

Dark Matter Out of Seclusion *

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- arXiv:0903.0363

- arXiv:0903.3396

Overview

- Secluded WIMPs
- Out of seclusion via
 - Colliders
 - * Higgs-strahlung at B-factories
 - Direct detection experiments
 - * Endothermic and exothermic inelastic scattering
 - * Higher order elastic scattering

Secluded Dark Matter

Pospelov, Ritz, Voloshin '07

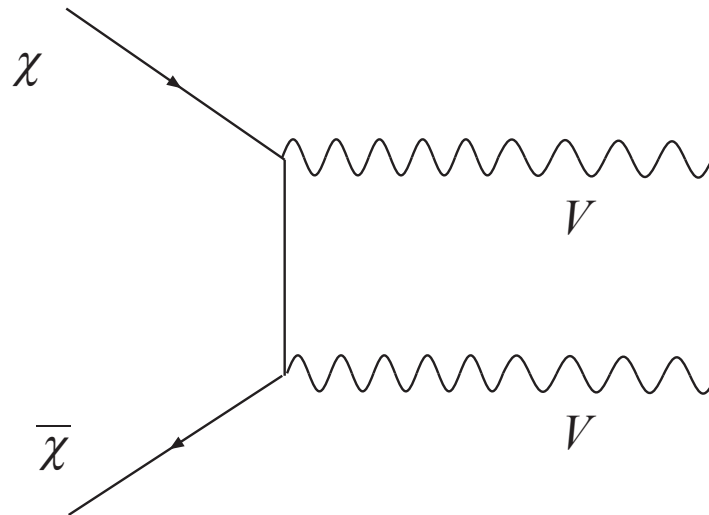
$$\chi \text{ ----- } V_{\mu\nu} B^{\mu\nu} \text{ ----- SM}$$

Holdom '86

“Secluded” regime

- $m_\chi > m_V$
 - Annihilation via $\bar{\chi}\chi \rightarrow VV$
 - Relic abundance independent of kinetic mixing
 - Makes direct detection, collider signatures tricky if mixing small.

Can't rule out WIMP hypotheses in principle if secluded



Astrophysical signatures

1) Light mediator \longrightarrow enhanced galactic annihilation cross section

- e.g. Sommerfeld enhancement: long range force distort wavefunctions near the origin from the plane wave to the Coulomb-type.

Arkani-Hamed, Finkbeiner, Slatyer, Weiner '08

Pospelov, Ritz '08

$$\mathcal{N} \sim \frac{\pi\alpha'}{v}$$

2) $m_V \leq \text{GeV} \implies$ vector won't decay to (anti-)protons by kinematics

Connection with PAMELA, ATIC, FERMI, HESS, ...?

B-factory Signatures

BB, Pospelov, Ritz '09;
Essig, Schuster, Toro '09;
Reece, Wang '09;

B-factories

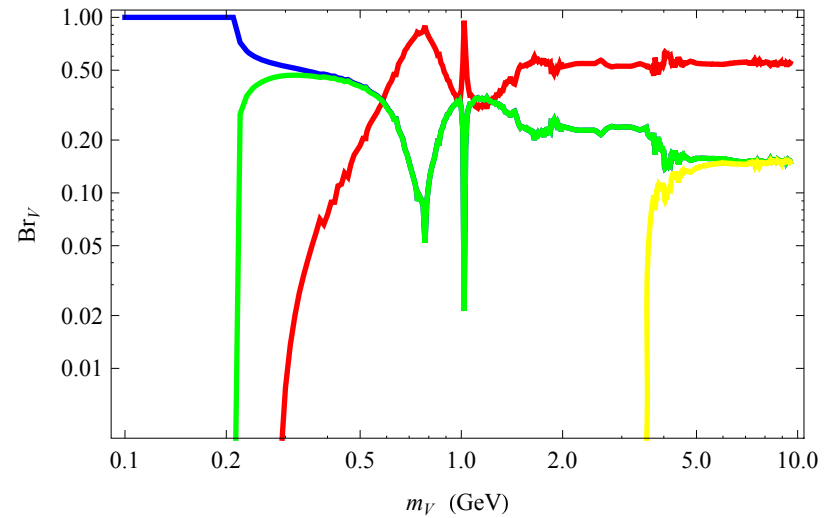
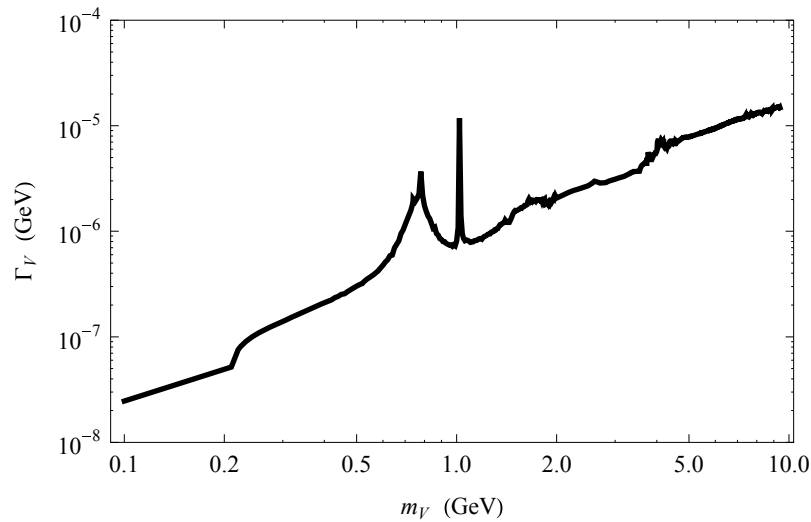
B-factory (BaBar and Belle) advantages:

- Large data sets $\sim 500 \text{ fb}^{-1}$
- COM energy $\sqrt{s} \simeq 10 \text{ GeV}$, close to masses of new particle

$$\mathcal{L}_{int} = -\frac{\kappa}{2} V_{\mu\nu} F^{\mu\nu} + \frac{m_V^2}{v'} h' V_\mu^2$$

- Higgs'-strahlung: $e^+e^- \rightarrow \gamma^*, V^* \rightarrow h'V$
- Cross section $\sigma \sim 20$ fb
- Leads to 6 lepton final state

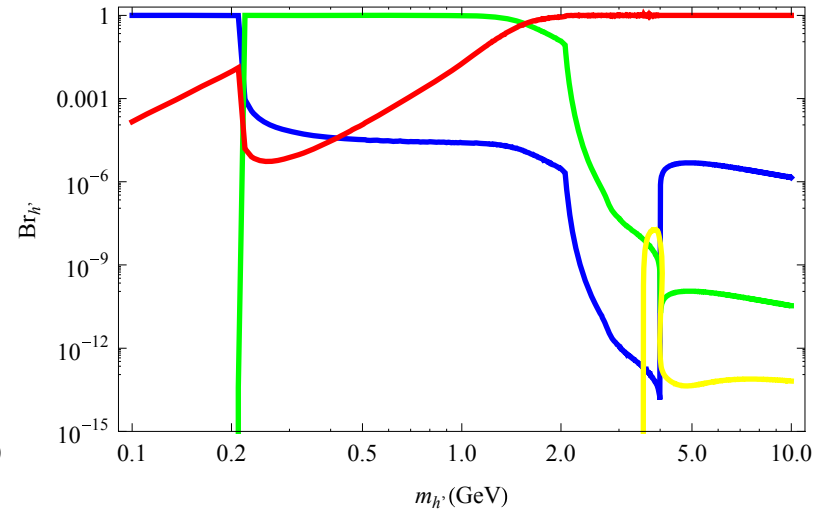
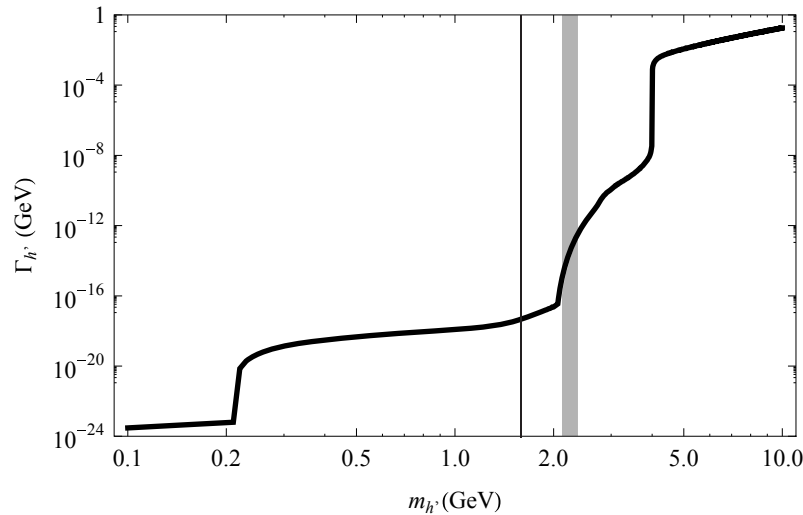
V_μ decays



- $V \rightarrow \text{hadrons}$ (Red)
- $V \rightarrow e^+e^-$ (Blue)
- $V \rightarrow \mu^+\mu^-$ (Green)

V_μ always has a significant branching to leptons

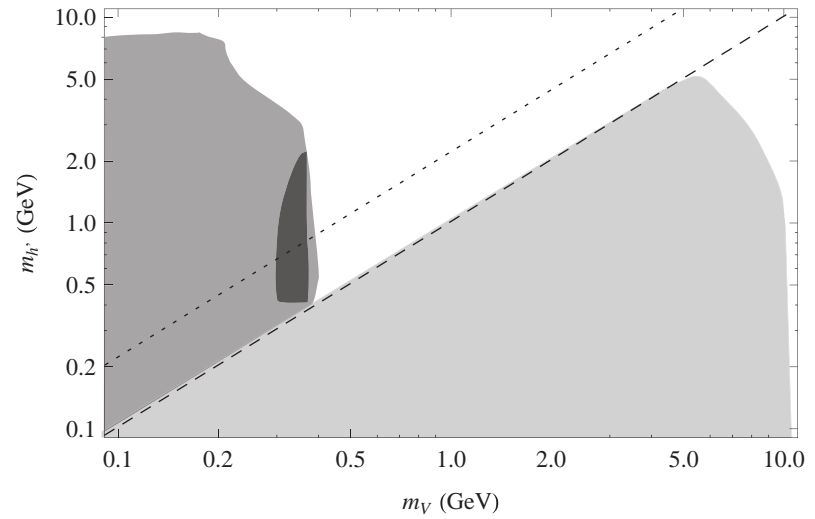
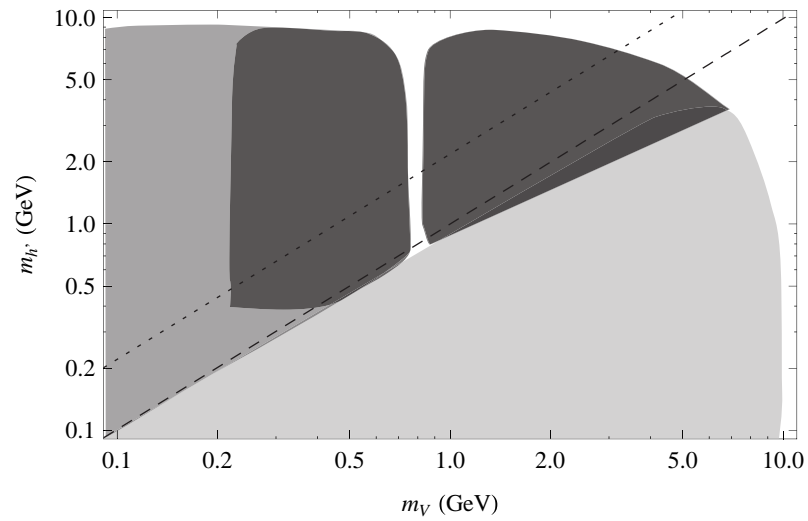
h' decays



- $h' \rightarrow VV$ (Red)
- $h' \rightarrow e^+e^-$ (Blue)
- $h' \rightarrow \mu^+\mu^-$ (Green)

-Decays of h' depend on whether $m_{h'} > m_V$ or $m_{h'} < m_V$

Sensitivity



- $2l+$ missing energy (light)
- $6e$ (medium)
- 6μ (dark)

Direct Detection

Finkbeiner, Slatyer, Weiner, Yavin '09;
BB, Pospelov, Ritz '09;

Multi-component WIMPs

- If WIMP is a complex scalar or Dirac fermion, the real components χ, χ' can be split after $U(1)_S$ is broken so that $\Delta m = m_{\chi_2} - m_{\chi_1}$

$$\mathcal{L} = \sum_{i=1,2} \left(i\bar{\chi}_i \bar{\sigma}_\mu \partial_\mu \chi_i - \frac{1}{2}(m\chi_i\chi_i + \text{h.c.}) \right) - ie'V_\mu(\bar{\chi}_1\bar{\sigma}_\mu\chi_2 - \bar{\chi}_2\bar{\sigma}_\mu\chi_1)$$

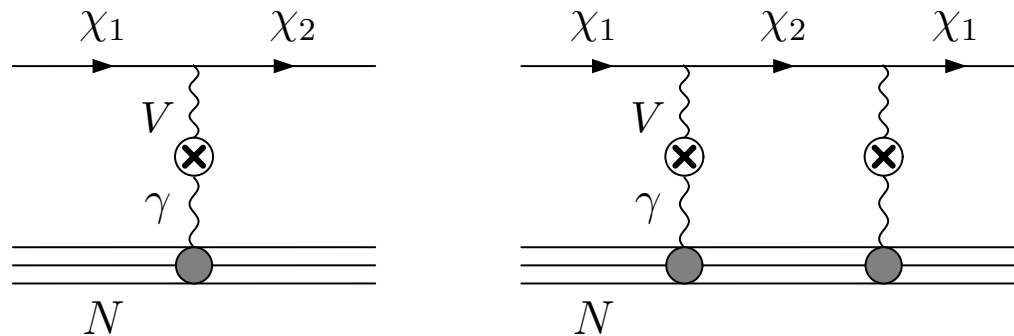
- No tree level elastic scattering process $\chi_1 N \rightarrow \chi_1 N$
- Example of *inelastic dark matter*

Smith, Weiner '01

- Direct detection depends sensitively on Δm

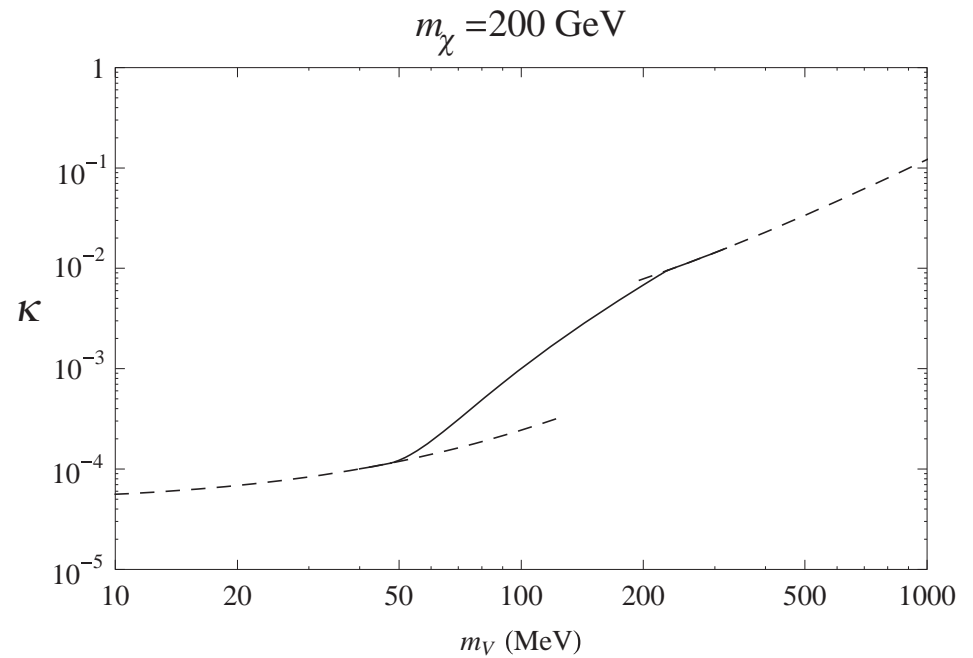
Rich scattering structure

- *Endothermic inelastic:* $\chi_1 N \rightarrow \chi_2 N$
- *Exothermic inelastic:* $\chi_2 N \rightarrow \chi_1 N$ (depends on χ_2 population)
- *Second-order elastic:* $\chi_1 N \rightarrow \text{virtual states} \rightarrow \chi_1 N$



Elastic: $\chi_1 N \rightarrow \chi_1 N$

- CDMS constraints
($\Delta m = 10$ MeV)



Sensitivity diminishes as WIMP probes nucleus

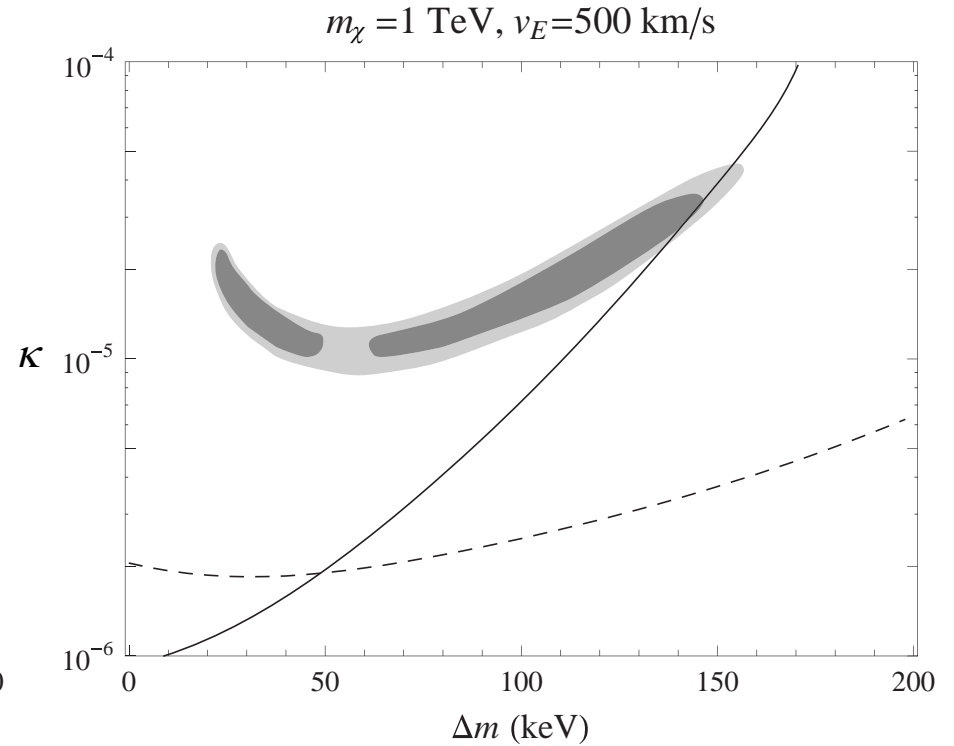
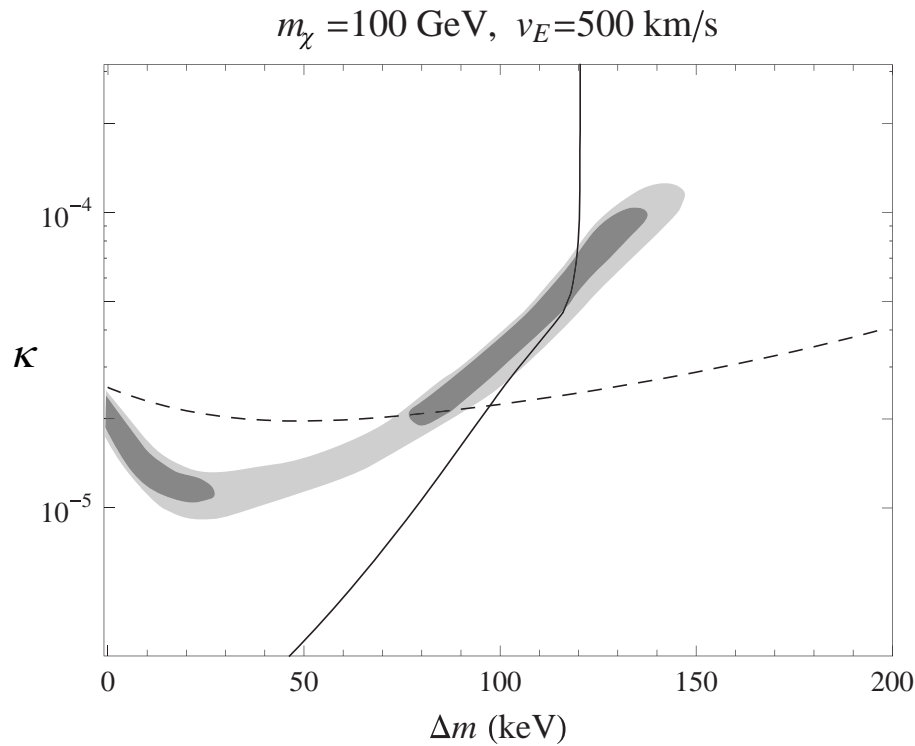
Exothermic: $\chi_2 N \rightarrow \chi_1 N$

- Requires substantial population of excited WIMPs χ_2
- Decay of χ_2 in 'minimal' $U(1)_S$ model:
 - $\Delta m > 2m_e \implies$ rapid decay $\chi_2 \rightarrow \chi_1 + e^+e^-$
 - $\Delta m < 2m_e \implies$ Loop induced $\chi_2 \rightarrow \chi_1 + 3\gamma$

$$\tau > 4 \times 10^{-47} \text{ GeV} \times \left(\frac{\kappa}{10^{-3}} \right)^2 \left(\frac{\Delta m}{100 \text{ keV}} \right)^{13} \left(\frac{100 \text{ MeV}}{m_V} \right)^4$$

- χ_2 lifetime longer than the age of the universe for small Δm .

Exothermic: $\chi_2 N \rightarrow \chi_1 N$



- DAMA preferred region 90(99)% CL - dark(light)
 - endothermic (solid)
 - exothermic (dashed)

Conclusions

- Secluded WIMPS
- B -factory signals:
 - Higgs'-strahlung \rightarrow multi-lepton final state
 - Probe $U(1)_S$ couplings $\kappa \sim O(10^{-2} - 10^{-3})$
- Rich direct detection phenomenology
 - Endothermic, exothermic, and (2nd order) elastic scattering provide sensitivity for different ranges of parameter space