



First MINOS Results from the NuMI Beam

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for the MINOS Collaboration

Argonne • Athens • Benedictine • Brookhaven • Caltech • Cambridge • Campinas • Fermilab
College de France • Harvard • IIT • Indiana • IFFP-Moscow • Lebedev • Livermore
Minnesota-Twin Cities • Minnesota-Duluth • Oxford • Pittsburgh • Protvino • Rutherford
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Oscillation Physics

- There is experimental evidence that neutrino flavor is not conserved and thus **neutrinos have non-zero masses**.
- A neutrino of flavor $\alpha = e, \mu, \tau$ is a superposition of mass eigenstates described by the lepton mixing matrix U :

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \quad s_{ij} = \sin\theta_{ij}, \quad c_{ij} = \cos\theta_{ij}$$

$$|v_\alpha\rangle = \sum_i U_{\alpha i} |v_i\rangle; \alpha = e, \mu, \tau, \dots; i = 1, 2, 3, \dots$$

$|U_{\alpha i}|^2$: probability that v_i interaction will produce v of flavor α

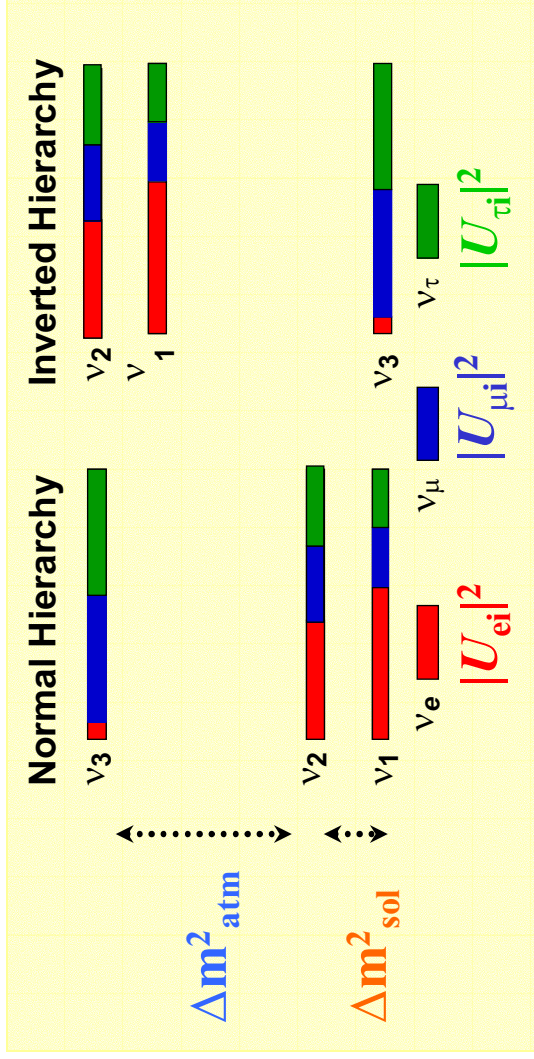
- In the two-flavor approximation:

$$P(v_\alpha \rightarrow v_\beta) = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 (eV^2) L (km)}{E_\nu (GeV)} \right), \alpha \neq \beta$$

→ **L/E determines what Δm^2 is probed**

Neutrino Mass Spectrum

(3 mass eigenstates - assume no ν_{sterile})



Two mass splittings:

Large (Δm^2_{atm}):

Atmospheric (up-down asymmetry)
Long baseline

Small (Δm^2_{sol}):

Solar + reactor

- $|U_{e3}|^2$ is not known (bounded by reactor expts) and neither is the mass hierarchy.
- Solar data suggests $|U_{e2}|^2 \sim 1/3$, and thus $|U_{e1}|^2 \sim 2/3$ (unitarity and small $|U_{e3}|^2$).
- Also, data suggests $\nu_\mu \rightarrow \nu_\tau$ dominant at Δm^2_{atm} and maximal mixing angle,

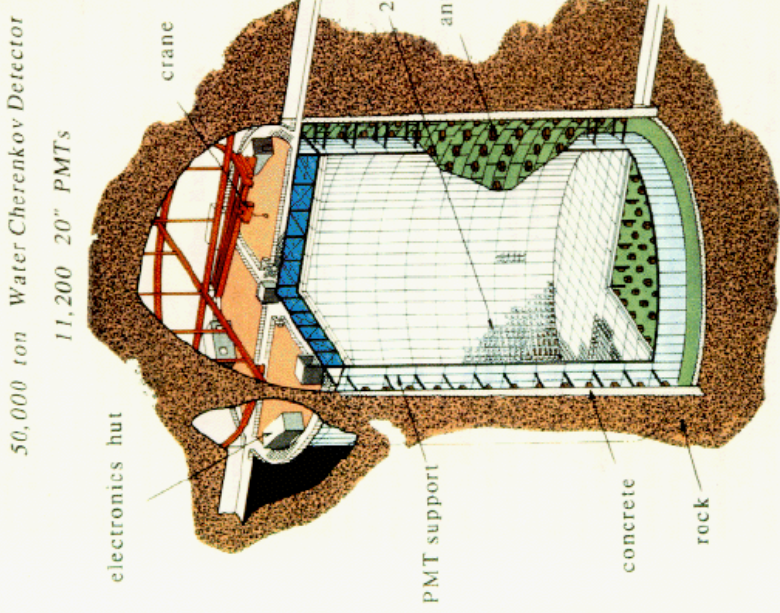
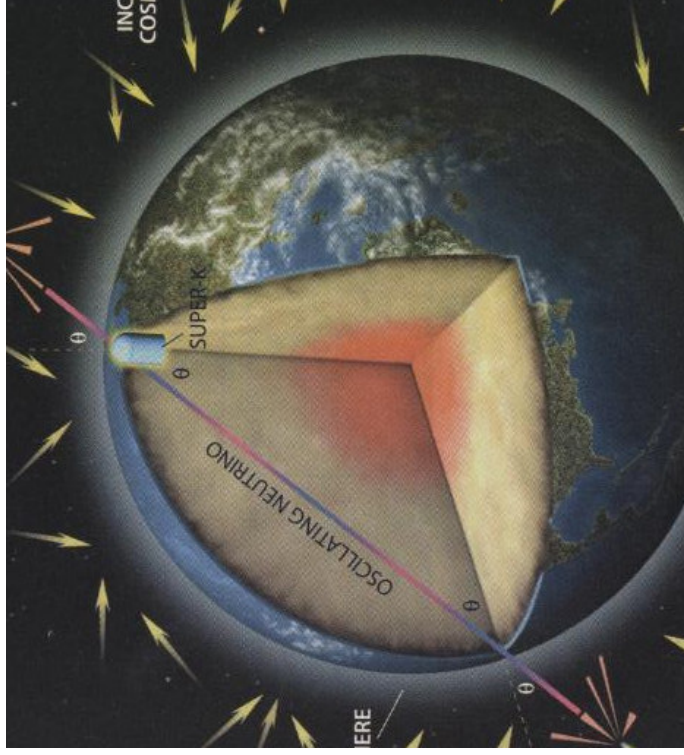
$$|U_{\mu i}|^2 \sim |U_{\tau i}|^2$$

With $L/E \sim 10^3$ MINOS is sensitive to Δm^2_{atm}
This scale will be the focus of this talk

Super-Kamiokande

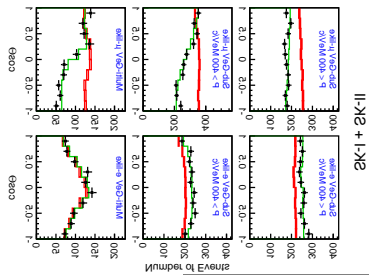
Ref. hep-ex/0501064, T. Kajita WHEPP9

- Neutrinos produced by the interaction of primary cosmic rays in the Earth's atmosphere. Wide range of energies and travel distances.
- Measurement relies on knowledge of flux w/o oscillations

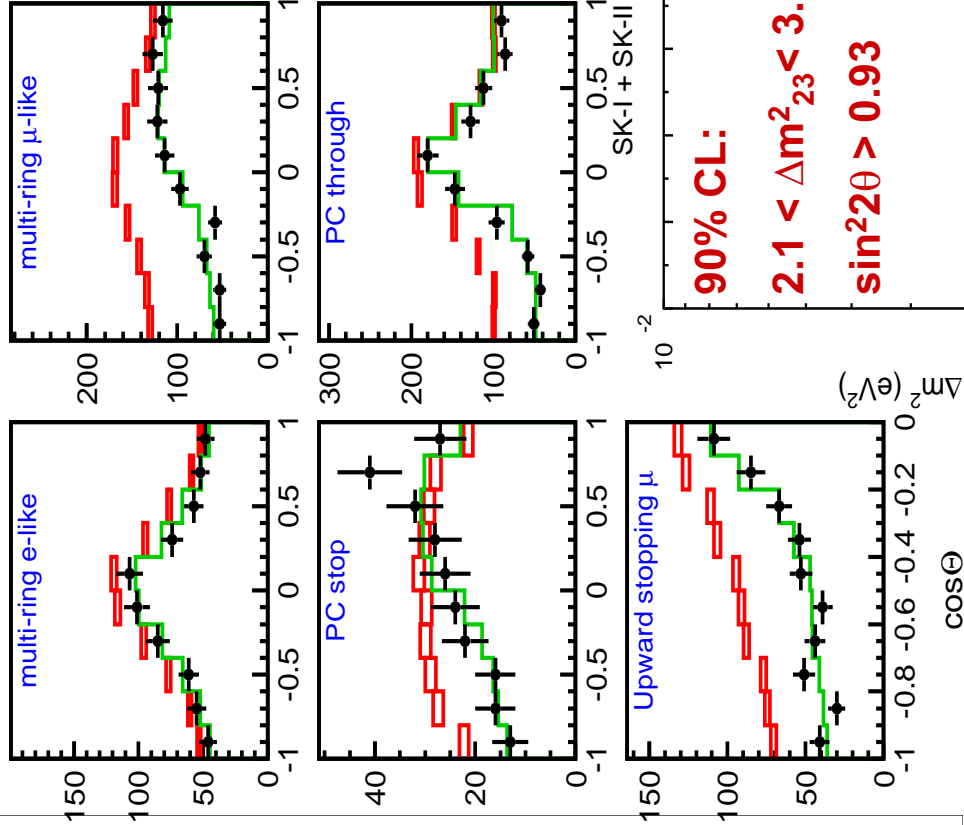


2,700 m.w.e. overburden
50 kT water Cherenkov detector

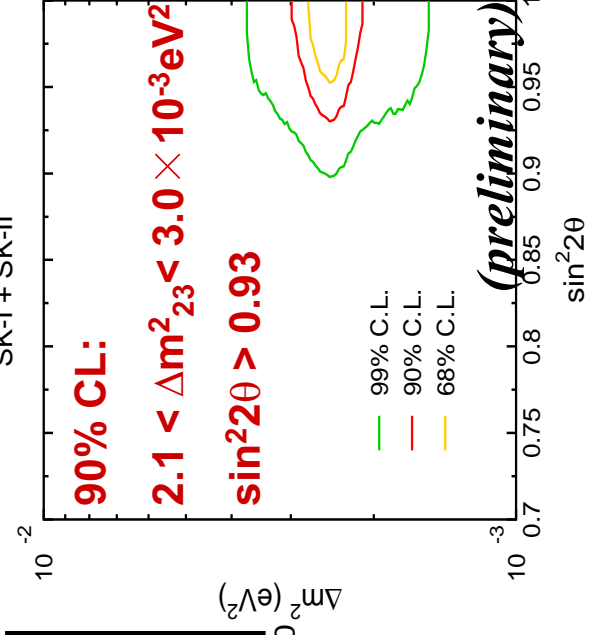




Super-K (I+II) - Zenith



Up Down

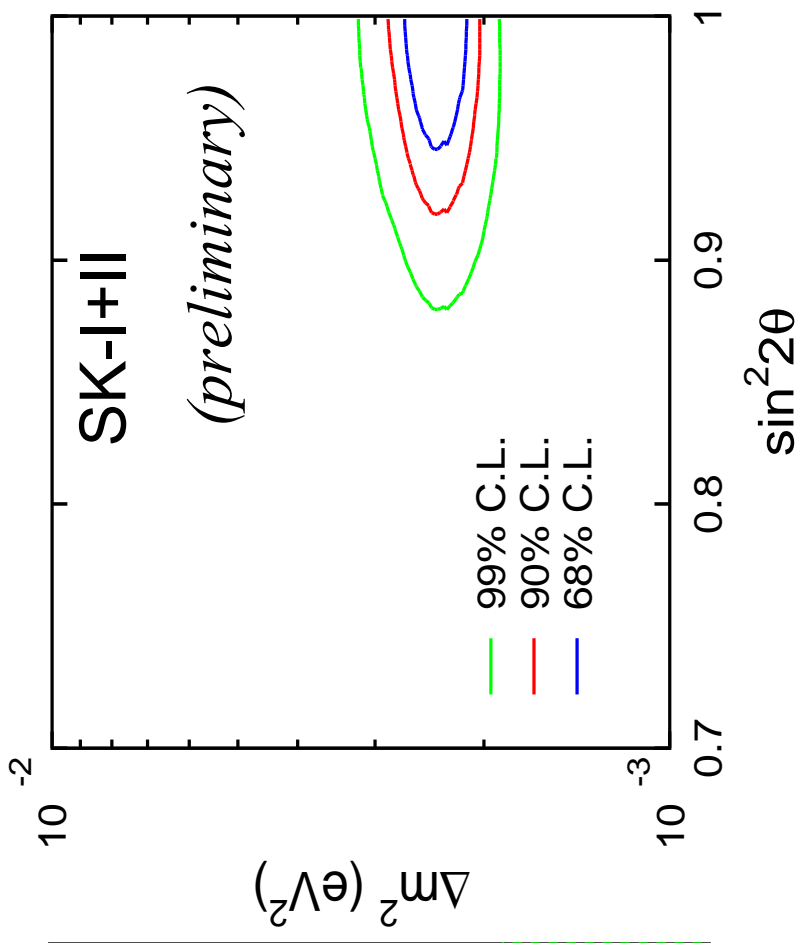
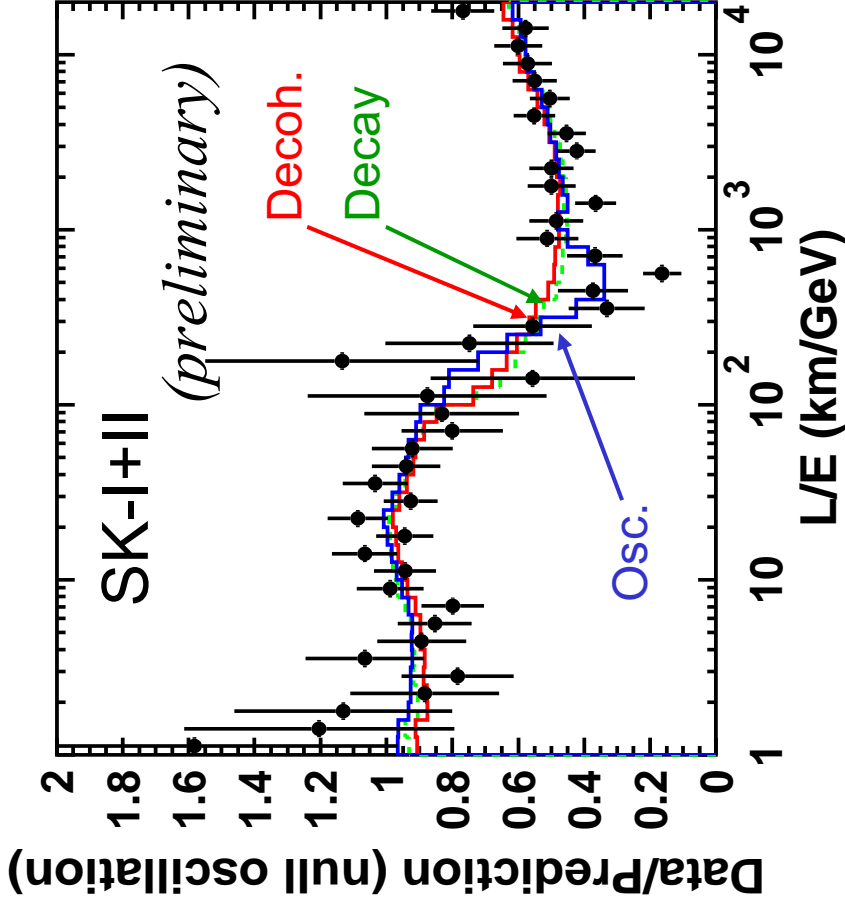


Unoscillated MC
Best fit: $\Delta m^2_{23} = 2.5 \times 10^{-3} \text{eV}^2, \sin^2 2\theta_{23} = 1$
Data



Super-K (I+II) – L/E

- Consistent with Zenith angle analysis results.
- At 90% CL: $2.0 < \Delta m^2_{23} < 2.9 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{23} > 0.92$

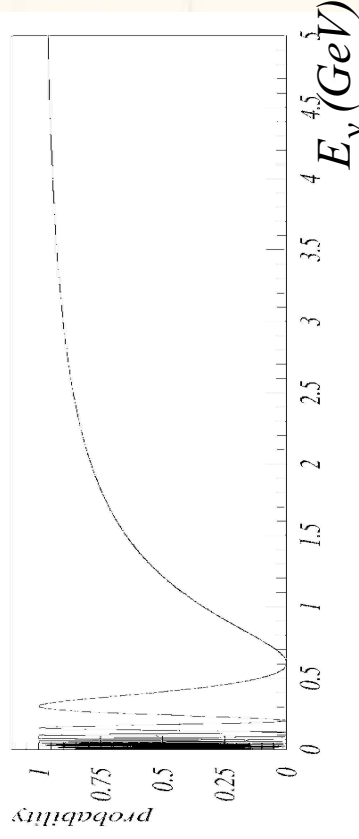


K2K

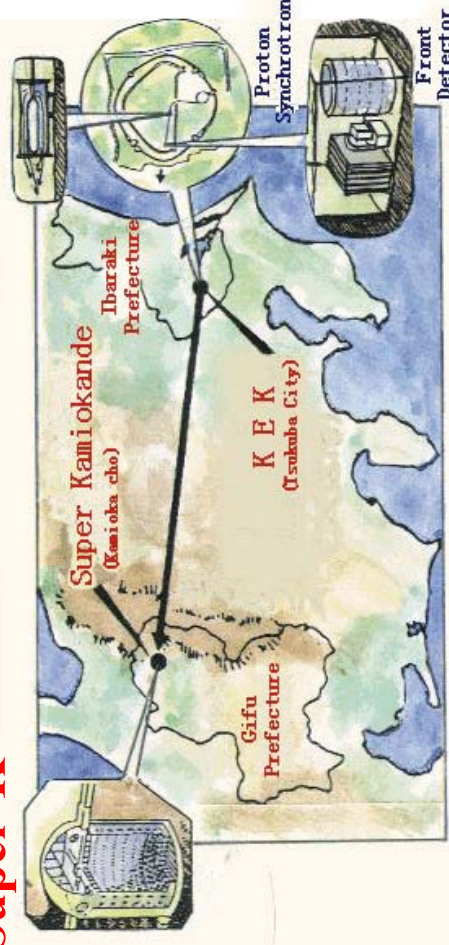
Ref. hep-ex/0411038, hep-ex/0512041

- First **long baseline** accelerator neutrino experiment.
- Accelerator produced ν_μ beam (**1.3 GeV mean energy**) travels **250 km** toward Super-K. The near detector is located 300 m from the production target.
- **Observe energy dependent ν_μ disappearance.**

$$\Delta m^2 = 0.003 \text{eV}^2, L = 250 \text{km}$$



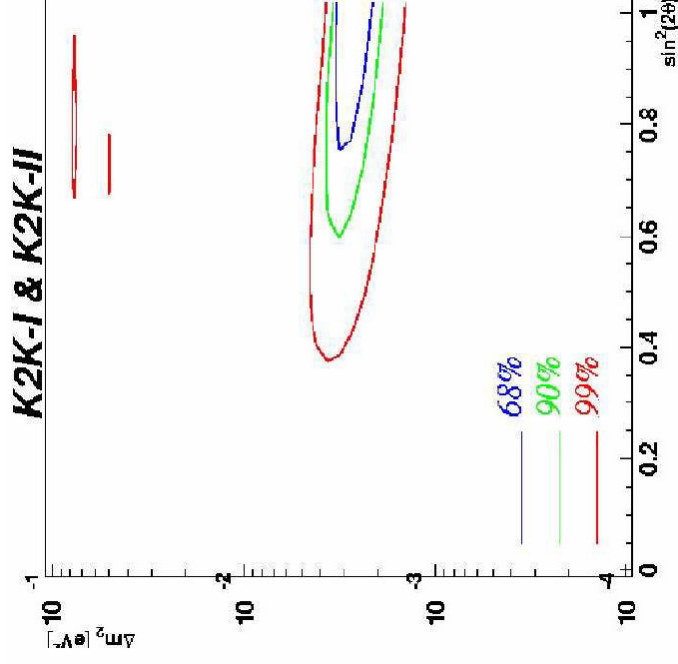
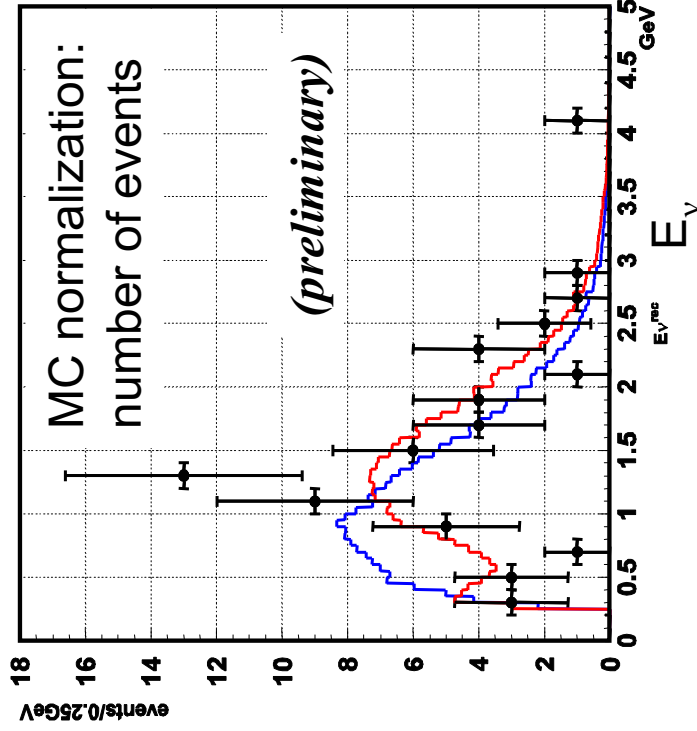
Super-K



**Near Detectors (SciBar,
1 kton water Cherenkov)**

K2K - Results

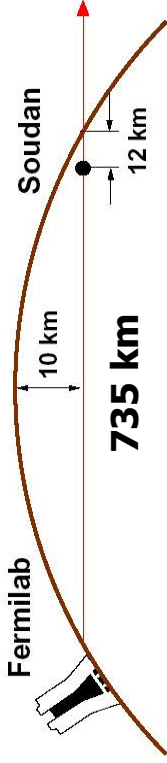
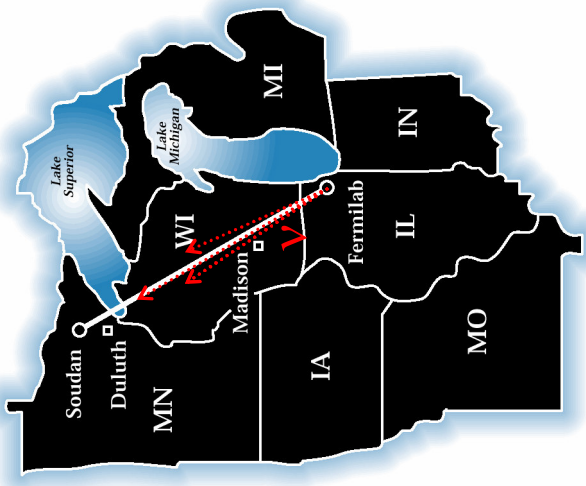
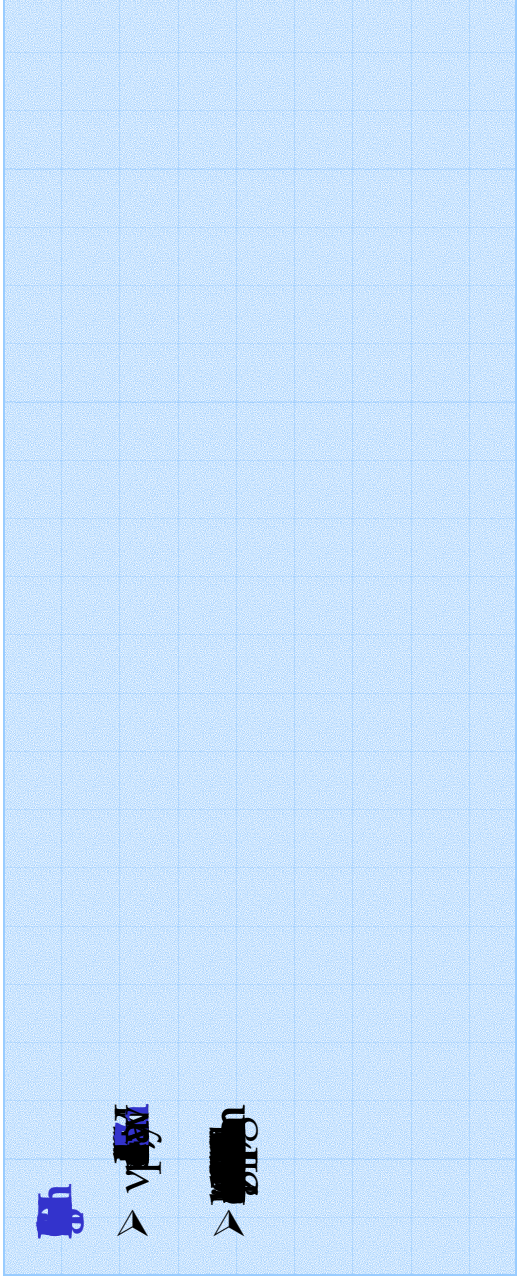
- Observe deficit of ν_μ at SK and distortion in the energy spectrum:
112 events at SK, expect $155.9^{+13.6}_{-15.6}$
- Well described by $\nu_\mu \rightarrow \nu_\tau$ oscillation



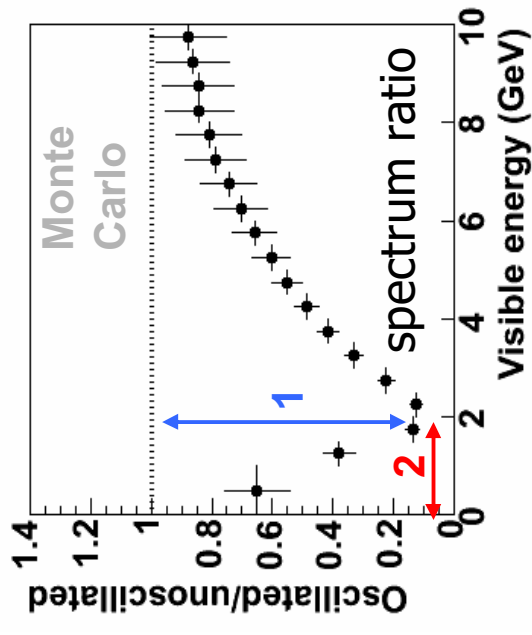
Best fit: $\Delta m^2_{23} = 2.76 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{23} = 1$
 For $\sin^2 2\theta_{23} = 1$: 90% CL $1.9 < \Delta m^2_{23} < 3.5 \times 10^{-3} \text{ eV}^2$

MINOS

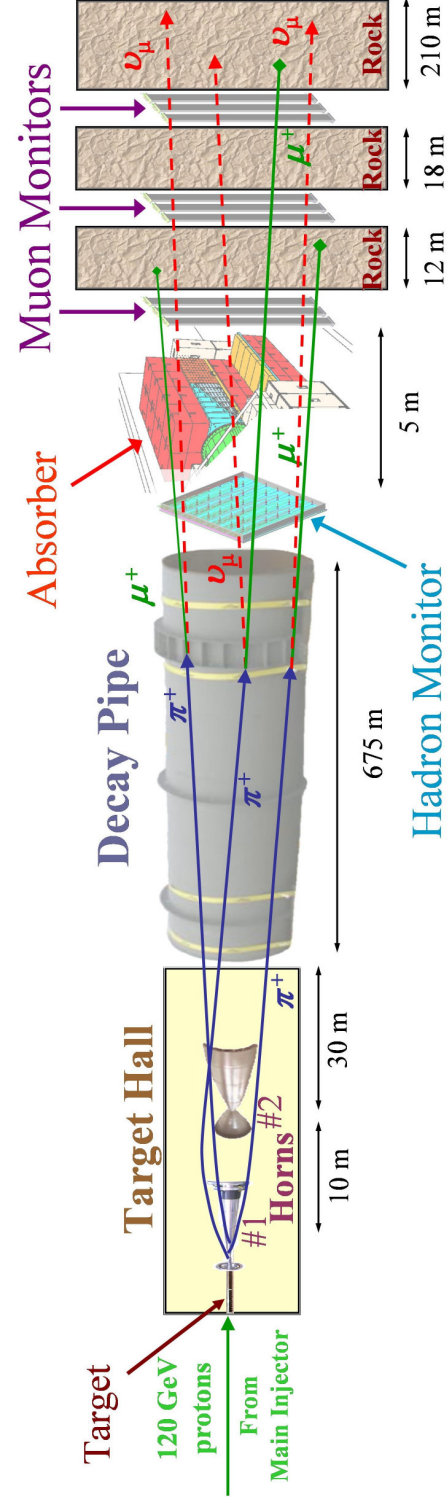
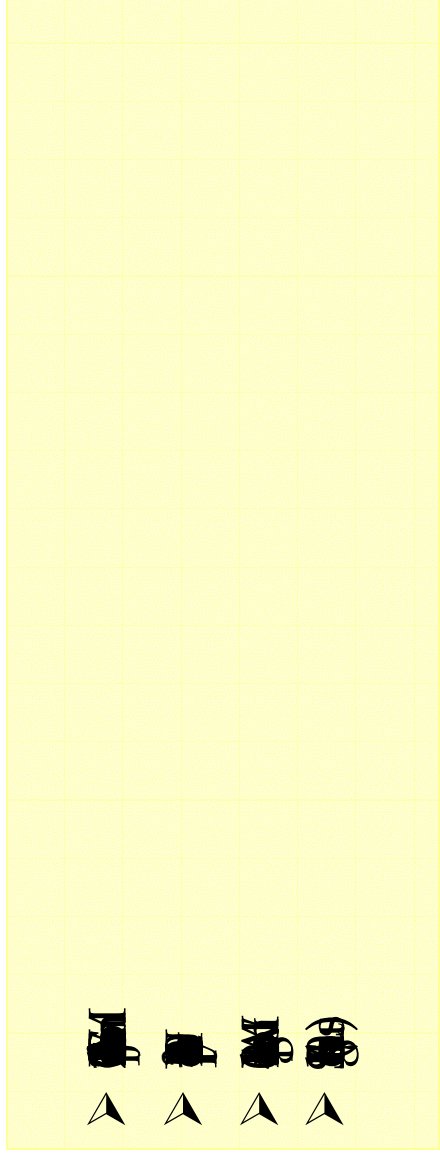
Main Injector Neutrino Oscillation Search



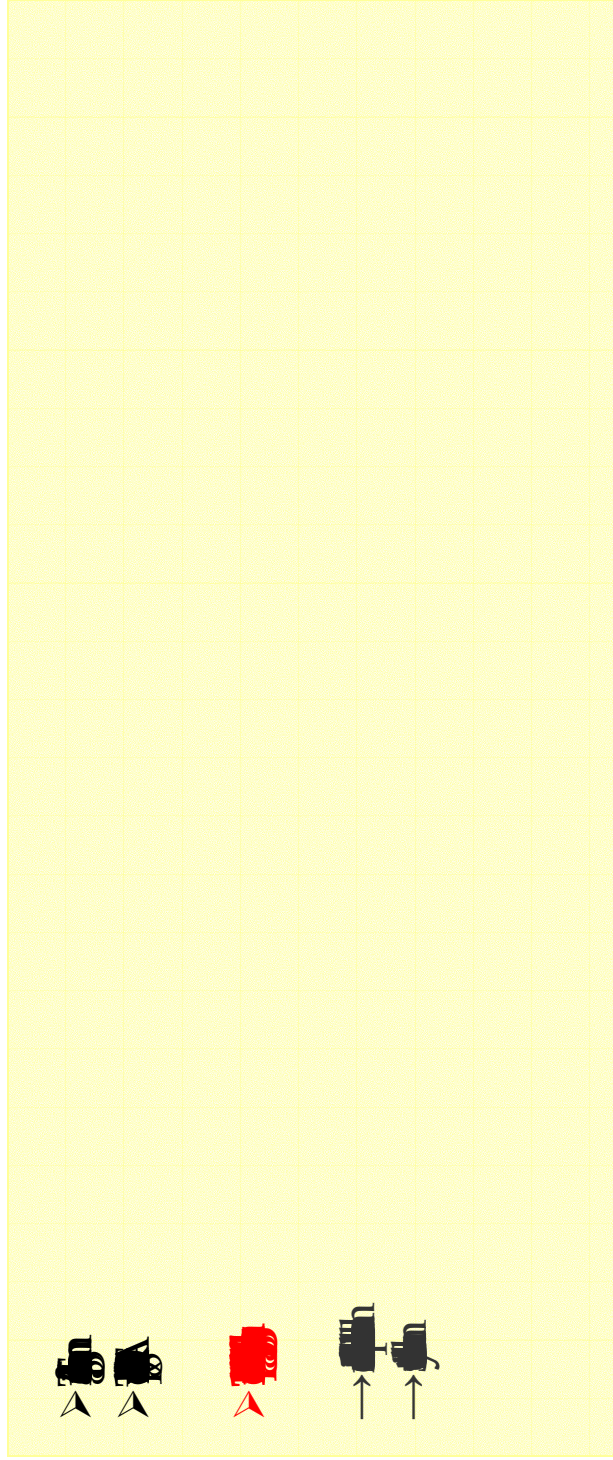
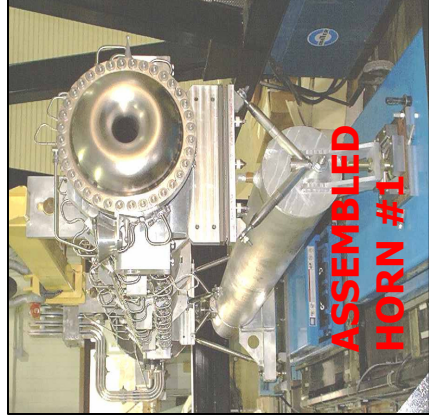
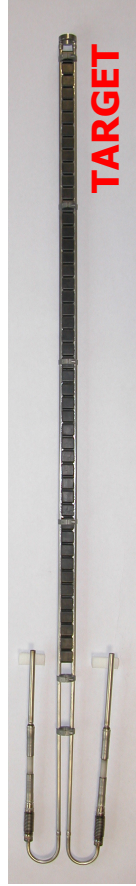
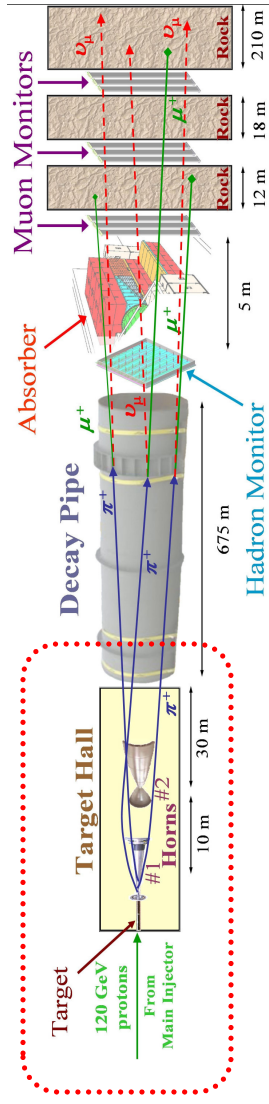
$$P(\nu_{\mu} \rightarrow \nu_{\mu}) = 1 - \sin^2 2\theta_{23} \sin^2 \left(1.27 \frac{\Delta m_{23}^2 (eV^2) L(km)}{E_{\nu} (GeV)} \right)$$



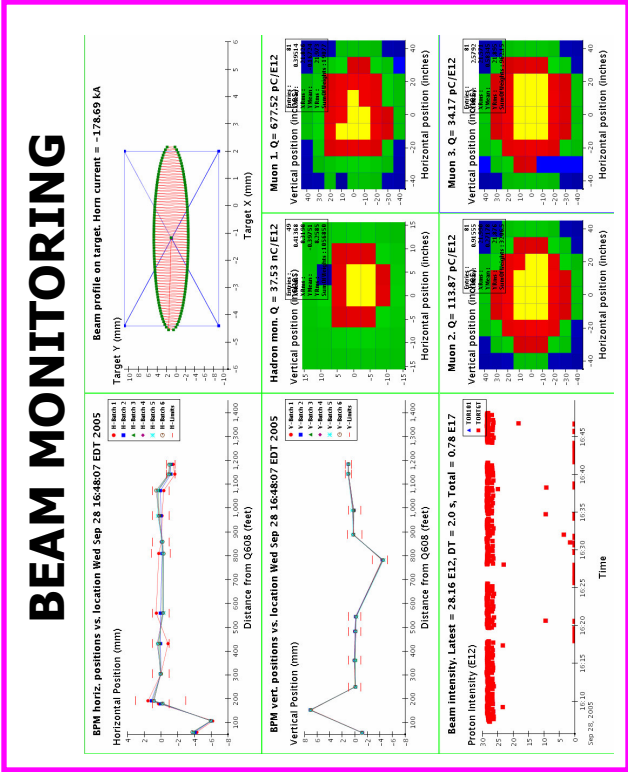
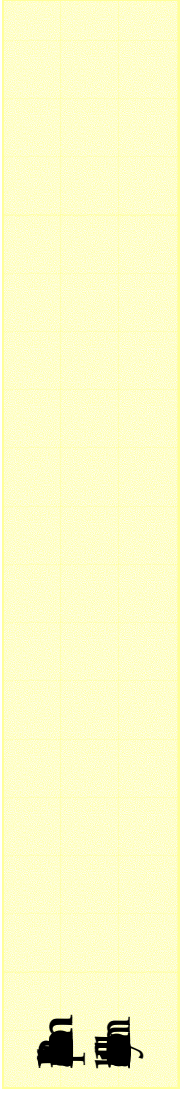
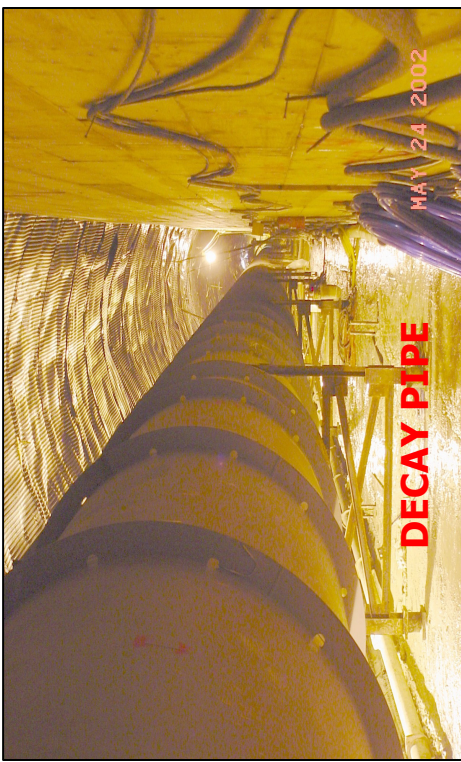
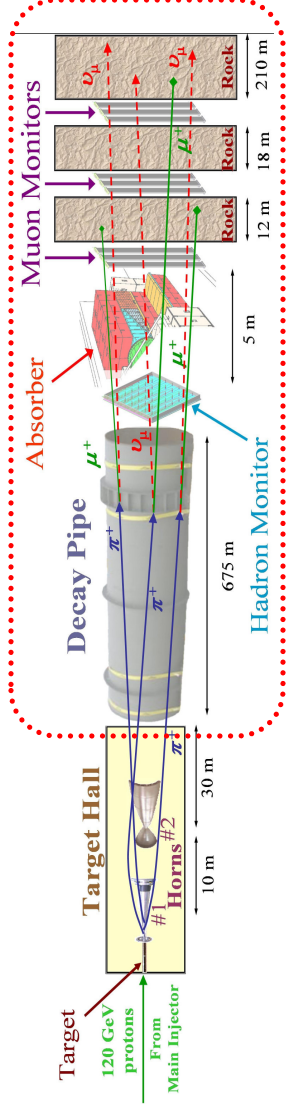
NuMI Beamline



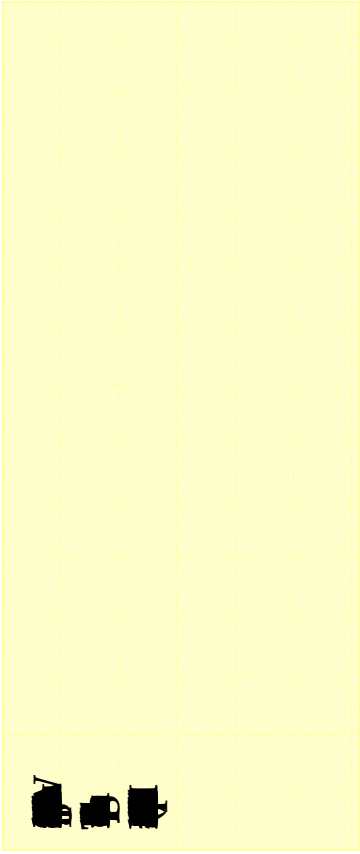
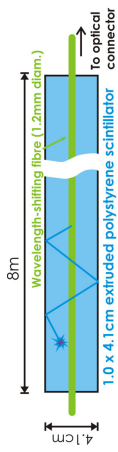
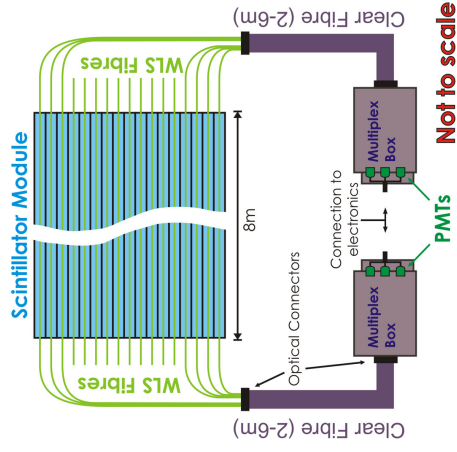
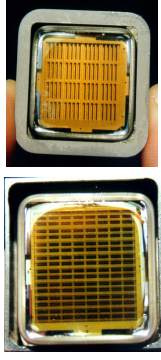
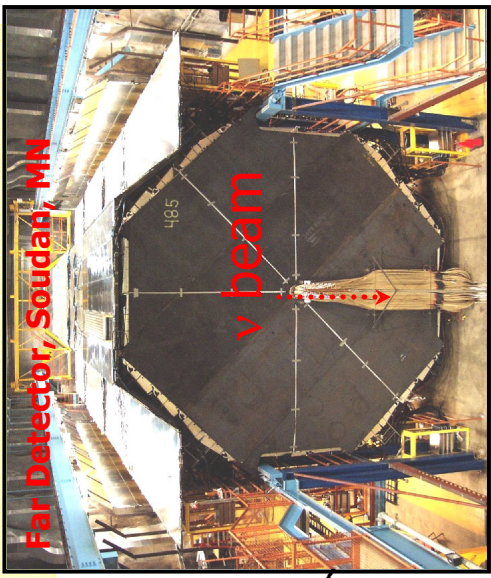
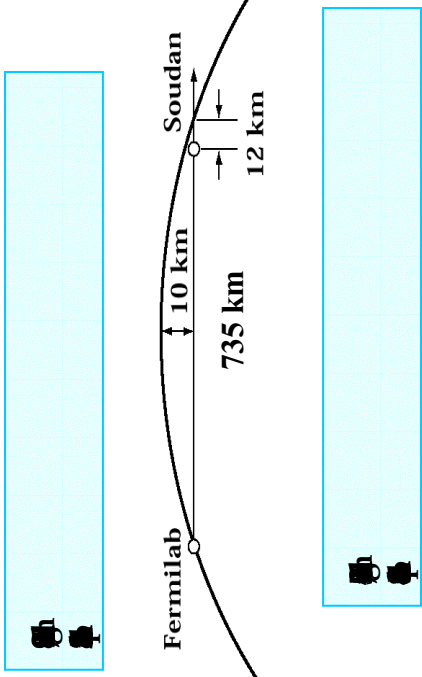
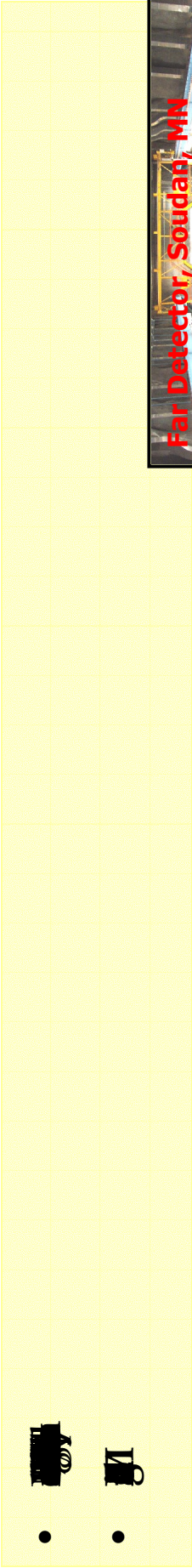
NuMI Target and Horns



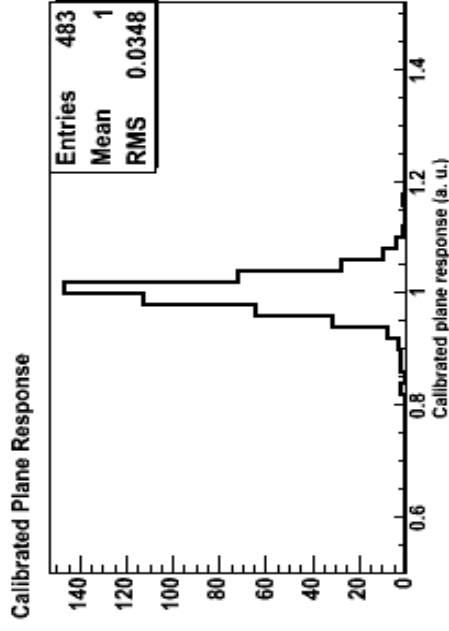
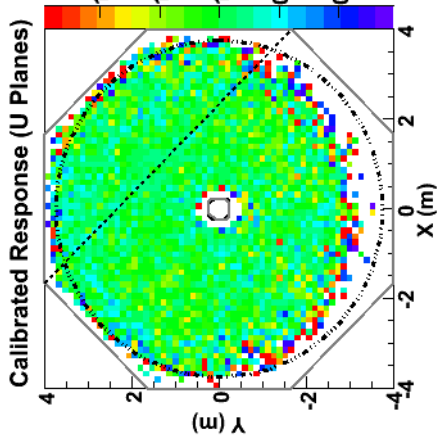
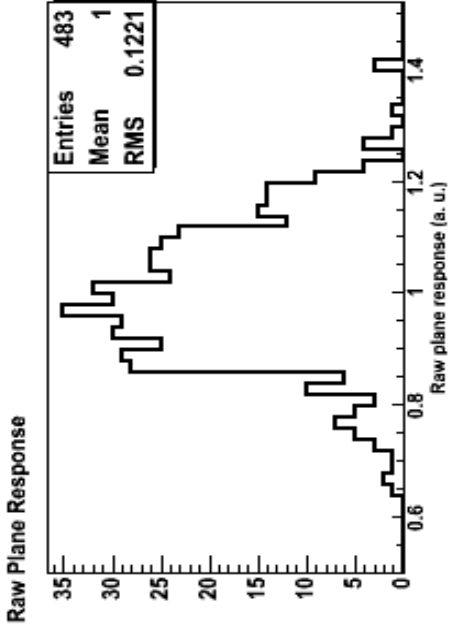
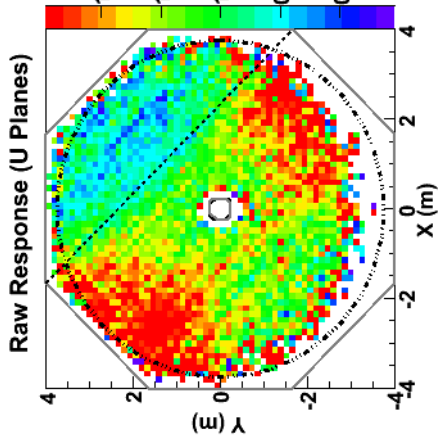
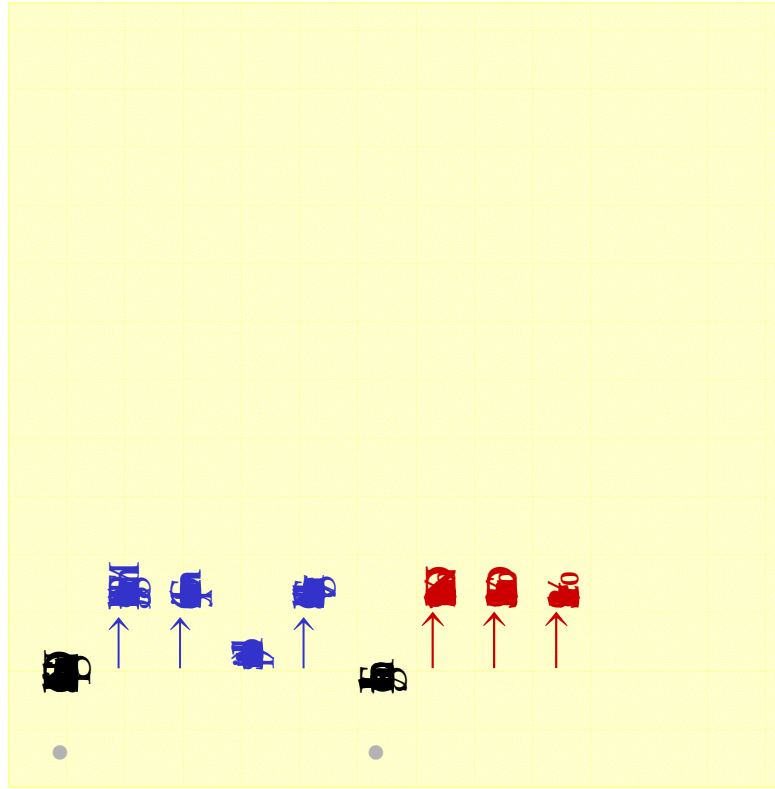
Hadron and Muon Monitors



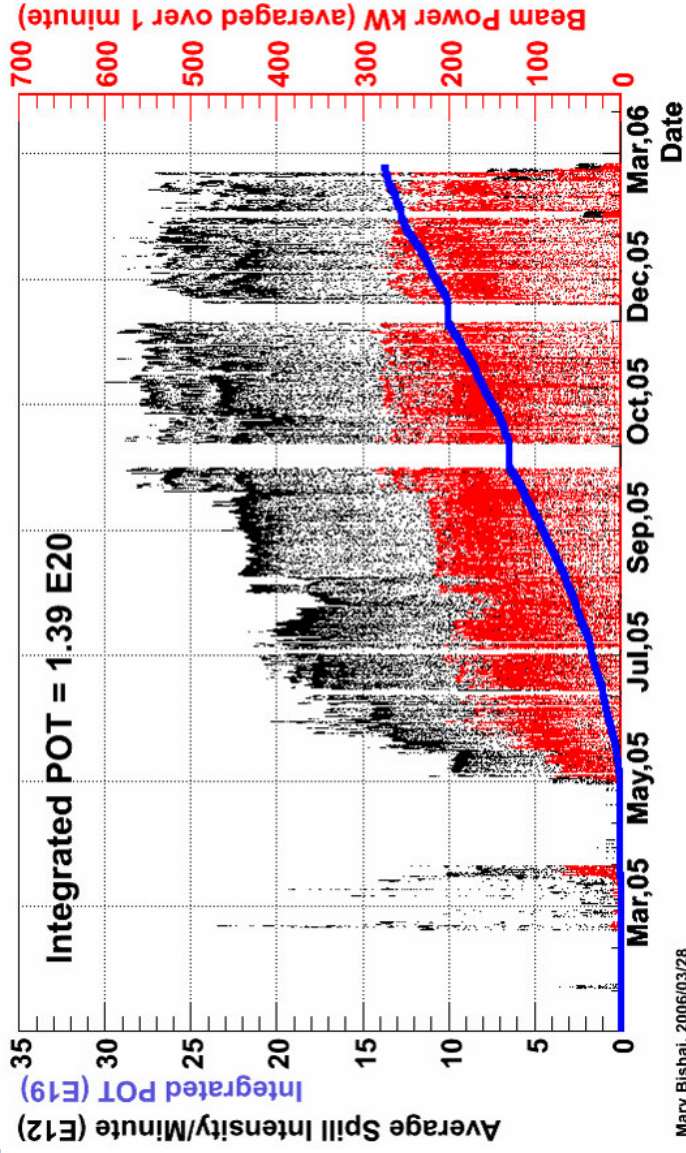
The MINOS Detectors



Calibration



Start-Up and Running



Mary Bishai, 2006/03/28

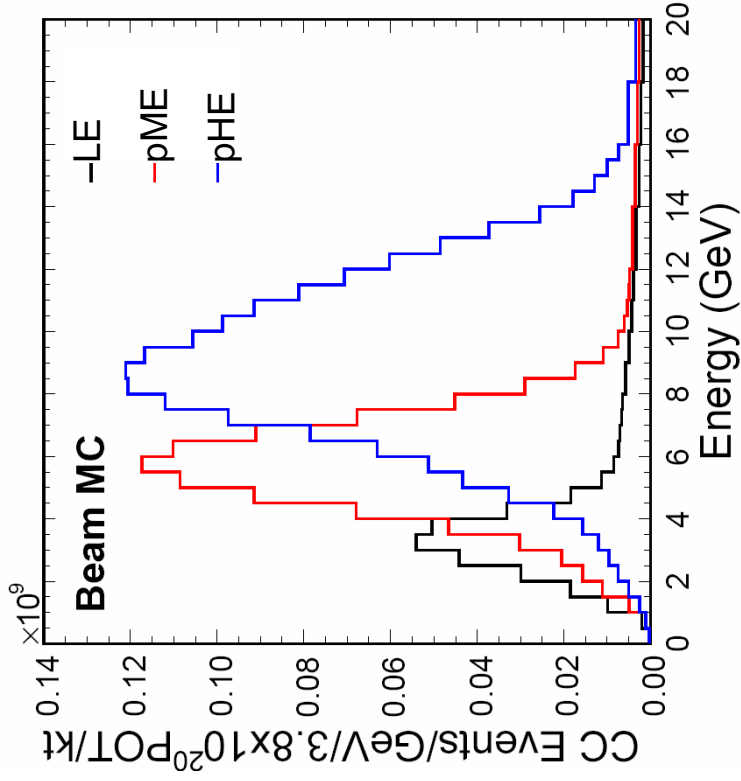


Neutrino Energy Spectrum



Δm^2_{23}

8.5% $\nu_\mu + \nu_\mu$ (6.5% ν_μ), 1.5% $\nu_e + \nu_e$

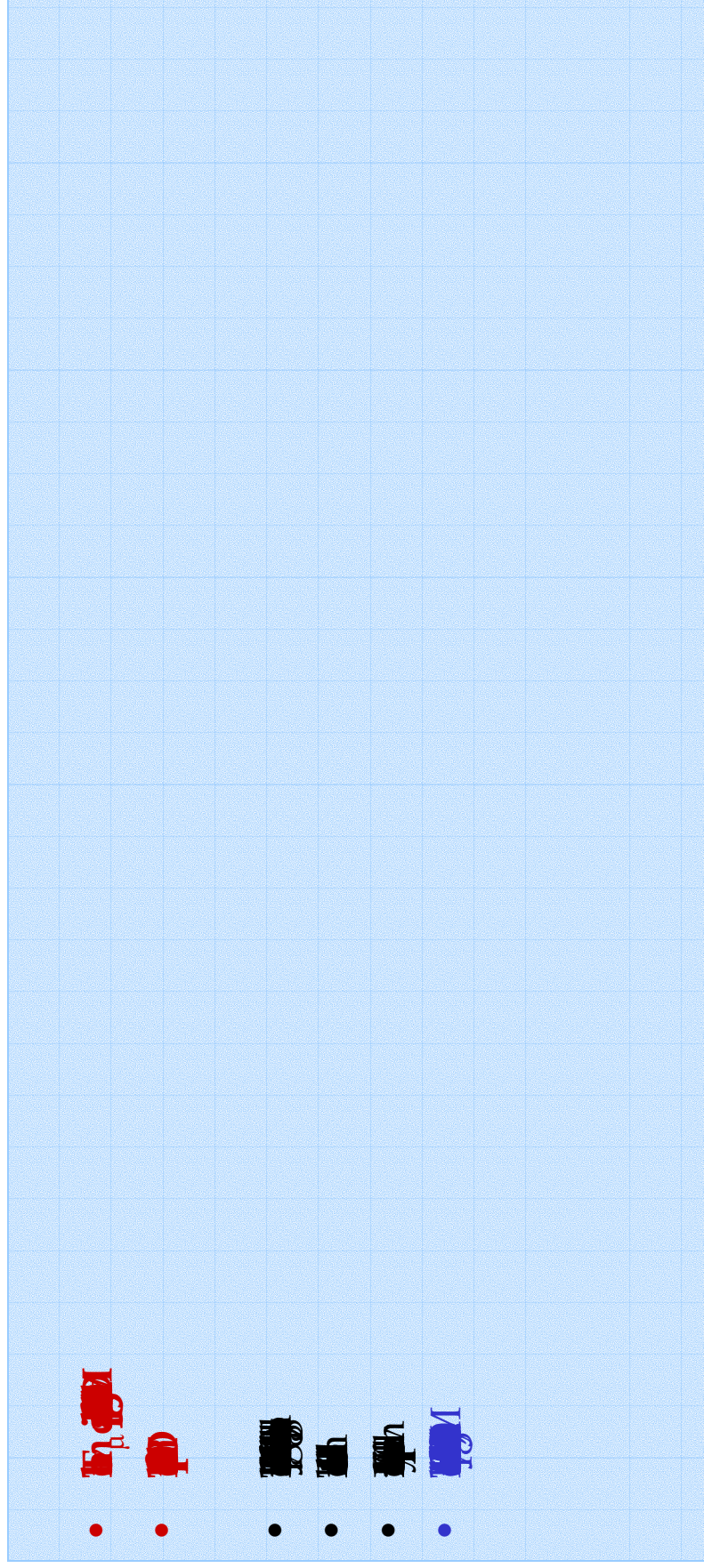


Beam	Target z position (cm)	FD Events (no osc) per 1 x 10 ²⁰ POT
LE-10	-10	390
pME	-100	970
pHE	-250	1340

Events in fiducial volume



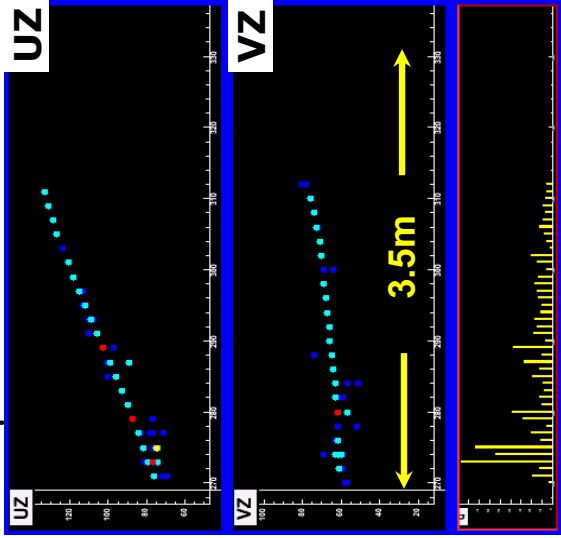
Oscillation Analysis



Event Topologies

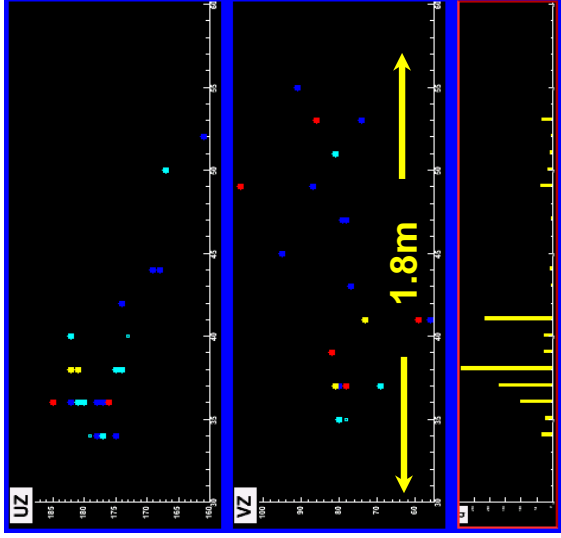
Monte Carlo

ν_{μ} CC Event



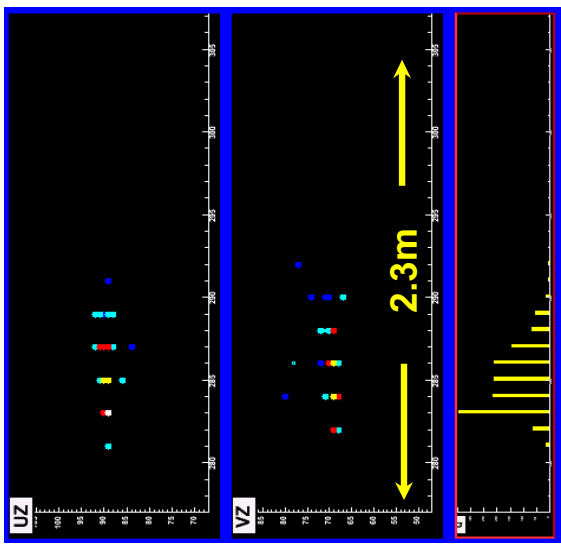
ν_{μ} CC

NC Event

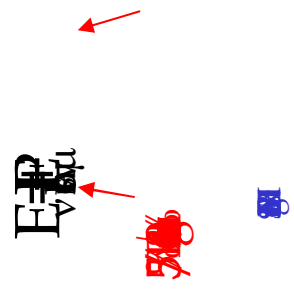


NC

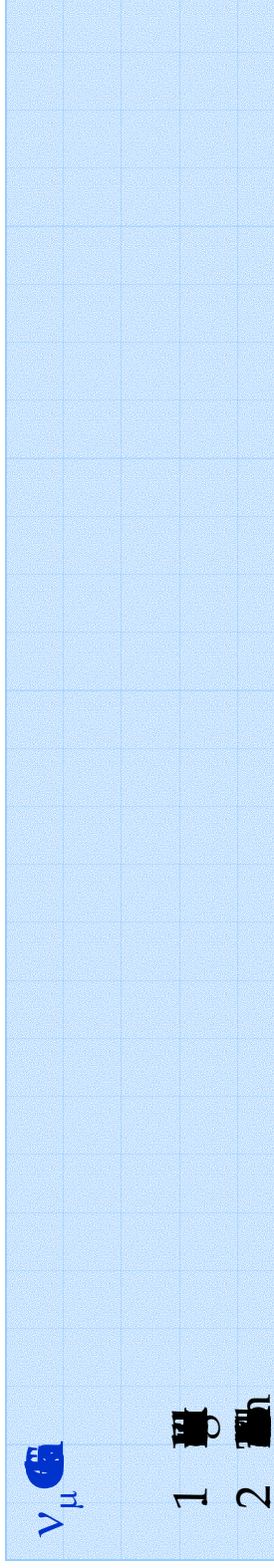
ν_e CC Event



ν_e CC



Data Selection



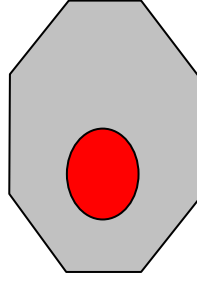
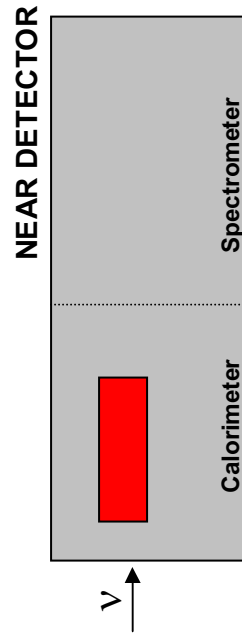
NEAR: $1\text{m} < z < 5\text{m}$ (z from front face)

$R < 1\text{m}$ from beam center

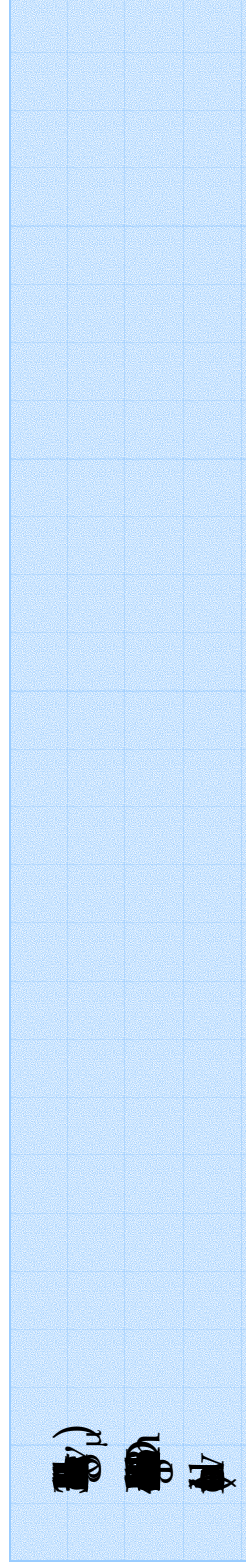
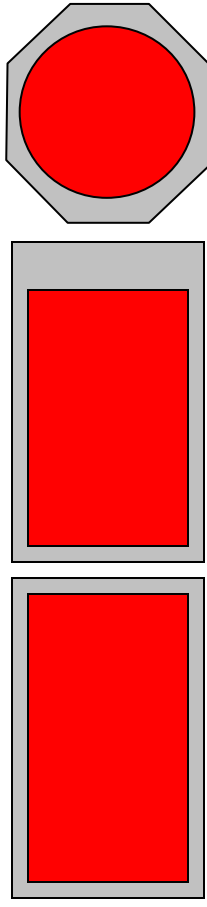
FAR: $z > 50\text{cm}$ from front face

$z > 2\text{m}$ from rear face

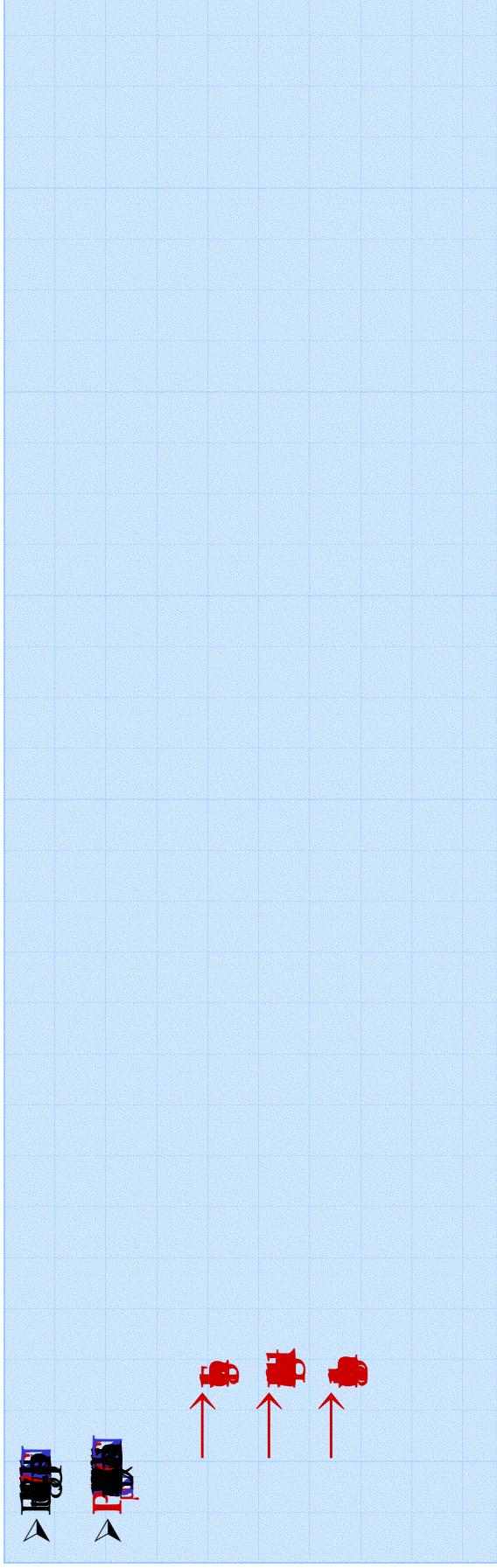
$R < 3.7\text{m}$ from center of detector



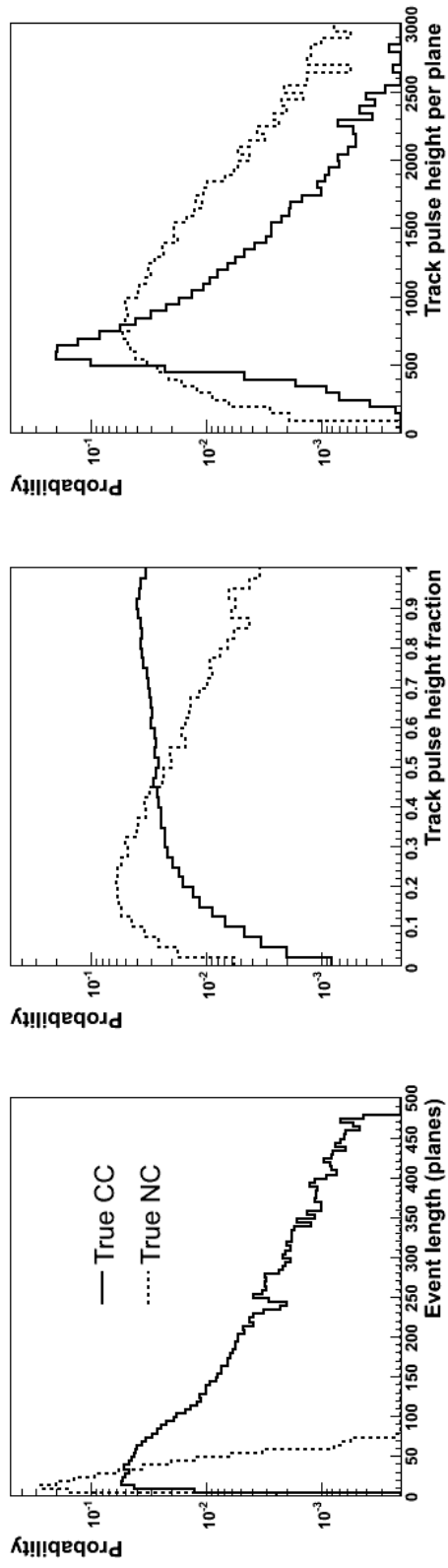
Fiducial Volume



Selecting CC events



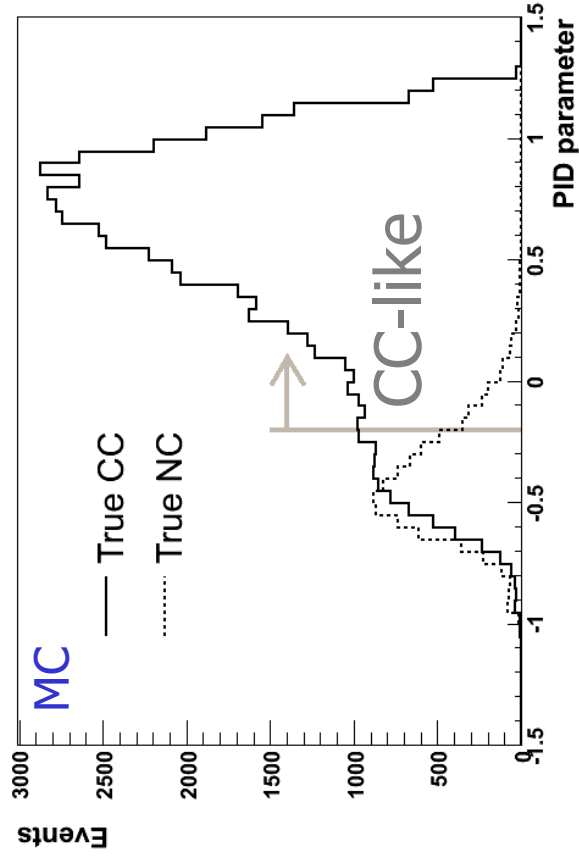
Input variables for PDF-based event selection



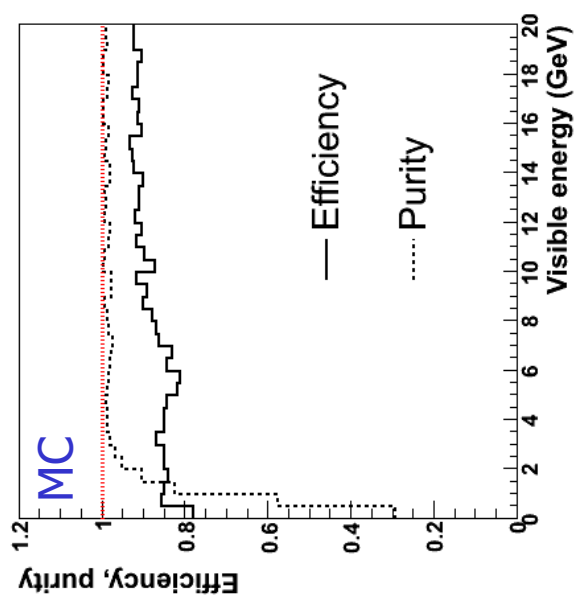
CC Selection Efficiencies

$$PID = -(\sqrt{-\log(P_\mu)} - \sqrt{-\log(P_{NC})})$$

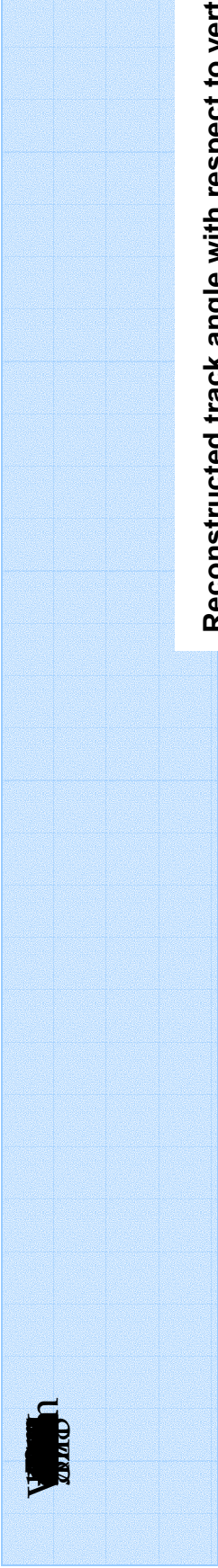
PDF PID parameter distribution for true CC and NC events



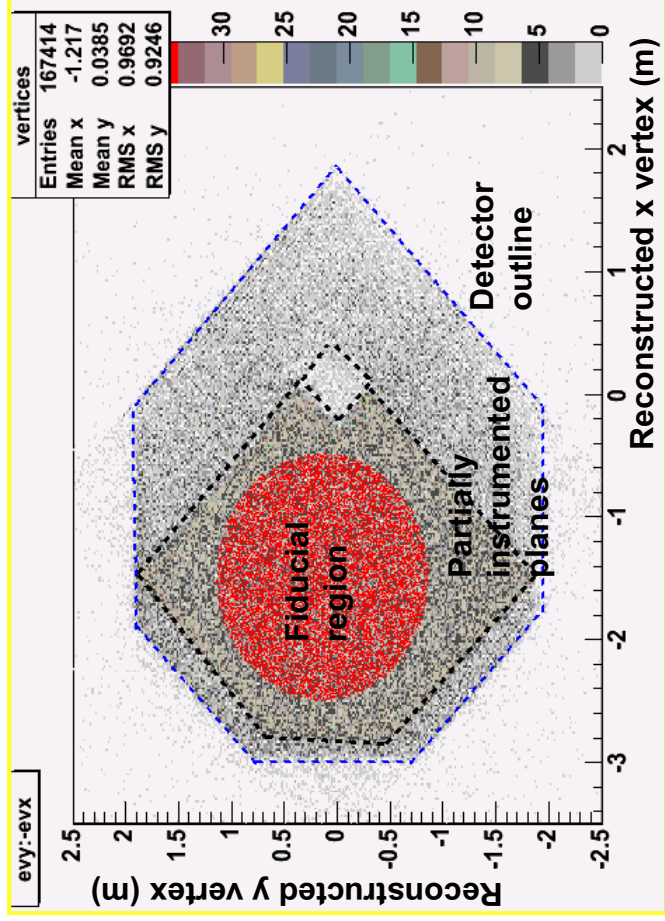
CC selection efficiencies and purities



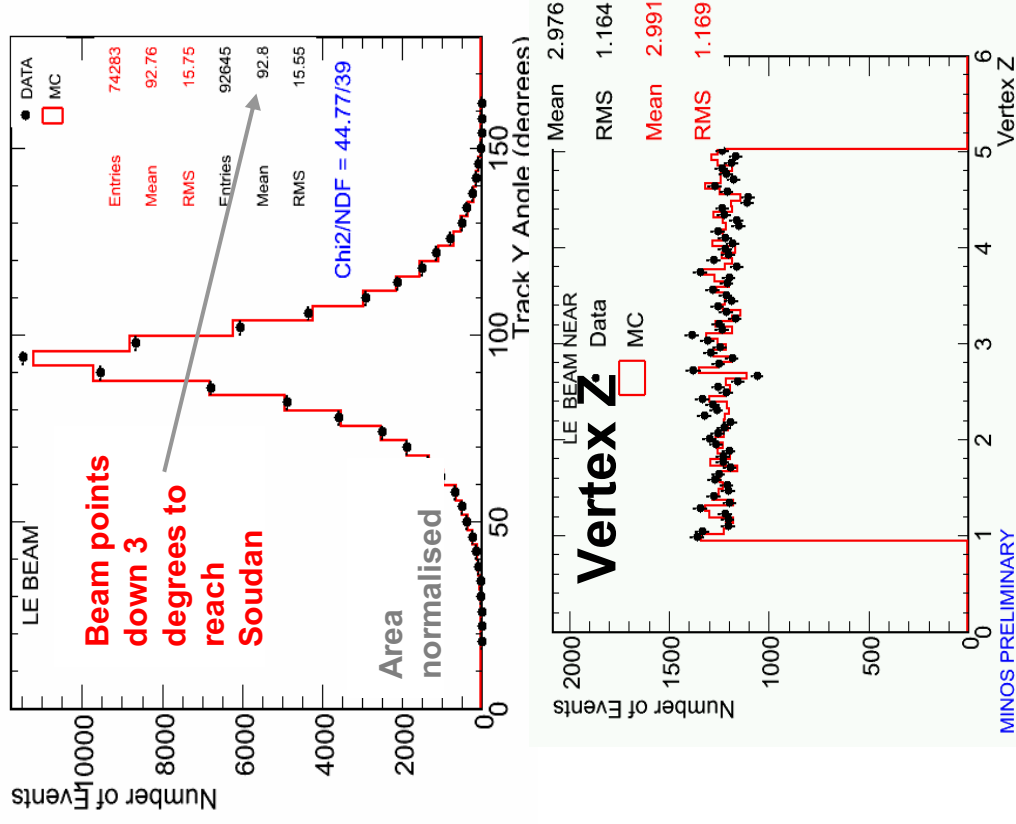
Near Detector Distributions



Distribution of reconstructed event vertices in the x-y plane

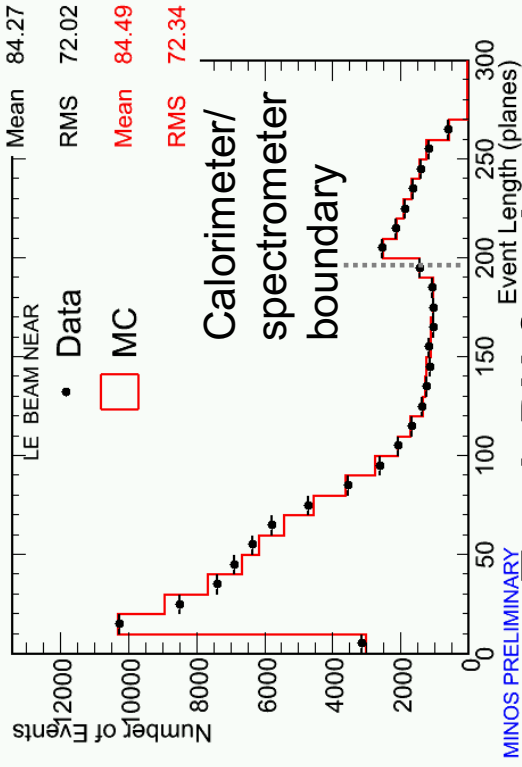


Reconstructed track angle with respect to vertical

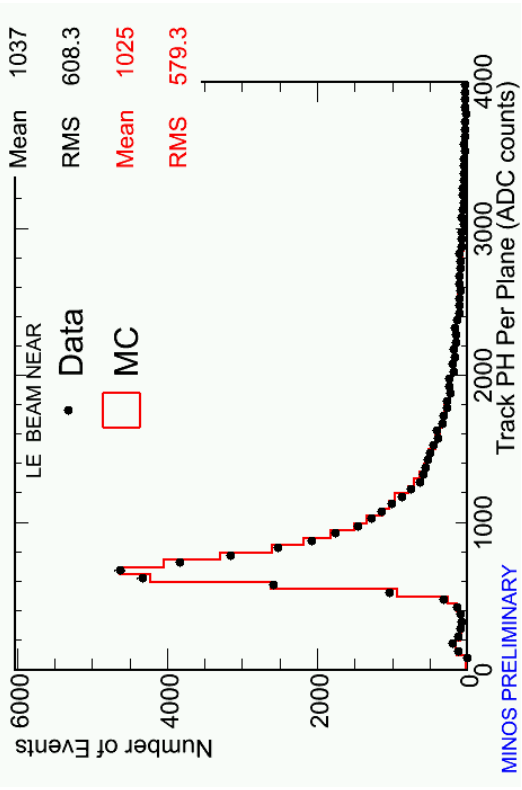


Near Detector - PID Distributions

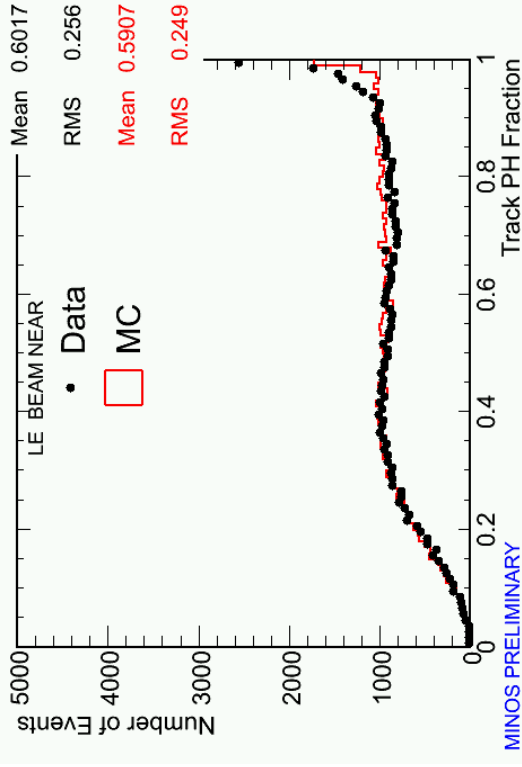
Event length



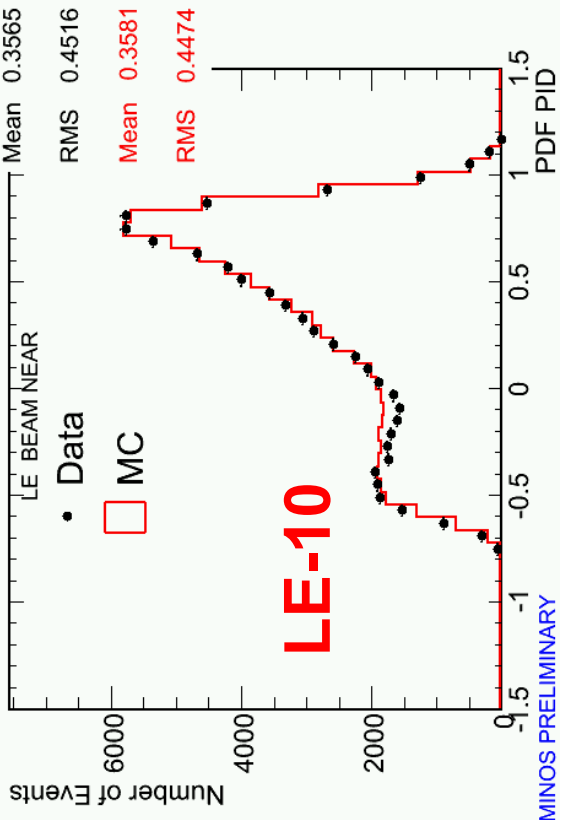
Track PH per plane



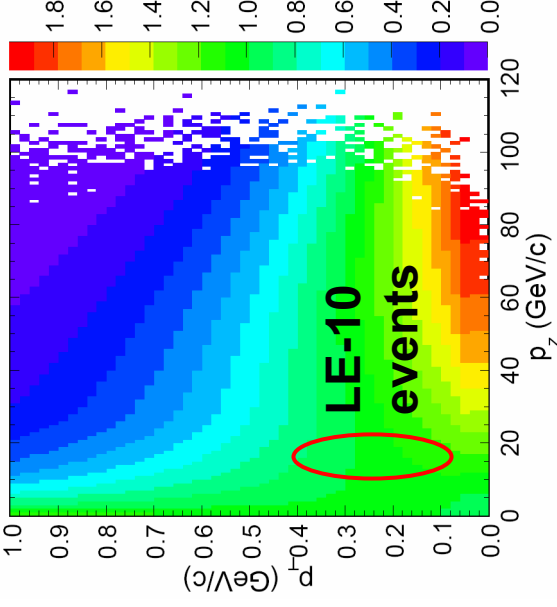
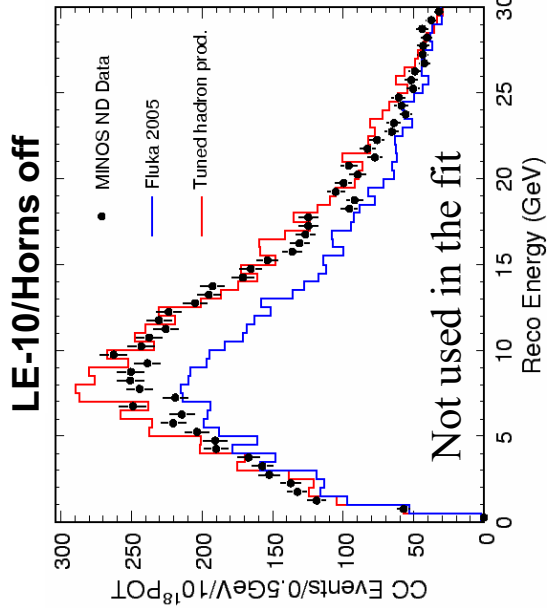
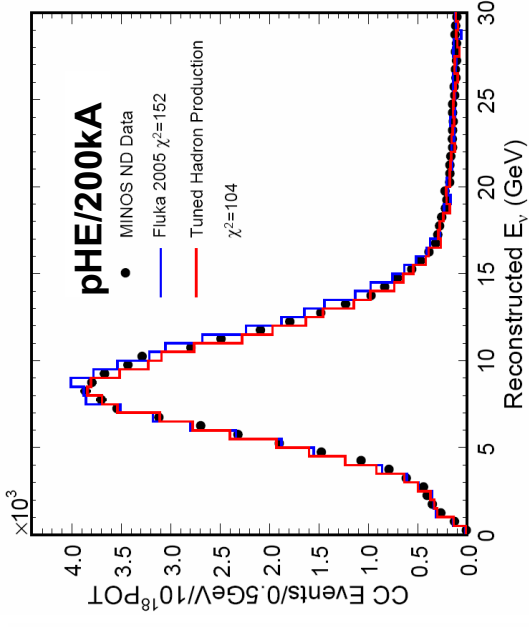
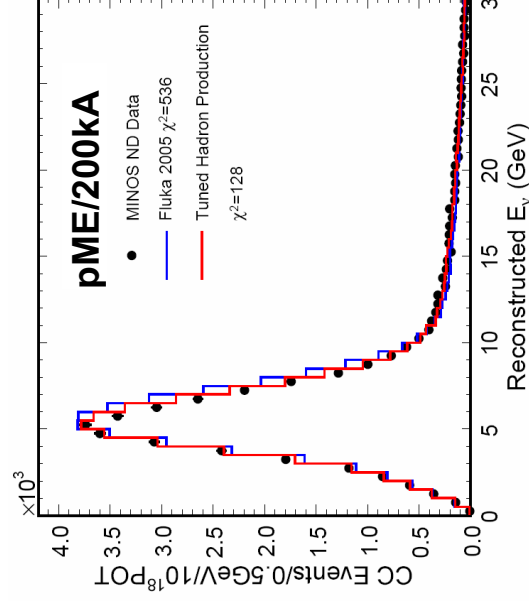
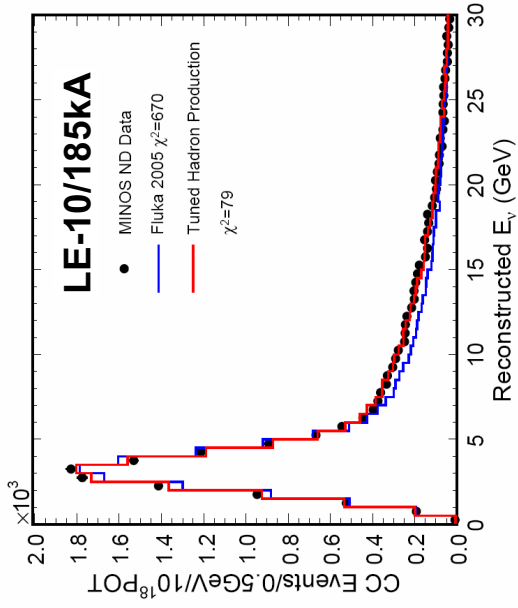
Track PH fraction



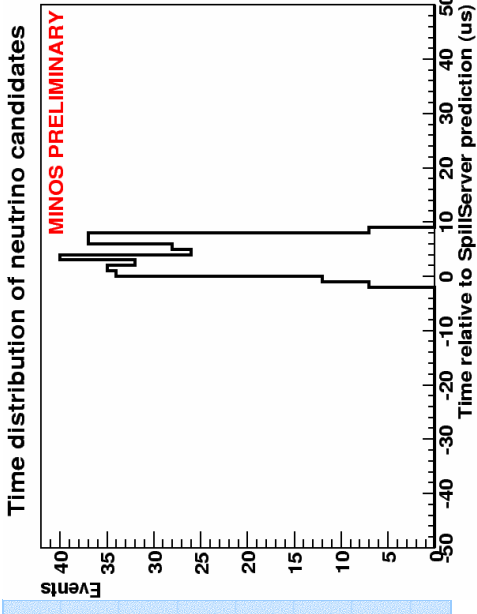
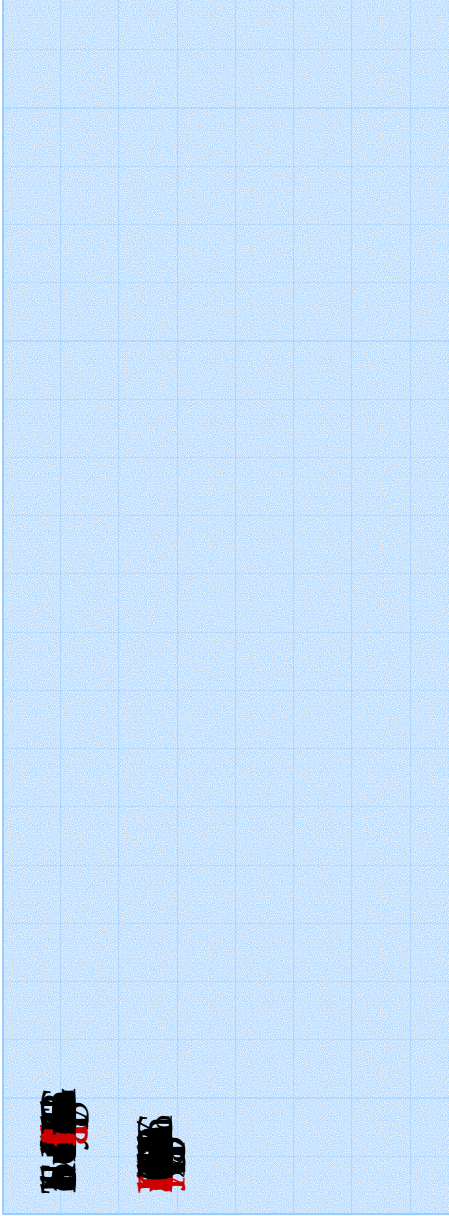
LE-10



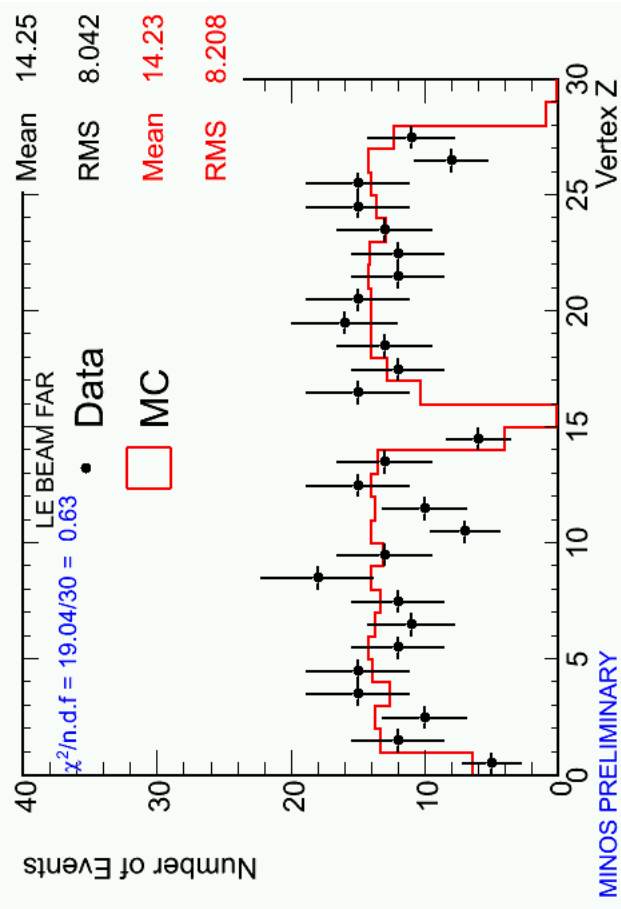
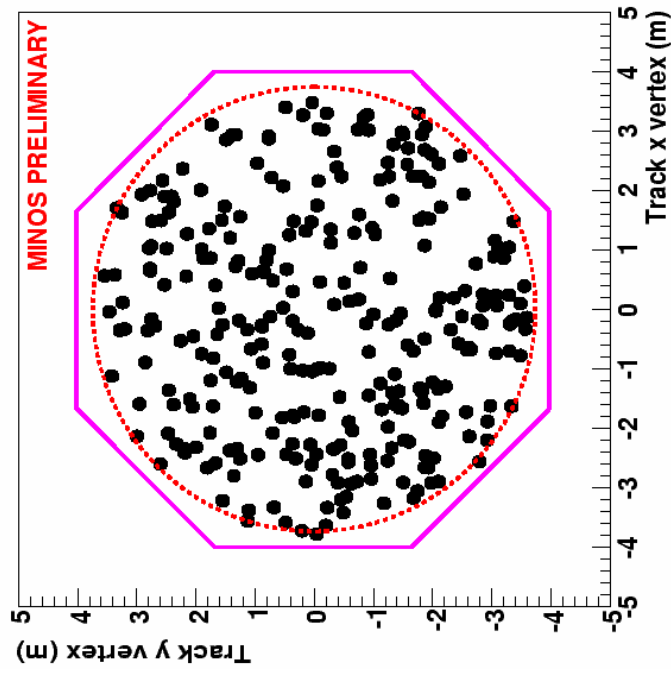
Hadron Production Tuning



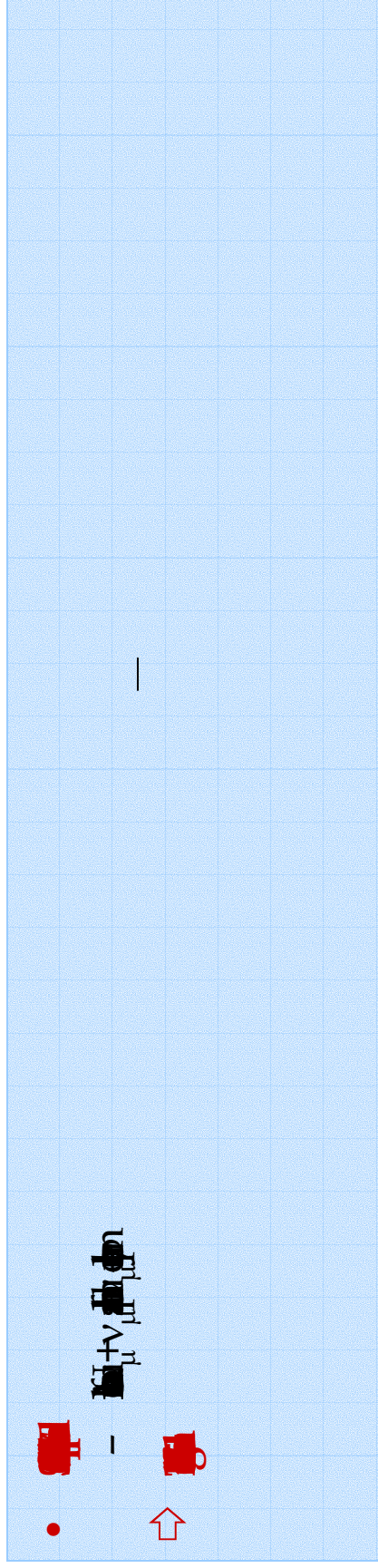
Far Detector Data



Reconstructed track vertices of neutrino candidates



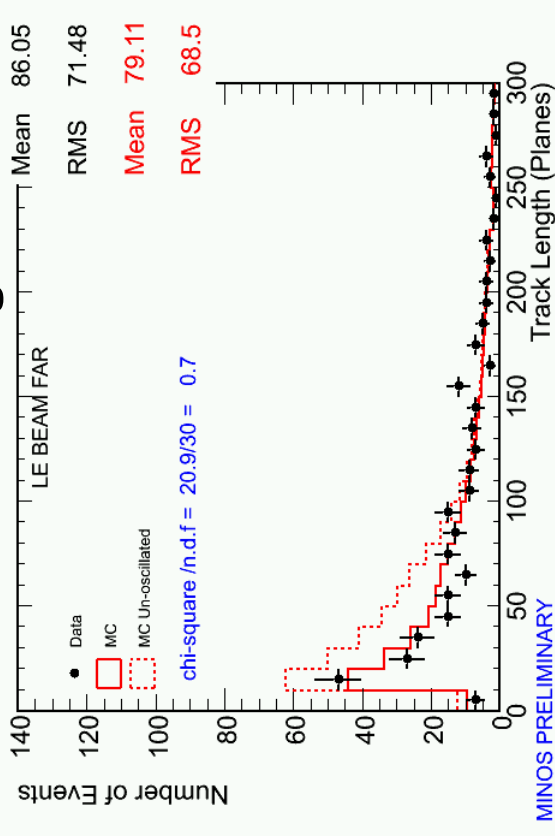
Observed/Expected Rates



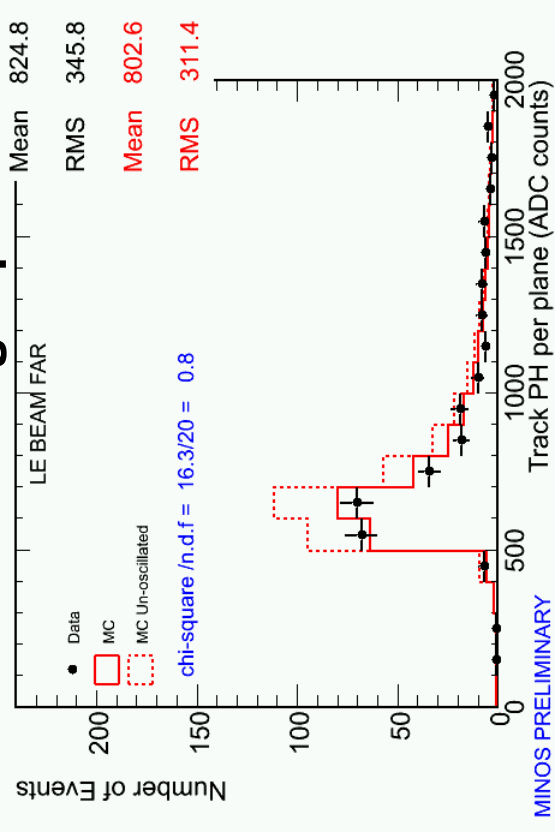
Data sample	observed	expected	ratio	significance
All CC-like events ($\nu_\mu + \bar{\nu}_\mu$)	204	298±15	0.69	4.1σ
ν_μ only (<30 GeV)	166	249±14	0.67	4.0σ
ν_μ only (<10 GeV)	92	177±11	0.52	5.0σ

Far Detector - PID Distributions

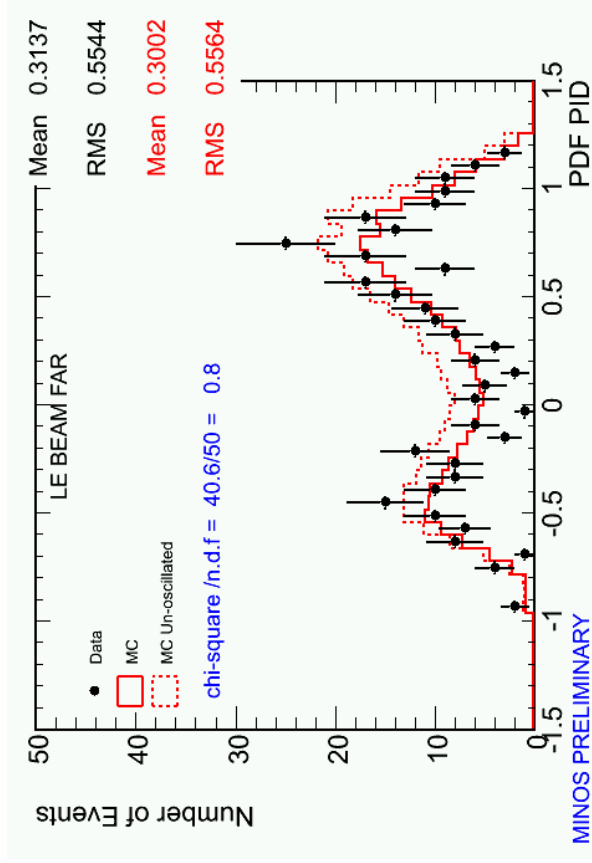
Track Length



Track Pulse Height per Plane

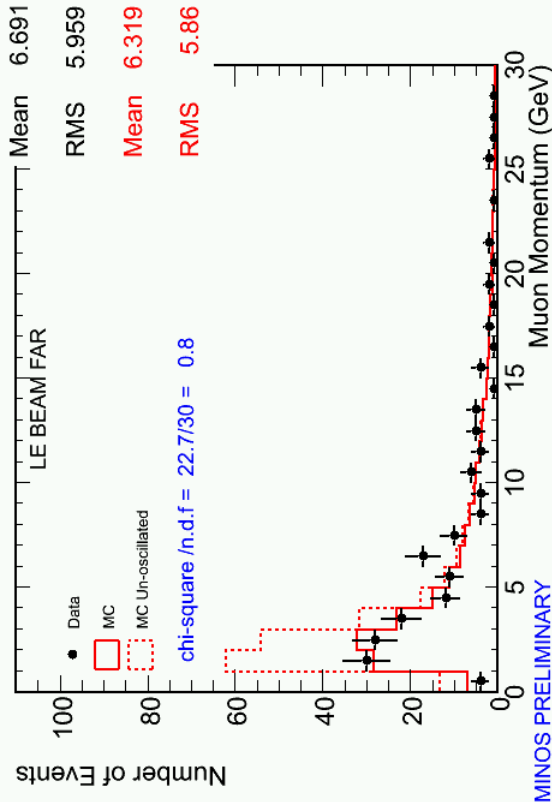


Particle Identification Parameter

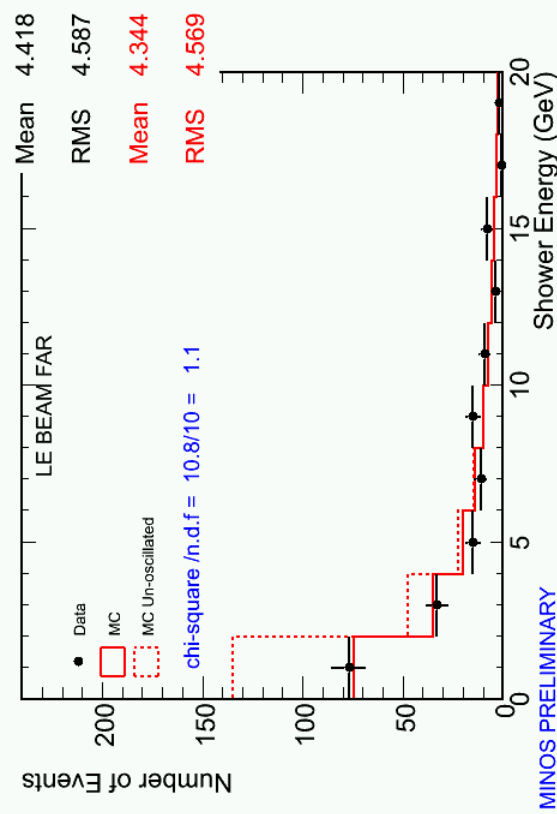


Far Detector - Physics Distributions

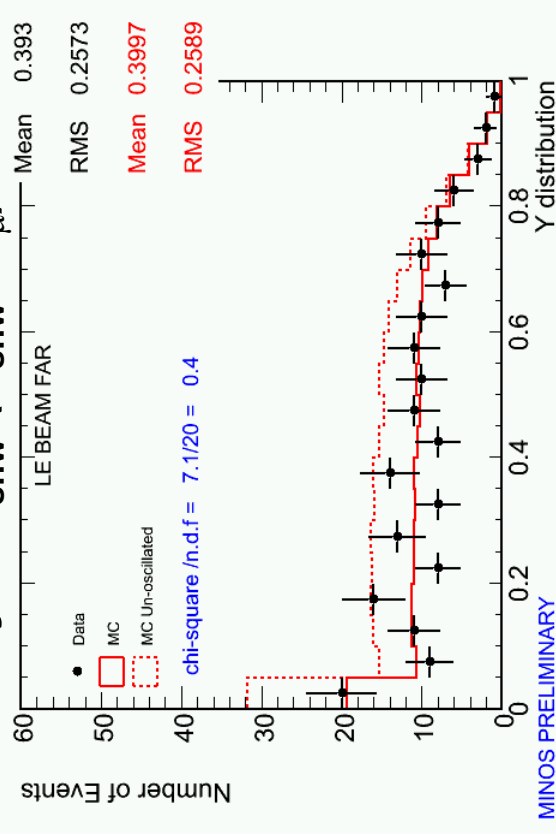
Muon Momentum (GeV/c)



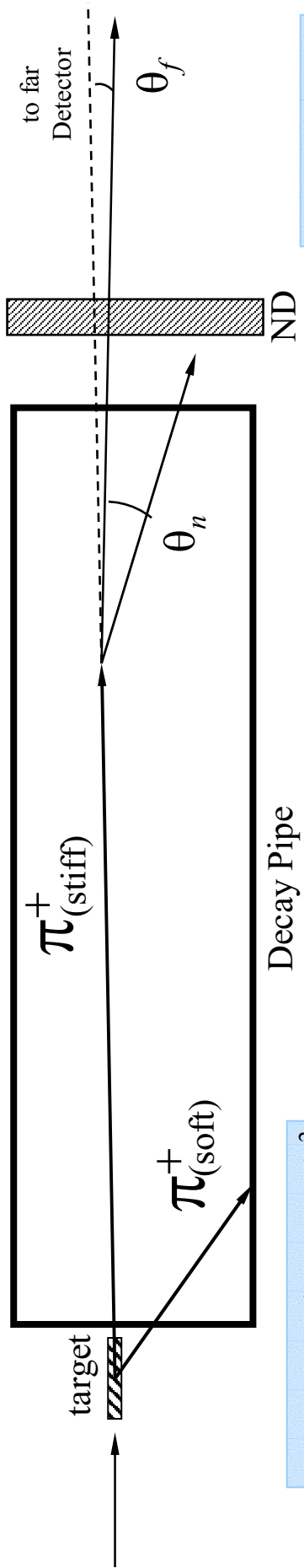
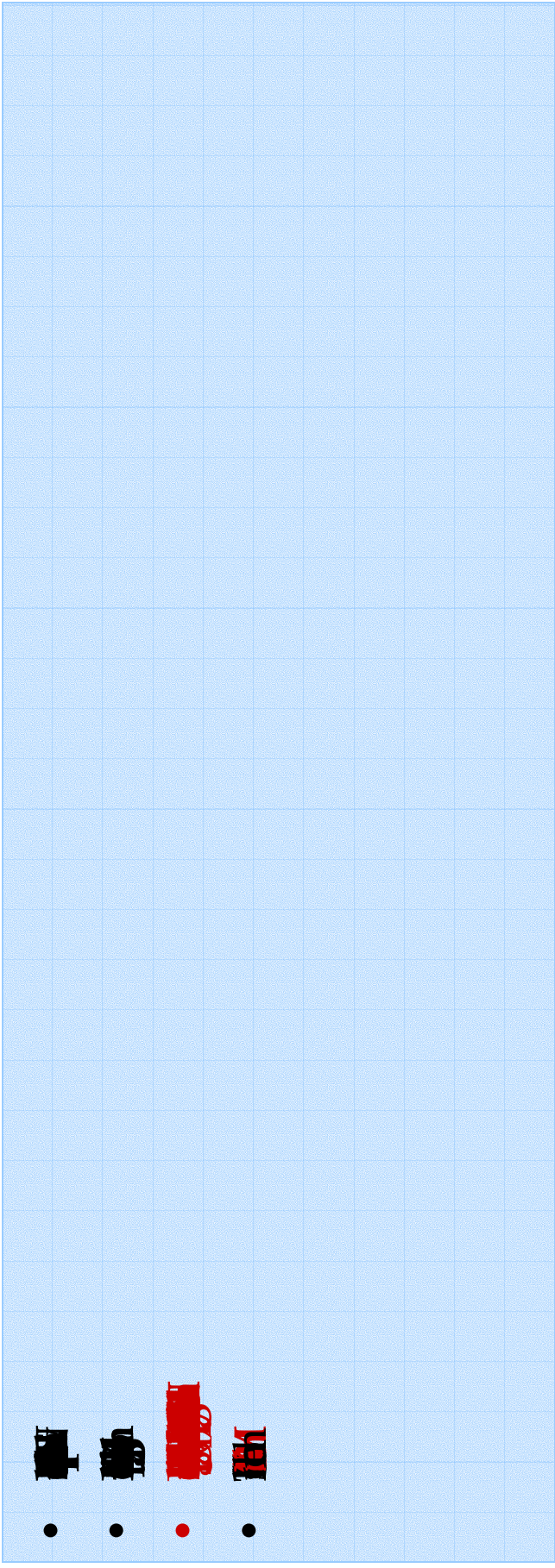
Shower Energy (GeV)



$$y = E_{shw} / (E_{shw} + P_{\mu})$$



Predicting the Far Detector Spectrum

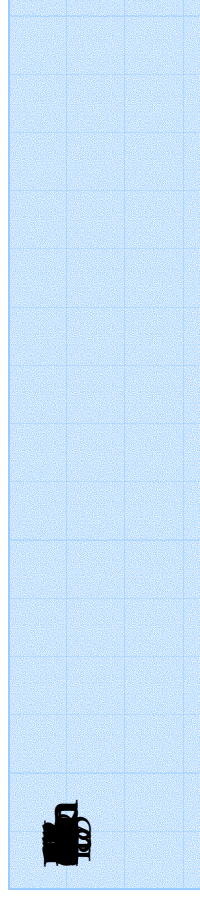
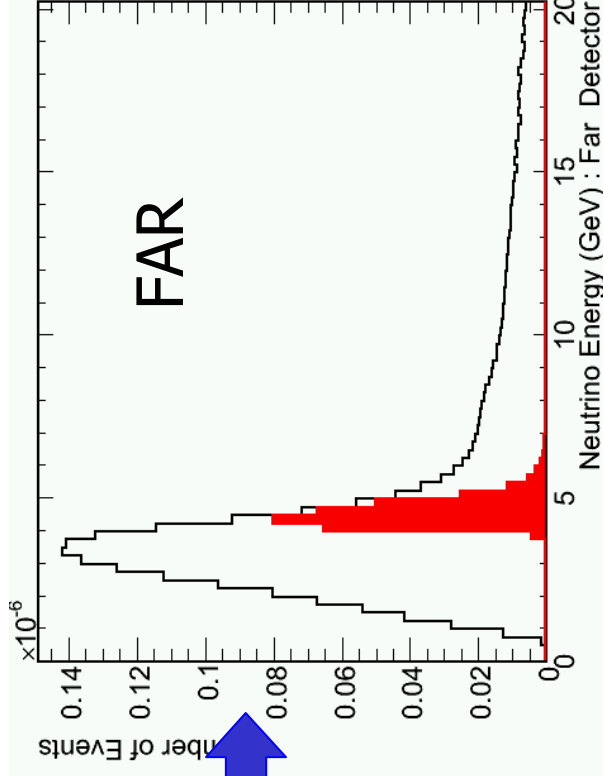
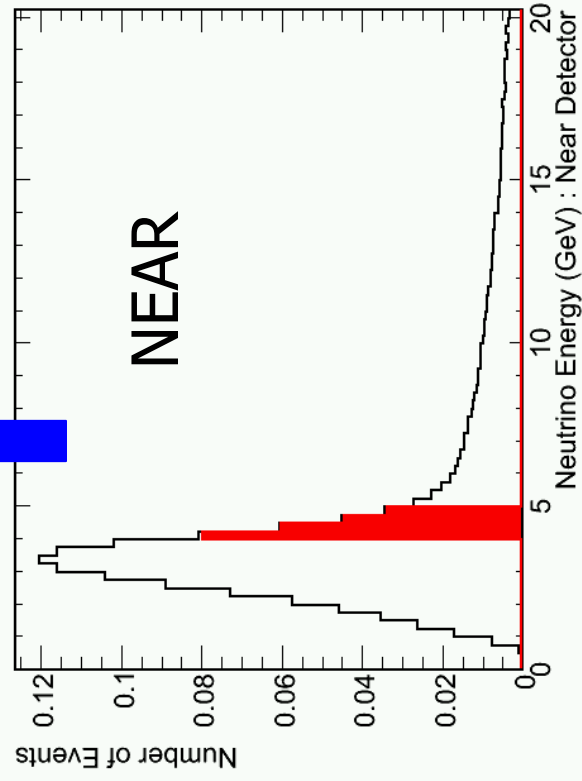
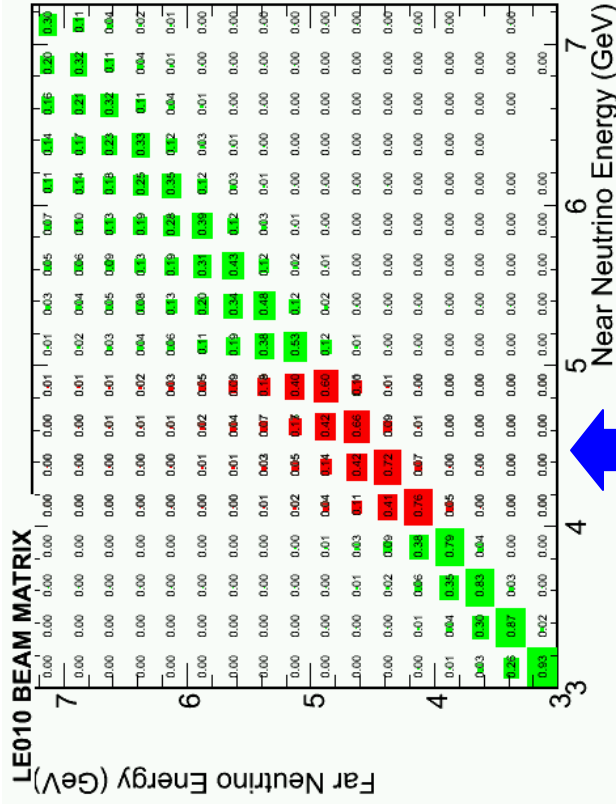


$$Flux \propto \frac{1}{L^2} \left(\frac{1}{1 + \gamma^2 \theta^2} \right)^2$$

$$E_\nu = \frac{0.43 E_\pi}{1 + \gamma^2 \theta^2}$$

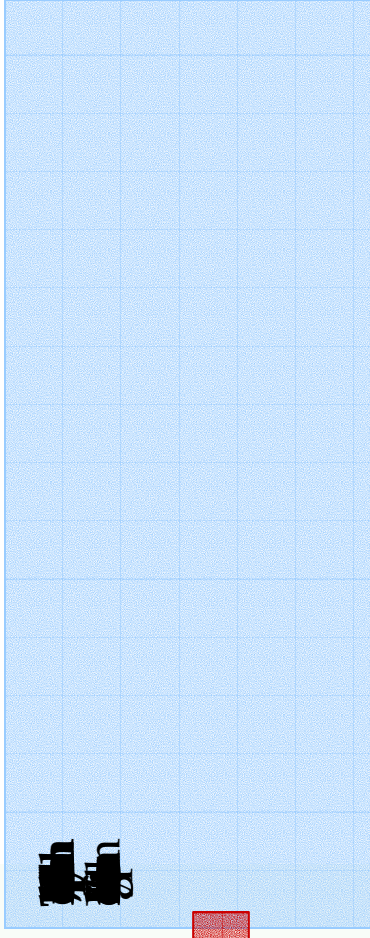
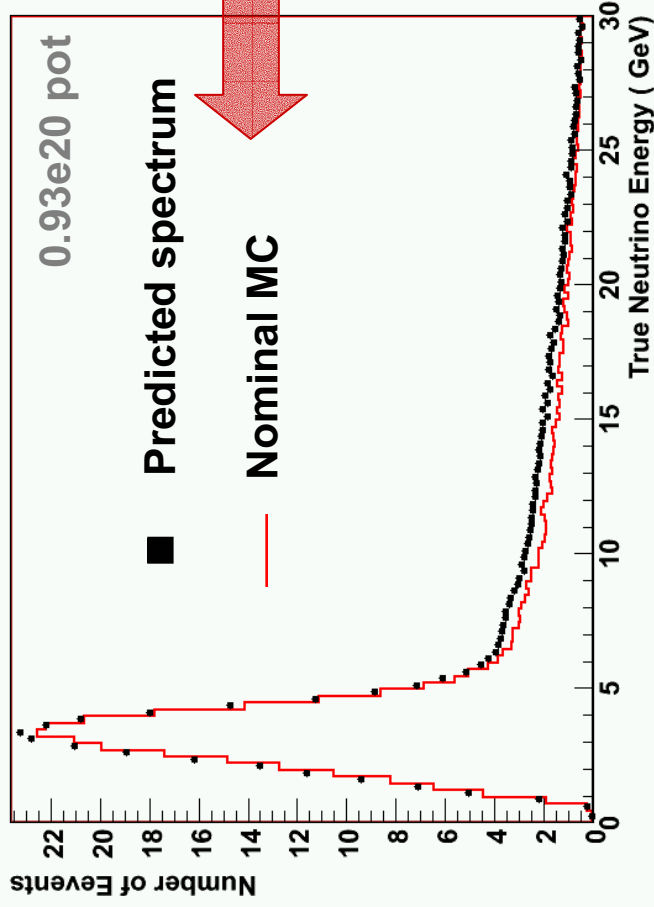


Beam Matrix Method

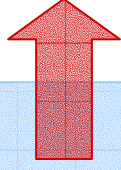
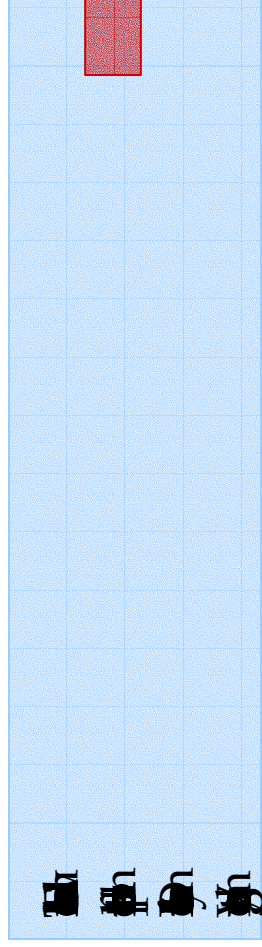
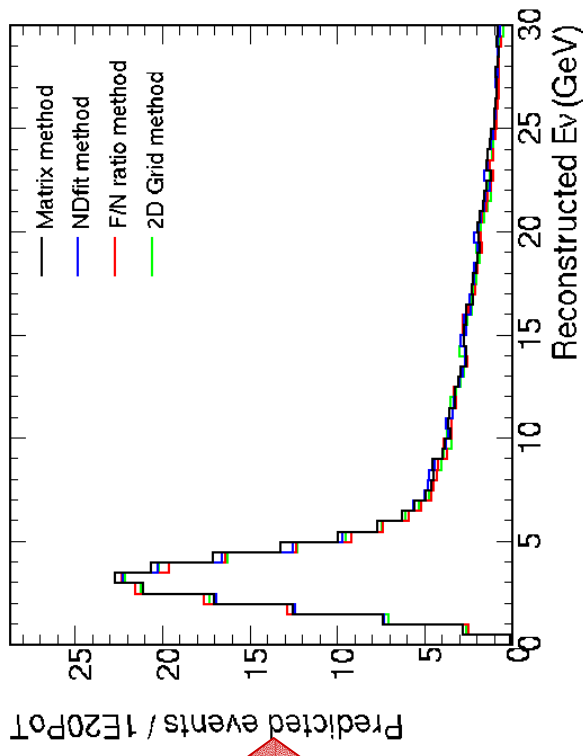


Predicted Spectrum

Predicted FD unoscillated spectra

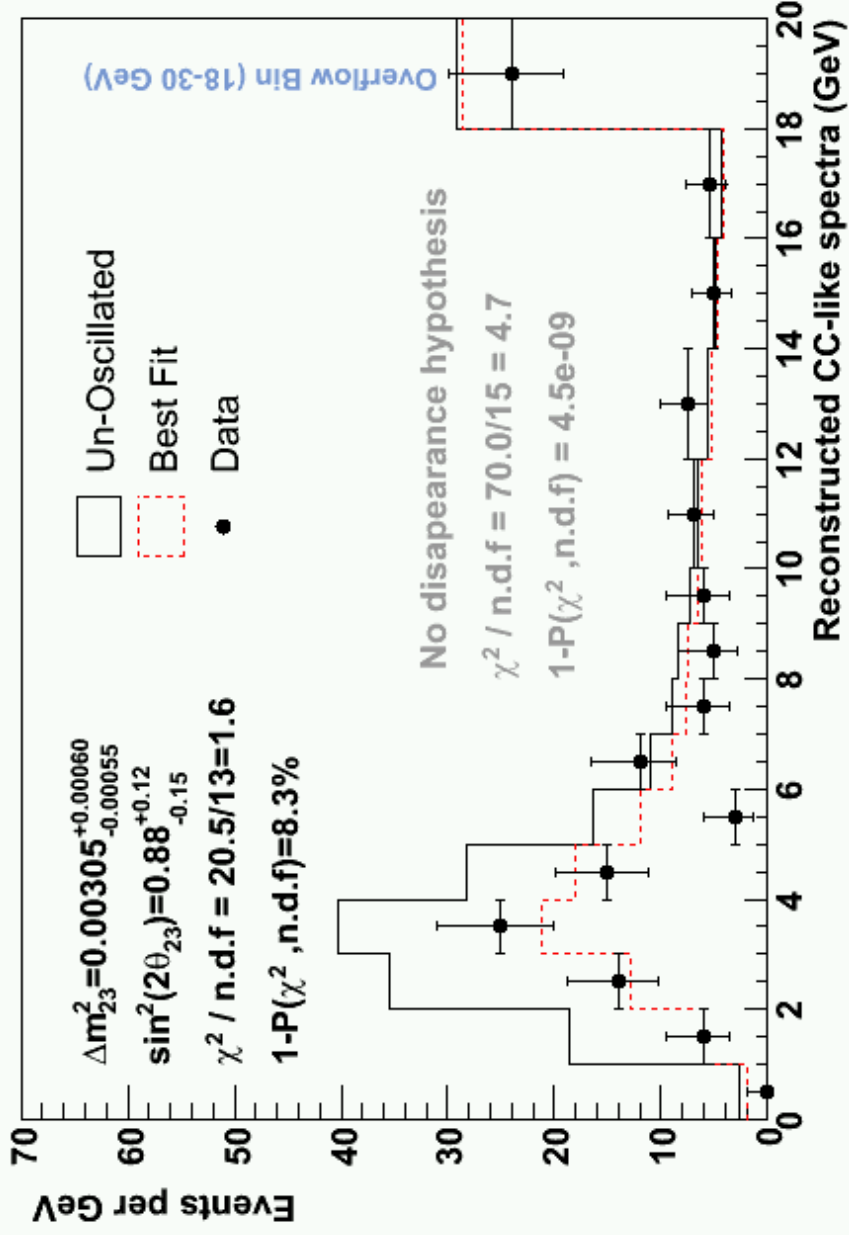


Predicted FD unoscillated spectra



Oscillation Parameters

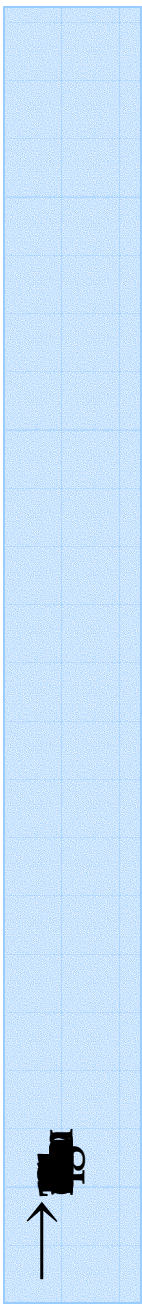
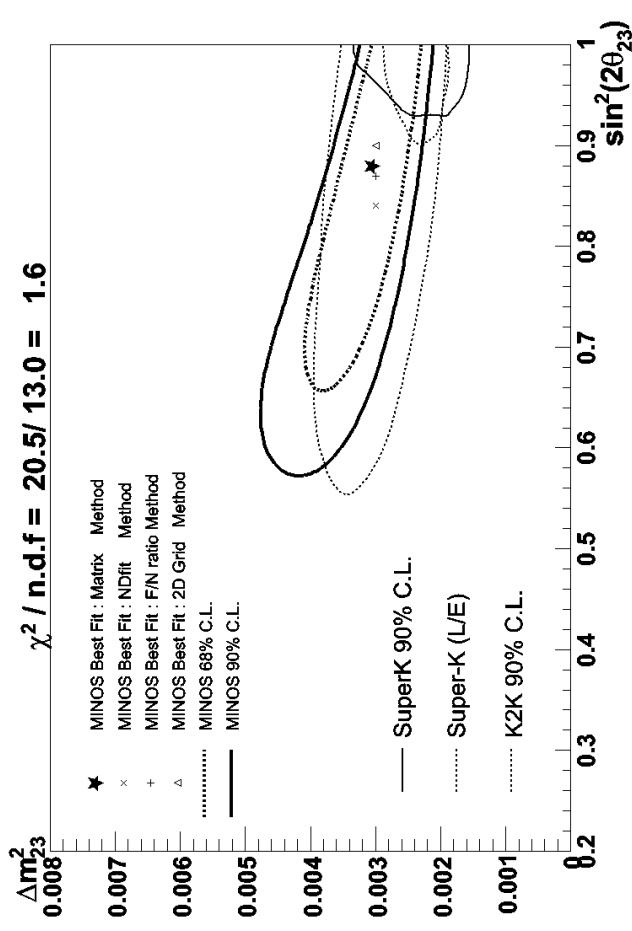
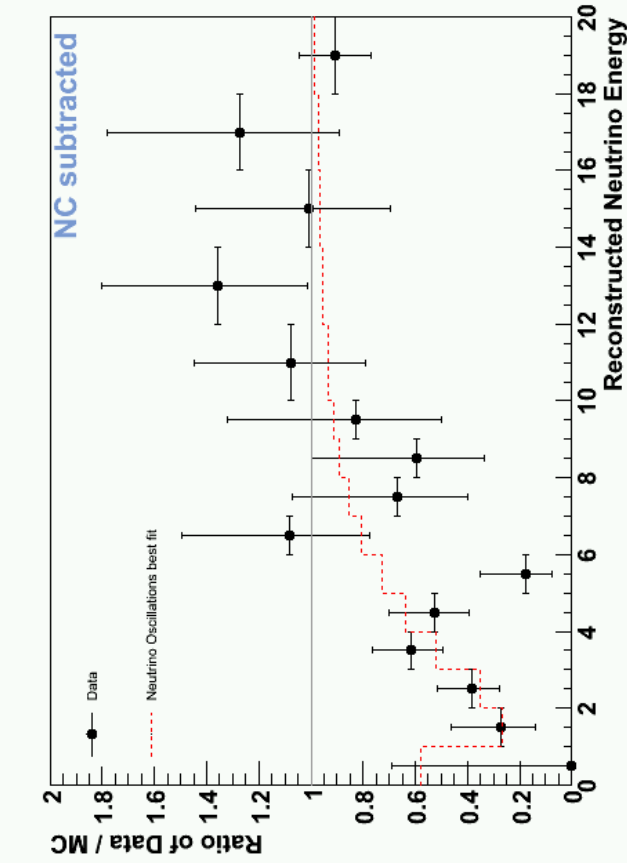
Oscillation Results for 0.93E20 p.o.t



$$\chi^2(\Delta m^2, \sin^2 2\theta) = \sum_{i=1}^{nbins} 2(e_i - o_i) + 2o_i \ln(o_i / e_i)$$

$o_i = \text{observed}$
 $e_i = \text{expected}$

Ratio Data/MC and Allowed Regions



Systematic Errors

- Systematic shifts in the fitted parameters have been computed with MC “fake data” samples for $\Delta m^2=0.003$ eV², $\sin^2 2\theta=0.9$ for the following uncertainties

Uncertainty	Δm^2 shift (eV ²)	$\sin^2 2\theta$ shift
Normalisation +/- 4%	0.63e-4	0.025
Muon energy scale +/- 2%	0.14e-4	0.020
Relative Shower energy scale +/- 3%	0.27e-4	0.020
NC contamination +/- 30%	0.77e-4	0.035
CC cross-section uncertainties	0.50e-4	0.016
Beam uncertainty	0.13e-4	0.012
Intranuclear re-scattering	0.27e-4	0.030
Total (sum in quadrature)	1.19e-4	0.063
Statistical error (data)	6.4e-4	0.15

Outlook

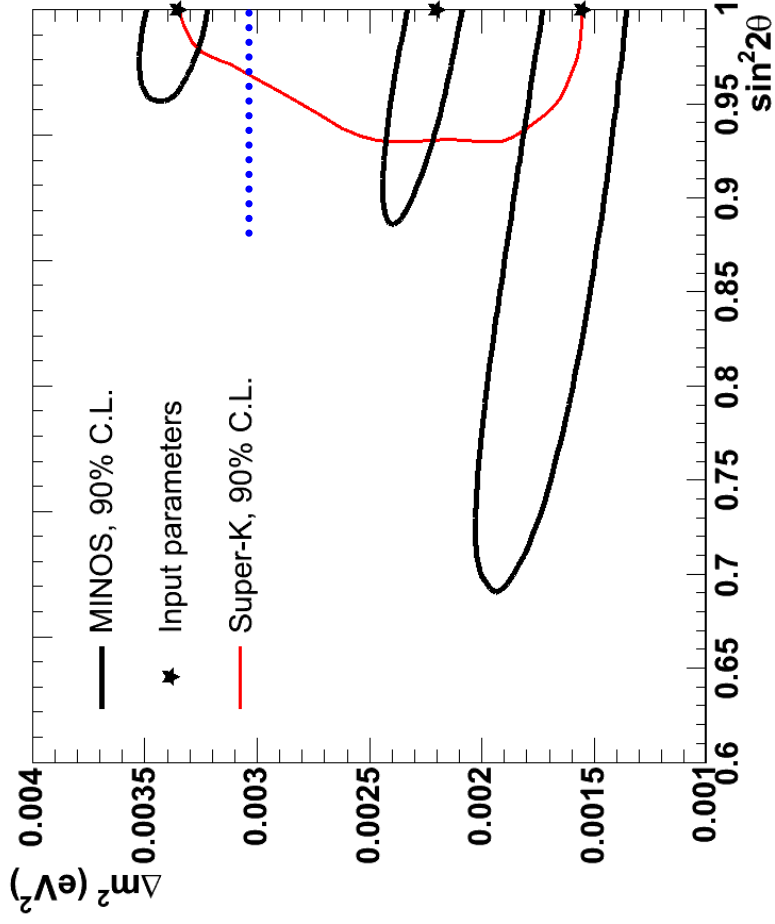
Δm_{23}^2

$\nu_{\mu} \rightarrow \nu_e$

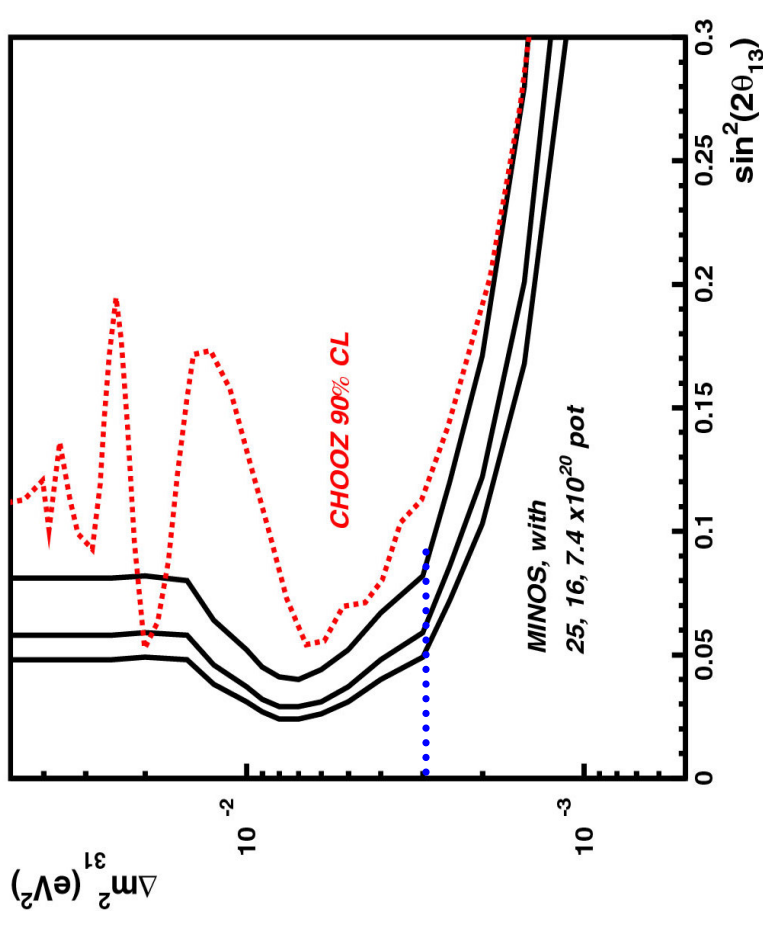


ν_{μ} disappearance




MINOS sensitivity, 16×10^{20} p.o.t.



$\nu_{\mu} \rightarrow \nu_e$ 3 σ Contours




Summary

- 
- $\delta m_{\mu\tau}^2$ 
- $\delta m_{\mu\nu}^2$ 

$$\Delta m_{23}^2 = 3.05_{-0.55}^{+0.60} (stat) \pm 0.12(syst) \times 10^{-3} eV^2$$
$$\sin^2 2\theta_{23} = 0.88_{-0.15}^{+0.12} (stat) \pm 0.06(syst)$$

- 

$$\Delta m_{23}^2 = 2.70_{-0.305}^{+0.263} \times 10^{-3} eV^2$$

- 
- 