#### Measuring the Dark Force at the LHC

Zhenyu Han UC Davis

# Reference: arXiv 0902.0006 (with Yang Bai @ Fermilab)

05/12/2009 @ PHENO 09

# Outline

#### Motivation

- Positron/electron excesses in cosmic rays, PAMELA, ATIC, FERMI...
- A class of dark matter models with a "dark force" mediated by a light (~GeV) particle: *a*
- Signature and measurements at the LHC
  - Lepton jets
  - Measurements: *M<sub>a</sub>*, *M<sub>DM</sub>*, *g*
  - A simple model as an illustration

Conclusion

**Positron/electron excesses in cosmic rays** 

- PAMELA: positron excess 10-100 GeV
- ATIC: positron/electron excess 300-800 GeV
- Fermi LAT: disfavors ATIC, but hint of positron/electron excess (?)

Dark Matter annihilation to positron/electrons? What is the implication for the LHC?

#### **DM annihilation** (Arkani-Hamed, Finkbeiner, Slatyer and Weiner)

- Dark matter annihilates to a light particle *a*
- The particle *a* dominantly decays to leptons
- Sommerfeld enhancement to give a large cross-section



# Major ingredients for the LHC

- The relevant ingredients for the LHC:
  - An O(GeV) light particle *a* couples to the DM with order one coupling constant, mediating a "dark force"
  - The light particle *a* dominantly decays to leptons
- Two extra assumptions
  - DM (or other particles charged under the dark force) is produced at the LHC
  - The particle *a* decays within the detector
    - Collinear leptons, "lepton jets" signature

### **Measuring the Dark Force**

• What to measure:  $M_{DM}$ ,  $M_a$ , g

- Crucial for calculating DM-DM annihilation rate.
- Consistent with PAMELA, ATIC, Fermi ...?
- Give the right relic density?



# The strategies

- $M_{a}$ : measuring the invariant mass of the "lepton jet".
- **M**<sub>DM</sub> : model-dependent, edges, m<sub>T2</sub>, kinematic constraints.....
- The coupling *g*: dark radiation
  - For any process containing a DM, there is another one with an extra *a* radiated (extra "lepton jet")
  - Significant rate:  $g \sim O(1)$ ,  $Ma \sim O(GeV)$



# A simple model with hidden U(1)

- A usual MSSM sector + hidden sector
- MSSM has a bino-like (N)LSP
- The dark sector: a supersymmetric (broken) U(1) gauge theory with Higgsino-like LSP, lighter than bino.
- DM: dark Higgsino, mediator: dark photon *a*<sub>dark=</sub>*a*
- Gauge mixing: *a* decays to leptons, MSSM bino decays to dark Higgsino +  $h_{dark}$



### **Benchmark numbers**

#### MSSM

- MSSM LSP: 700 GeV
- Squark 1000 GeV, gluino 1200 GeV, 0.84 pb for squark/gluino production. 8400 events for 10 inverse fb
- All squarks decays directly to bino + quark

Dark sector

- $M_{DM} = 600 \ GeV$
- $M_a = 1 \text{ GeV}, M_h = 3 \text{ GeV}$  (fine tuned), h > aa > 4l ("h-jet")
- Coupling g = 0.41 (to give the correct relic density)

Generated events with Madgraph/BRIDGE/Calchep/PYTHIA/PGS

# **Identify the lepton jets**

Group muons in small cones

- All muons are sorted by  $p_T$
- Take the highest  $p_T$  muon as the seed of a lepton jet.
- Add muons within 0.2 *rad* of the seed muon to the jet.
  Remove used muons from the list.
- Repeat until all muons used

Tag the jets

- 4 or 3 muons: *h-jet*
- 2 muons: *a-jet*
- 5 or 6 muons? *a* and *h* tends to be colinear: "*h*&*a*-*jet*"

M<sub>h</sub>, M<sub>a</sub> Measurements





2-muon jets and 4-muon jets

J/psi, dimuon, CMS Technical Design Report

Resolution ~ m/100

Resolution  $\sim m/30$ 

More precise than the other measurements

# Determining *m<sub>DM</sub>* using kinematic constraints



Count the number of events consistent with assumed masses.

 $m_{\tilde{\chi}_{\text{dark}}} = 616 \pm 12 \text{GeV}, \ m_{\tilde{\chi}_{1}^{0}} - m_{\tilde{\chi}_{\text{dark}}} = 101.6 \pm 0.6 \text{GeV}$ 

# Measuring the coupling g

Dark radiation: 2-body decay vs 3 body-decay



# Determining g

- For *g*=0.41, 10 inverse fb, expect 230 *2h1a* (three-body decay) events, 4k *2h* events. *2h1a* dominates the error.
- Count the number of events with 2h-jets + 1 a-jet or 1h-jet + 1 h&a-jet (5 or 6 muons): 70 events identified

 $g=0.40\pm0.03$ 

# **Relic density**

Calculate the DM relic density (10 inverse fb)

 $\Omega_{DM}h^2 = 0.119 \pm 0.033$ Compare with WMAP error:

 $\Omega_{DM}h^2 = 0.113 \pm 0.0034$ 

Not as precise, but encouraging if consistent

## **Conclusion and outlook**

- It is possible to measure the dark force at the LHC
  - Important for calculating DM annihilation rate
  - Illustrated with a simple model
- Many theoretical possibilities unexplored
- Electrons experimentally more challenging





Approximate formula for  $M_h = 3M_a$ :

$$R \approx \frac{11 g^2}{120 \pi^2} \left[ 4((\log (1 - r_{\rm DM}) - 4 \log 2 - 2) \log (1 - r_{\rm DM}) + \log^2 r_a - (4 \log (1 - r_{\rm DM}) - 8 \log 2 - 4) \log r_a + 3 \log^2 2 + 4 \log 2 + 2) \right], \qquad (2)$$

# The electrons

"Electron jets" characterized by ECAL energy deposit, no/small HCAL energy deposit and multiple tracks from the interacting vertex.

- Contamination from converted photons?
- What's the efficiency for identifying electron jets? electon+muon jets?
- To measure the invariant mass, have to measure individual electrons' momentum in a jet. Prefer relatively soft electrons:
  - Can be separated by the magnetic field before they hit the ECAL (~10GeV)
  - Better measured by the tracker.