

Testing Gaugino Mass Hypotheses at the LHC

Brent D. Nelson



with B. Altunkaynak, M. Holmes,
P. Grajek and G. Kane

[hep-ph/0901.1145](#)
[JHEP 04 \(2009\) 114](#)

Pheno Symposium 2009

May 11, 2009

- **Q:** LHC era is (almost) here – *what should theorists be doing now?*

- **Q:** LHC era is (almost) here – *what should theorists be doing now?*
- **A1:** Short run: improving the toolkit
 - ★ Boosted tops
 - ★ M_{T2} , $M_{T\text{Gen}}$, etc.
 - ★ Jet reconstruction algorithms
 - ★ Spin measurement techniques
 - ★ Standard Model backgrounds
 - ★ ...

- **Q:** LHC era is (almost) here – *what should theorists be doing now?*
- **A1:** Short run: improving the toolkit
 - ★ Boosted tops
 - ★ M_{T2} , $M_{T\text{Gen}}$, etc.
 - ★ Jet reconstruction algorithms
 - ★ Spin measurement techniques
 - ★ Standard Model backgrounds
 - ★ ...
- **A2:** Medium run: think two to three years out.
Post-discovery, what will be the big issues?
 - ★ Assume SUSY is established early on by standard methods – *what next?*
 - ★ Will want to connect observations to some underlying framework
 - ★ For me: an underlying *string* framework (hopefully)

- **Q:** How is this best accomplished?

- **Q:** How is this best accomplished?
- **A1:** Through fits to *simple benchmark models* Allanach et al., Ellis et al.
 - ★ For example, mSUGRA at first
 - ★ Then perturbed sequentially using goodness-of-fit as our guide
 - ★ Effectively a “directed walk through theory space”
 - ★ **Pro:** Completely and utterly obvious thing to do
 - ★ **Pro:** Quasi-model independent
 - ★ **Con:** Computationally intensive
 - ★ **Con:** May saturate its effectiveness early on in LHC era
 - ★ **Con:** Suffers from the “**LHC inverse problem**”
Arkani-Hamed et al., JHEP 0608 (2006) 070

- **A2:** Via the method of *model footprints* Kane et al.
 - ★ Pick models and map their parameter space onto signature space
 - ★ **Pro:** *Guaranteed* to have consistent high-scale model
 - *assuming LHC data falls in at least one footprint! (See below)*
 - ★ **Pro:** Requisite analysis can be done once and for all “offline,” ahead of the experiment
 - ★ **Pro:** LHC inverse problem less of an issue
 - ★ **Con:** Goodness-of-fit trickier to determine
 - ★ **Con:** “**Straw-man effect**”: method best at discriminating between models, not so much establishing a model
 - ★ **Con:** Maximally model-dependent: only as good as our imagination in dreaming up models (the problem with all top-down methods)

- **A3:** Try to extract *broad characteristics* of the underlying theory
 - ★ Measurement of $m_{\tilde{N}_2} - m_{\tilde{N}_1}$
 - ★ Measurement of $m_{\tilde{N}_2}$ itself
 - ★ Extraction of the values of $M_2, \mu, \tan \beta, \dots$
 - ★ Evidence of how the μ -term was generated in the first place

- **A3:** Try to extract *broad characteristics* of the underlying theory
 - ★ Measurement of $m_{\tilde{N}_2} - m_{\tilde{N}_1}$
 - ★ Measurement of $m_{\tilde{N}_2}$ itself
 - ★ Extraction of the values of $M_2, \mu, \tan \beta, \dots$
 - ★ Evidence of how the μ -term was generated in the first place
- ★ **Pro:** LHC inverse problem completely avoided
- ★ **Pro:** Directly probe the things theorists most care about
(e.g. [gaugino mass unification](#)) [Binetruy, Kane, Lykken and BDN, J. Phys. G32 \(2006\) 129](#)
- ★ **Pro/Con:** Never been tried before...
- ★ **Con:** ...because it's outrageously difficult!
- Want to know this independent of everything else that's going on with the supersymmetry breaking Lagrangian (if possible)
- Big job: need a **tractable and concrete** starting point

- Mirage pattern of gaugino masses – a one-parameter family:

$$M_1 : M_2 : M_3 \simeq (1 + 0.66\alpha) : (2 + 0.2\alpha) : (6 - 1.8\alpha)$$

- A logical first step
 - ★ Easy to understand and visualize
 - ★ Interpolates between mSUGRA ($\alpha = 0$) and AMSB limit ($\alpha \rightarrow \infty$)
 - ★ Motivated by a variety of constructions, including string theory (heterotic and Type II) as well as “deflected” AMSB
 - ★ Disadvantage: Only one-parameter family of models \Rightarrow not fully general
- **All** values of α correspond to a unified pattern – the only issue is at which *energy scale* they unify Choi & Nilles, JHEP 0704 (2007) 006
 - ★ When $\alpha = 0$ gaugino masses unify at $M_{\text{GUT}} \simeq 2 \times 10^{16}$ GeV
 - ★ Other α values give effective unification scale elsewhere (hence “mirage”)
 - ★ Example: $\alpha = 2$ gives $M_1 \simeq M_2 \simeq M_3$ at low-energy scale
 - ★ Scale dependent! Coefficients change with scale (here 1 TeV)

A Quick Derivation of the Mirage Pattern (I)

6

⇒ High scale: universal and anomaly-induced piece to gaugino masses

$$M_a(\Lambda_{\text{UV}}) = M_a^{\text{univ}}(\Lambda_{\text{UV}}) + M_a^{\text{anom}}(\Lambda_{\text{UV}}) = M_u + g_a^2(\Lambda_{\text{UV}}) \frac{b_a}{16\pi^2} M_g$$

- Gauge couplings continue to unify at the $\Lambda_{\text{UV}} = \Lambda_{\text{GUT}}$ scale

$$g_1^2(\Lambda_{\text{UV}}) = g_2^2(\Lambda_{\text{UV}}) = g_3^2(\Lambda_{\text{UV}}) = g_{\text{GUT}}^2 \simeq \frac{1}{2}$$

- Anomaly piece is proportional to SM beta-function coefficients

$$b_a = -(3C_a - \sum_i C_a^i) \Rightarrow \{b_1, b_2, b_3\} = \left\{ \frac{33}{5}, 1, -3 \right\}$$

- If these are going to be competitive you need $M_g \gtrsim 30 M_u$

⇒ Now evolve to electroweak scale using one-loop RGEs

$$M_a(\Lambda_{\text{EW}}) = M_u \left\{ 1 - g_a^2(\Lambda_{\text{EW}}) \frac{b_a}{8\pi^2} \ln \left(\frac{\Lambda_{\text{UV}}}{\Lambda_{\text{EW}}} \right) \left[1 - \frac{1}{2} \frac{M_g}{M_u \ln \left(\frac{\Lambda_{\text{UV}}}{\Lambda_{\text{EW}}} \right)} \right] \right\}$$

A Quick Derivation of the Mirage Pattern (II)

7

⇒ Introduce the parameter $\alpha = \frac{M_g}{M_u \ln(\Lambda_{UV}/\Lambda_{EW})}$

$$M_a(\Lambda_{EW}) = M_u \left[1 - \left(1 - \frac{\alpha}{2} \right) g_a^2(\Lambda_{EW}) \frac{b_a}{8\pi^2} \ln \left(\frac{\Lambda_{UV}}{\Lambda_{EW}} \right) \right]$$

• Some notable properties of this solution

- ★ If you can engineer $M_g \sim 30M_u$ then you obtain $\alpha \sim 1$
- ★ When $\alpha = 2$ gaugino masses universal *at the electroweak scale*
- ★ Take $\Lambda_{EW} = 1000 \text{ GeV}$, $\Lambda_{UV} = \Lambda_{GUT}$ and divide through by $M_1(\Lambda_{EW})|_{\alpha=0}$

$$M_1 : M_2 : M_3 = (1.0 + 0.66\alpha) : (1.93 + 0.19\alpha) : (5.87 - 1.76\alpha)$$

⇒ Finding the scale of “mirage unification”: redefine $\alpha \equiv \frac{M_g}{M_u \ln(M_{PL}/M_g)}$

$$M_a(\Lambda_{EW}) = M_u \left\{ 1 - g_a^2(\Lambda_{EW}) \frac{b_a}{8\pi^2} \left[\ln \left(\frac{\Lambda_{UV} (M_g/M_{PL})^{\alpha/2}}{\Lambda_{EW}} \right) \right] \right\}$$

• Effective unification scale is now at

$$\Lambda_{\text{mir}} = \Lambda_{GUT} \left(\frac{M_g}{M_{PL}} \right)^{\alpha/2}$$

• Model A

M.K. Gaillard and BDN, *Int. J. Mod. Phys. A* **22** (2007) 1451

- ★ Based on heterotic string theory
- ★ Dilaton stabilized with non-perturbative corrections to the Kähler potential
- ★ Stabilization mechanism causes $M_g \sim 30M_u$
- ★ Scalar masses generally universal
- ★ Absolute prediction: $\alpha \gtrsim 0.12$

• Model B

Choi, Falkowski, Nilles, Olechowski, *NPB* **718** (2005) 113

Falkowski, Lebedev, Mambrini, *JHEP* **0511** (2005) 034

- ★ Based on Type II string theory
- ★ Includes internal fluxes for moduli stabilization
- ★ Large warping in compact space produces $M_g \sim 30M_u$
- ★ AMSB plays a large role in all soft terms
- ★ Basic model predicts $\alpha \simeq 1$

Point	A	B
α	0.3	1.0
$\tan \beta$	10	10
Λ_{mir}	2.0×10^{14}	1.5×10^9
M_1	198.7	851.6
M_2	172.1	553.3
M_3	154.6	339.1
A_t	193.0	1309
A_b	205.3	1084
A_τ	188.4	1248
$m_{Q_3}^2$	$(1507)^2$	$(430.9)^2$
$m_{U_3}^2$	$(1504)^2$	$(610.3)^2$
$m_{D_3}^2$	$(1505)^2$	$(352.2)^2$
$m_{L_3}^2$	$(1503)^2$	$(381.6)^2$
$m_{E_3}^2$	$(1502)^2$	$(407.9)^2$
$m_{Q_{1,2}}^2$	$(1508)^2$	$(208.4)^2$
$m_{U_{1,2}}^2$	$(1506)^2$	$(302.7)^2$
$m_{D_{1,2}}^2$	$(1505)^2$	$(347.0)^2$
$m_{L_{1,2}}^2$	$(1503)^2$	$(379.8)^2$
$m_{E_{1,2}}^2$	$(1502)^2$	$(404.5)^2$
$m_{H_u}^2$	$(1500)^2$	$(752.0)^2$
$m_{H_d}^2$	$(1503)^2$	$(388.7)^2$

All values in GeV

Physical Spectra for Benchmark Models

9

Parameter	Point A	Point B		Parameter	Point A	Point B
$m_{\tilde{N}_1}$	85.5	338.7		$m_{\tilde{t}_1}$	844.7	379.9
$m_{\tilde{N}_2}$	147.9	440.2		$m_{\tilde{t}_2}$	1232	739.1
$m_{\tilde{N}_3}$	485.3	622.8		$m_{\tilde{c}_L}, m_{\tilde{u}_L}$	1518	811.7
$m_{\tilde{N}_4}$	494.0	634.3		$m_{\tilde{c}_R}, m_{\tilde{u}_R}$	1520	793.3
$m_{\tilde{C}_1^\pm}$	147.7	440.1		$m_{\tilde{b}_1}$	1224	676.8
$m_{\tilde{C}_2^\pm}$	494.9	635.0		$m_{\tilde{b}_2}$	1507	782.4
$m_{\tilde{g}}$	510.0	818.0		$m_{\tilde{s}_L}, m_{\tilde{d}_L}$	1520	815.4
μ	476.1	625.2		$m_{\tilde{s}_R}, m_{\tilde{d}_R}$	1520	793.5
m_h	115.2	119.5		$m_{\tilde{\tau}_1}$	1487	500.4
m_A	1557	807.4		$m_{\tilde{\tau}_2}$	1495	540.4
m_{H^0}	1557	806.8		$m_{\tilde{\mu}_L}, m_{\tilde{e}_L}$	1500	545.1
m_{H^\pm}	1559	811.1		$m_{\tilde{\mu}_R}, m_{\tilde{e}_R}$	1501	514.6

Low Energy Physical Masses for Benchmark Points

- ⇒ Our goal is to ask how well we can determine α at the LHC using only **actual observations**
- Most importantly, can we demonstrate $\alpha \neq 0$?
 - Want to do this independent of any particular model
 - Not going to assume reconstruction any sparticle masses
 - We *will* assume we know all other inputs for the Monte Carlo comparison to data – unrealistic but this is a first step
- ⇒ Basic idea: use an ensemble of signatures wisely chosen to perform a fit of Monte Carlo to “data”
- We break the problem into a “base model” specified by the parameters

$$\left\{ \begin{array}{c} \tan \beta, m_{H_u}^2, m_{H_d}^2 \\ M_3, A_t, A_b, A_\tau \\ m_{Q_{1,2}}, m_{U_{1,2}}, m_{D_{1,2}}, m_{L_{1,2}}, m_{E_{1,2}} \\ m_{Q_3}, m_{U_3}, m_{D_3}, m_{L_3}, m_{E_3} \end{array} \right\}$$

and a value of α which determines the three gaugino masses
(with overall scale set by M_3)

- Given a model we construct a ***model line*** by varying α while keeping the base model fixed
- For each point we generate data using PYTHIA + PGS4 and construct our signatures
- Analysis is performed using a modification of ROOT generated by Baris Altunkaynak at Northeastern

<http://www.atsweb.neu.edu/ialtunkaynak/heptools.html#parvicursor>

- How do we determine the value of α ? We compare Monte Carlo predictions for our signatures against the “data”
- For example, we can ask whether we can distinguish the prediction for the case $\alpha = 0$ from the data we simulate at $\alpha \neq 0$

⇒ We want to distinguish models A and B using the n (counting) signatures S_i

- Define a measure in *signature space* analogous to a chi-squared variable

$$(\Delta S_{AB})^2 = \frac{1}{n} \sum_i \left[\frac{S_i^A - S_i^B}{\delta S_i^{AB}} \right]^2$$

- Convert to effective cross-sections via $\bar{\sigma}_i = S_i/L$ and assuming errors are purely statistical (\sqrt{N})

$$(\Delta S_{AB})^2 = \frac{1}{n} \sum_i \left[\frac{\bar{\sigma}_i^A - \bar{\sigma}_i^B}{\sqrt{\bar{\sigma}_i^A/L_A + \bar{\sigma}_i^B/L_B}} \right]^2$$

- We always include the Standard Model background so that $\bar{\sigma}_i = \bar{\sigma}_i^{\text{SUSY}} + \bar{\sigma}^{\text{SM}}$

⇒ We want to distinguish models A and B using the n (counting) signatures S_i

- Define a measure in *signature space* analogous to a chi-squared variable

$$(\Delta S_{AB})^2 = \frac{1}{n} \sum_i \left[\frac{S_i^A - S_i^B}{\delta S_i^{AB}} \right]^2$$

- Convert to effective cross-sections via $\bar{\sigma}_i = S_i/L$ and assuming errors are purely statistical (\sqrt{N})

$$(\Delta S_{AB})^2 = \frac{1}{n} \sum_i \left[\frac{\bar{\sigma}_i^A - \bar{\sigma}_i^B}{\sqrt{\bar{\sigma}_i^A/L_A + \bar{\sigma}_i^B/L_B}} \right]^2$$

- We always include the Standard Model background so that $\bar{\sigma}_i = \bar{\sigma}_i^{\text{SUSY}} + \bar{\sigma}^{\text{SM}}$
- So how big should $(\Delta S_{AB})^2$ be to say models A and B are distinguished from one another?
- LHC Inverse criterion: this number needs to be at *least* as big as the value induced by quantum fluctuations

- Effect of fluctuations estimated by comparing the same single model to itself many times and computing $(\Delta S_{AA})^2|_{95}$
- But this really depends on the model point and (especially) the signature list you choose to consider

- Effect of fluctuations estimated by comparing the same single model to itself many times and computing $(\Delta S_{AA})^2|_{95}$
- But this really depends on the model point and (especially) the signature list you choose to consider
- We can obtain an analytic answer valid for any model pair and any signature list provided
 - ★ Fluctuations for each signature are assumed to be **uncorrelated**
 - ★ We assume that our extracted $\bar{\sigma}_i$ are very close to the true cross-section values σ_i
 - ★ We assume the count rates form normal distributions
- Under these assumptions $(\Delta S_{AB})^2$ is a randomly-distributed variable with a probability distribution of

$$P(\Delta S^2) = n \chi_{n,\lambda}^2(n\Delta S^2)$$

where $\chi_{n,\lambda}^2$ is the **non-central chi-squared distribution** for n degrees of freedom

⇒ $(\Delta S_{AB})^2$ distributed according to a non-central chi-square distribution

- The non-centrality parameter λ is given by

$$\lambda = \sum_i \frac{(\sigma_i^A - \sigma_i^B)^2}{\sigma_i^A / L_A + \sigma_i^B / L_B}$$

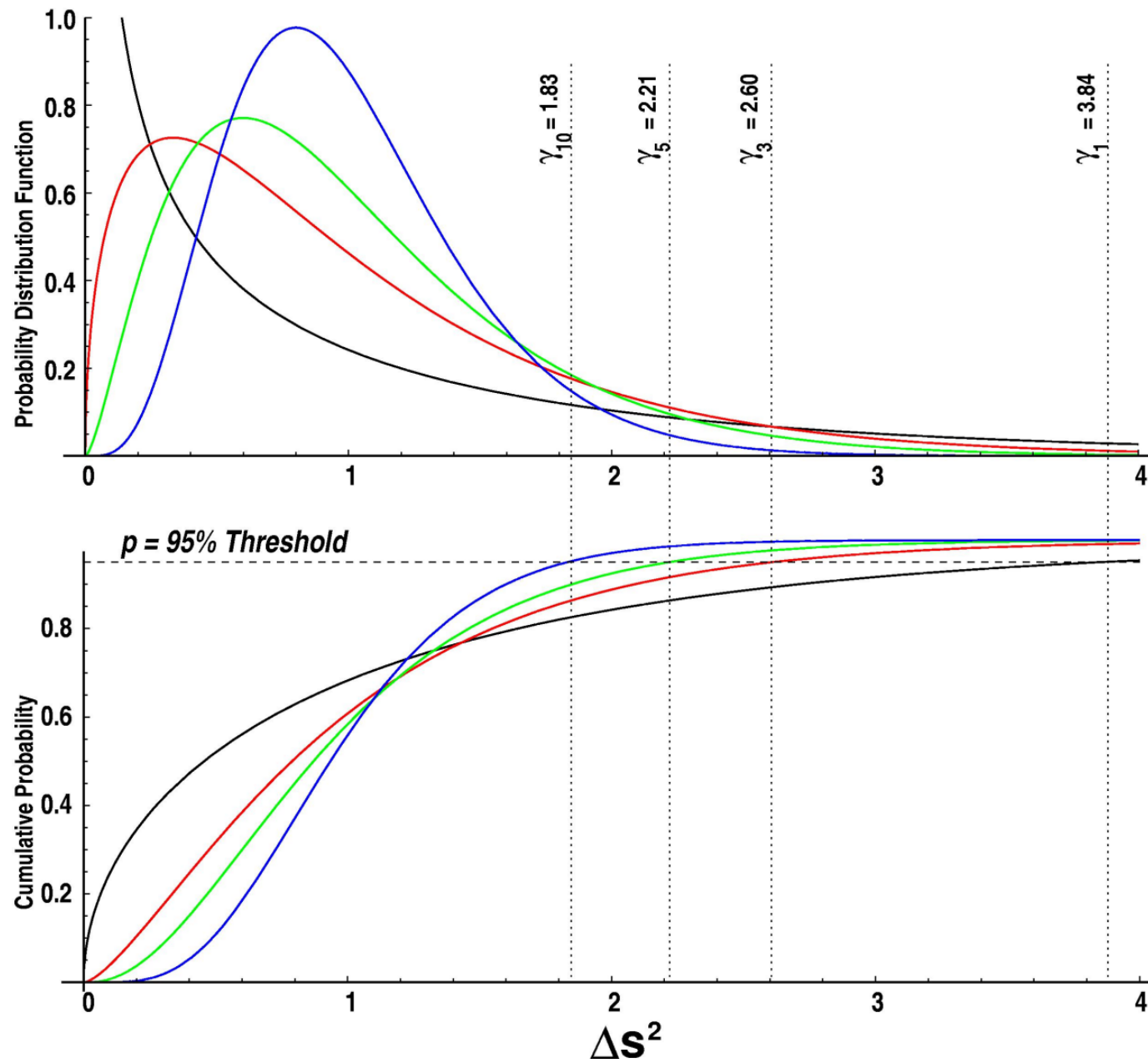
- Taking $\lambda = 0$ gives distribution for $(\Delta S_{AA})^2$
- Can now solve analytically for $(\Delta S_{AA})^2|_p \equiv \gamma_n(p)$ for any confidence level p as a function of the number of signatures n
- Having $(\Delta S_{AB})^2 > (\Delta S_{AA})^2|_{95}$ may be thought of as a *necessary* condition, but it is not *sufficient* to distinguish models A and B

⇒ For two models that truly are different we expect $\lambda \neq 0$

⇒ We want to quantify the probability that two truly distinct models undergo a fluctuation such that their measured $(\Delta S_{AB})^2$ is a very low value

Non-central Chi-Square Distribution & $\gamma_n(p)$

15

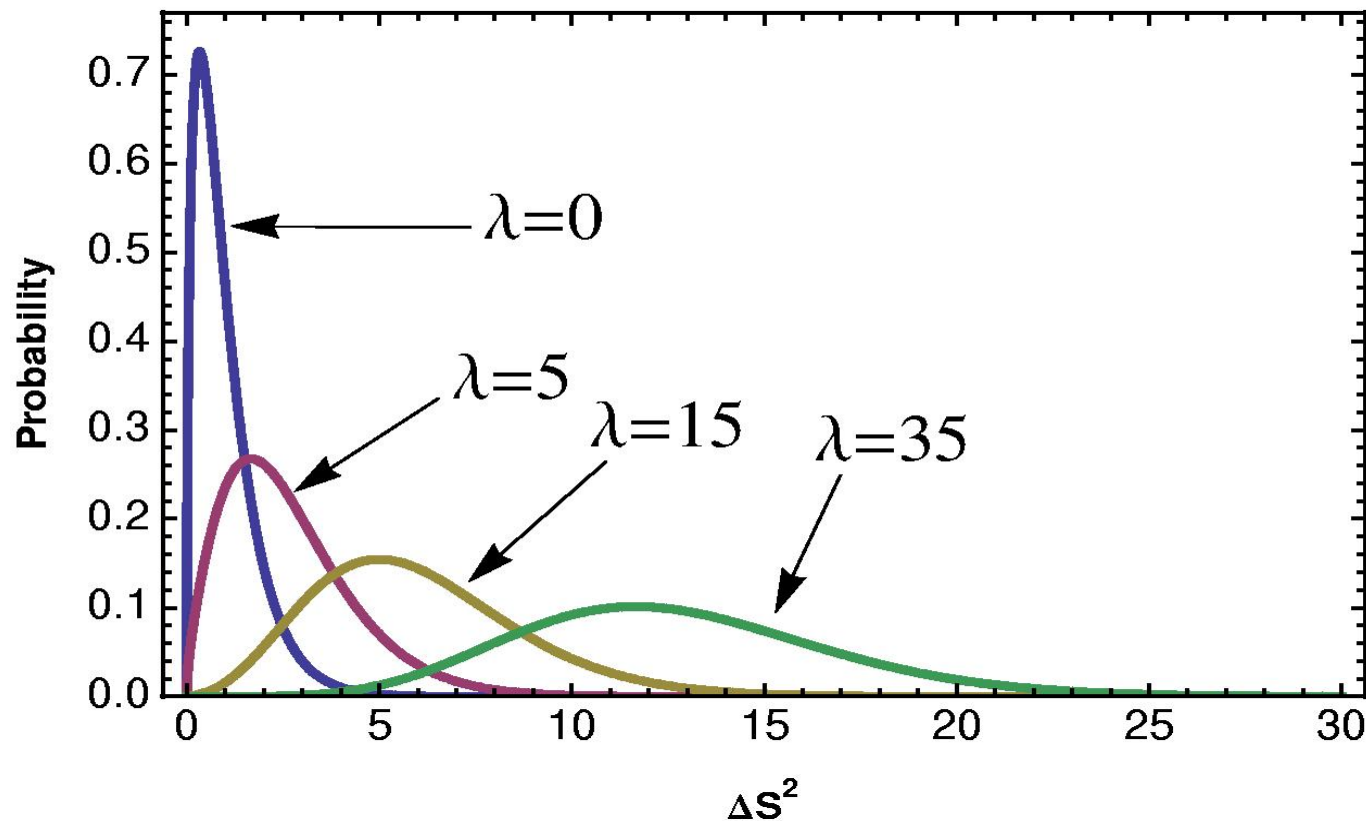


Our Distinguishability Criterion

⇒ Want the probability for $(\Delta S_{AB})^2$ to fluctuate below $\gamma_n(p)$ to be less than 5%

$$P = \int_{\gamma_n(p)}^{\infty} n \chi_{n,\lambda}^2(n\Delta S_{AB}^2) d(\Delta S_{AB}^2) = \int_{n\gamma_n(p)}^{\infty} \chi_{n,\lambda}^2(y) dy \geq 0.95$$

- Value of this integral decreases monotonically as λ increases
- When $P = 0.95$ we have found the minimum value $\lambda_{\min}(n)$ for the non-centrality parameter



- Any combination of n -parameters yielding $\lambda > \lambda_{\min}(n)$ will be effective in demonstrating that the two models are indeed distinct, 95% of the time, with a confidence level of 95%
- The value of λ is proportional to integrated luminosity

$$L_{\min} = \frac{\lambda_{\min}(n)}{R_{AB}} \quad \text{with} \quad R_{AB} = \sum_i (R_{AB})_i = \sum_i \frac{(\sigma_i^A - \sigma_i^B)^2}{\sigma_i^A + \sigma_i^B}$$

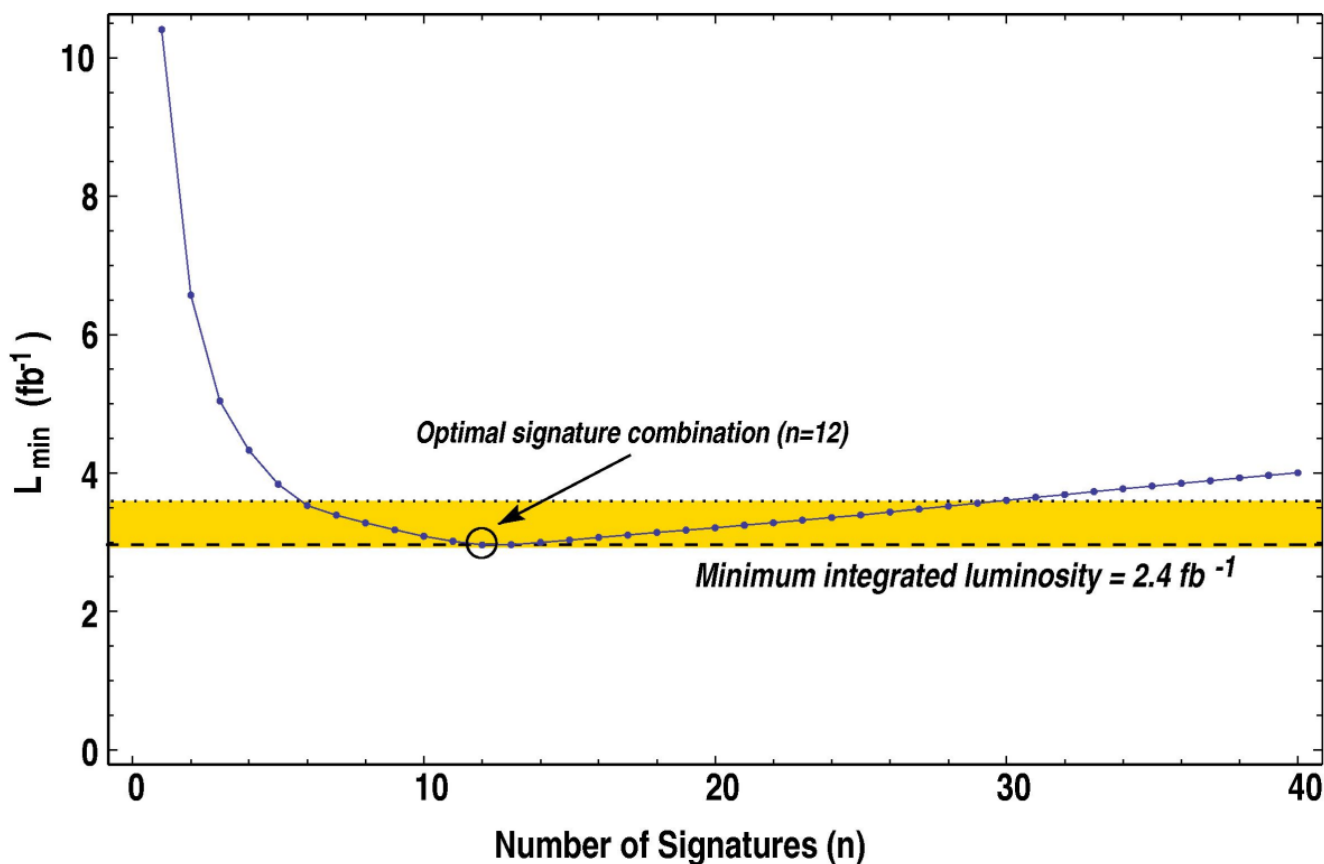
⇒ All the physics of the specific signature list is contained in R_{AB} !

- This just says given *any* signature list there is always some minimal luminosity that will distinguish the models
- Now the goal is clear: choose your signature list so as to maximize R_{AB} , with as few signatures as possible so as to minimize $\lambda_{\min}(n)$

Choosing an Optimal Signature List

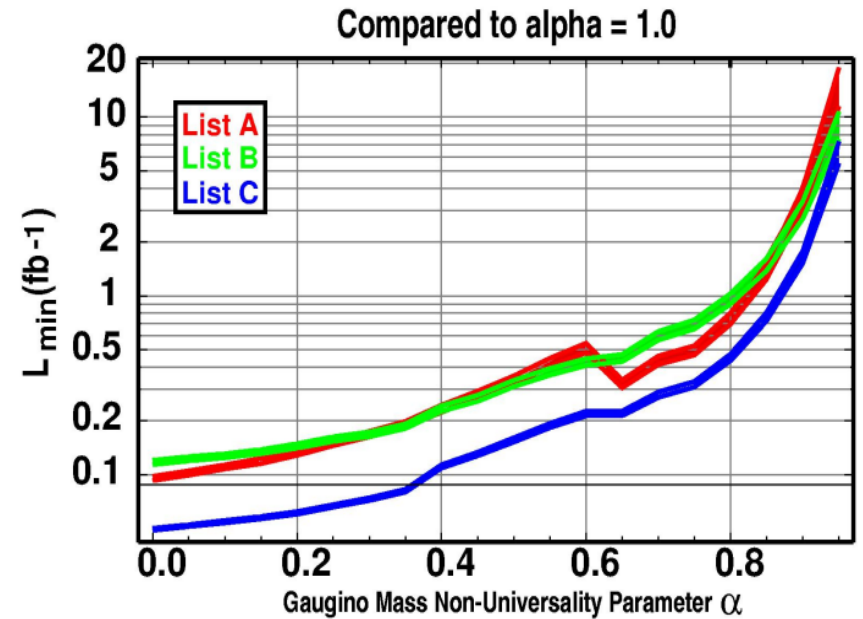
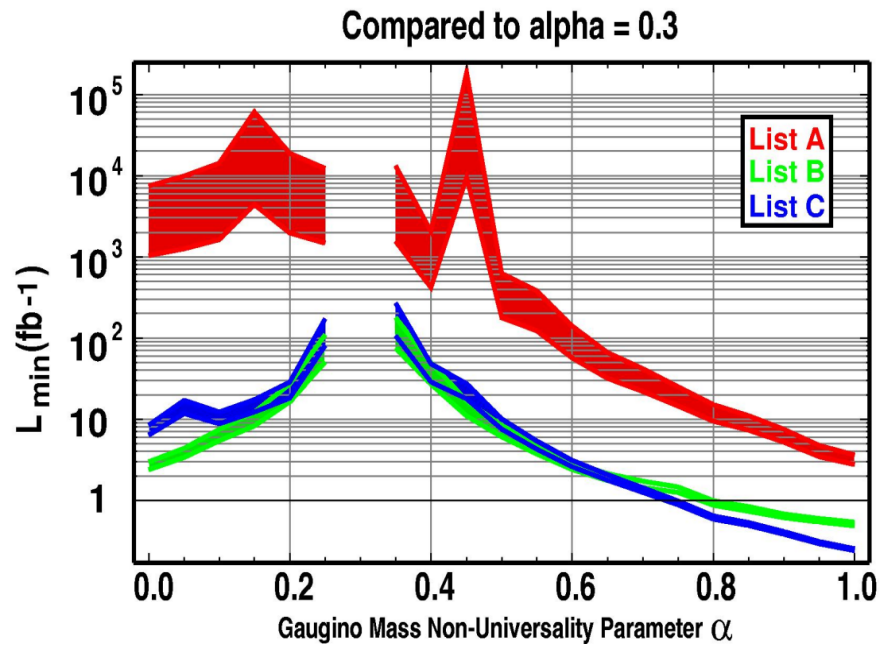
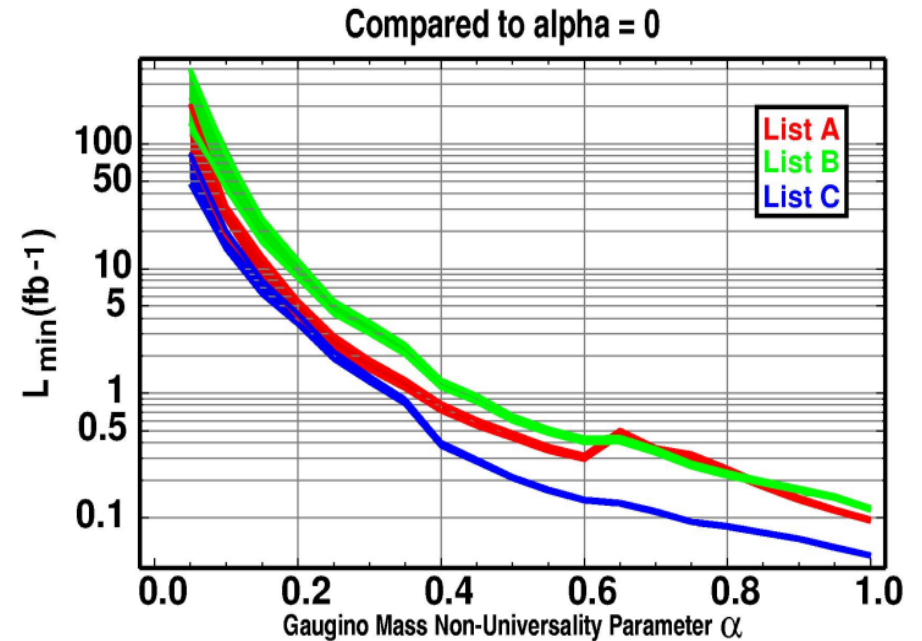
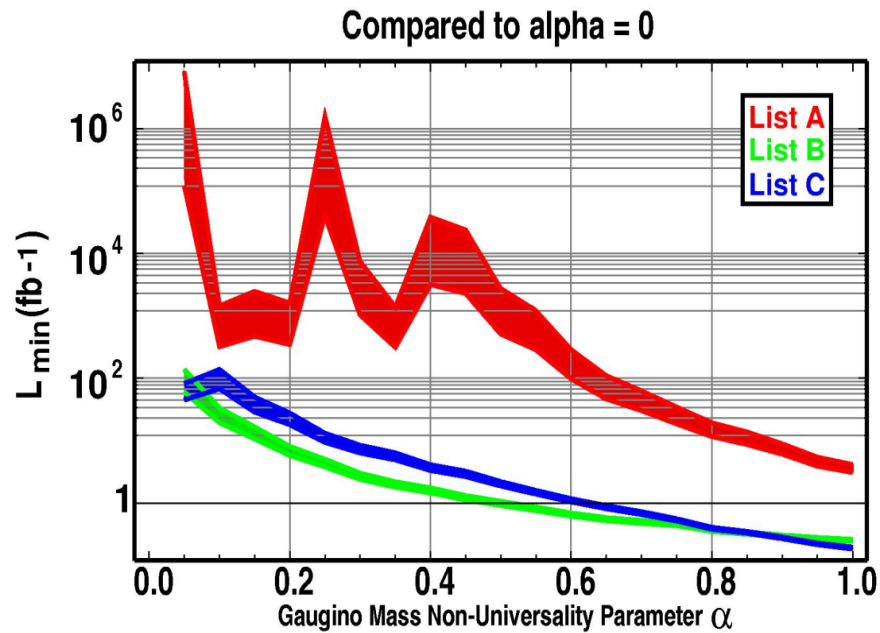
18

- Given a model pair A and B compute the absolute quantity $(R_{AB})_i$ for all of the possible signatures you can imagine
- Now order them from highest R_i value (smallest L_{\min}) to smallest R_i value (largest L_{\min}) – what fraction of the list should you employ?



- Given a model pair A and B compute the absolute quantity $(R_{AB})_i$ for all of the possible signatures you can imagine
- Now order them from highest R_i value (smallest L_{\min}) to smallest R_i value (largest L_{\min}) – what fraction of the list should you employ?
- No cheating! Can't use your best signature N times... (correlations)
- Kitchen sink method is not ideal!
⇒ Take a big hit since $\lambda(n)$ eventually grows faster than $\sum_i R_i$
- For any *particular pair* of models you can optimize this choice
- But once you average over a large ensemble of models the list will now only be (at best) **close to optimal** for any model

Benchmark Results



Model A

Model B

- LHC v2.0 will be about **synthesis** of multiple measurements
⇒ *Theorists can and will play a major role here*
- Bigger is not necessarily better when using LHC observations!
⇒ *Experimentalists know this well – theorists sometimes less so*
- Rather than fit to **models** can we fit to **characteristics**?
⇒ *Yes, at least in this (artificial) first step*