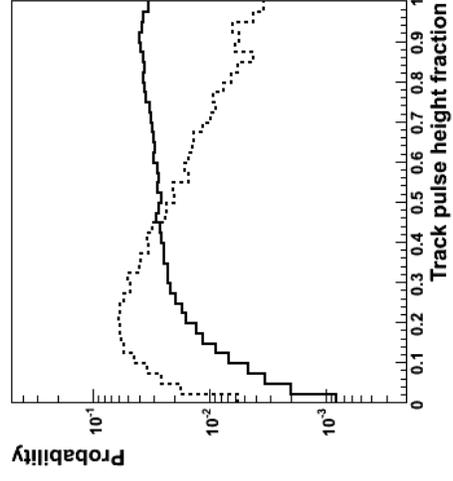
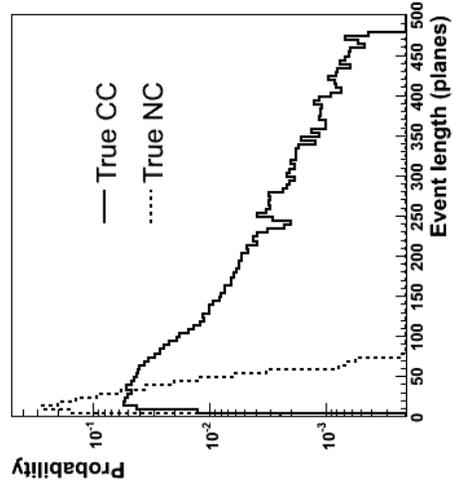
FAR DETECTOR



# Selecting CC events



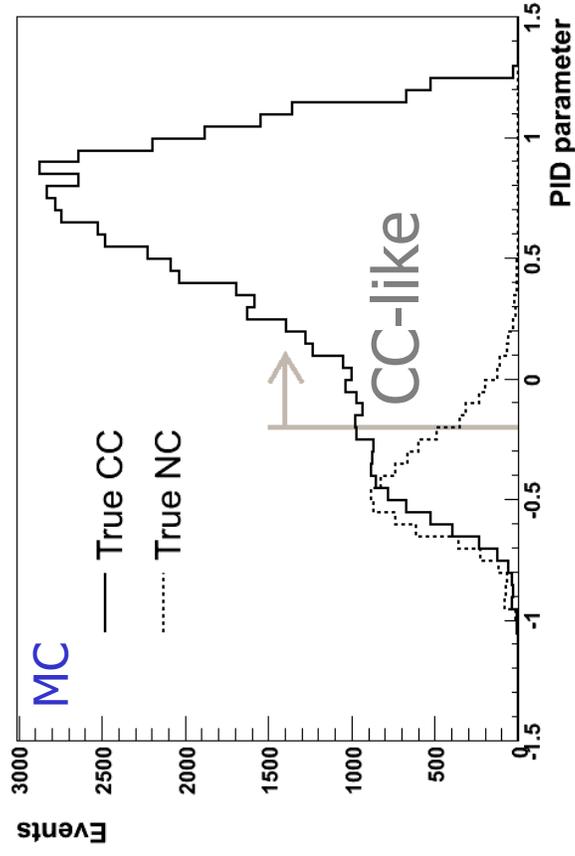
Input variables for PDF-based event selection



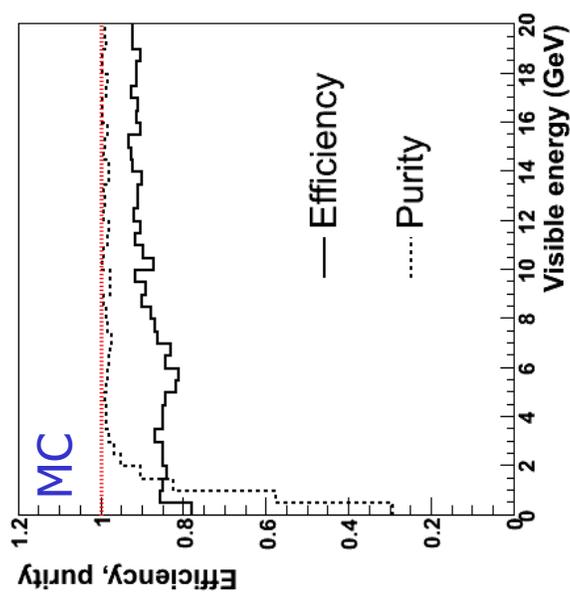
# CC Selection Efficiencies

$$PID = -(\sqrt{-\log(P_\mu)} - \sqrt{-\log(P_{NC})})$$

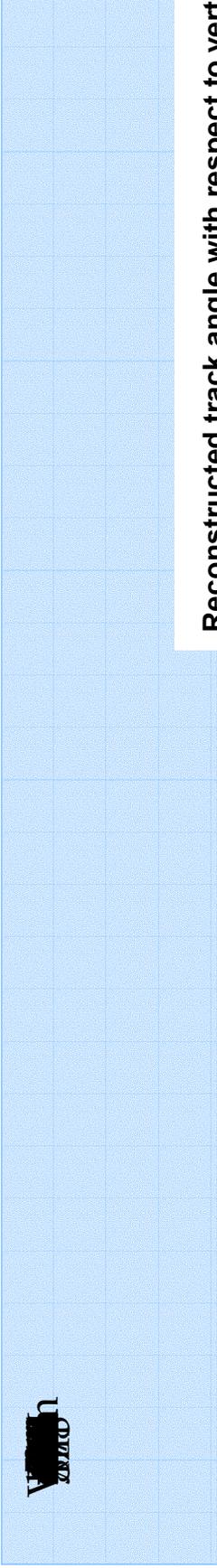
PDF PID parameter distribution for true CC and NC events



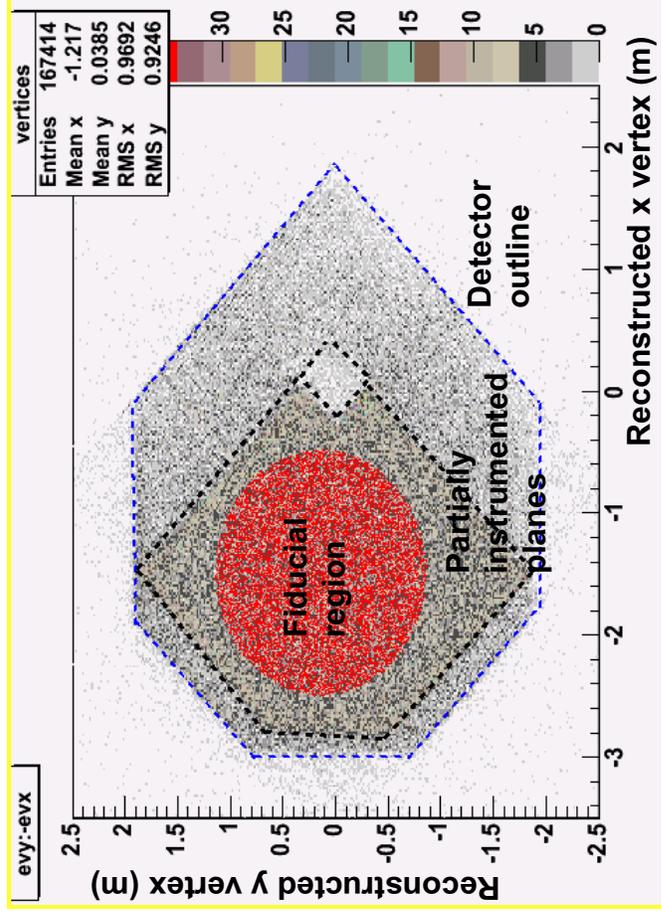
CC selection efficiencies and purities



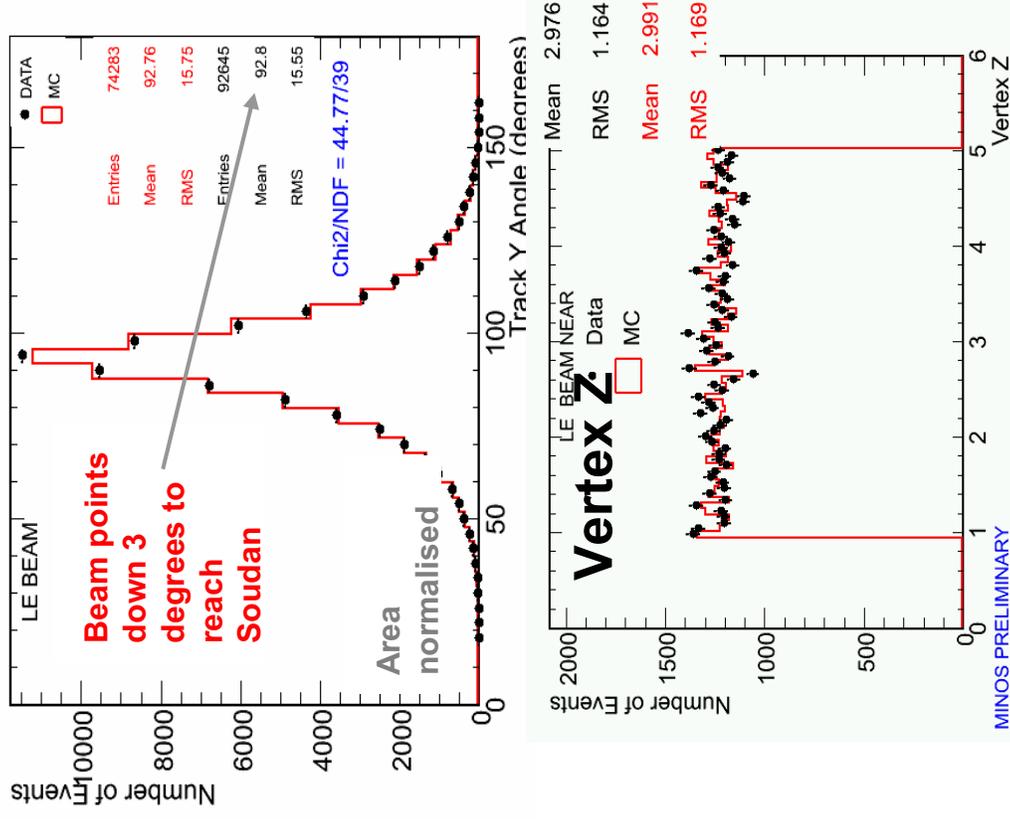
# Near Detector Distributions



Distribution of reconstructed event vertices in the x-y plane

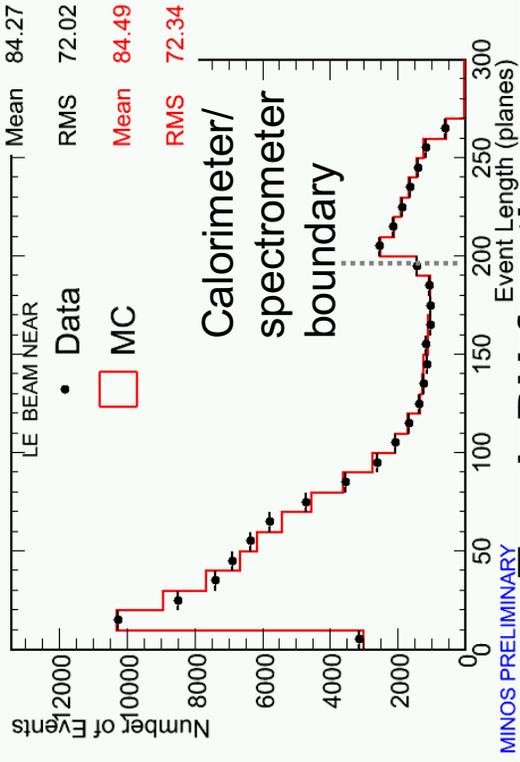


Reconstructed track angle with respect to vertical

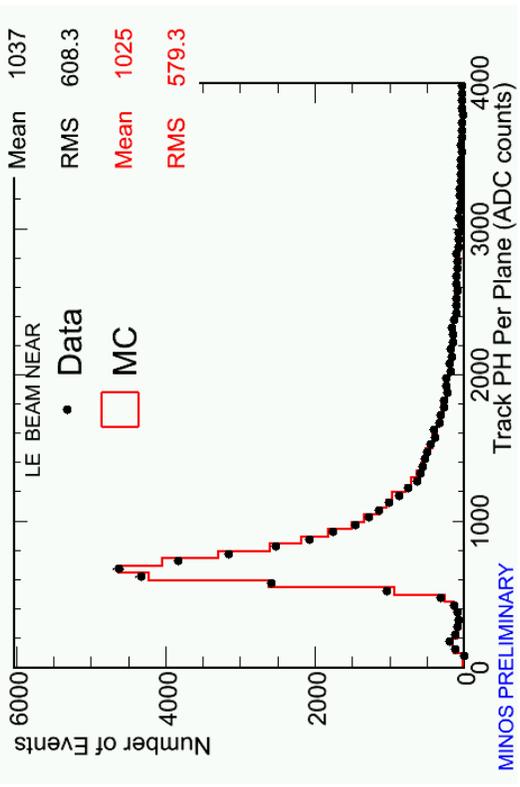


# Near Detector - PID Distributions

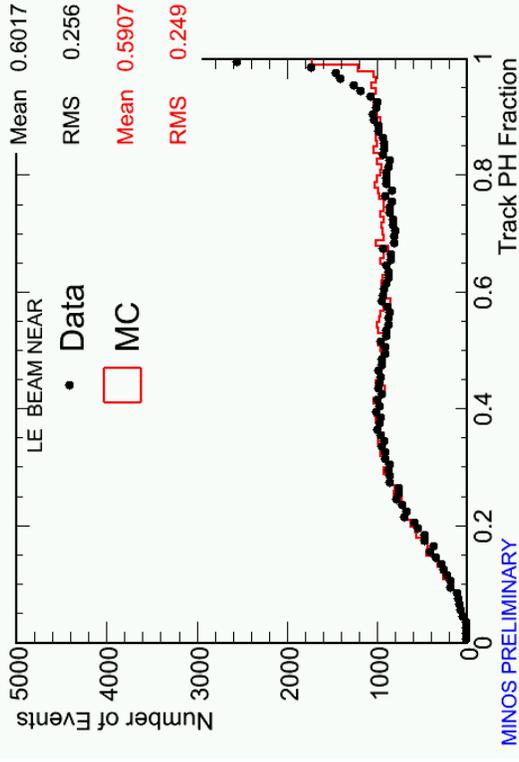
## Event length



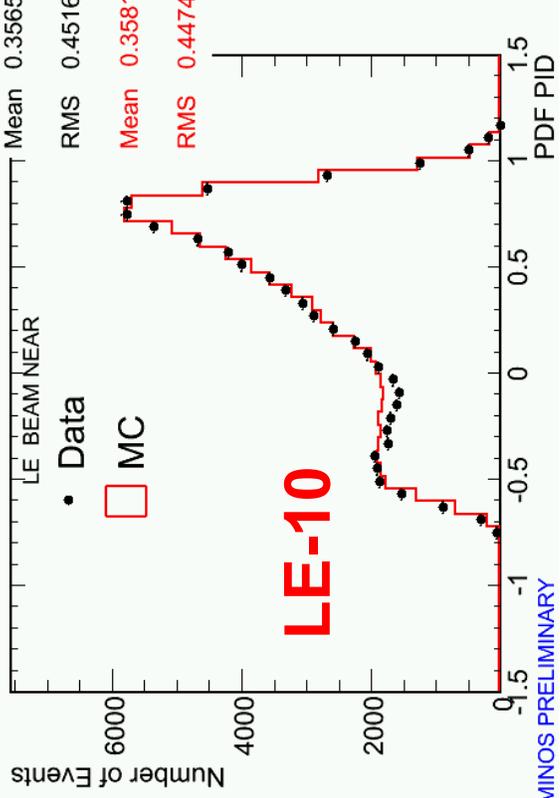
## Track PH per plane



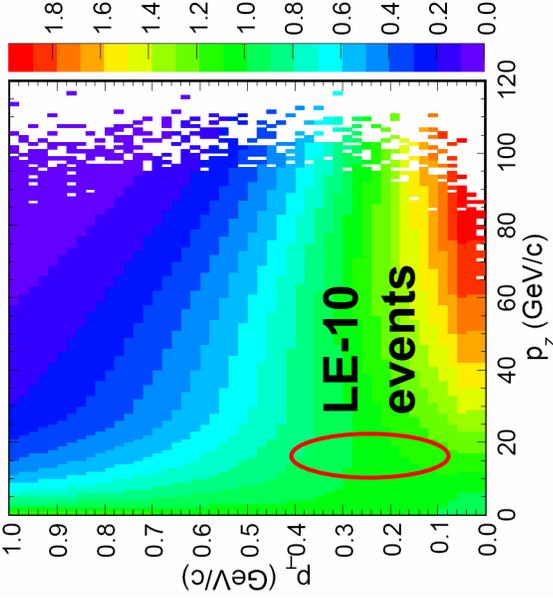
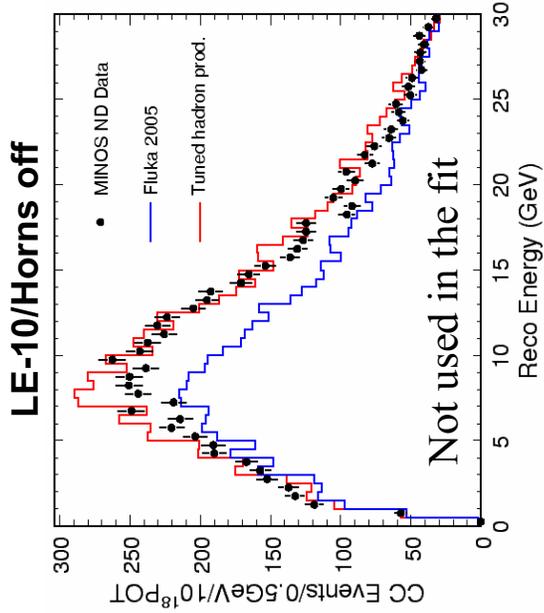
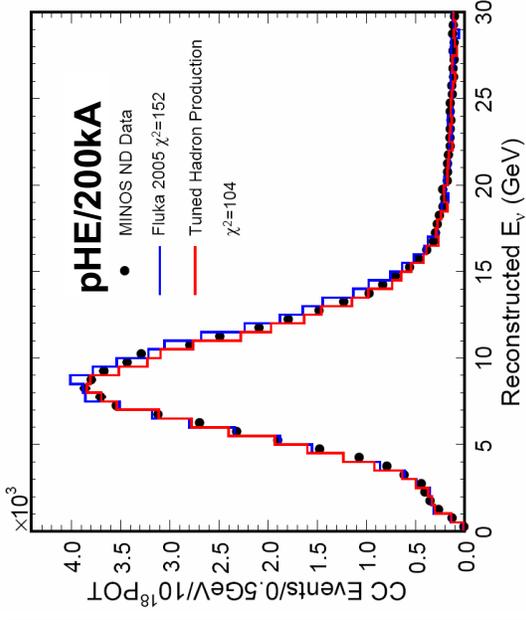
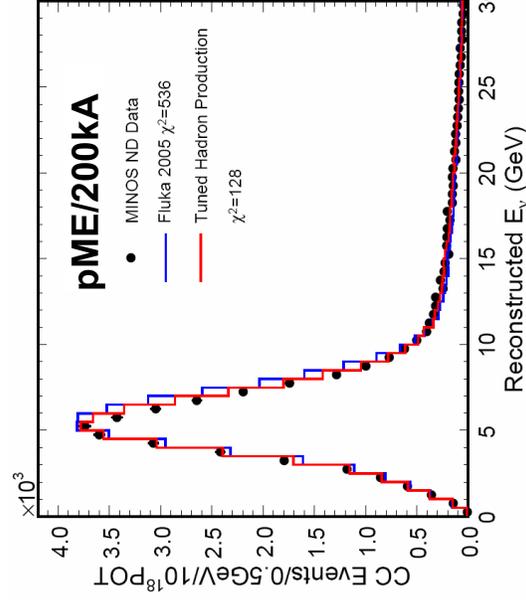
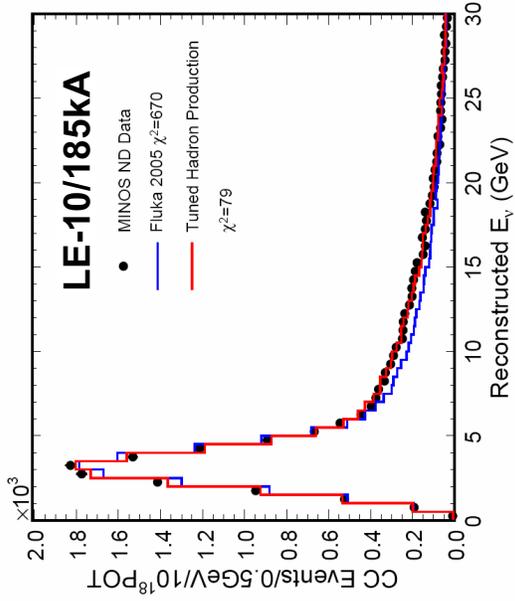
## Track PH fraction



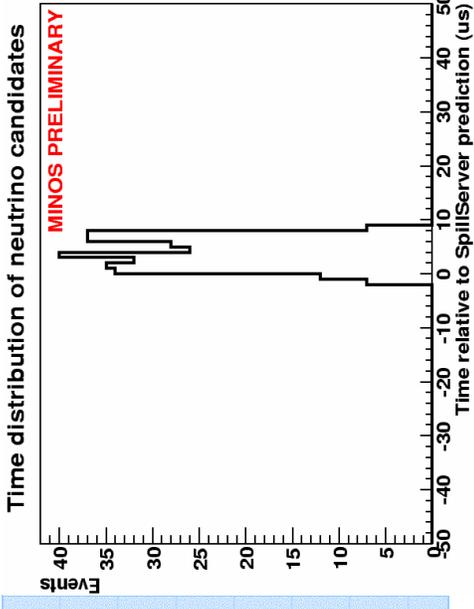
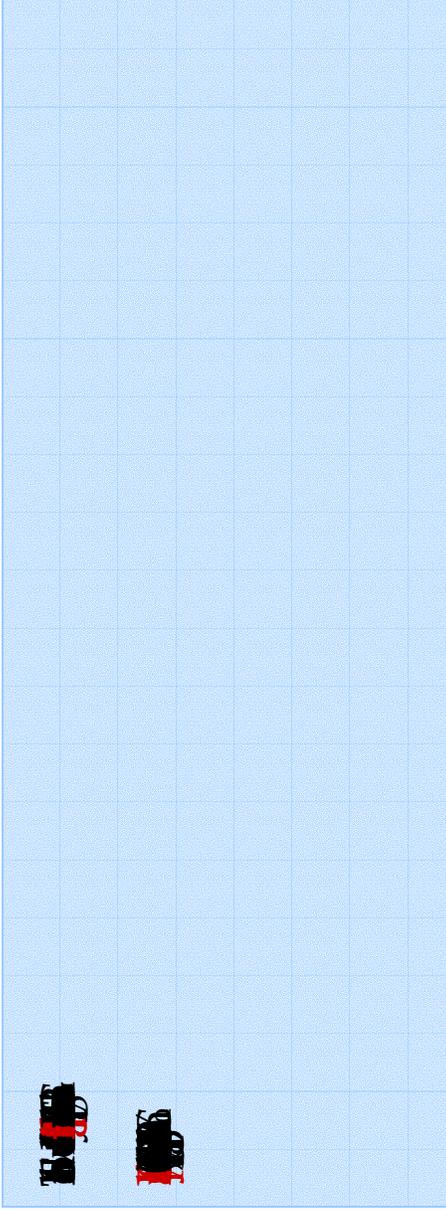
**LE-10**



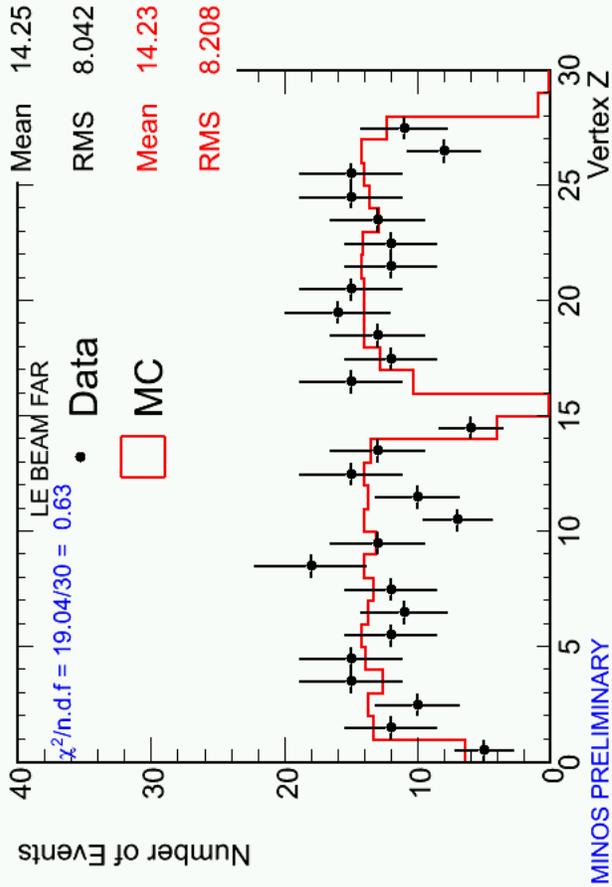
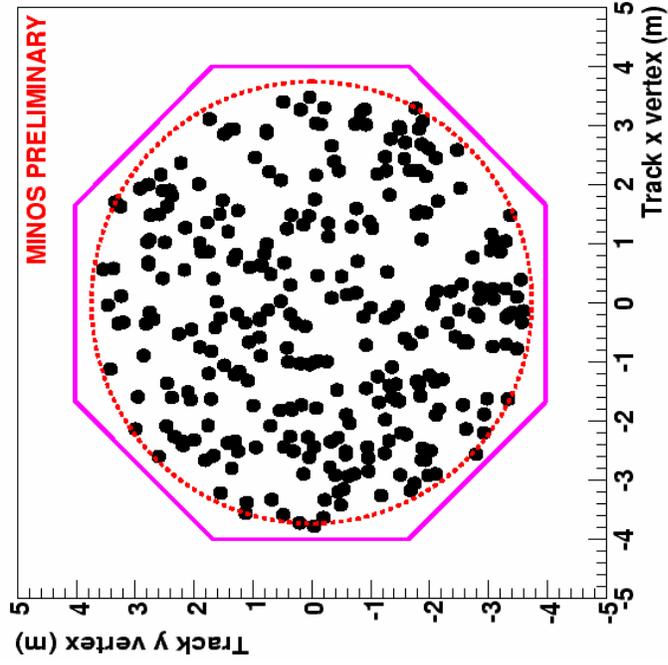
# Hadron Production Tuning



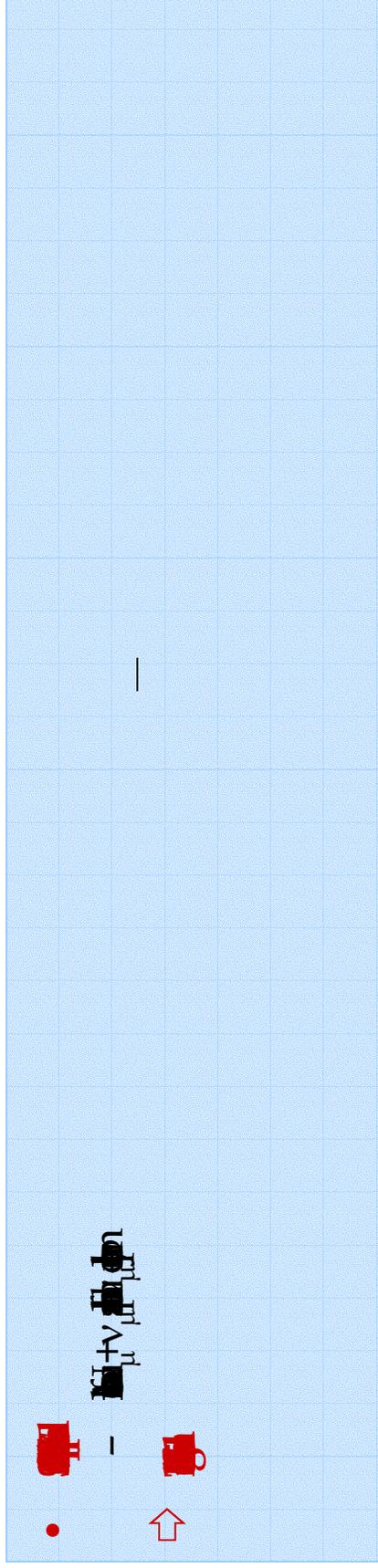
# Far Detector Data



Reconstructed track vertices of neutrino candidates



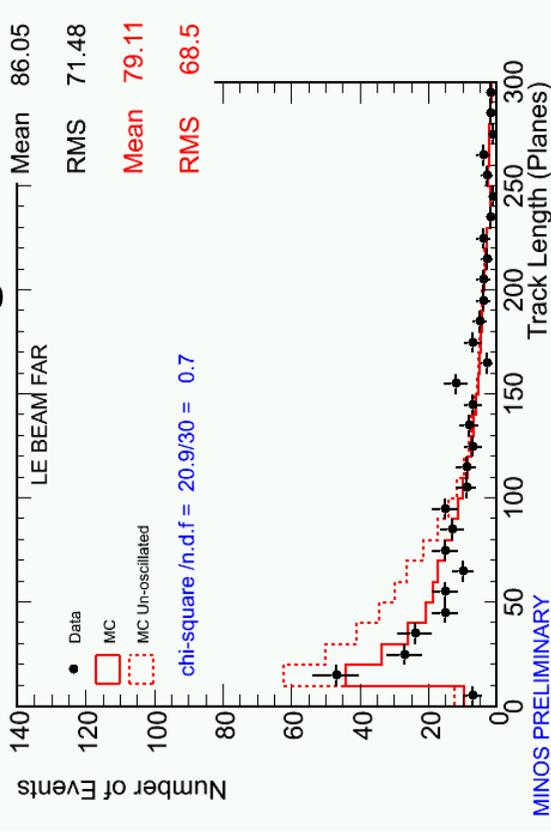
# Observed/Expected Rates



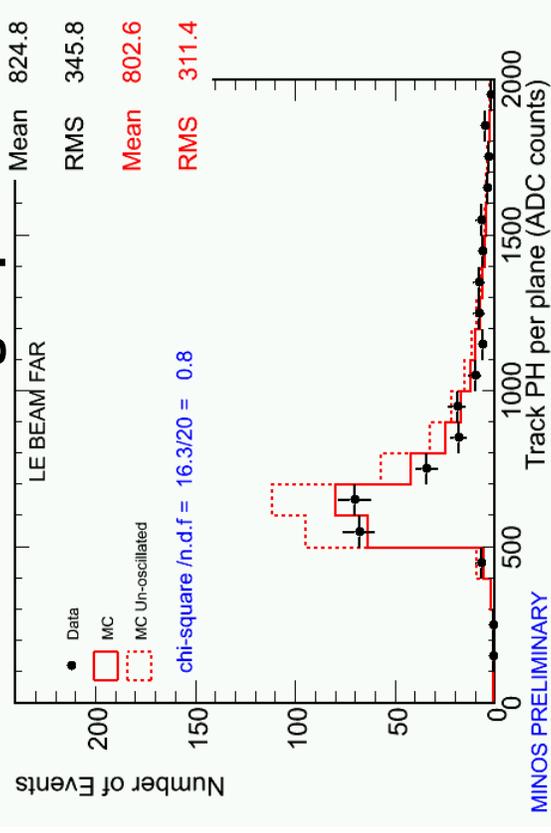
Data sample	observed	expected	ratio	significance
All CC-like events ( $\nu_{\mu} + \bar{\nu}_{\mu}$ )	204	298±15	0.69	4.1σ
$\nu_{\mu}$ only (<30 GeV)	166	249±14	0.67	4.0σ
$\nu_{\mu}$ only (<10 GeV)	92	177±11	0.52	5.0σ

# Far Detector - PID Distributions

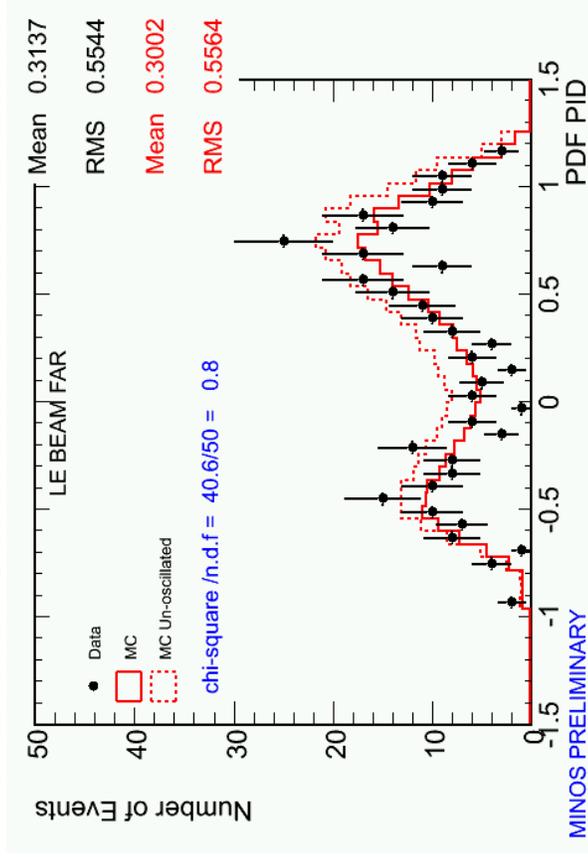
## Track Length



## Track Pulse Height per Plane

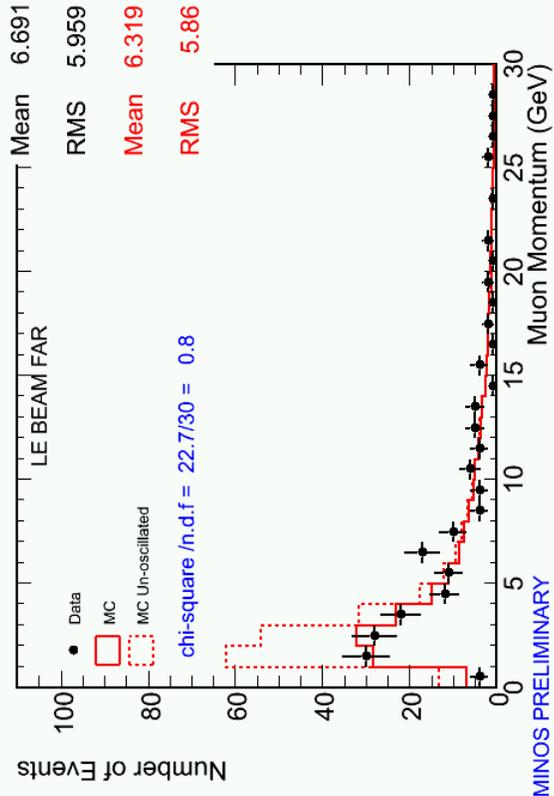


## Particle Identification Parameter

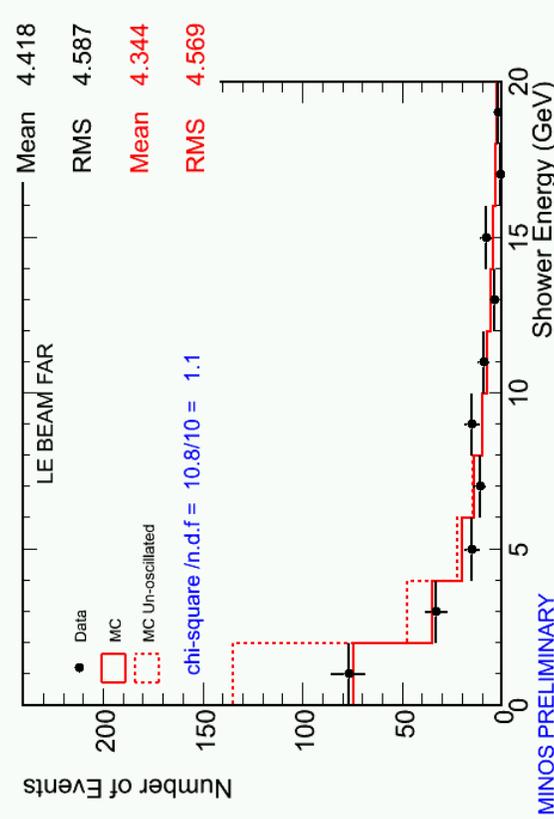


# Far Detector - Physics Distributions

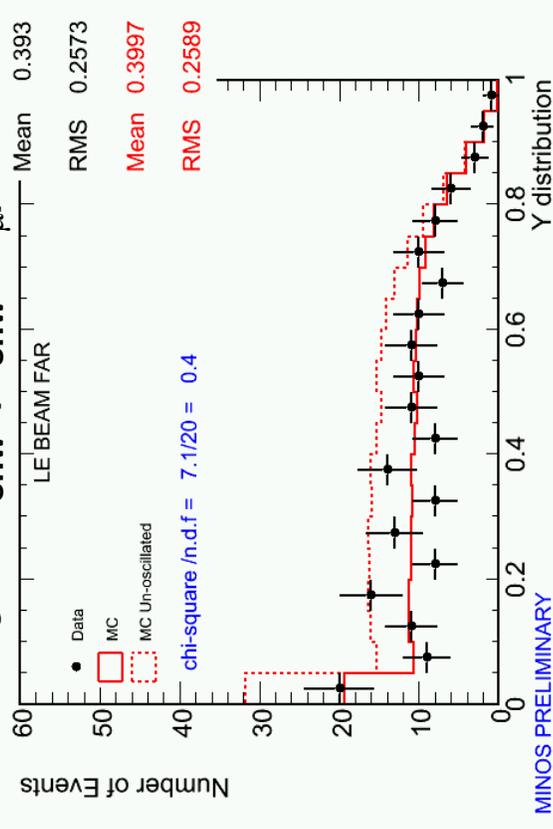
## Muon Momentum (GeV/c)



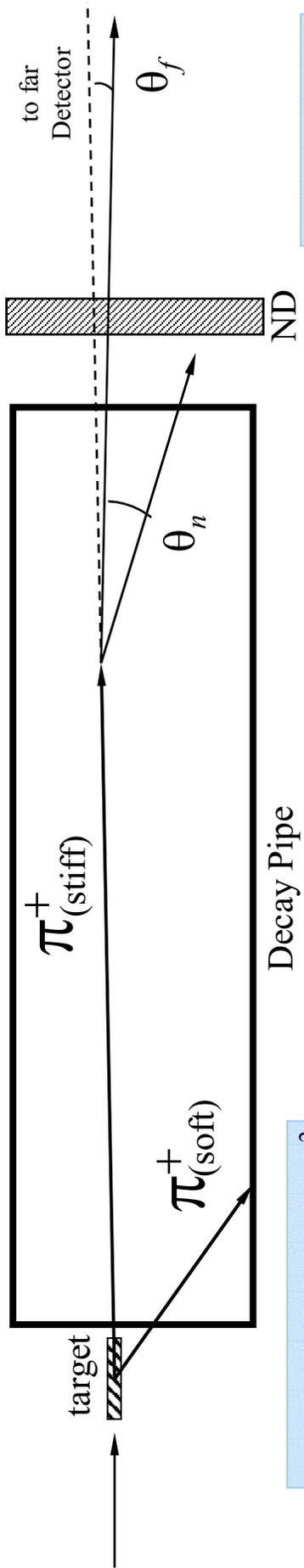
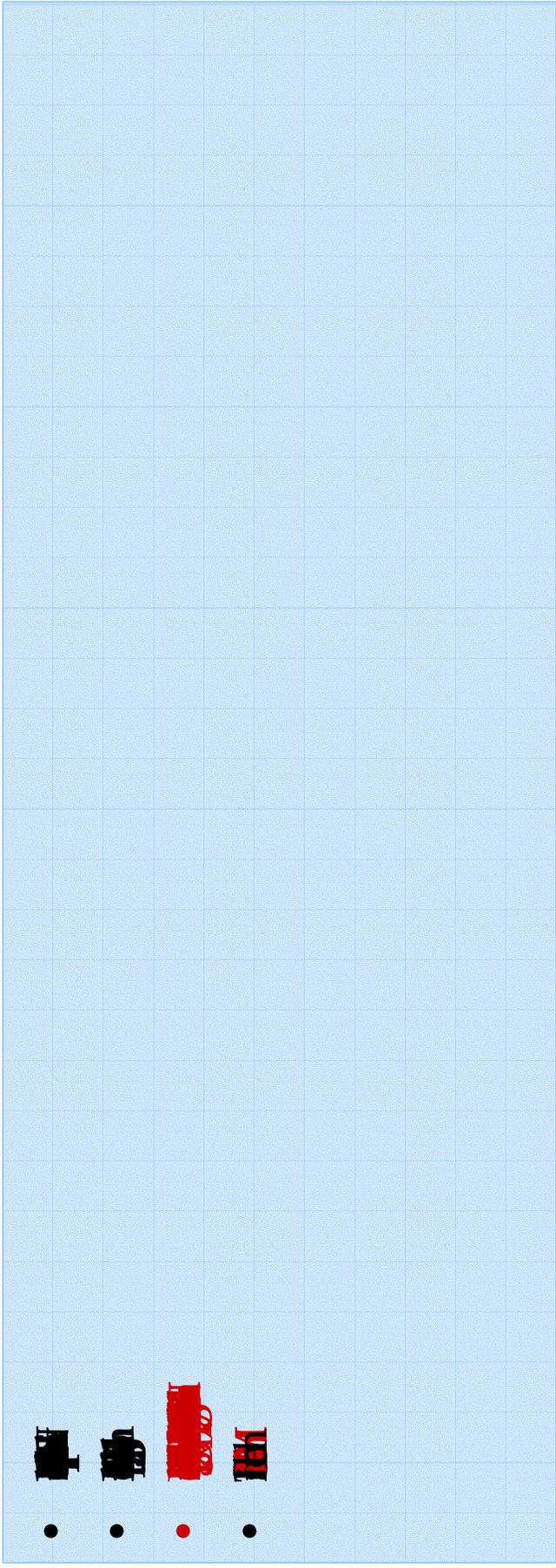
## Shower Energy (GeV)



$$y = E_{shw} / (E_{shw} + P_{\mu})$$



# Predicting the Far Detector Spectrum

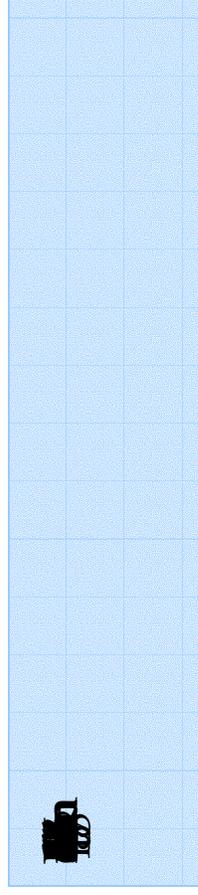
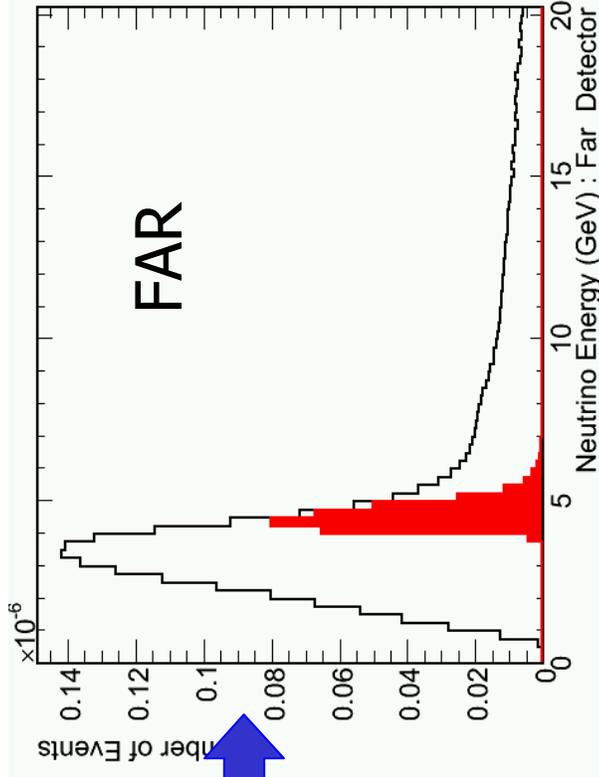
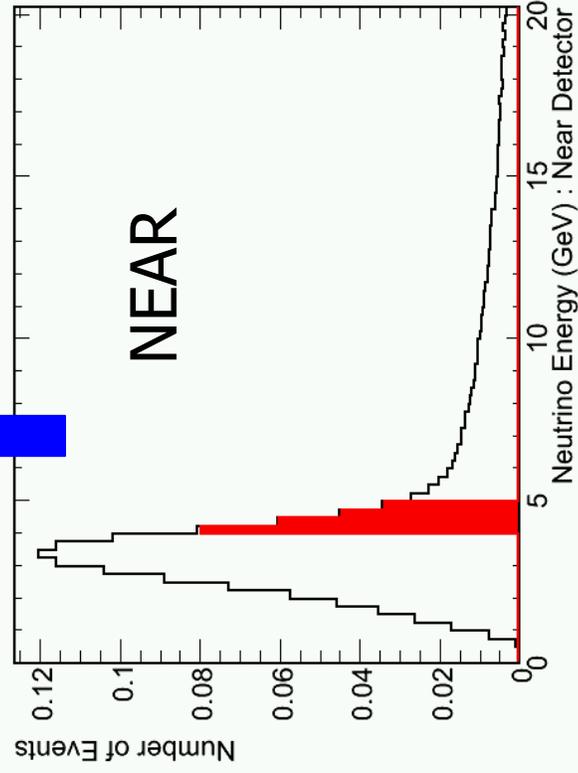
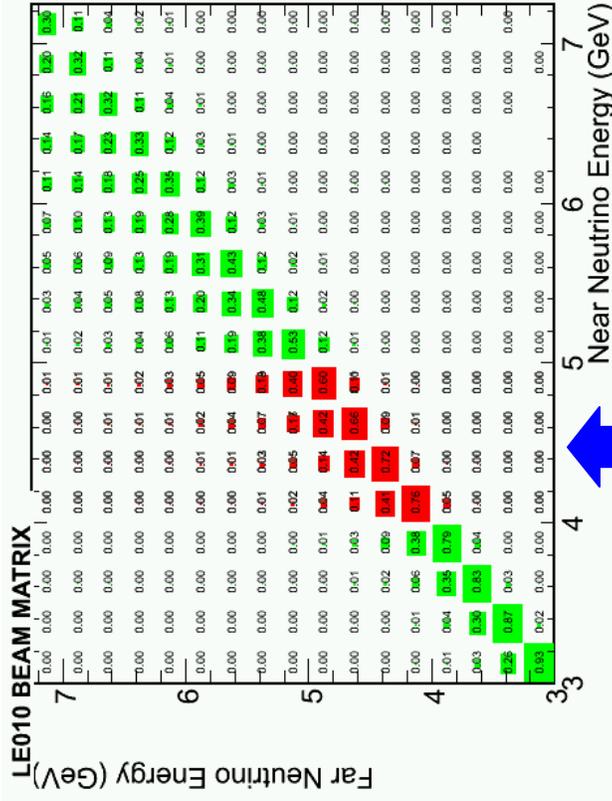


$$Flux \propto \frac{1}{L^2} \left( \frac{1}{1 + \gamma^2 \theta^2} \right)^2$$

$$E_\nu = \frac{0.43 E_\pi}{1 + \gamma^2 \theta^2}$$

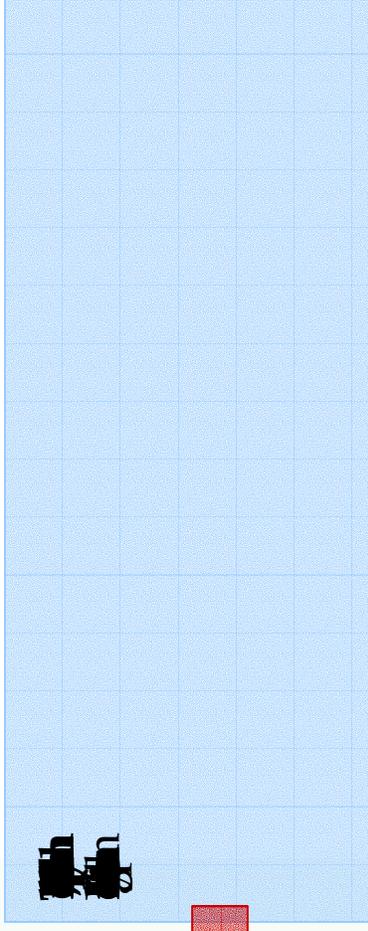
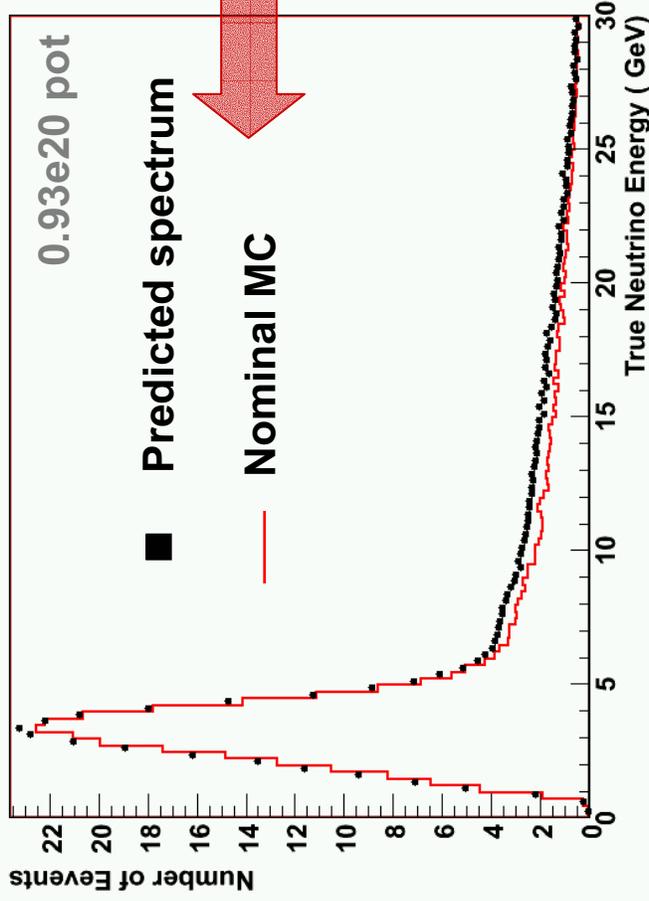


# Beam Matrix Method

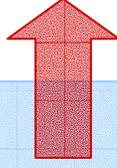
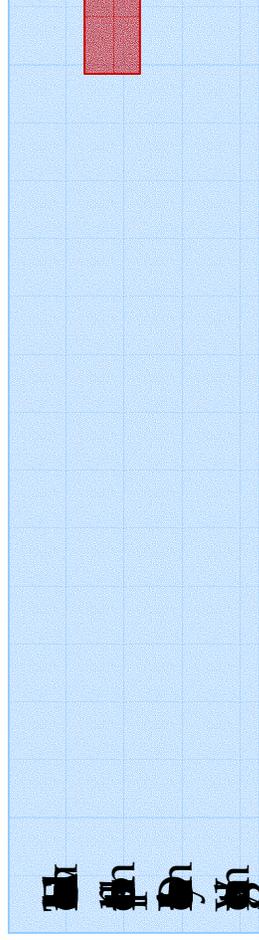
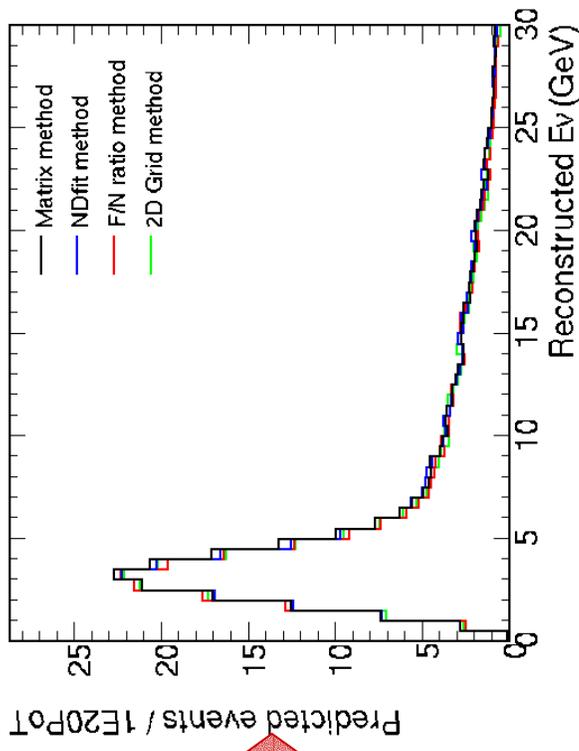


# Predicted Spectrum

Predicted FD unoscillated spectra

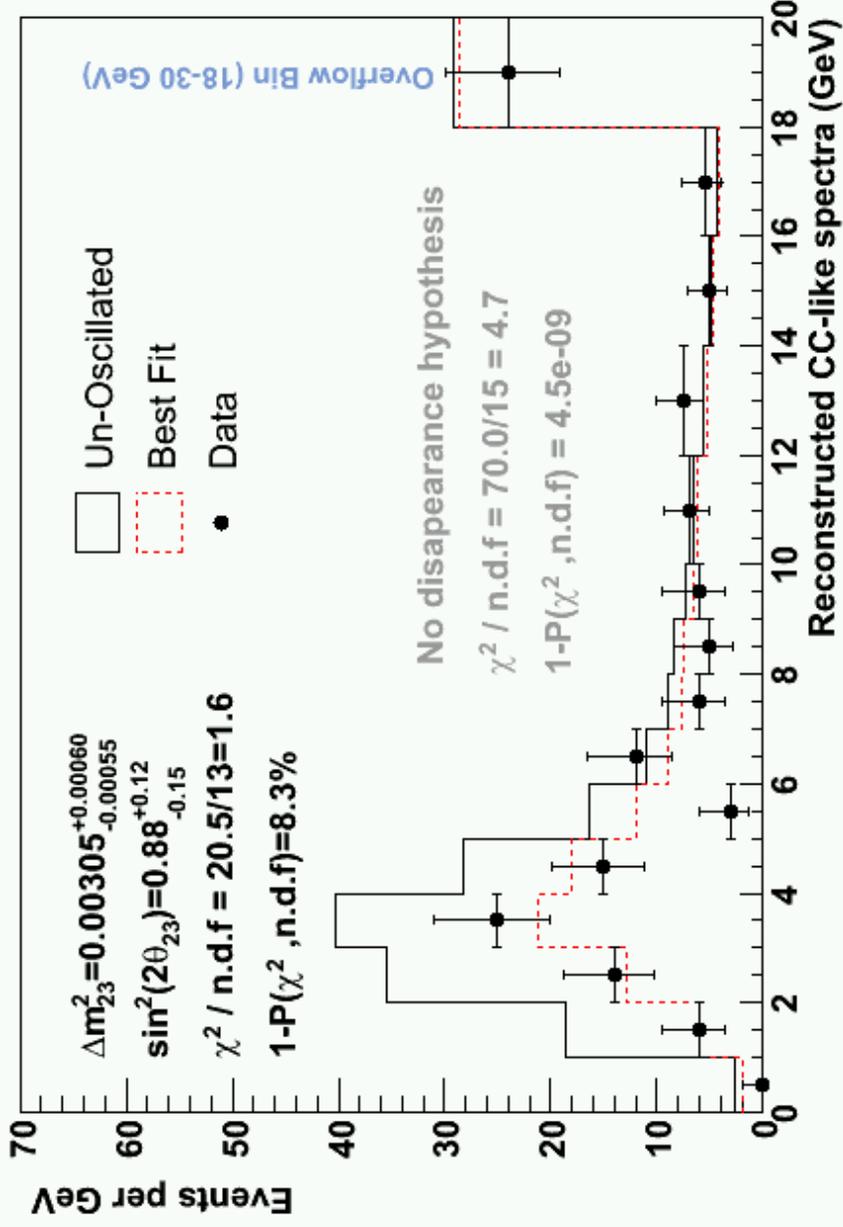


Predicted FD unoscillated spectra



# Oscillation Parameters

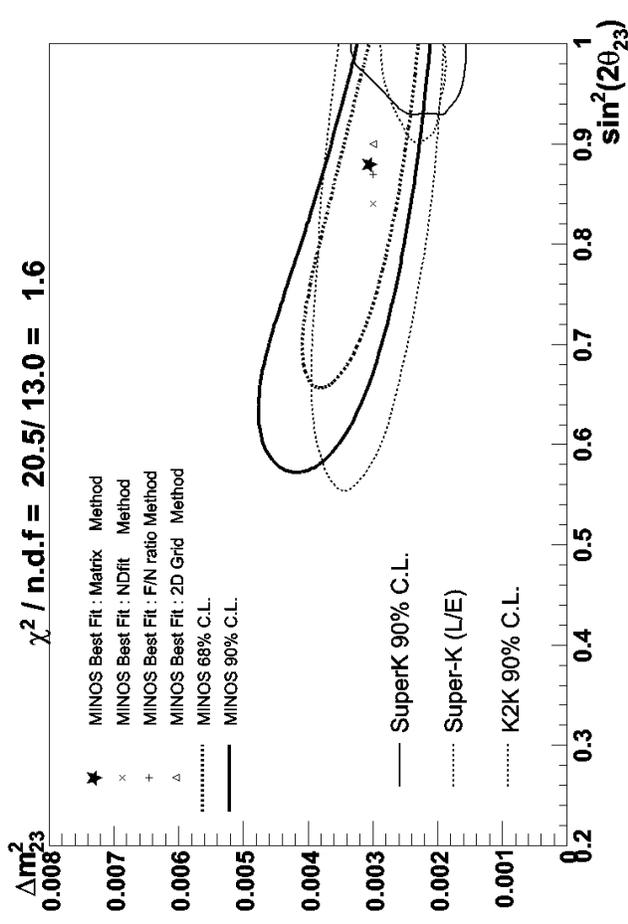
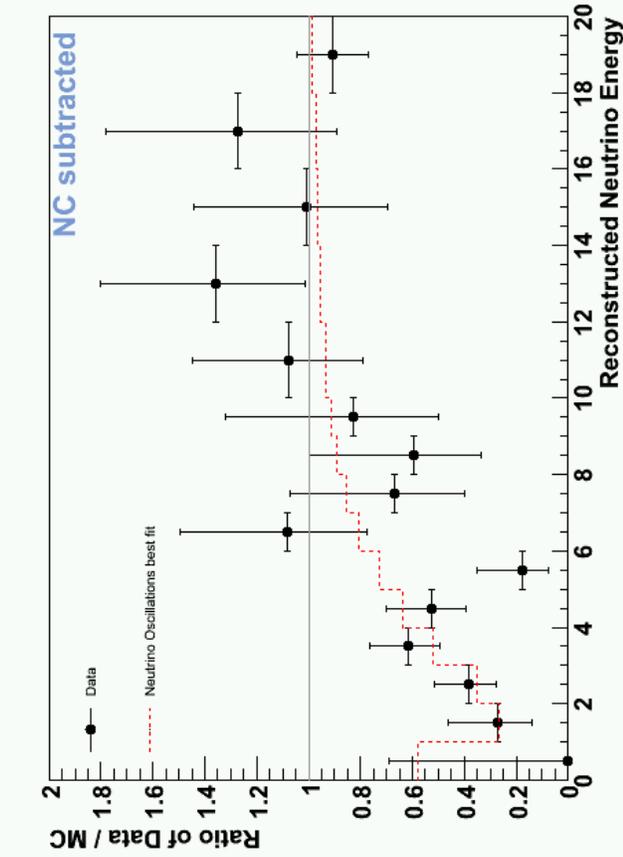
Oscillation Results for 0.93E20 p.o.t



$$\chi^2(\Delta m^2, \sin^2 2\theta) = \sum_{i=1}^{nbins} 2(e_i - o_i) + 2o_i \ln(o_i / e_i)$$

$o_i = \text{observed}$   
 $e_i = \text{expected}$

# Ratio Data/MC and Allowed Regions



# Systematic Errors

- Systematic shifts in the fitted parameters have been computed with MC “fake data” samples for  $\Delta m^2=0.003$  eV<sup>2</sup>,  $\sin^2 2\theta=0.9$  for the following uncertainties

Uncertainty	$\Delta m^2$ shift (eV <sup>2</sup> )	$\sin^2 2\theta$ shift
Normalisation +/- 4%	0.63e-4	0.025
Muon energy scale +/- 2%	0.14e-4	0.020
Relative Shower energy scale +/- 3%	0.27e-4	0.020
NC contamination +/- 30%	0.77e-4	0.035
CC cross-section uncertainties	0.50e-4	0.016
Beam uncertainty	0.13e-4	0.012
Intranuclear re-scattering	0.27e-4	0.030
<b>Total (sum in quadrature)</b>	<b>1.19e-4</b>	<b>0.063</b>
<b>Statistical error (data)</b>	<b>6.4e-4</b>	<b>0.15</b>

# Outlook

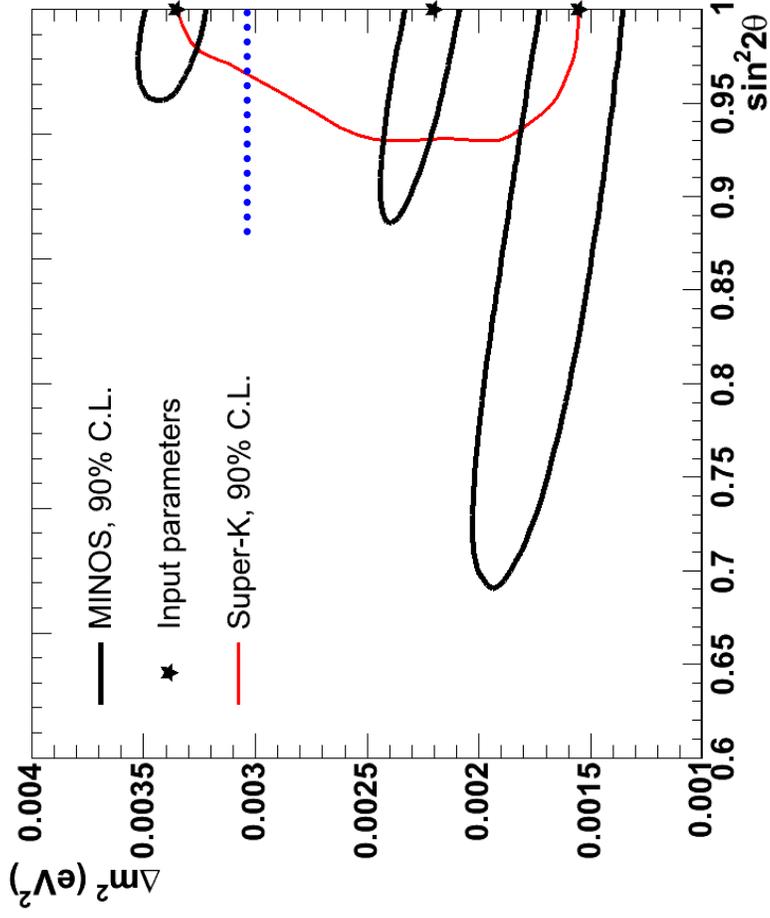
↑  $\Delta m_{23}^2$

↑  $\nu_{\mu} \rightarrow \nu_e$

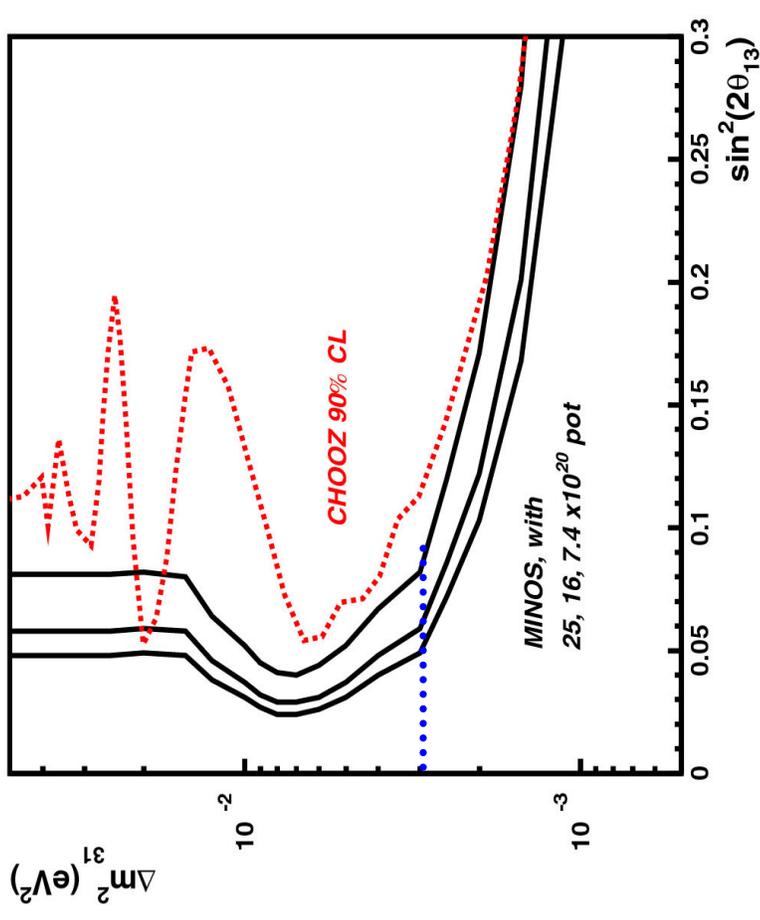


## $\nu_{\mu}$ disappearance

MINOS sensitivity,  $16 \times 10^{20}$  p.o.t.



## $\nu_{\mu} \rightarrow \nu_e$ 3 $\sigma$ Contours



# Summary

- 
- $\delta m_{\mu\tau}^2$  
- $\delta m_{\mu e}^2$  

$$\Delta m_{23}^2 = 3.05_{-0.55}^{+0.60} (stat) \pm 0.12(syst) \times 10^{-3} eV^2$$
$$\sin^2 2\theta_{23} = 0.88_{-0.15}^{+0.12} (stat) \pm 0.06(syst)$$

- 

$$\Delta m_{23}^2 = 2.70_{-0.305}^{+0.263} \times 10^{-3} eV^2$$

- 
- 