Pseudo-Dirac See-Saw Neutrinos

Northwestern University Wei-Chih Huang 05.12.2009 Pheno 2009, Madison, May 11-13

Motivations

- What are the current experimental constraints on right-handed neutrino masses?
- What are the consequences of right-handed neutrinos masses much smaller than Dirac neutrino masses?
- What are the consequences of the first Borexino experiment measurement of the ⁷Be neutrino flux?

What is the Pseudo-Dirac Neutrino?

 In normal see-saw mechanism, the mass matrix for neutrino is:

 $\left(\begin{array}{cc} 0 & yv \\ (yv)^T & M \end{array} \right)$

Where y is the Yukawa coupling, v is the expectation value of the Higgs field and M is Majarona mass which is much larger than y*v to account for tiny neutrino mass.

 In the Pseudo-Dirac case, M is much smaller than yv; therefore, the mass spectrum are nearly degenerate.

• General n active and n sterile neutrinos case

$$M_{\nu} = \left(\begin{array}{cc} 0 & m \\ m^T & \epsilon_m \end{array}\right)$$

$$M_{\nu} \simeq \begin{pmatrix} 1 & -\delta^* \\ \delta^T & 1 \end{pmatrix} \frac{1}{\sqrt{2}} \begin{pmatrix} U^* & -U^* \\ 1 & 1 \end{pmatrix} \begin{pmatrix} m^D(1+\epsilon^D) & 0 \\ 0 & -m^D(1-\epsilon^D) \end{pmatrix} \frac{1}{\sqrt{2}} \begin{pmatrix} U^{\dagger} & 1 \\ -U^{\dagger} & 1 \end{pmatrix} \begin{pmatrix} 1 & \delta \\ -\delta^{\dagger} & 1 \end{pmatrix}$$

$$m = U^* m^D$$
 $\delta = U\left(\frac{\epsilon^D}{2} + \epsilon\right),$

$$\epsilon_m = 2\epsilon^D m^D + \varepsilon^T m^D + m^D \varepsilon,$$

and $m^D \varepsilon^T = -\varepsilon m^D.$

$$\begin{aligned} \mathbf{The } 2\mathbf{+1 } \mathbf{case} \\ M_{\nu} &= \begin{pmatrix} 0 & 0 & m \sin \theta \\ 0 & 0 & m \cos \theta \\ m \sin \theta & m \cos \theta & \epsilon_m \end{pmatrix}, \\ V^T M_{\nu} V &= \begin{pmatrix} 0 & 0 & 0 \\ 0 & m \left(1 + \frac{\epsilon}{2}\right) & 0 \\ 0 & 0 & -m \left(1 - \frac{\epsilon}{2}\right) \end{pmatrix}, \\ \mathbf{\epsilon}_m &= m * \epsilon \end{aligned}$$
$$V &= \begin{pmatrix} V_{e1} & V_{e2} & V_{e2'} \\ V_{a1} & V_{a2} & V_{a2'} \\ V_{s1} & V_{s2} & V_{s2'} \end{pmatrix} = \begin{pmatrix} \cos \theta & \frac{\sin \theta}{\sqrt{2}} \left(1 - \frac{\epsilon}{4}\right) & \frac{\sin \theta}{\sqrt{2}} \left(1 + \frac{\epsilon}{4}\right) \\ -\sin \theta & \frac{\cos \theta}{\sqrt{2}} \left(1 - \frac{\epsilon}{4}\right) & \frac{\cos \theta}{\sqrt{2}} \left(1 + \frac{\epsilon}{4}\right) \\ 0 & -\frac{1}{\sqrt{2}} \left(1 + \frac{\epsilon}{4}\right) & \frac{1}{\sqrt{2}} \left(1 - \frac{\epsilon}{4}\right) \end{cases} \end{aligned}$$
$$1 - P_{ee} &= \sin^2 2\theta \sin^2 \left(\frac{m^2 L}{4E}\right) + \mathcal{O}(\epsilon^2). \end{aligned}$$
$$P_{\mu s} &= \cos^2 \theta \sin^2 \left(\frac{2m^2 \epsilon L}{4E}\right), \\ P_{\mu s} &= \cos^2 \theta \sin^2 \left(\frac{2m^2 \epsilon L}{4E}\right). \end{aligned}$$

Neutrino Propagation

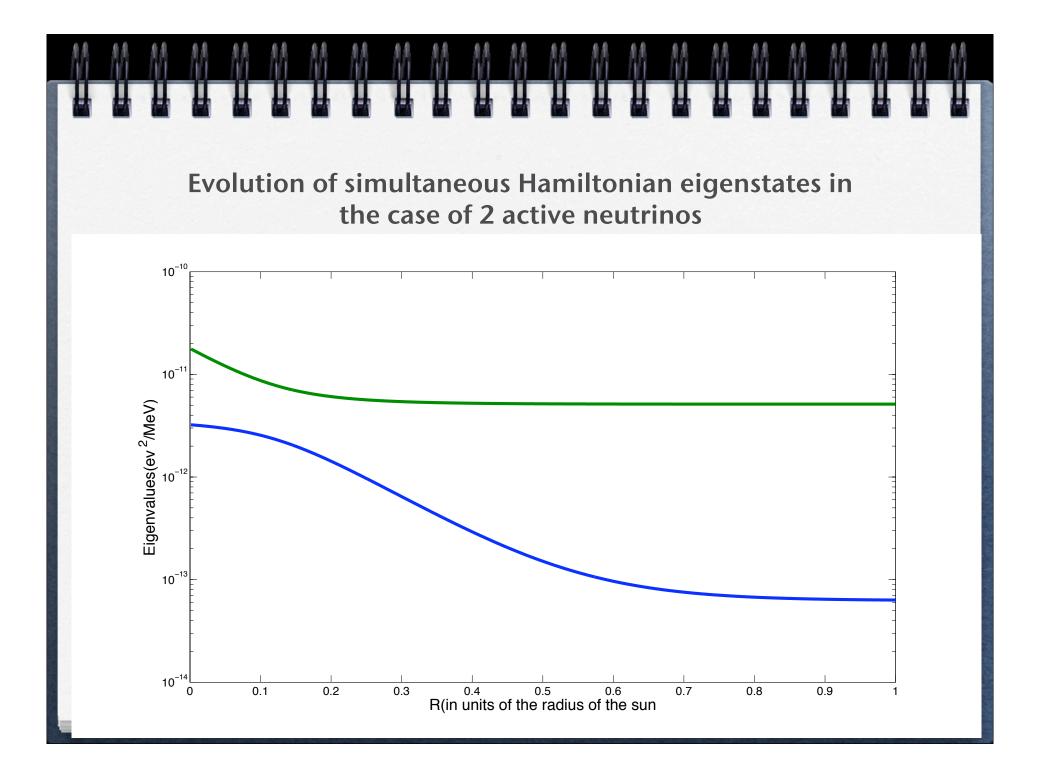
• In vacuum(for simplicity, just two active neutrinos)

$$\left| \begin{array}{c} \imath \frac{d}{dL} \left(\begin{array}{c} |\nu_e\rangle \\ |\nu_\mu\rangle \end{array} \right) = \frac{\Delta m^2}{2E} \left(\begin{array}{c} \sin^2\theta & \cos\theta\sin\theta \\ \cos\theta & \cos^2\theta \end{array} \right) \left(\begin{array}{c} |\nu_e\rangle \\ |\nu_\mu\rangle \end{array} \right)$$

where $\Delta m^2 = m_2^2 - m_1^2$, θ is the mixing angle and E is the neutrino energy

$$i\frac{d}{dL}\left(\begin{array}{c}|\nu_e\rangle\\|\nu_\mu\rangle\end{array}\right) = \left(\begin{array}{cc}A & \frac{\Delta}{2}\sin 2\theta\\\frac{\Delta}{2}\sin 2\theta & \Delta\cos 2\theta\end{array}\right)\left(\begin{array}{c}|\nu_e\rangle\\|\nu_\mu\rangle\end{array}\right)$$

where $\Delta = \frac{\Delta m^2}{2E}$ and A is the matter potential

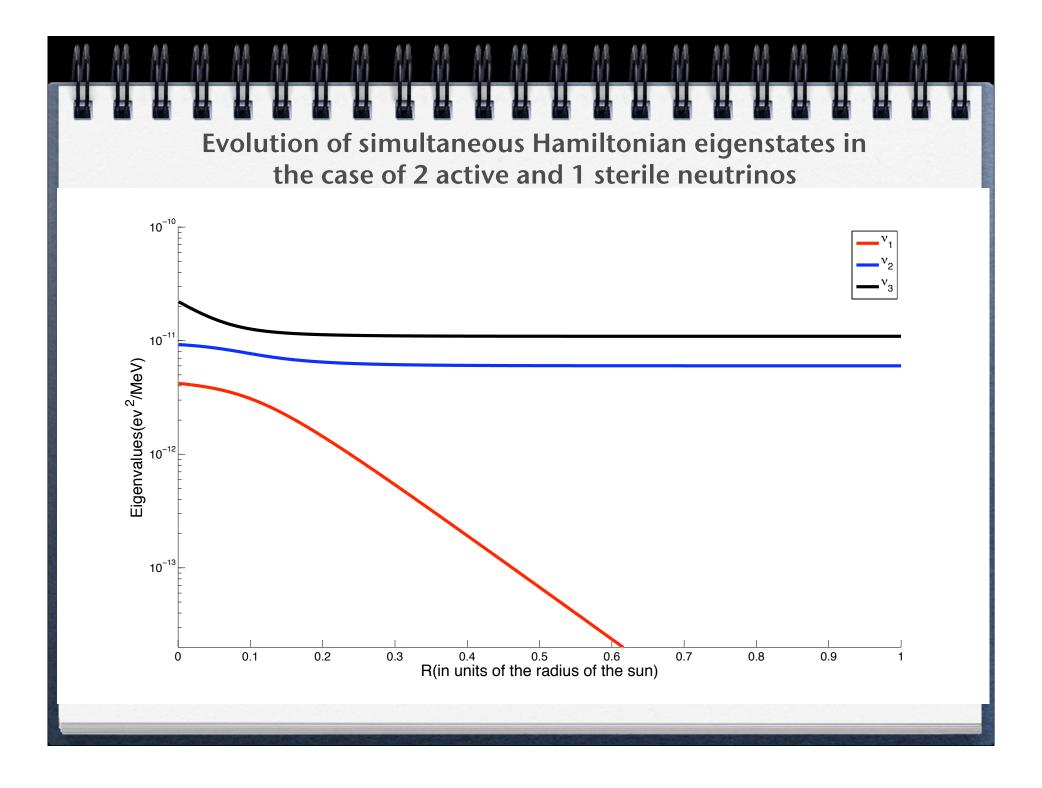


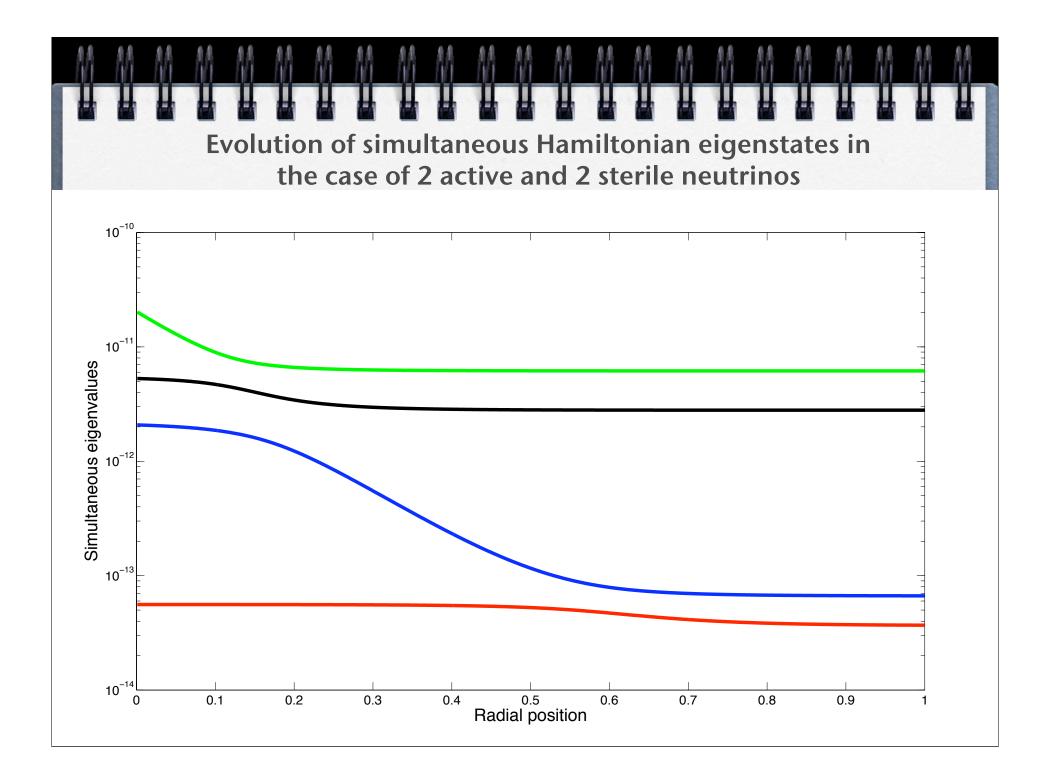
Transition Probability in Active Neutrinos only

For an exponential electron number density $N_e = N_{e0}e^{-L/r_0}$, P_c , also referred to as the crossing probability, is exactly calculable²⁹

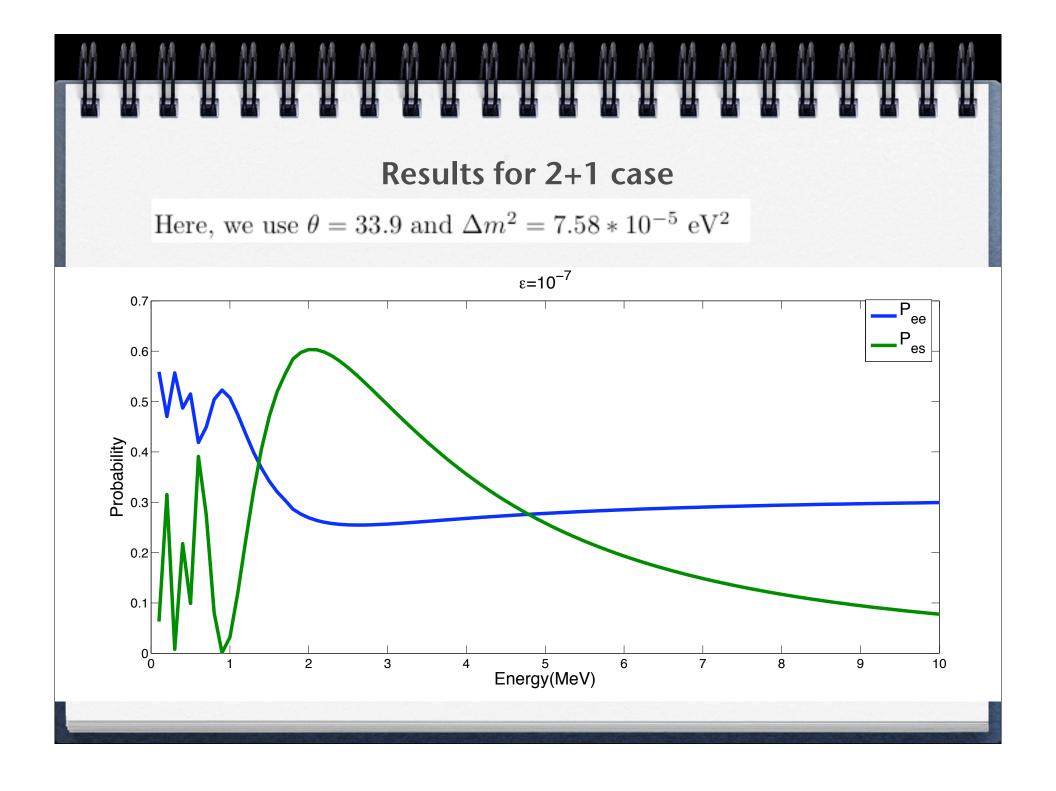
$$P_c = \frac{e^{-\gamma \sin^2 \theta} - e^{-\gamma}}{1 - e^{-\gamma}}, \quad \gamma = 2\pi r_0 \Delta.$$
(53)

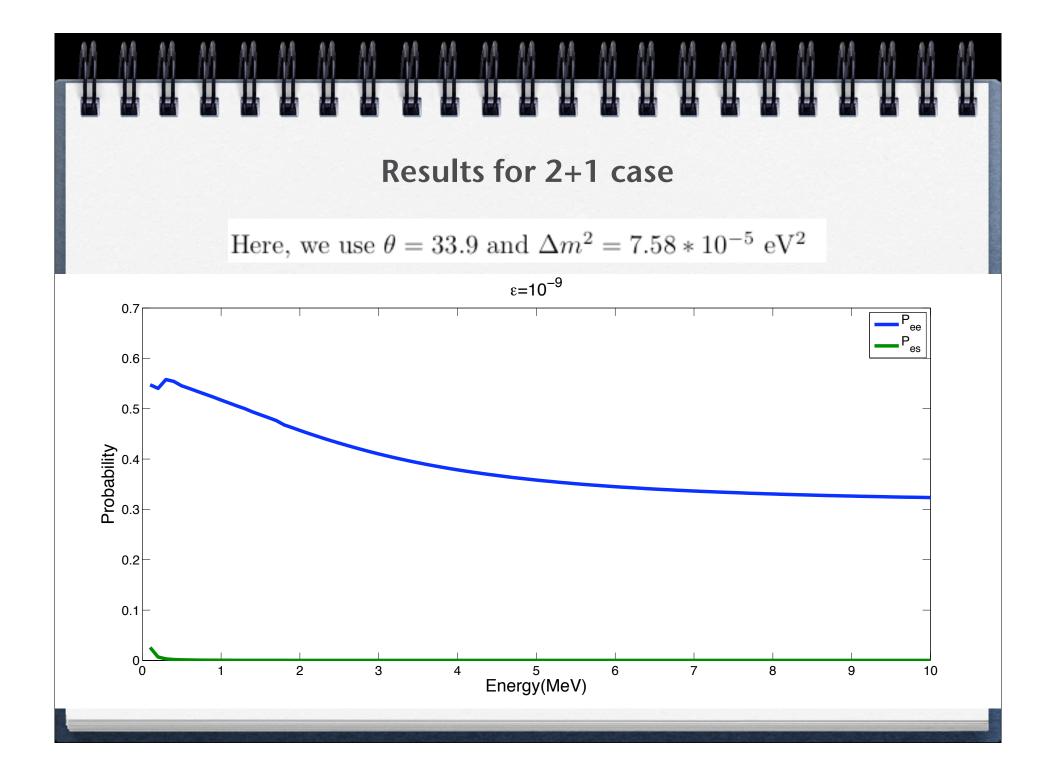
(from A. de Gouvêa, hep-ph/0411274)





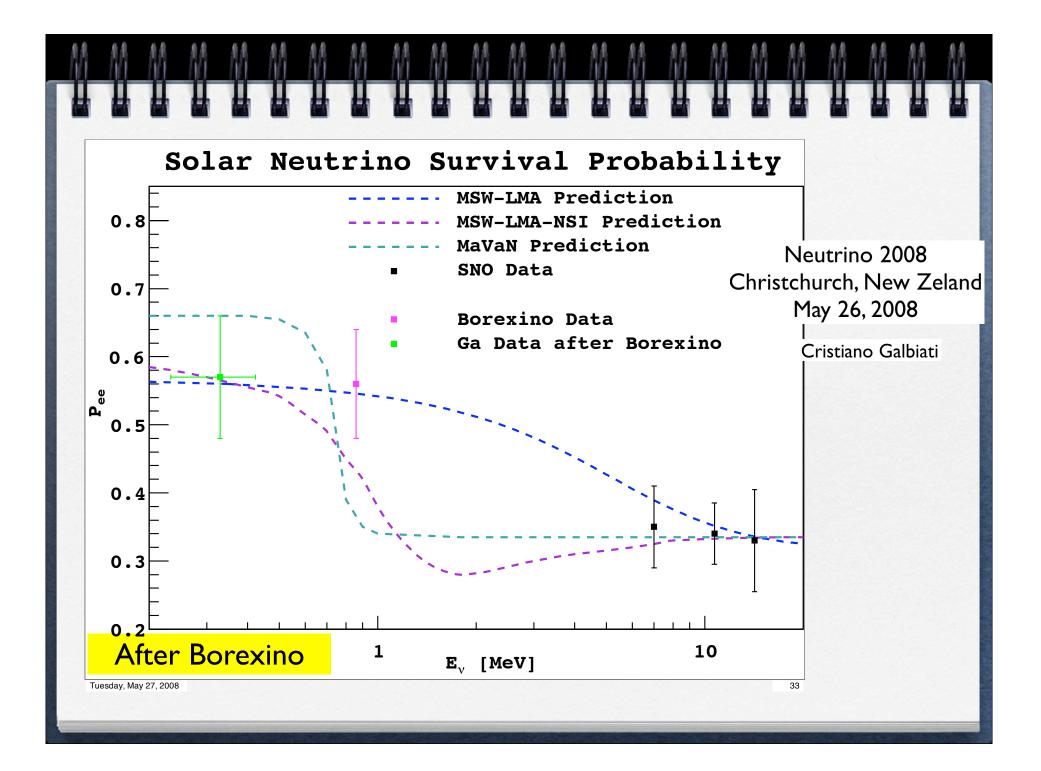
$$\begin{split} \mathbf{Transition \ Probability \ in \ 2+1 \ case} \\ H \longrightarrow V^T HV &= \begin{pmatrix} A\cos^2\theta & A\cos\theta\sin\theta & 0\\ A\cos\theta\sin\theta & A\sin^2\theta + \frac{m^2}{2E} & \frac{m\epsilon}{2E} \\ 0 & \frac{m\epsilon}{2E} & \frac{m^2}{2E} \end{pmatrix}. \\ P_c &= \frac{e^{-\gamma|V_{s2}|^2} - e^{-\gamma}}{1 - e^{-\gamma}}, \quad \gamma \simeq 9.8 \left(\frac{\epsilon}{10^{-4}}\right) \left(\frac{m^2}{8 \times 10^{-5} \ \text{eV}^2}\right) \left(\frac{0.862 \ \text{MeV}}{E}\right), \\ P_{ee} &= \left| \left(1 \ 0 \ 0 \right) V \left(\begin{array}{c} \exp(-i\phi_1') & 0 & 0\\ 0 & \exp(-i\phi_2') & 0\\ 0 & 0 & \exp(-i\phi_3') \end{array} \right) \left(\begin{array}{c} 1 & 0 & 0\\ 0 & \sqrt{1 - P_c} & \sqrt{P_c} \\ 0 & -\sqrt{P_c} & \sqrt{1 - P_c} \end{array} \right) \\ & \left(\begin{array}{c} \exp(-i\phi_1) & 0 & 0\\ 0 & \exp(-i\phi_3) & 0 \\ 0 & \exp(-i\phi_3) \end{array} \right) V_{\text{mat}}^{\dagger} \left(\begin{array}{c} 1\\ 0\\ 0 \end{array} \right) \right|^2, \end{split}$$



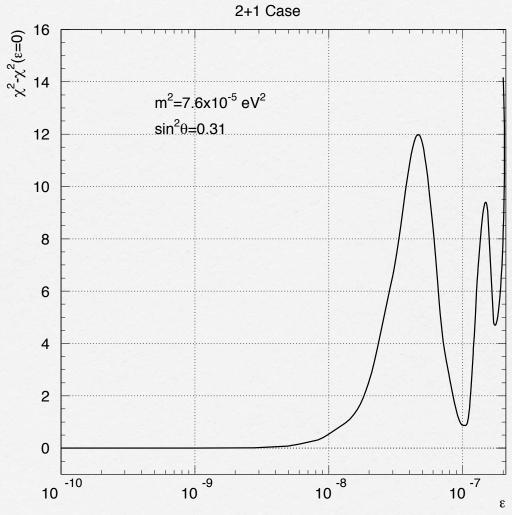


Experimental data

- Super-K elastic-scattering
- SNO phase-III charged-current and neutralcurrent
- Borexino
- GALLEX+GNO
- Homestake



Chi-Squared Fit



Conclusion

- We analytically diagonalized the neutrino mass matrix in the limit of the small Majorana mass term.
- By using a special basis, the resonance region could be identified and the transition probability could be computed analytically.
- Constraint on right-handed Majorana masses (from solar data): M<10⁻⁹ eV(for 2+1 case)