

Beryllium-7 Electron Capture in Superconducting Tunnel Junctions: The *BeEST* Sterile Neutrino Search



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for the BeEST Collaboration



LLNL-PRES-836503

This work was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract No. DE-AC52- 07NA27344.



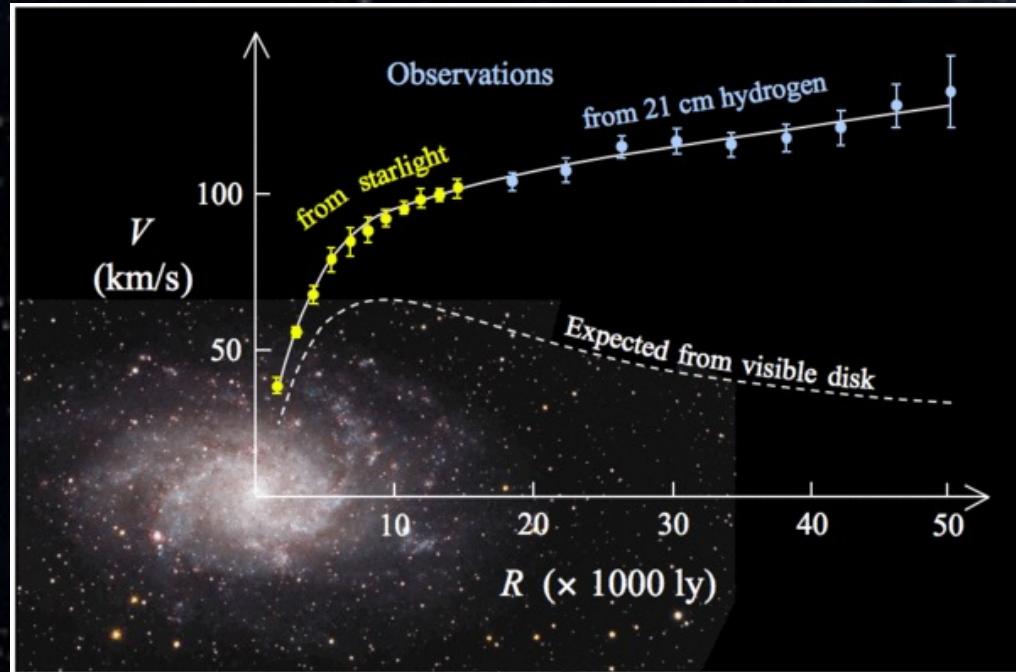
A Very Brief History Outline of Radiation Detection

- 1) Dark Matter and Sterile Neutrinos
- 2) Superconducting Tunnel Junction (STJ) Radiation Detectors
- 3) The BeEST Sterile Neutrino Search
- 4) Material Effects in Superconducting (STJ) Sensors
- 5) Next: Increased Sensitivity with Arrays and New Materials

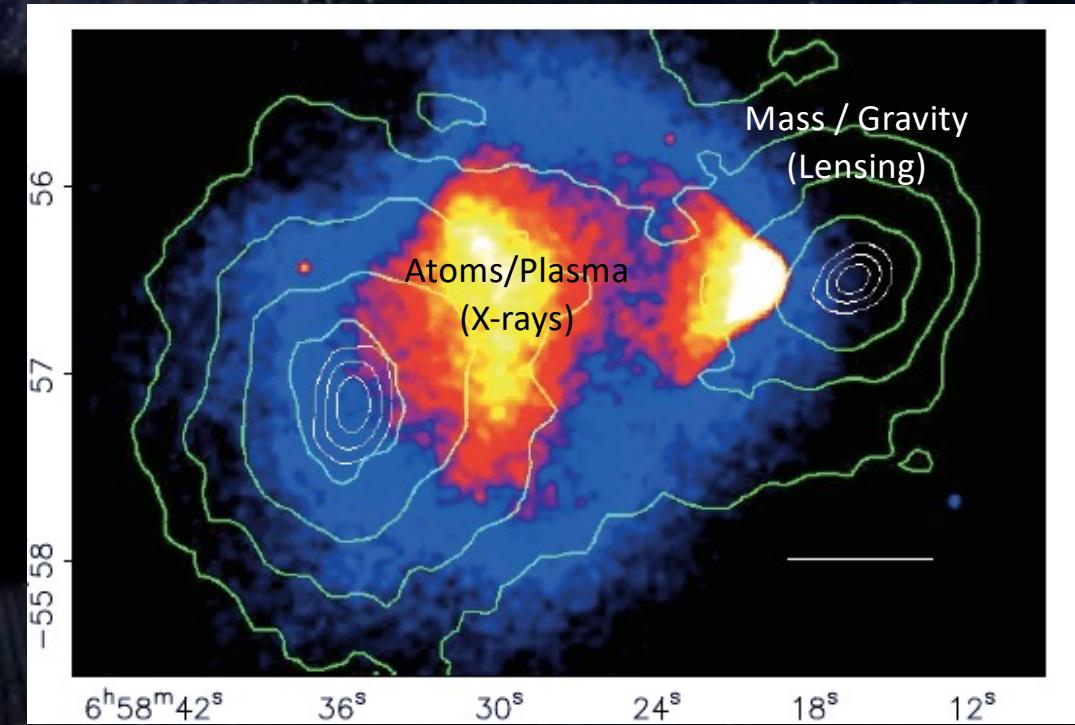
Tabletop BSM Physics.

A Very Brief History of Radiation Detection

Galaxy Rotation Curves



Merging “Bullet Cluster”



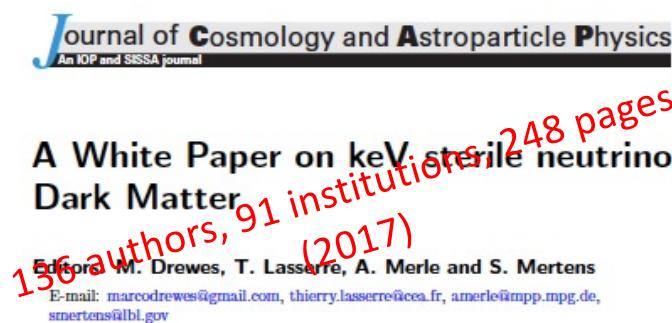
The composition of $\sim 85\%$ of the mass in the universe is unknown.



A Very Brief History of Sterile Neutrinos in Detection

Questions in Cosmology:

- What is dark matter?
- Why is there more matter than antimatter in the universe?



Questions in Particle Physics:

- What is nature of neutrinos mass?
- Do right-handed neutrinos exist?

		Fermions				Bosons	
		Quarks		Leptons			
		2.4 MeV $\frac{2}{3}$ u up Left Right	1.27 GeV $\frac{2}{3}$ c charm Left Right	171.2 GeV $\frac{2}{3}$ t top Left Right	0 0 g gluon Left Right		
		4.8 MeV $-\frac{1}{3}$ d down Left Right	104 MeV $-\frac{1}{3}$ s strange Left Right	4.2 GeV $-\frac{1}{3}$ b bottom Left Right	0 0 γ photon Left Right		
		0 eV electron neutrino Left Right	0 eV muon neutrino Left Right	0 eV tau neutrino Left Right	91.2 GeV 0 Z weak force Bosons (Forces) spin 1 Left Right		
		0.511 MeV -1 e electron Left Right	105.7 MeV -1 μ muon Left Right	1.777 GeV -1 τ tau Left Right	80.4 GeV ± 1 W weak force Bosons (Forces) spin 1 Left Right		
							>114 GeV 0 H Higgs boson spin 0

Sterile neutrinos are an appealing dark matter candidate.

But they are well-motivated without dark matter, too.

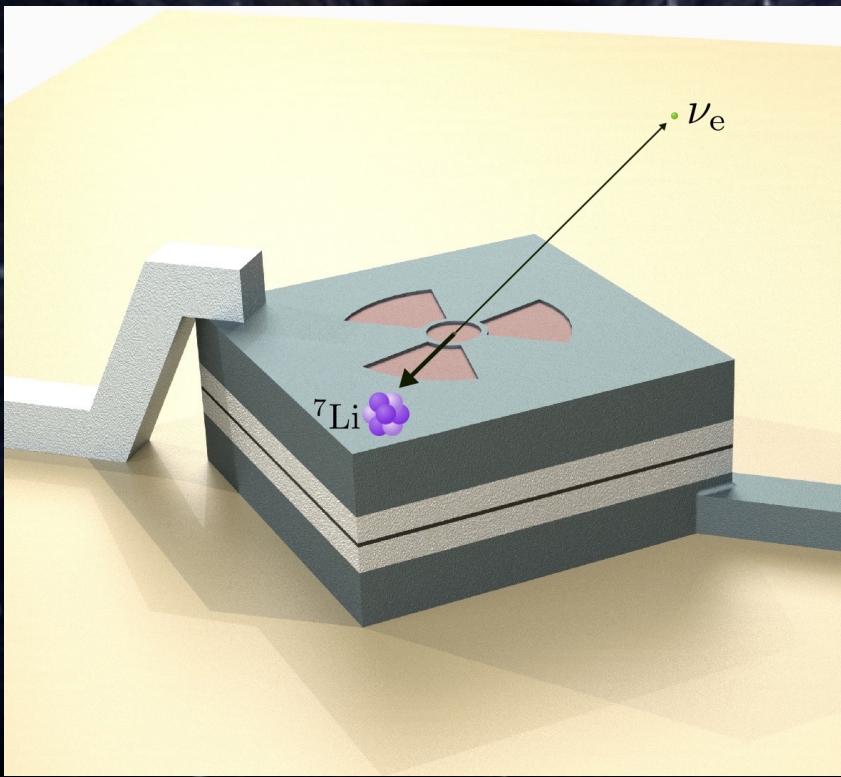
- 1) (Active) neutrinos exist
- 2) (Sterile) neutrinos could explain several big questions

A Very Brief History of Radiation Detection

- ## How to Find Something that
- 1) Doesn't emit light
 - 2) Doesn't absorb light
 - 3) Doesn't interact (except through its mass)?

A Very Brief History of Radiation Detection

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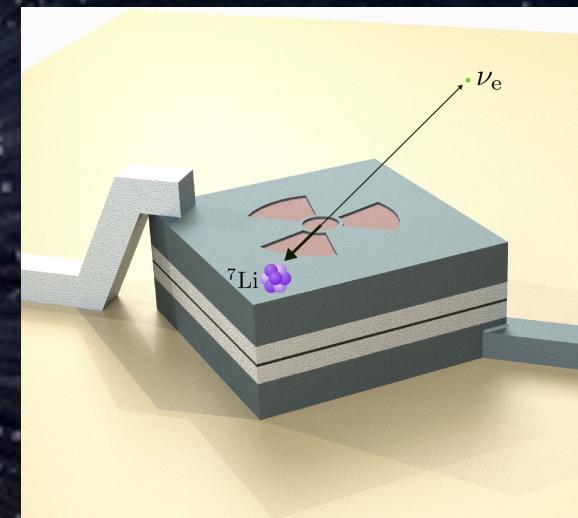
Use $|p_{\text{neutrino}}| = |p_{\text{Li-7}}|$ to search for missing momentum in nuclear decay of ${}^7\text{Be}$.

A Very Brief History of Why ${}^7\text{Be}$? Radiation Detection

Why ${}^7\text{Be}$?

- 1) 2-body Electron Capture decay ${}^7\text{Be} + e^- \Rightarrow {}^7\text{Li} + \nu_e$
⇒ Monochromatic recoil
- 2) Long half life of 53 days
⇒ Easy handling
- 3) Low Z = 7
⇒ Manageable calculations
- 4) Large Q = 861 keV
⇒ “Large” recoil signal E = 56 eV⁵⁶

Superconducting Tunnel Junction Detector



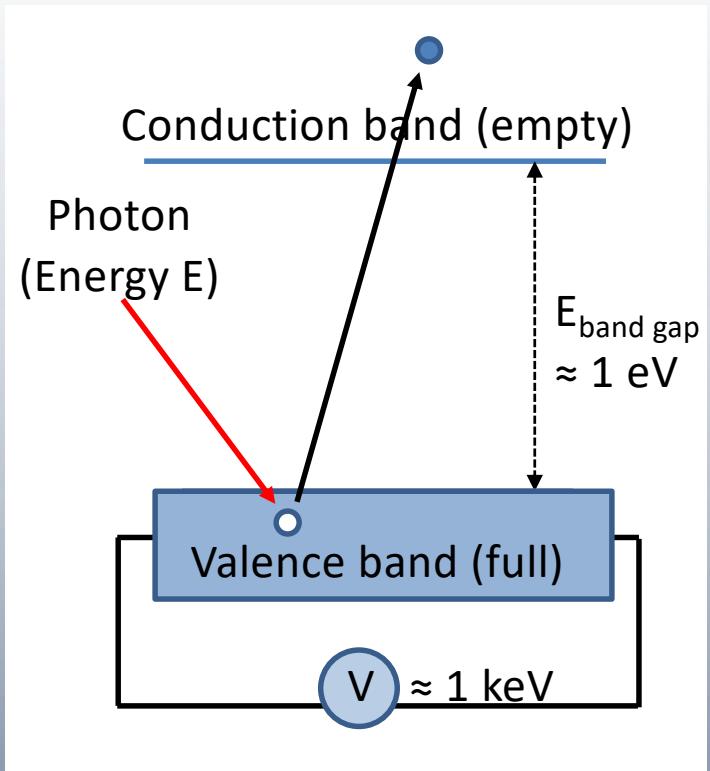
Why Not ${}^7\text{Be}$?

- 1) 10% EC decay into excited state ${}^7\text{Li}^*$
⇒ Doppler-broadened secondary peaks

- 1) High energy resolution
- 2) High speed, >1000 counts/s
- 3) Sensitive to phonons

Use $|p_{\text{neutrino}}| = |p_{\text{Li-7}}|$ to search for missing momentum in nuclear decay of ${}^7\text{Be}$.

2) Semiconducting vs Superconducting Detectors

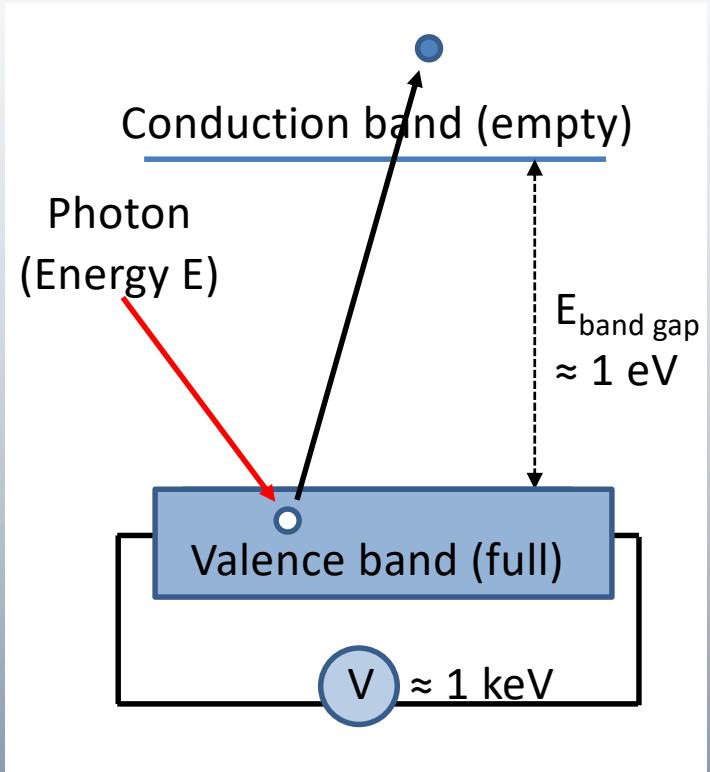


$$\text{Signal} = eN \approx e \frac{E}{3E_{\text{bandgap}}}$$

$$\text{Noise} = (FN)^{1/2} \text{ with } F \approx 0.1$$

$$\text{Signal/Noise} \propto 1/\sqrt{E_{\text{bandgap}}}$$

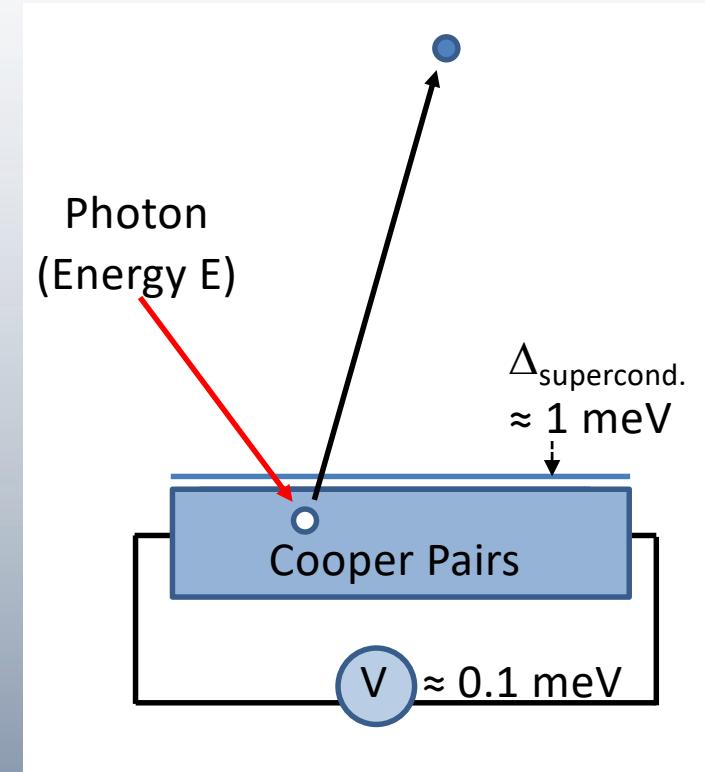
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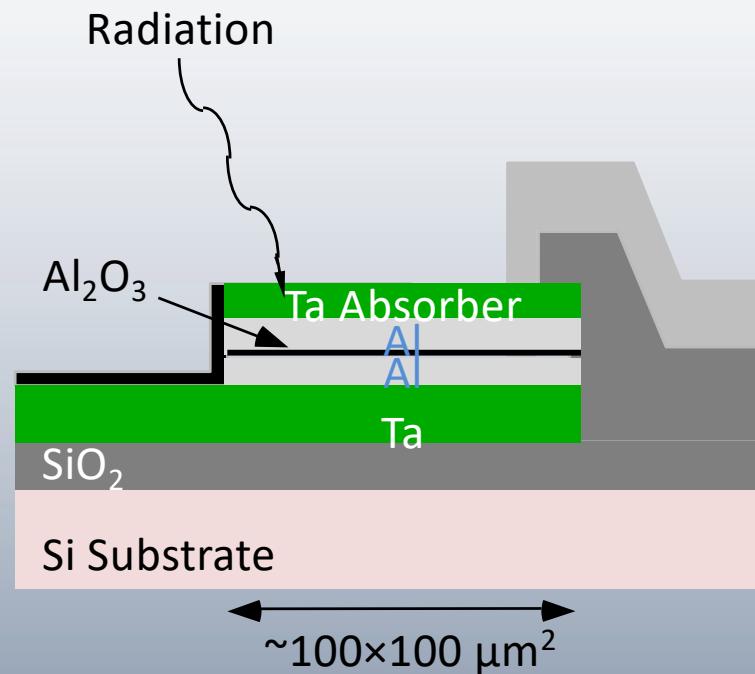
$$\text{Signal} = eN \approx e \frac{E}{1.7\Delta_{\text{supercond.}}} \rightarrow \sim 1000x \text{ larger !}$$

$$\text{Noise} = (FN)^{1/2} \text{ with } F \approx 0.2 \rightarrow \sim 30x \text{ larger}$$

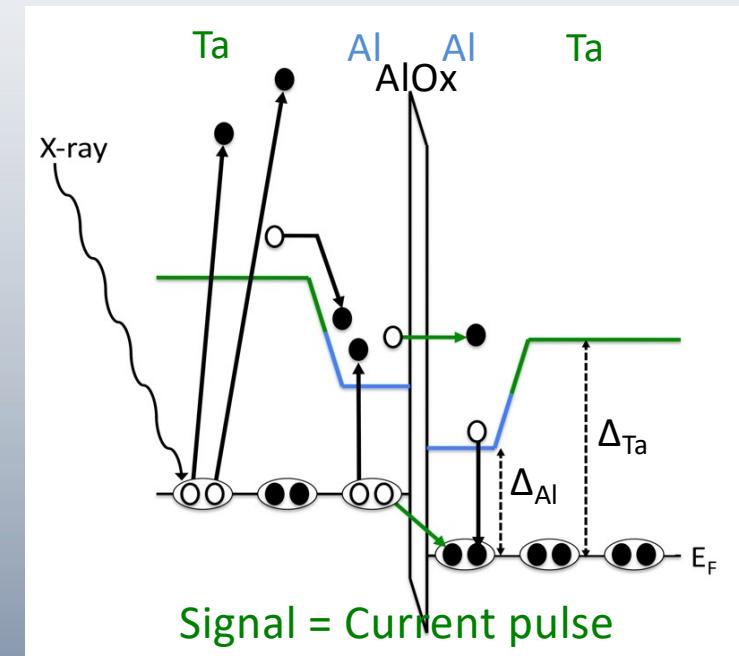
~30x higher energy resolution, no quenching.

Superconducting Tunnel Junction (STJ) Radiation Detectors

STJ Cross Section



STJ Operating Principle

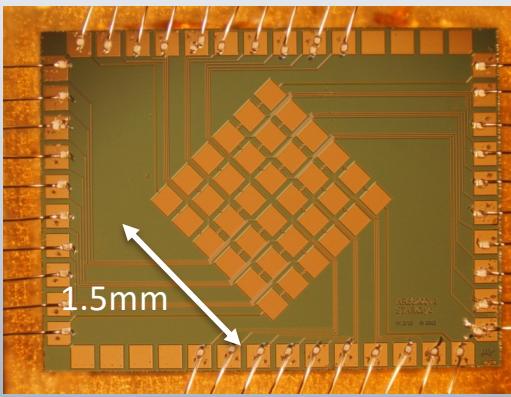


$$\text{Energy resolution } E_{\text{FWHM}} = 2.355 \sqrt{1.7\Delta \cdot E_{\text{X-ray}} \left(F + 1 + \frac{1}{n} \right)}$$

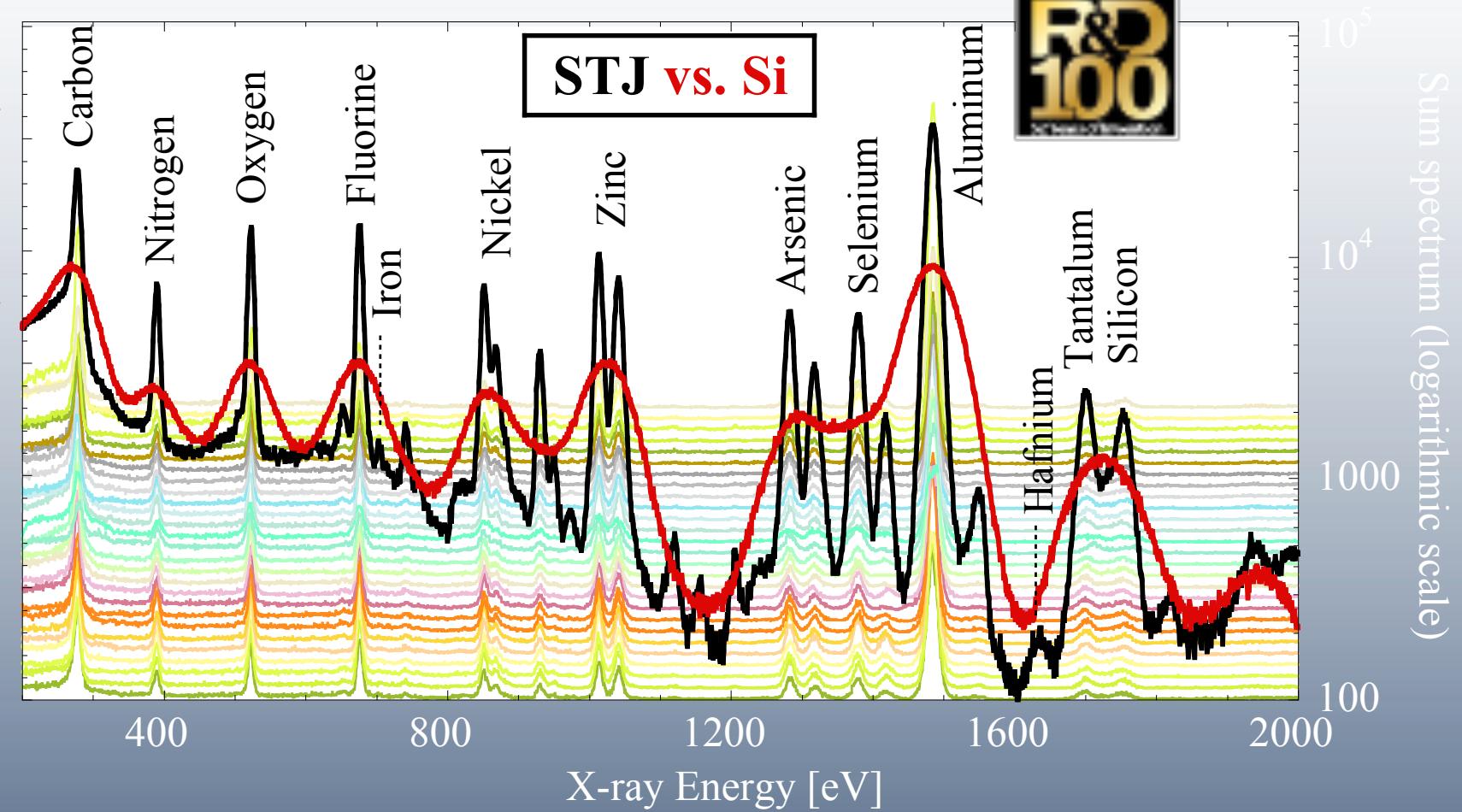
Small superconducting energy gap ($\Delta \approx 1\text{meV}$) \Rightarrow High resolution (<10 eV FWHM)
Short excess charge life-time ($\sim \mu\text{s}$) \Rightarrow High count rate (>1,000 counts/s/pixel)

STJ Soft X-ray Detectors

36-pixel STJ Array

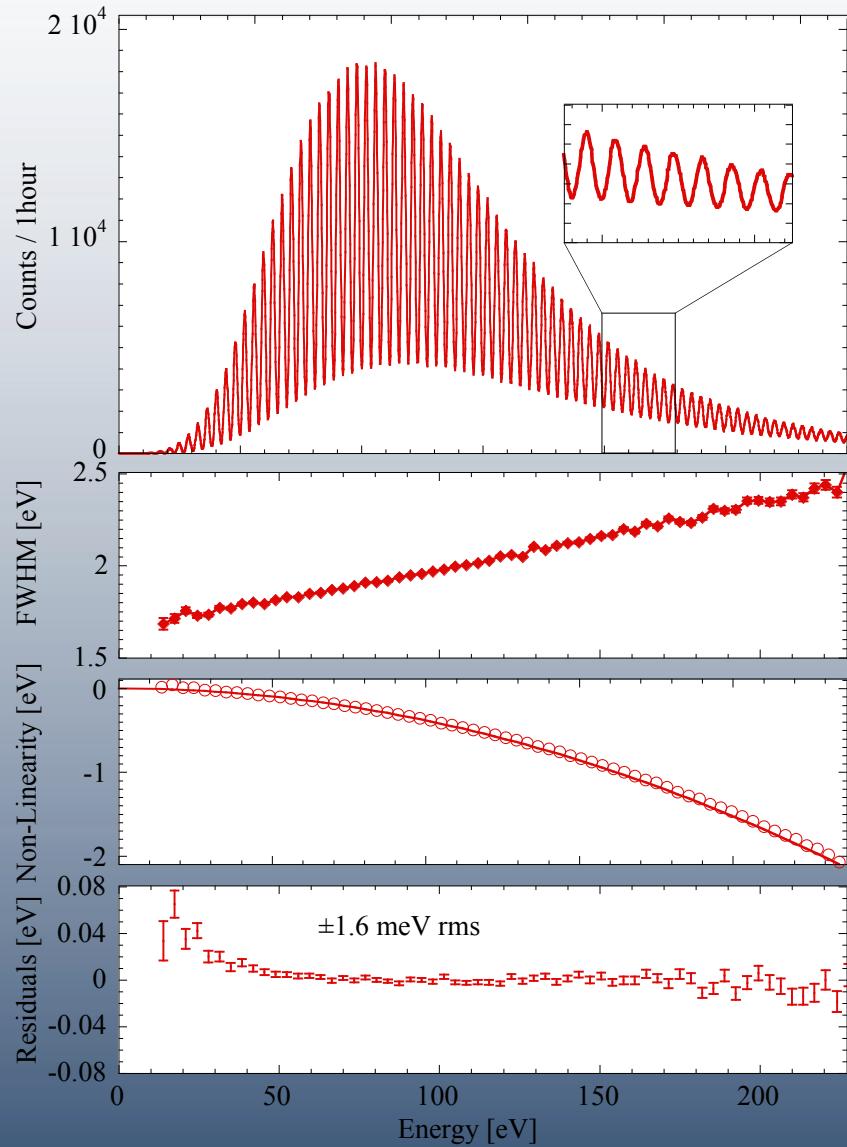


Individual sensors (linear scale)



STJs were initially developed for high-resolution spectroscopy with soft X-rays where lines are closely spaced.

STJ Characterization with Pulsed UV Laser



Pulsed 355 nm (3.5eV) laser at 5,000 Hz

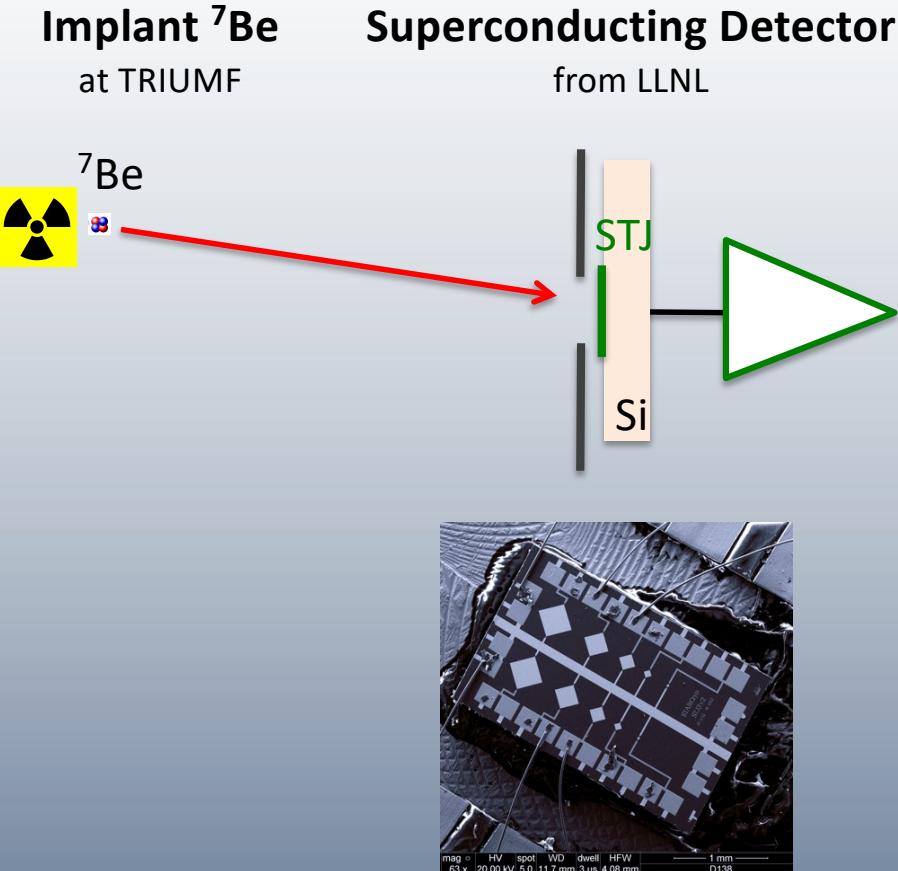
⇒ Comb of peaks at integer multiples of 3.5 eV

⇒ Energy resolution between ~1.5 and ~2.5 eV FWHM

⇒ Only quadratic non-linearity

⇒ Calibration accuracy of order ± 1 meV in 1 hour

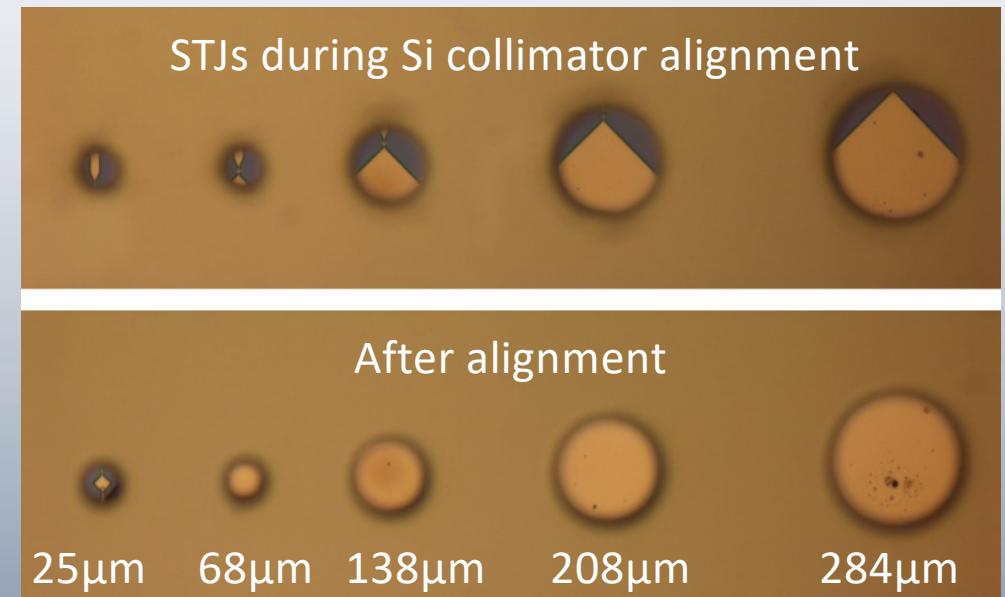
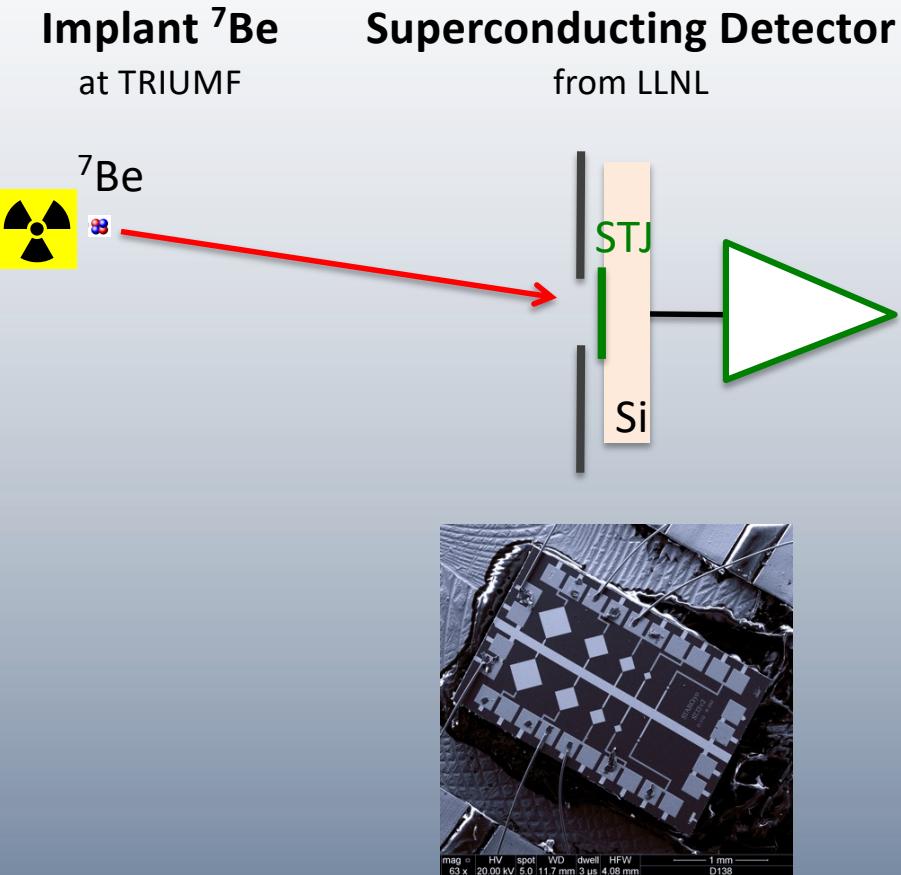
3) A Sterile Neutrino Search with Superconducting Tunnel Junctions



Implant ${}^7\text{Be}$ into STJ detectors.

$$\tau_{1/2} = 53 \text{ days}, Q = 861 \text{ keV}$$

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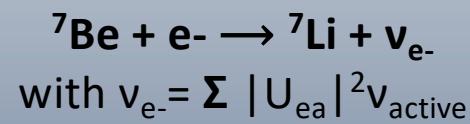
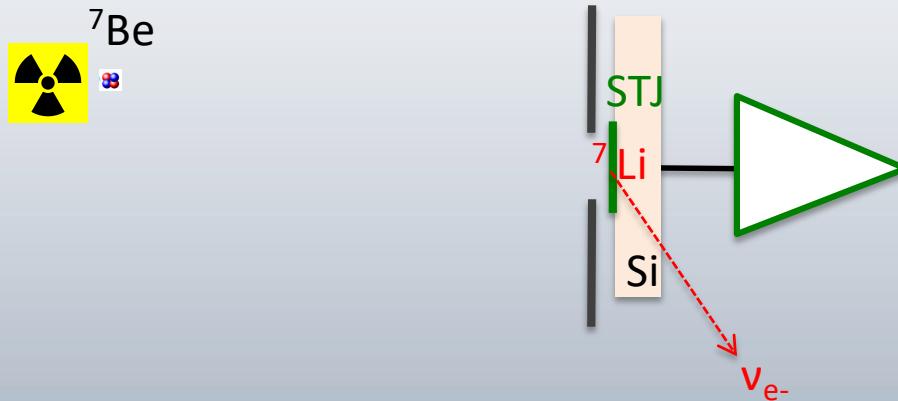
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Beryllium-7 Electron Capture in Superconducting Tunnel Junctions

Implant ${}^7\text{Be}$
at TRIUMF

Detect ${}^7\text{Be}$ Decay
at LLNL



Measure electron capture decay of ${}^7\text{Be}$ to ${}^7\text{Li}$.

2-body decay \Rightarrow Monochromatic recoil (in principle)

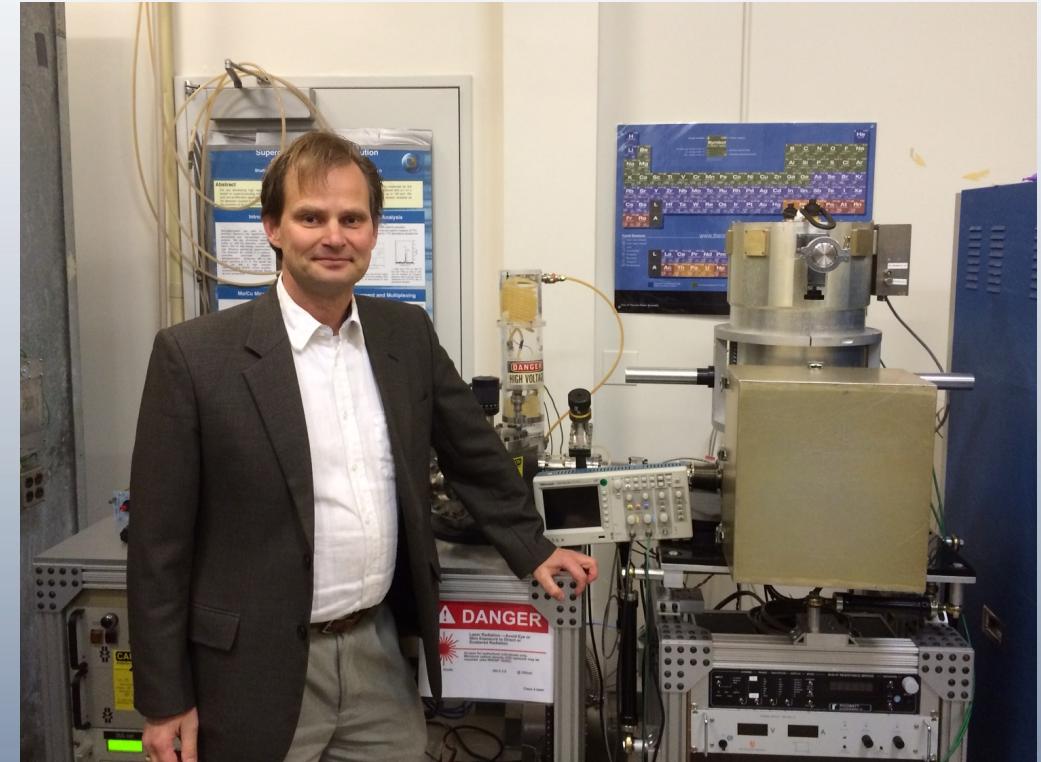
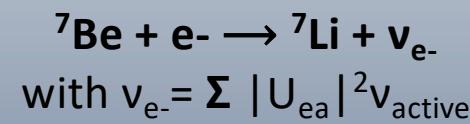
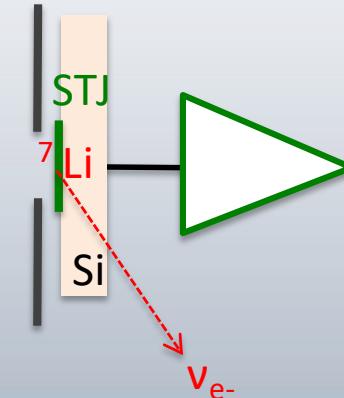


Beryllium-7 Electron Capture in Superconducting Tunnel Junctions

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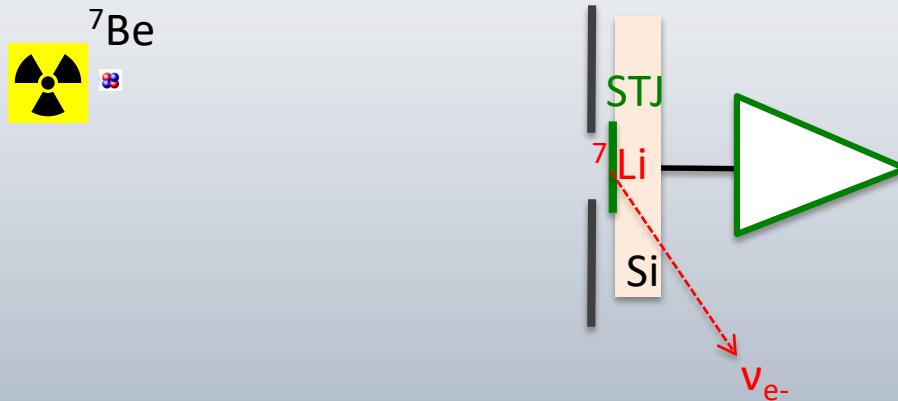
2-body decay \Rightarrow Monochromatic recoil (in principle)



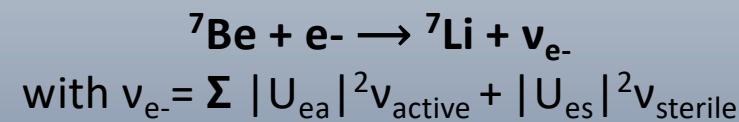
Be-7 Electron Capture in STJs: The BeEST Experiment

Implant ^7Be
at TRIUMF

Detect ^7Be Decay
at LLNL



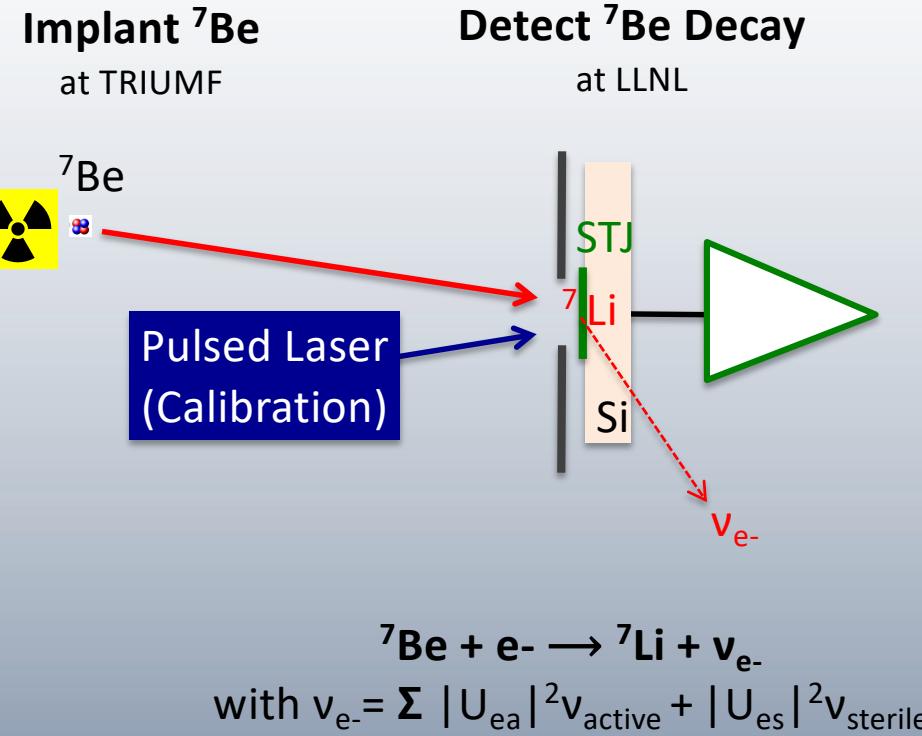
$$E_{^7\text{Li recoil}} = \frac{Q^2 - m_\nu^2 c^4}{2(Q - m_{^7\text{Li}} c^2)}$$
$$\rightarrow 56.826(9)\text{eV for } m_\nu \approx 0$$



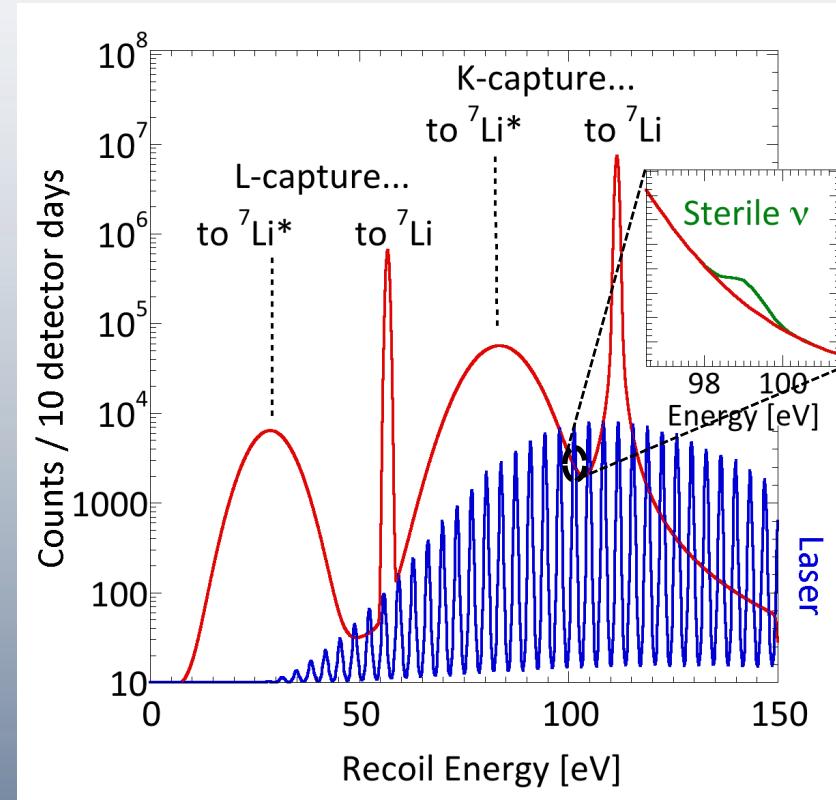
Heavy sterile neutrinos would reduce ^7Li recoil energy.

Look for shifted peaks in the recoil spectrum.

The BeEST Sterile Neutrino Experiment



High-Resolution Recoil Simulation



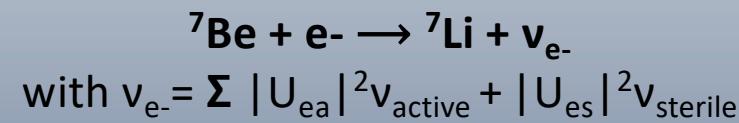
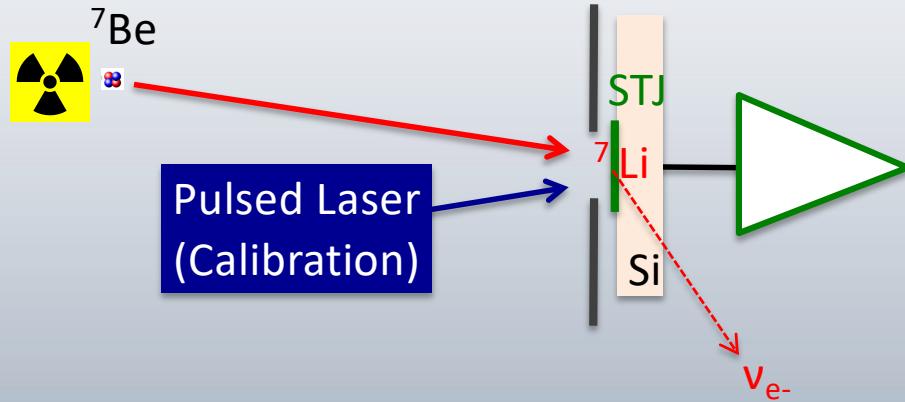
Calibrate STJ with pulsed laser.

Four peaks due to K- and L- capture into ${}^7\text{Li}$ ground and excited state

The BeEST Sterile Neutrino Experiment

Implant ${}^7\text{Be}$

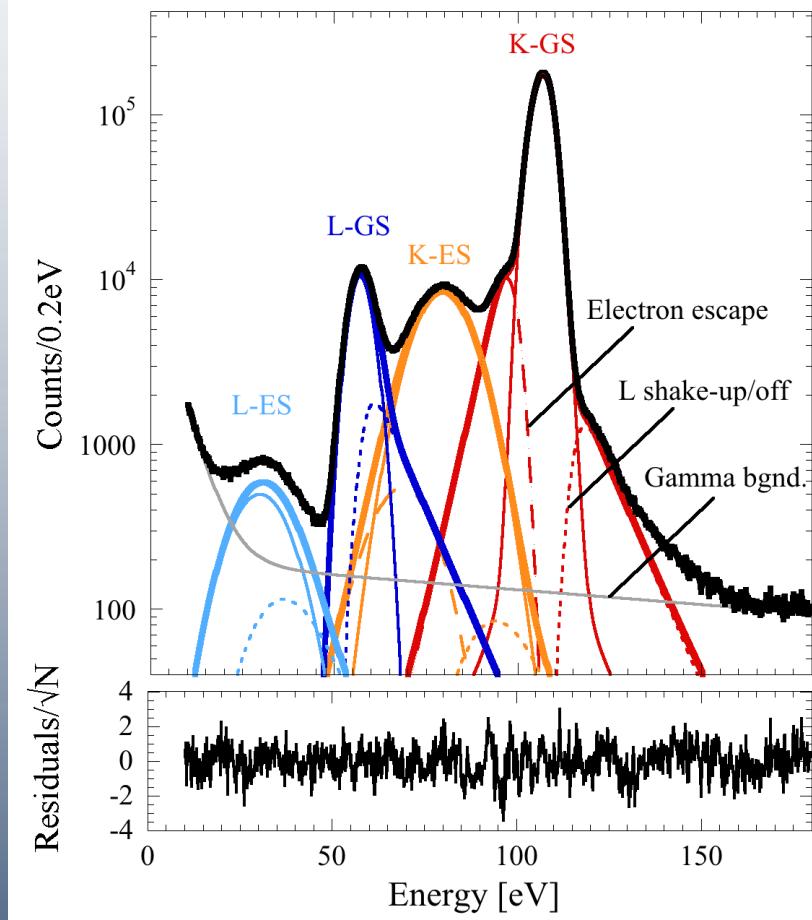
at TRIUMF



Detect ${}^7\text{Be}$ Decay

at LLNL

Data from 1 Ta-STJ Detector for 28 Days



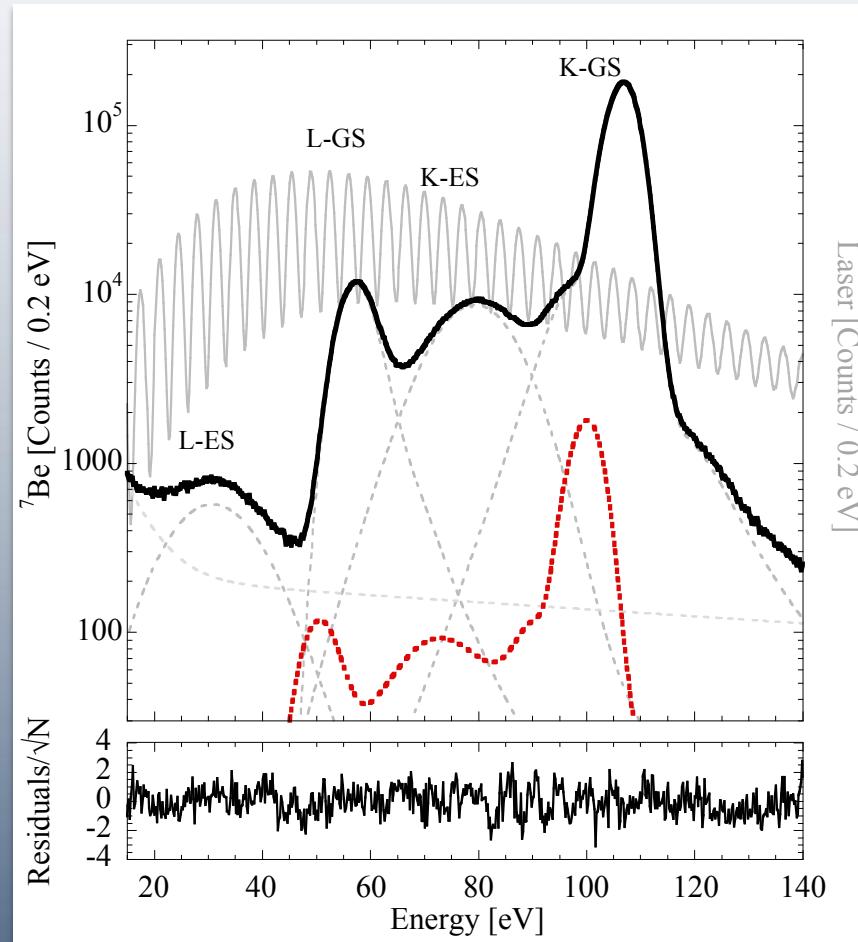
- 4 primary peaks
 - 2 x K-capture, 2x L-capture
 - to ${}^7\text{Li}$ ground state and to ${}^7\text{Li}^*$
- 4 high-energy tails
 - Shake-off effects
- 2 low-energy tails
 - (Partial) Auger e- energy loss
- 1 broad background
 - 478 keV γ 's in substrate

$$\text{L/K Ratio} = 0.070(7)$$

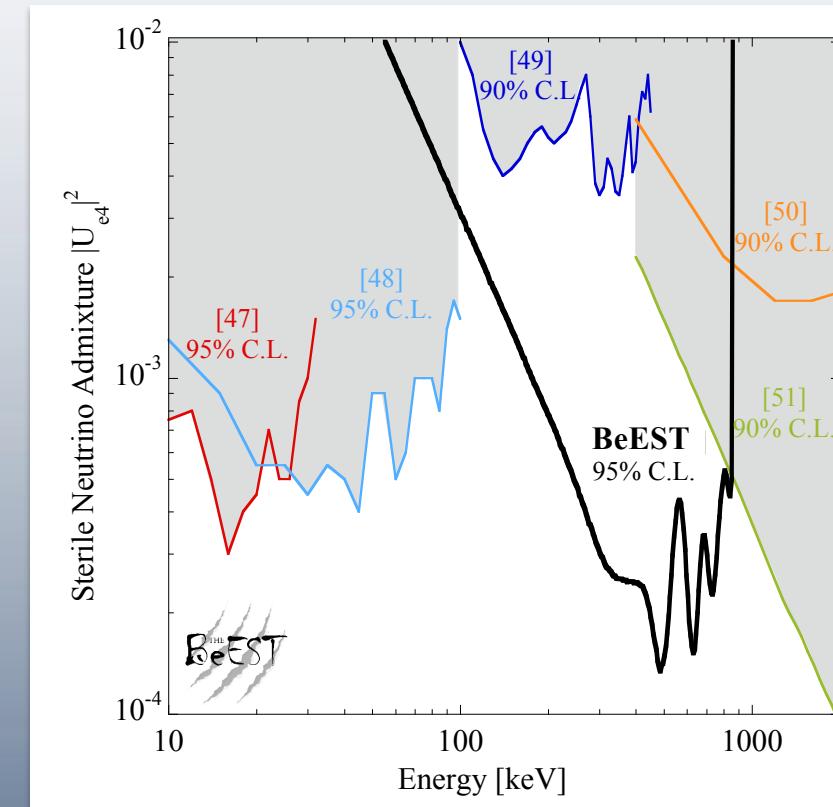
PRL 125, 032701 (2020)

The BeEST Sterile Neutrino Experiment

Data with Hypothetical Sterile ν Signal

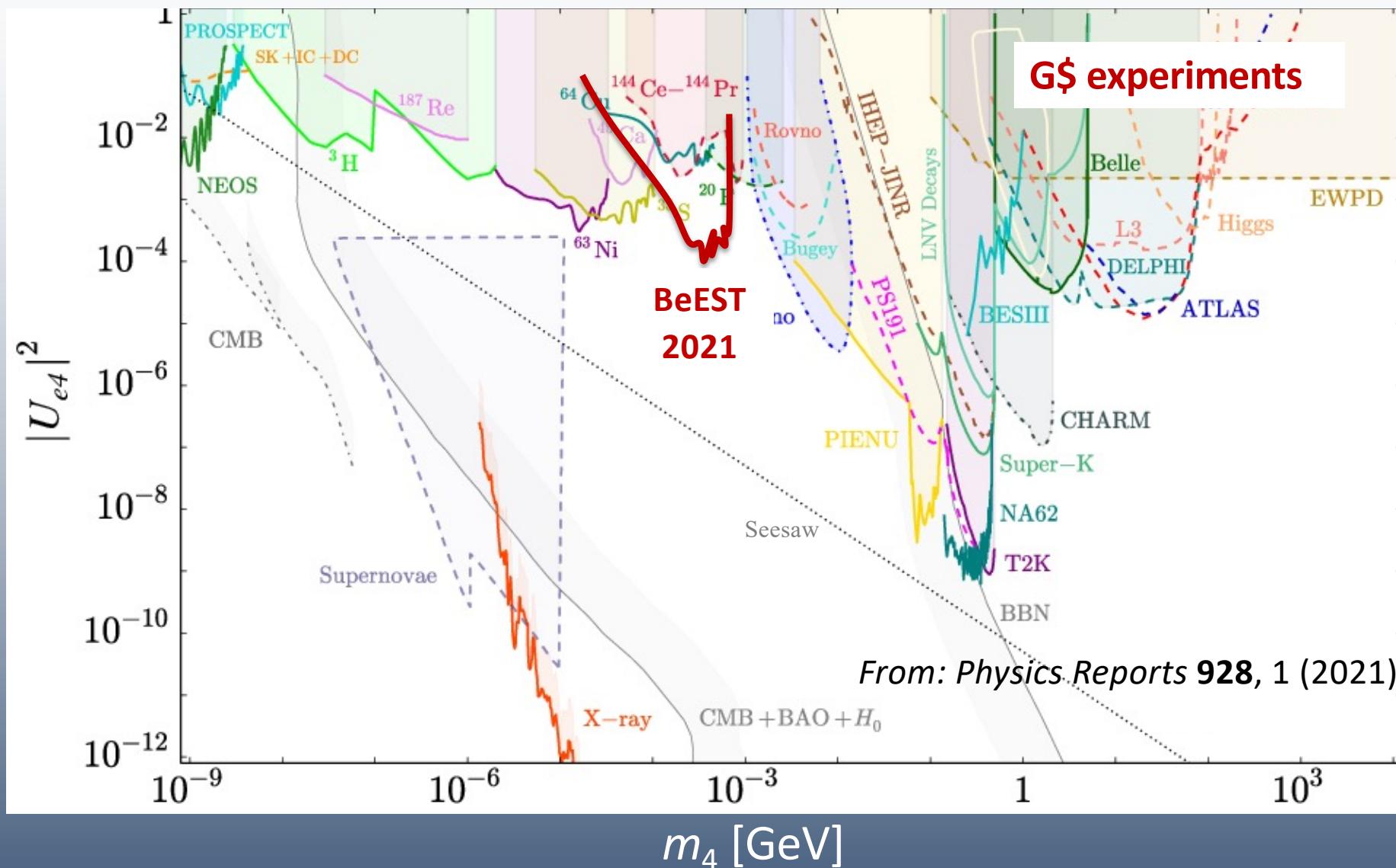


Exclusion Plot



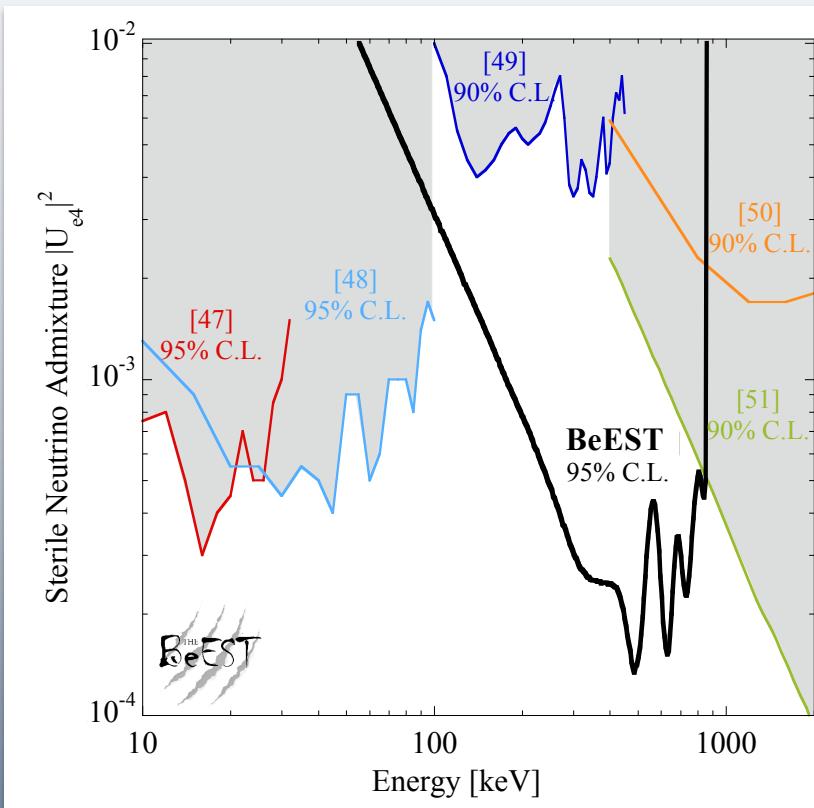
10x improvement with 1 pixel at 10 counts/s.

The BeEST in Context



4) Why Are The Capture Peaks so Broad?

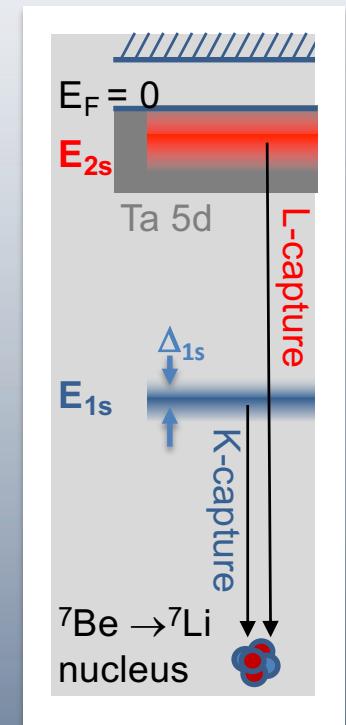
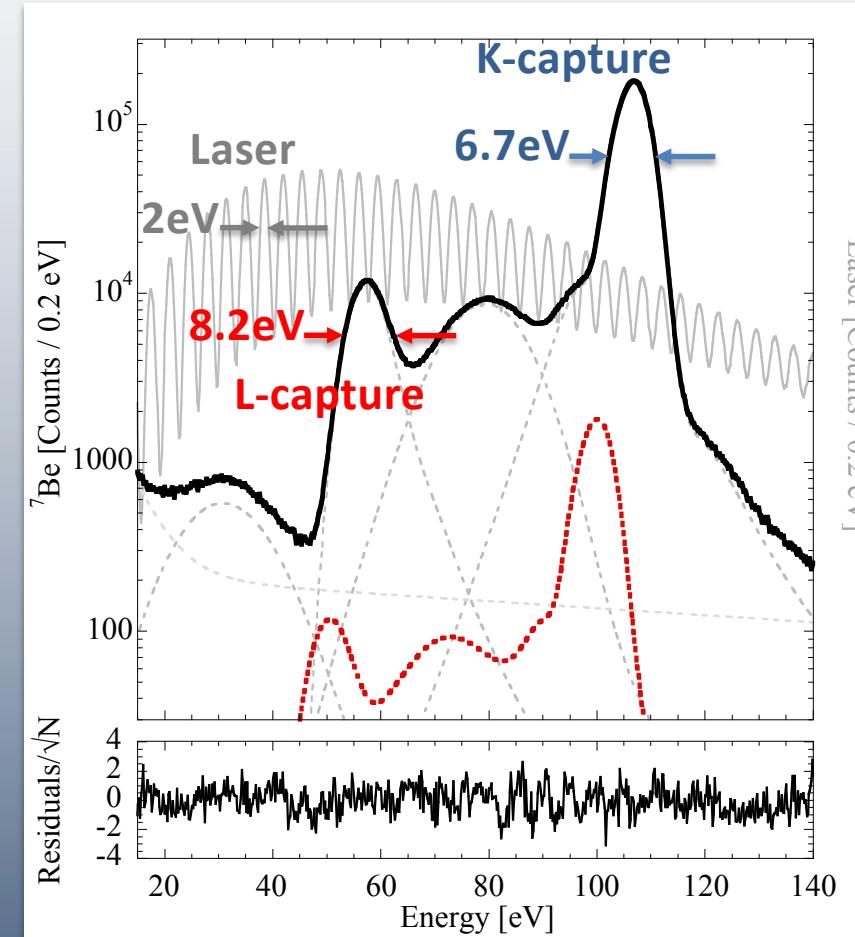
Exclusion Plot



10x improvement with 1 pixel at 10 counts/s.

PRL 126, 021803 (2021)

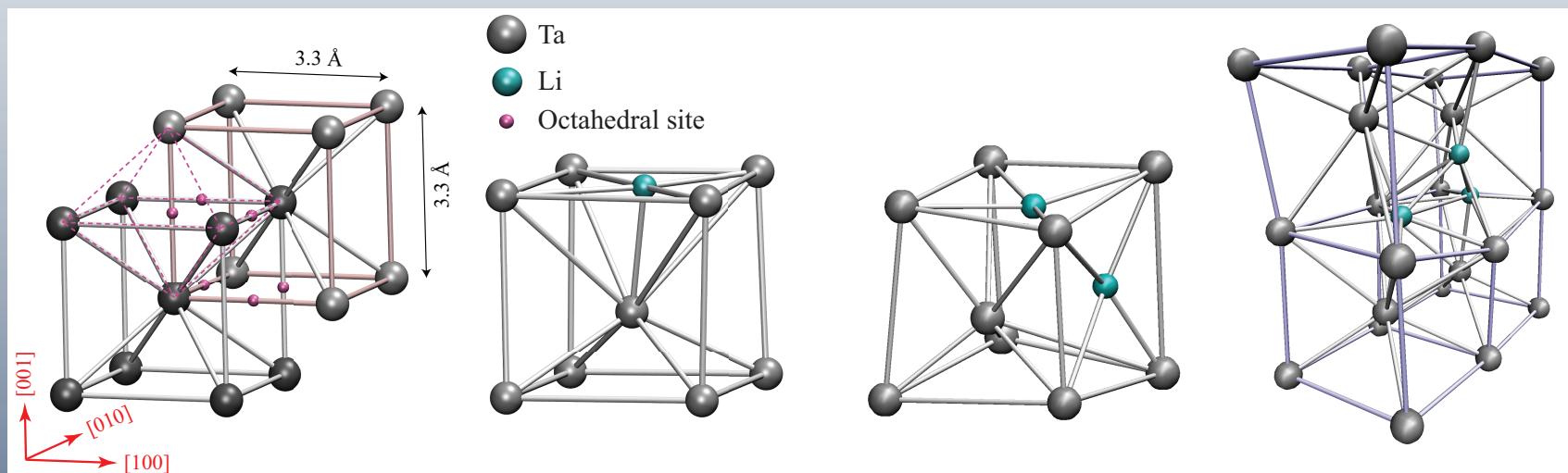
Data with Hypothetical Sterile ν Signal



Chemical Broadening?

Density Functional Theory (DFT) Simulations of Li 1s Energy Shift

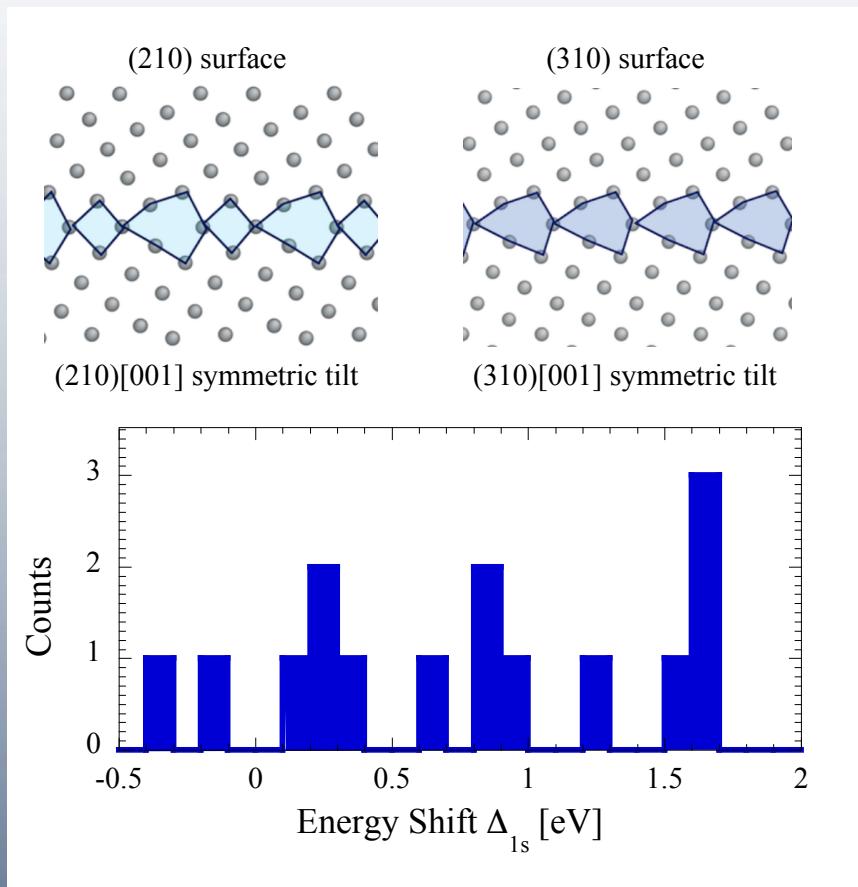
Li Configuration:	Substitutional Li	1 Li Intersitital	2 Li Intersititals	3 Li Intersititals	$\text{Li}_{\text{Ta}}\text{-V}_{\text{Ta}}$ Complex
1s Energy Shift:	0 (reference case)	1.18 eV	1.58 eV	1.30 eV	1.67 eV
Formation E:	1.23 eV	4.07 eV	7.63 eV	11.23 eV	2.50 eV
Binding E / Li atom vs. separate interstitials	–	–	-0.26 eV	-0.33 eV	-1.53 eV



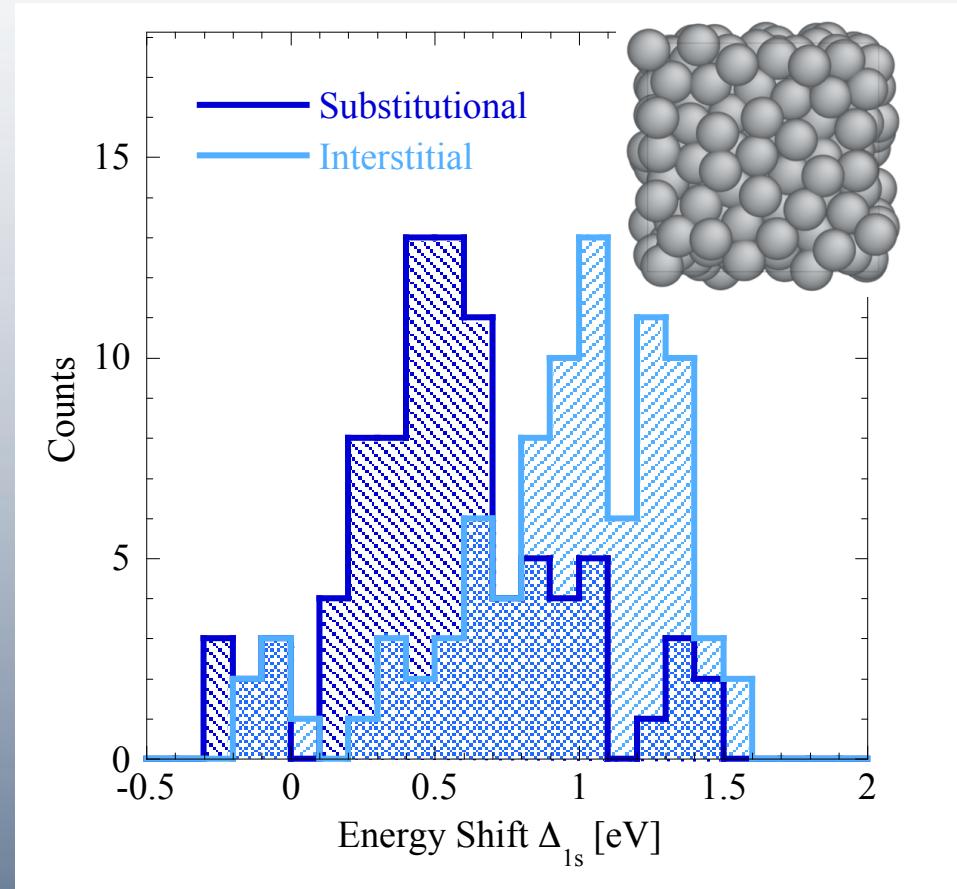
Simple configuration changes can alter Li 1s energy levels by >1 eV.

DFT Simulations in Disordered / Amorphous Ta

Li at Grain Boundaries



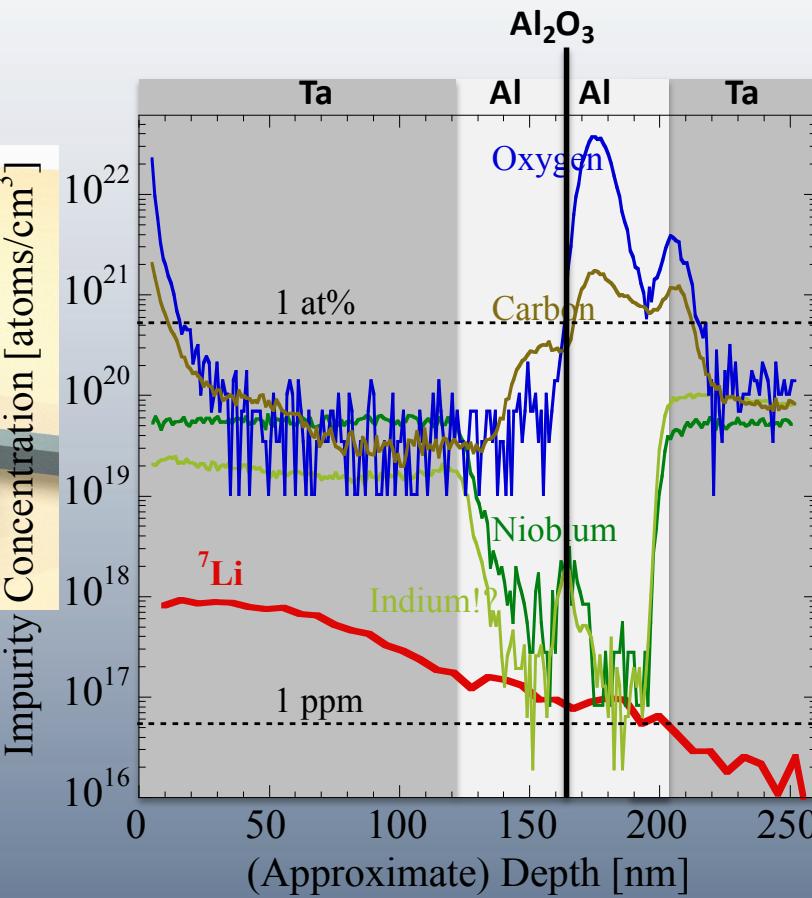
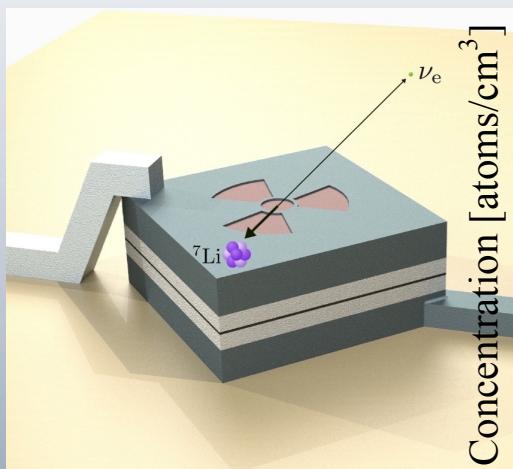
Li in in Amorphous Ta



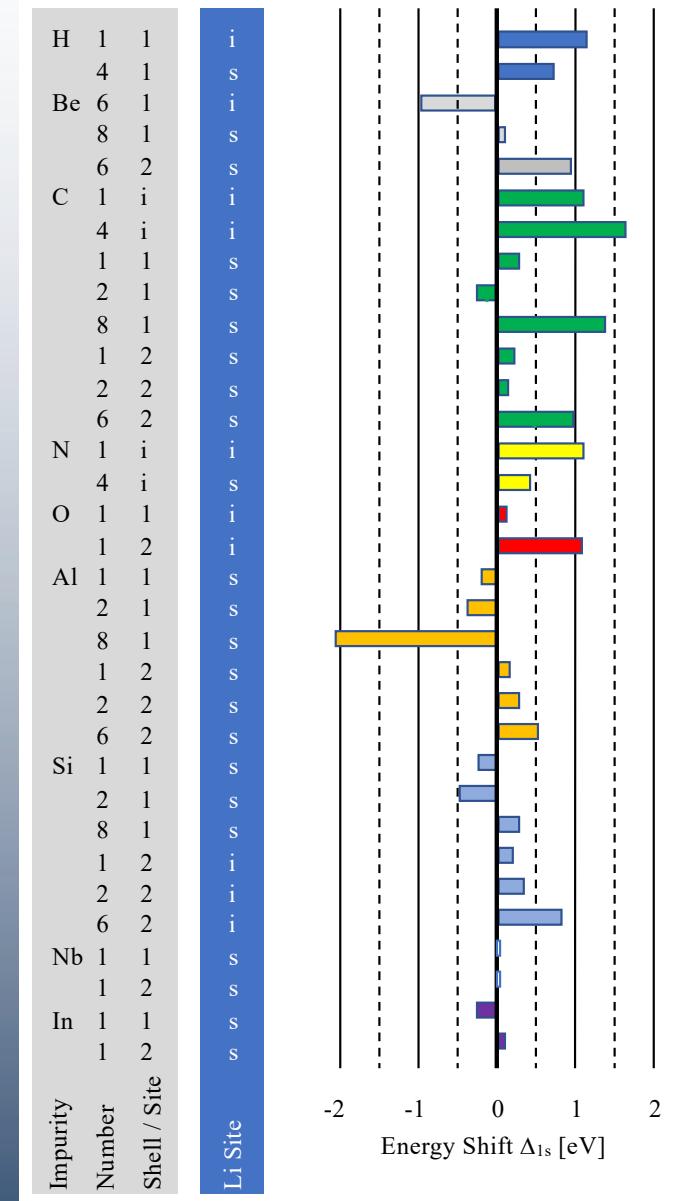
Even extreme disorder only shifts 1s levels by ~ 2 eV and cannot explain the observed 6.7 eV broadening

DFT Simulations of Li 1s Energy Shifts due to Impurities

TOF-SIMS Analysis of Impurities in STJ Detector

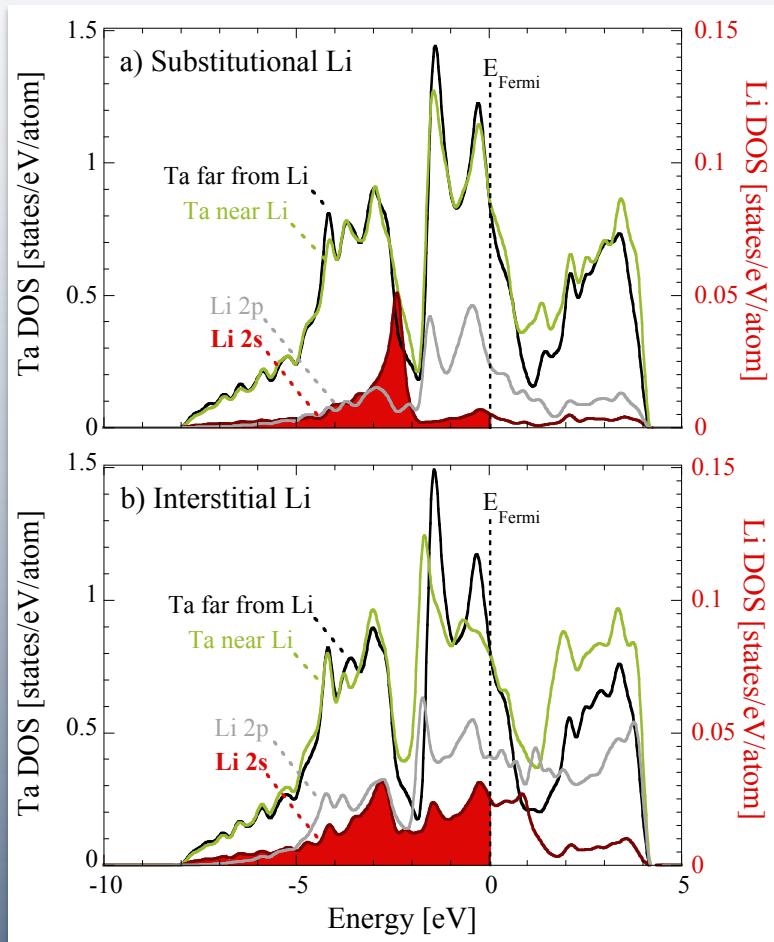


Impurities can change the Li 1s energy levels by ~2eV
(Again, much less than observed experimentally.)

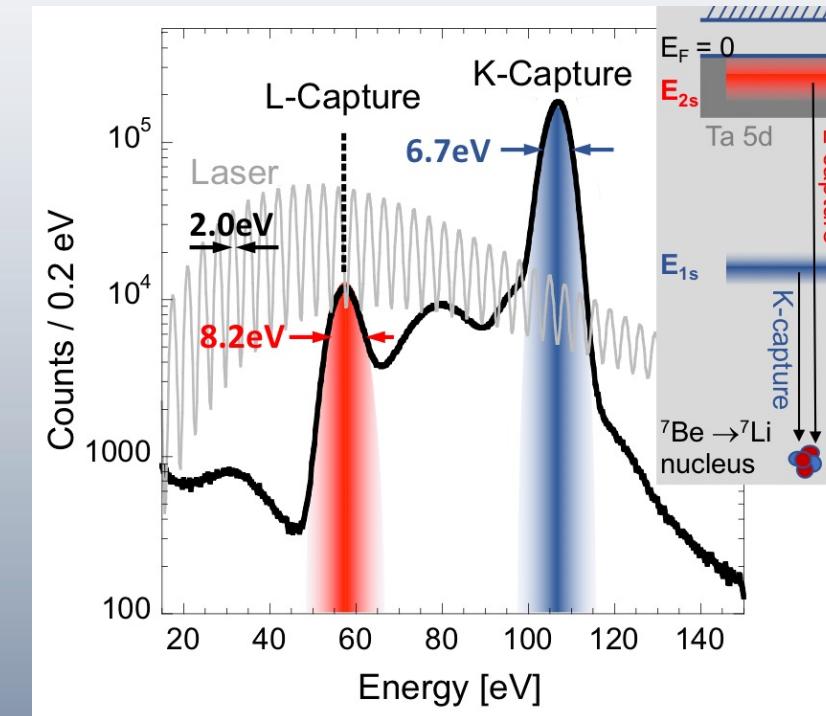


DFT Simulations of Li 2s Energies in bcc Ta

Li 2s Levels in Ta 5d Band



L-Capture is Broader than K-Capture Peak



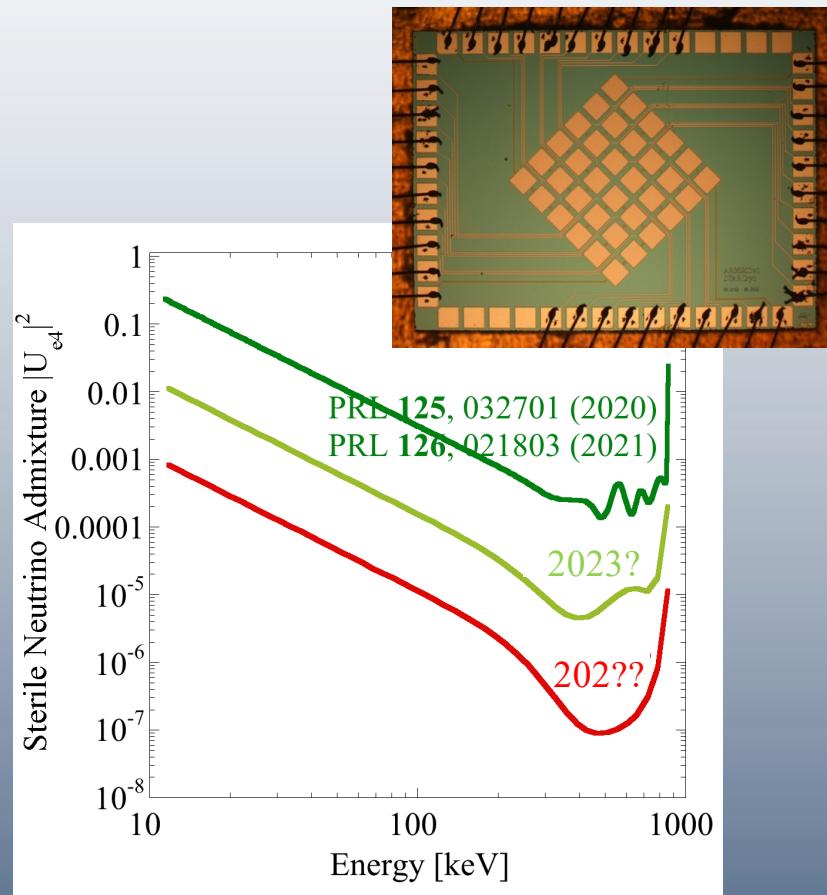
A. Samanta et al., arXiv: 2206.00150

Hybridization of Li 2s with Ta 5d-6s band with explains why L-capture peak is broader than K-capture peak.

Material effects matter but cannot explain entire observed broadening.

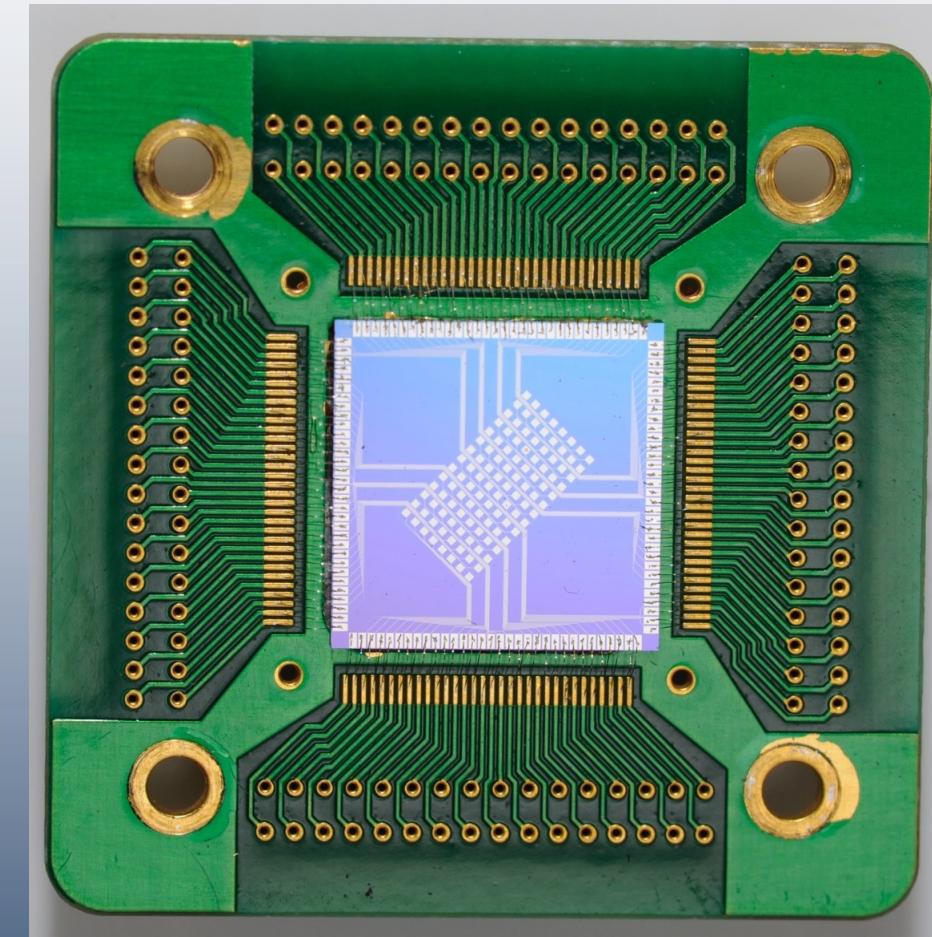
5) Next: Bigger Detector Arrays, Higher ${}^7\text{Be}$ Dose

Improved Sensitivity with Ta-STJ Arrays



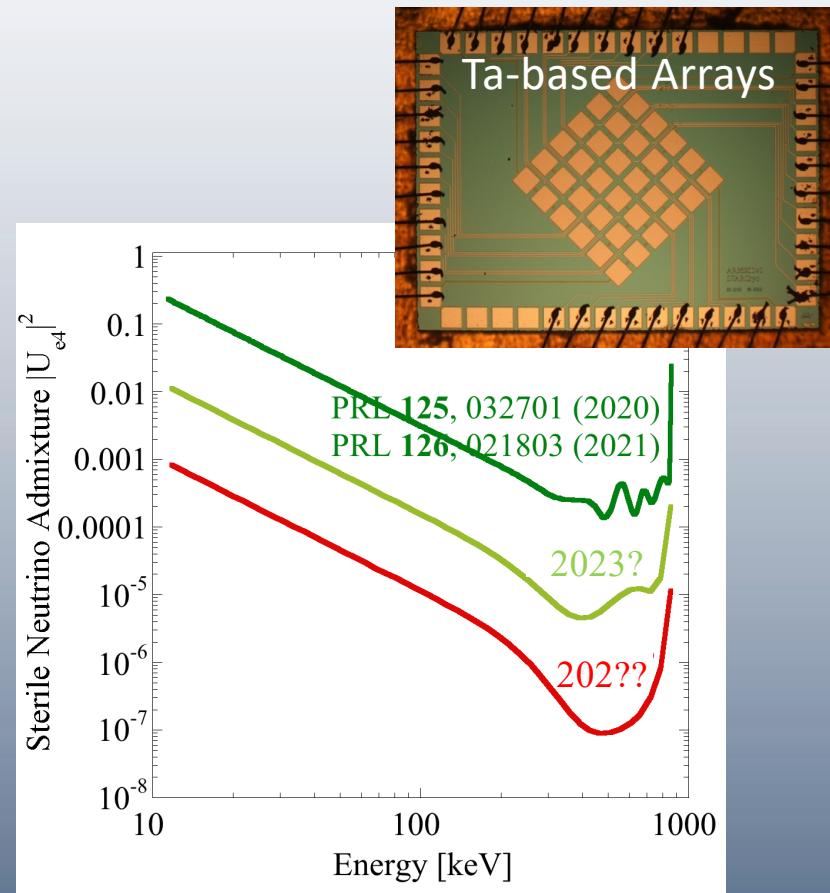
Next ${}^7\text{Be}$ implantation on 09/22/22

112-pixel Ta-STJ Array

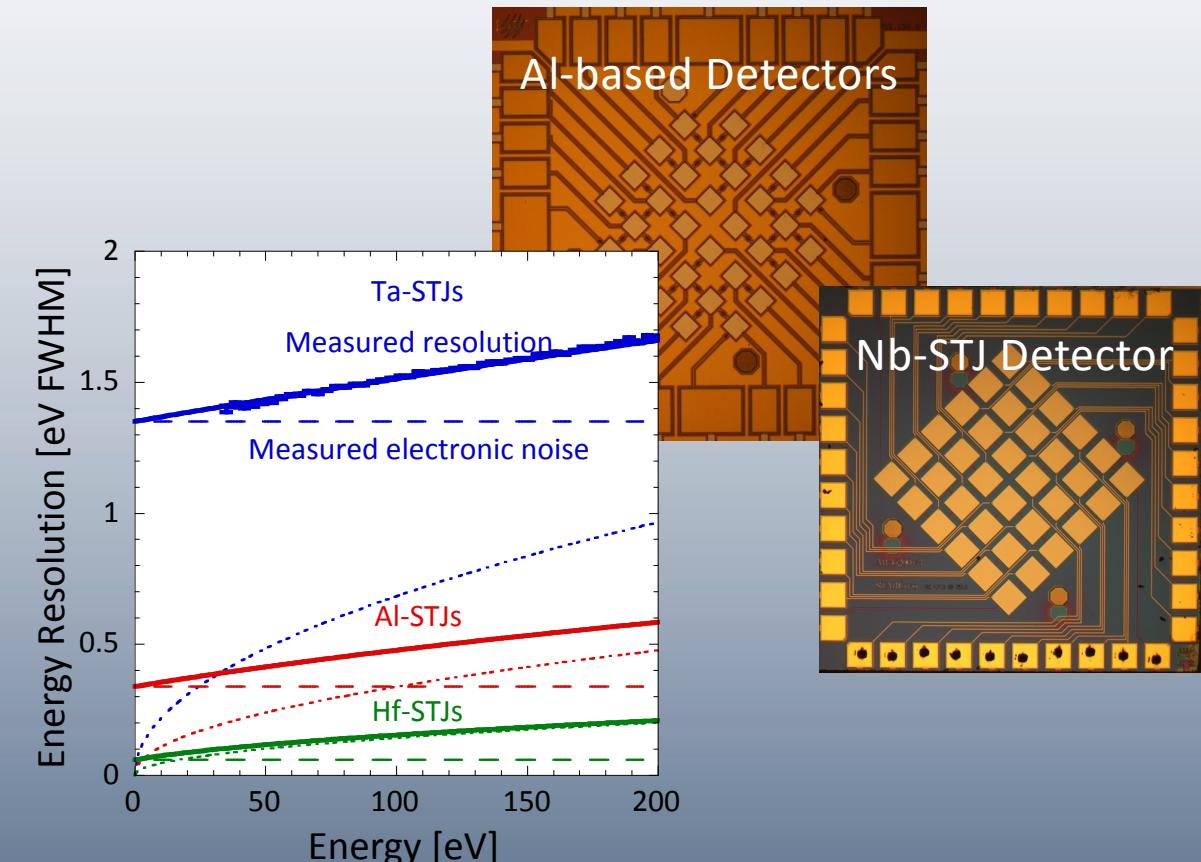


Next: New Materials

Improved Sensitivity with Ta-STJ Arrays



STJ Detectors from Different Materials

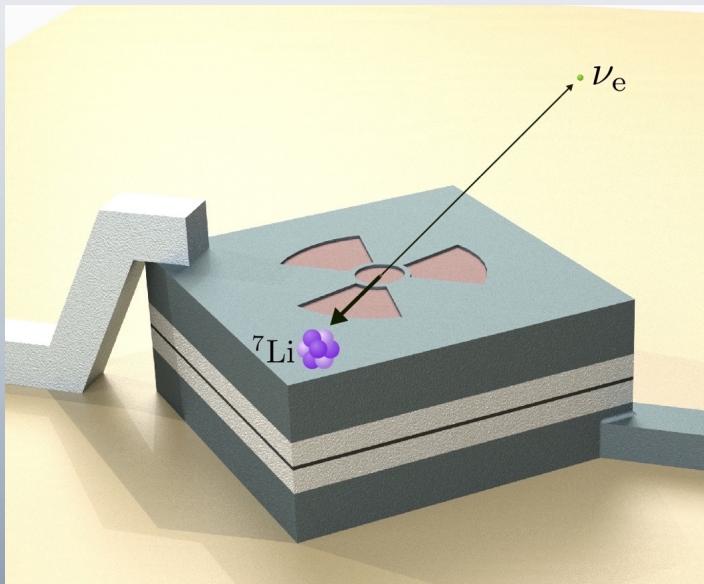


Next ${}^7\text{Be}$ implantation on 09/22/22

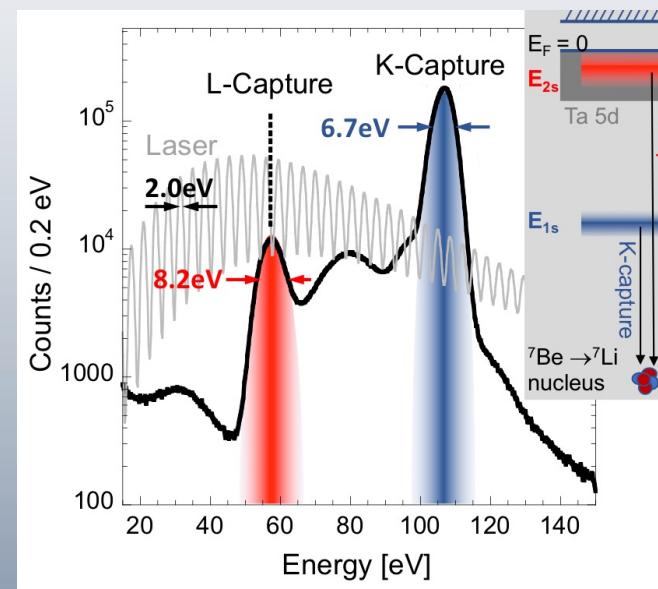
Distinguish material effects from BSM physics

The *BeEST* Sterile Neutrino Search

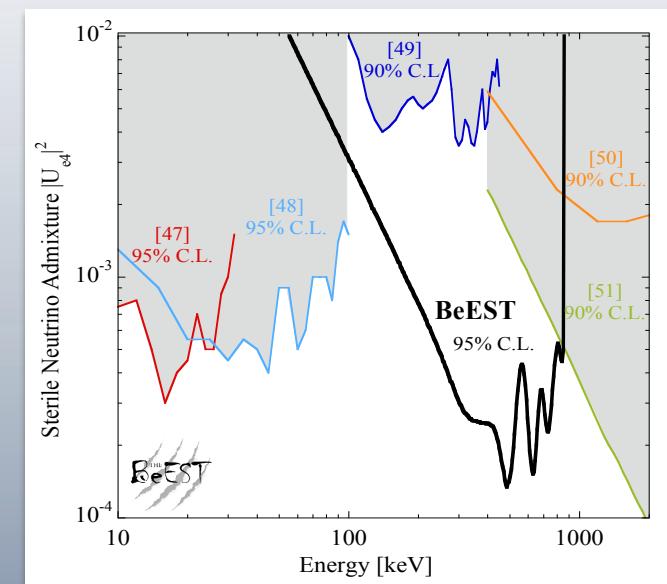
^7Be in STJ detectors
 \Rightarrow Missing Momentum Search



2 eV FWHM and 2-body decay
 \Rightarrow High Sensitivity



1STJ at 10 counts/s
 \Rightarrow Exclusion to $O(10^{-4})$



- High resolution: ~ 2 eV at 100 eV
- High speed: $> 1,000$ counts/s/pixel
- No quenching

- Accurate laser calibration
- Material effects non-negligible

- $O(10^{-7})$ with 128-pixel arrays
- New Al- and Nb-STJs \Rightarrow Separate BSM physics from material effects

Thank You!



STJ Fabrication
(STAR Cryo)
Robin Cantor
Ad Hall



STJ Experiments
(LLNL)
Stephan Friedrich
Inwook Kim



STJ Preamplifier
(XIA)
Bill Warburton
Jack Harris



Processing / Analysis
Francisco Ponce (PNNL)
Geon-Bo Kim (LLNL)
Spencer Fretwell



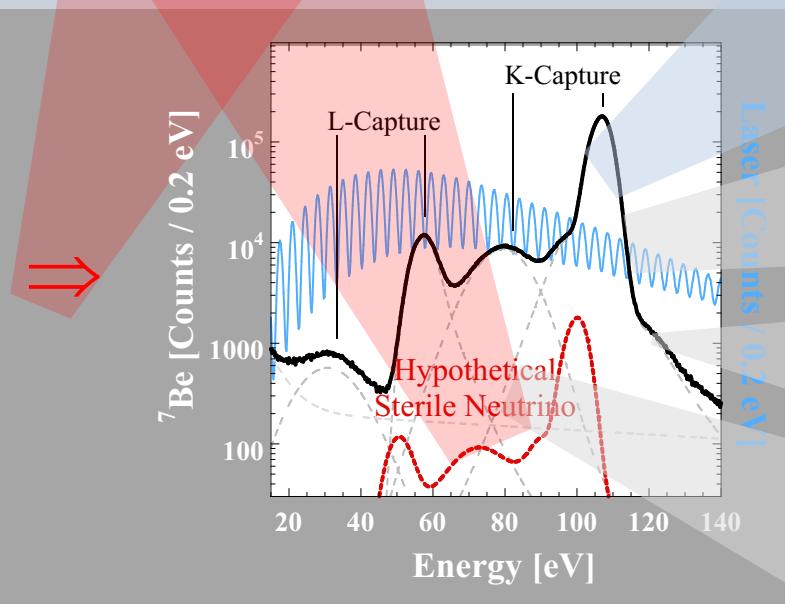
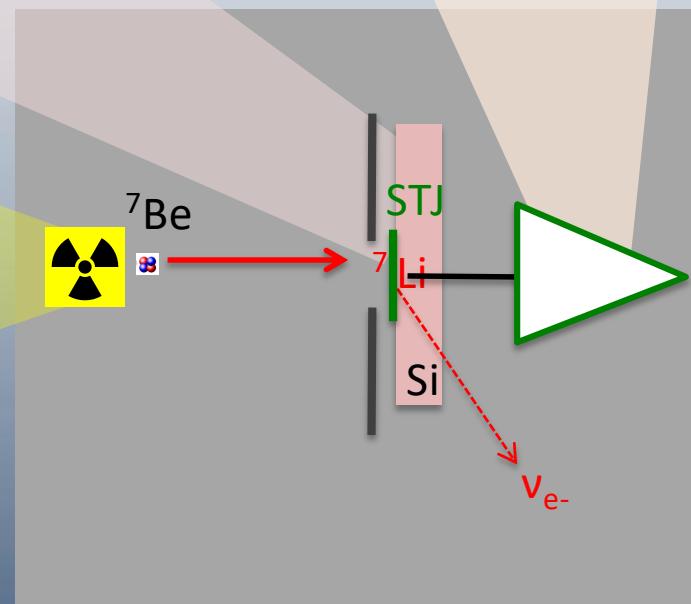
BSM Physics
(Mines)
Kyle Leach
Dave McKeen



Material Characterization
Dave Diercks (Mines)
Tony Li (LLNL)
Yi-De Chuang (ALS)
Cameron Harris



$^{7\text{Be}}$ Implantation
(TRIUMF)
Chris Ruiz
Annika Lennarz



Simulations
Broadening (LLNL)

Vince Lordi
Amit Samanta



Shake-up/-off (UNL)

Jose-Paolo Santos
Jorge Machado
Mauro Guerra
Padro Amaro



Escape Tails

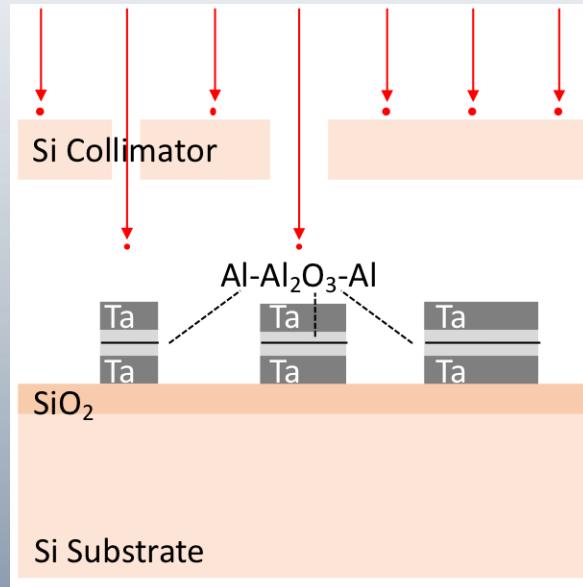
Connor Bray

Electron Capture
Xavier Mousseot (CEA)

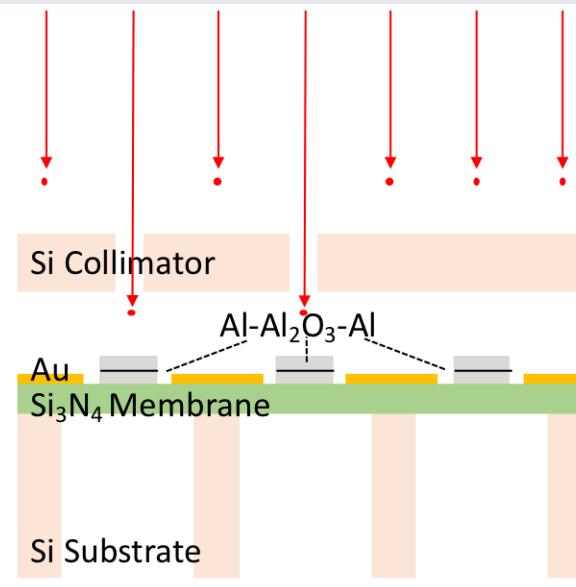
Back-up Slides

Different STJ Detectors

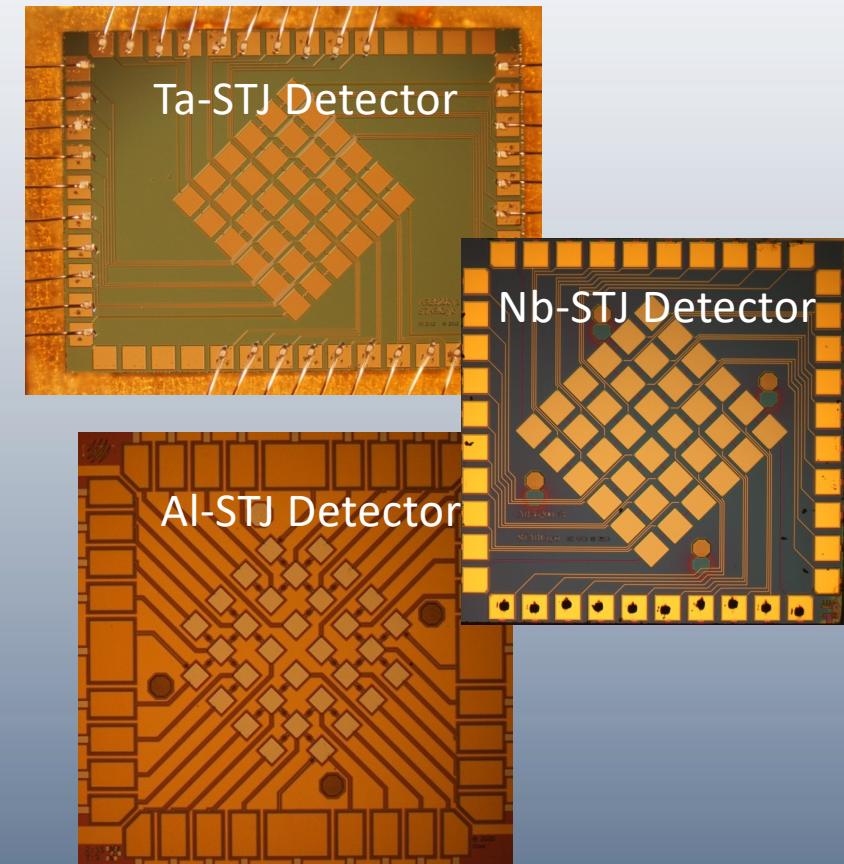
Current Geometry



Next Geometry



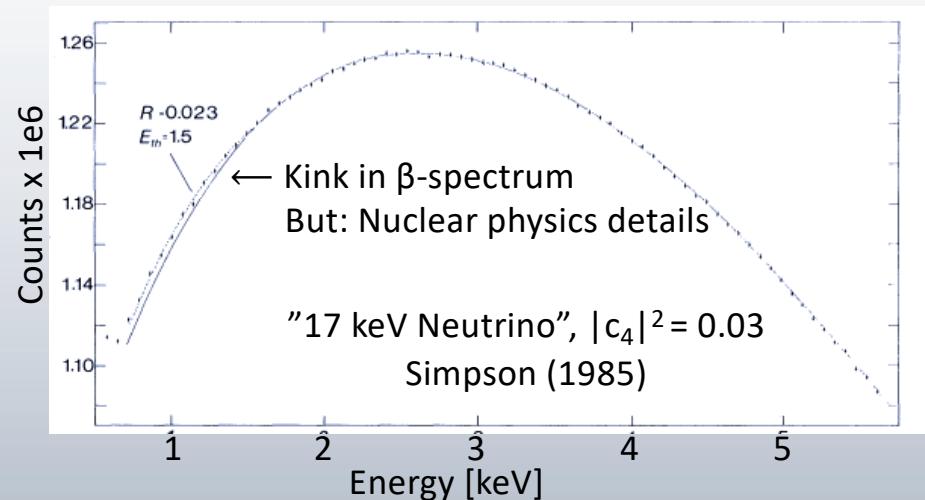
Different Materials



Complementary experiments and simulations of different materials might separate BSM physics from material effects.

Competing Sterile Neutrino Experiments

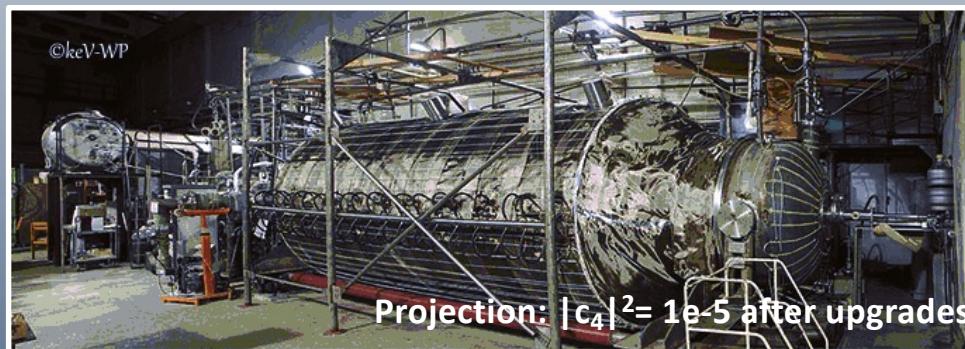
^3H Beta Decay



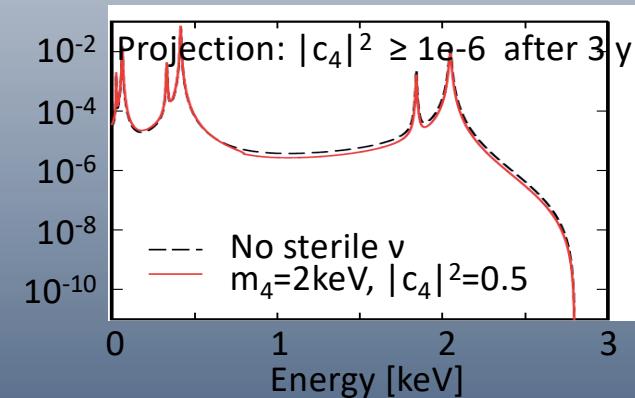
KATRIN = KArlsruhe TRItium Neutrino Expt



Troitsk "nuMass" ^3H



ECHo = Electron Capture in Ho

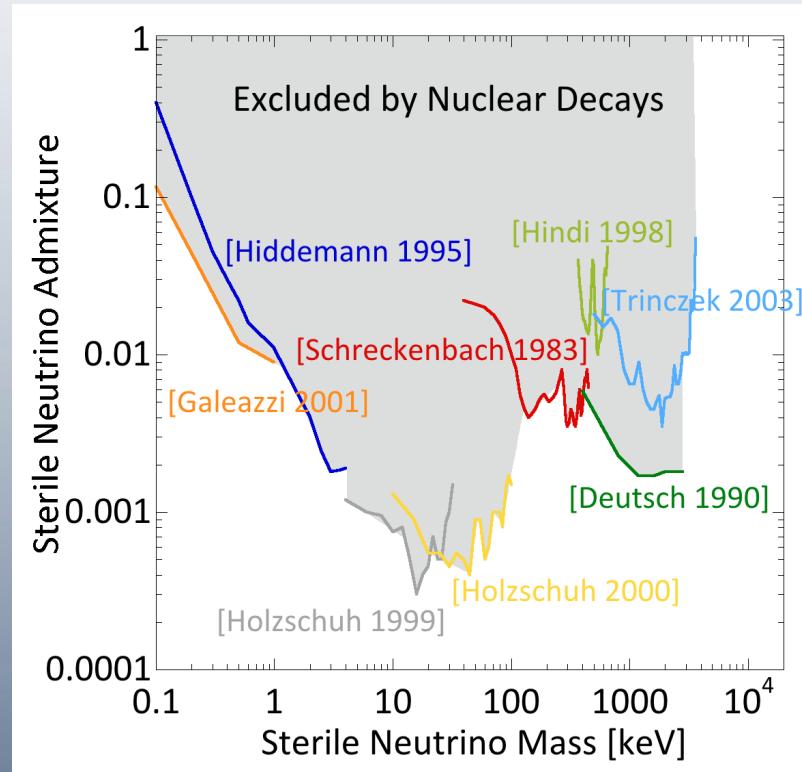


We are more sensitive and much cheaper than competing experiments
(which admittedly are not optimized for sterile neutrinos.)

Experimental Limits on Sterile Neutrino Dark Matter

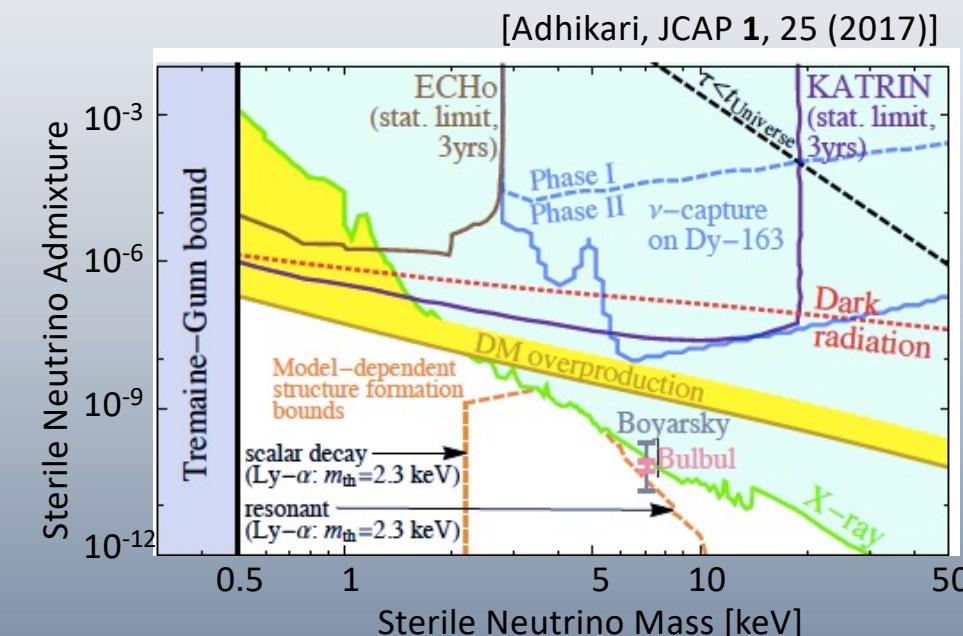
Limits from Nuclear Decays

Look (in vain) for kink in beta spectrum



Limits from Astronomy

Look (in vain) for decay X-rays from galaxies



So we're looking for ~keV neutrinos, 10^{-12} to 10^{-3} of which are mixed to active neutrinos.

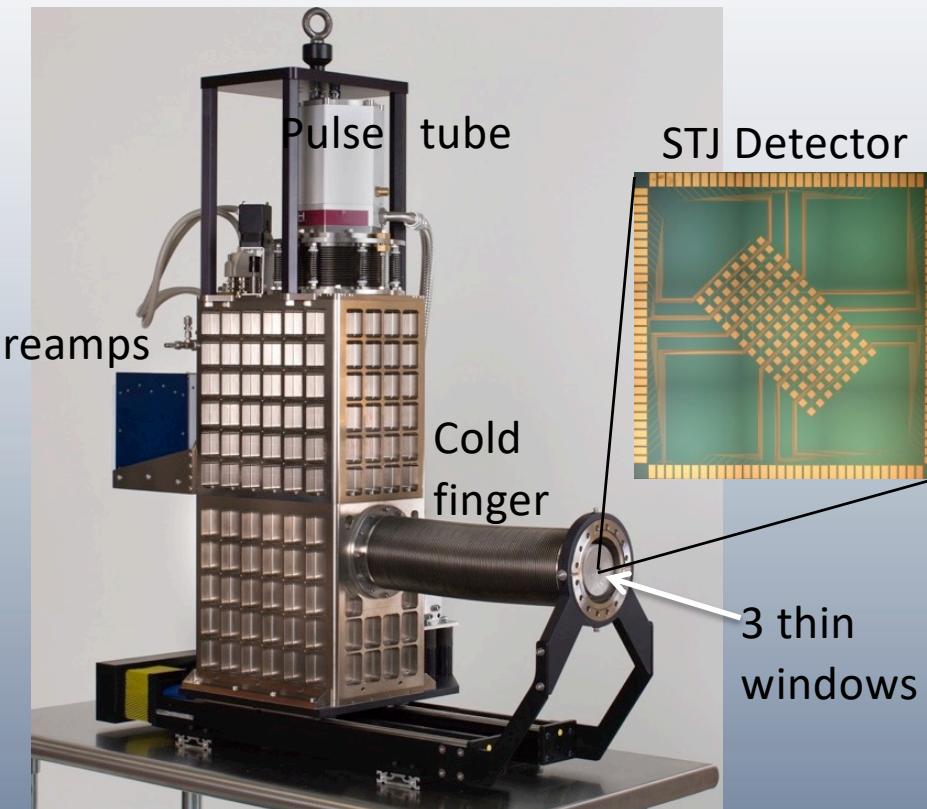
Astronomy limits depend on underlying assumptions and models.

Adiabatic Demagnetization Refrigerators (ADRs)

“Wet” Cryodetector ADR (2001)



Commercial “Dry” ADR (2016)

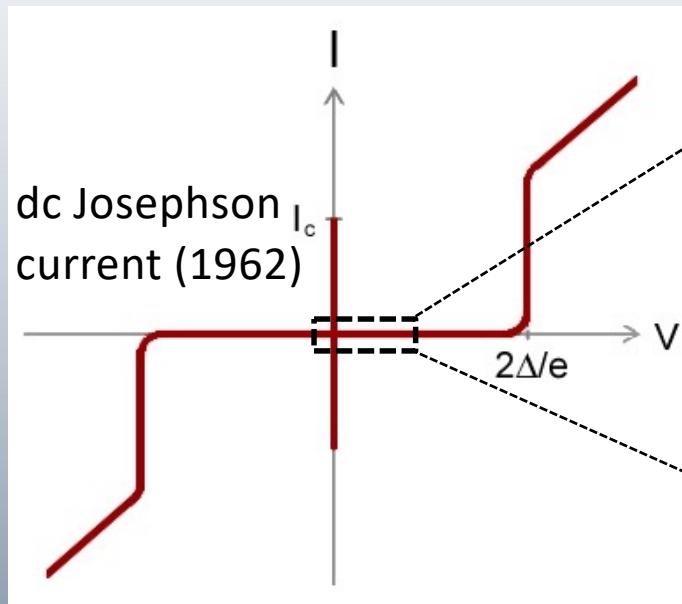


ADR refrigerators for automated cooldown to $<0.1\text{K}$ are commercially available ($\sim \$500\text{k}$).

STJ Electronics

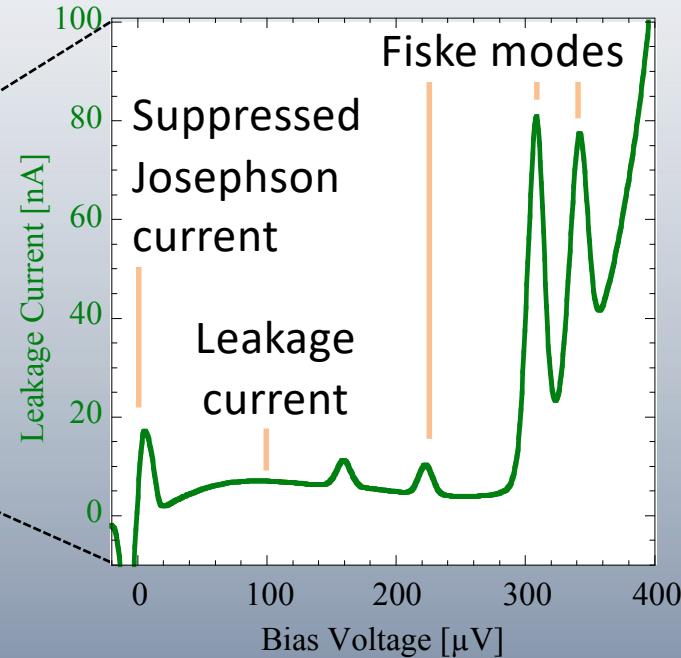
STJ Biasing

...in Textbooks



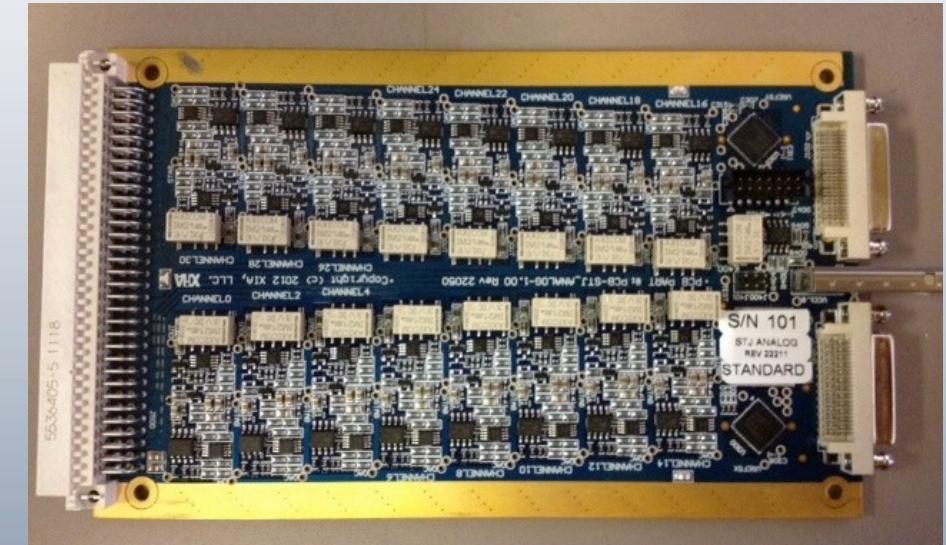
dc Josephson current (1962)

...in Real Life



- Low noise ($e_n \leq 1$ nV/ $\sqrt{\text{Hz}}$)
- Stable biasing (to \pm few μ V)
- Voltage bias (dc load line $< 100 \Omega$)

32-channel STJ Preamplifier (XIA)

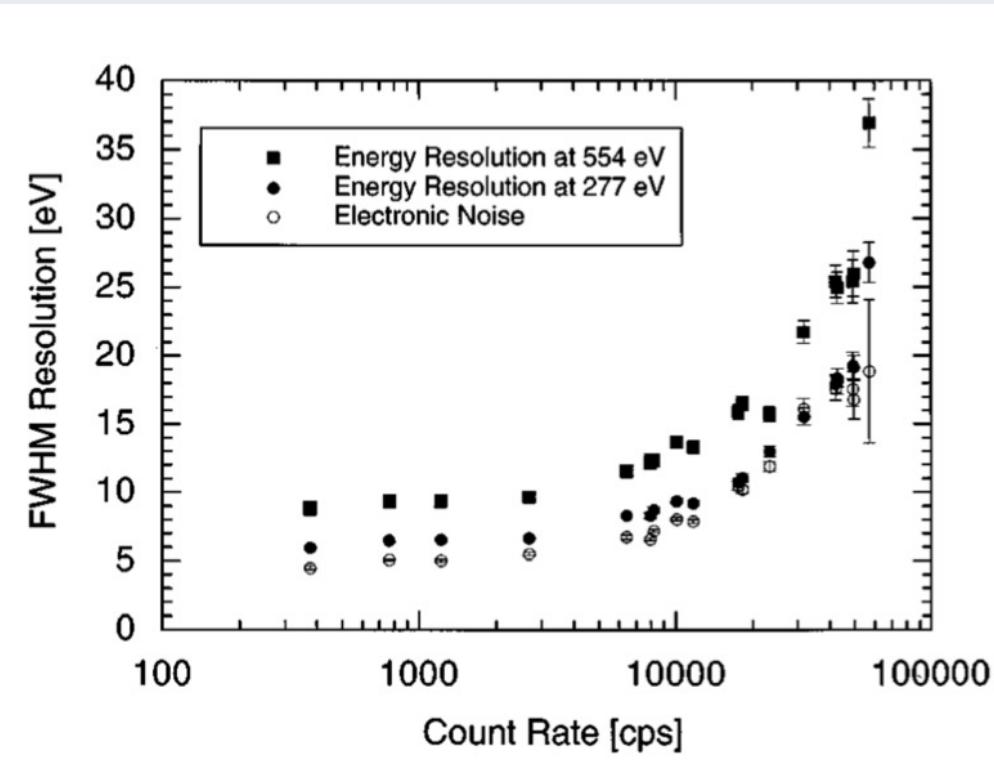


- Low noise, stable biasing, dc voltage bias
- Computer-controlled
- 32 preamplifiers per board

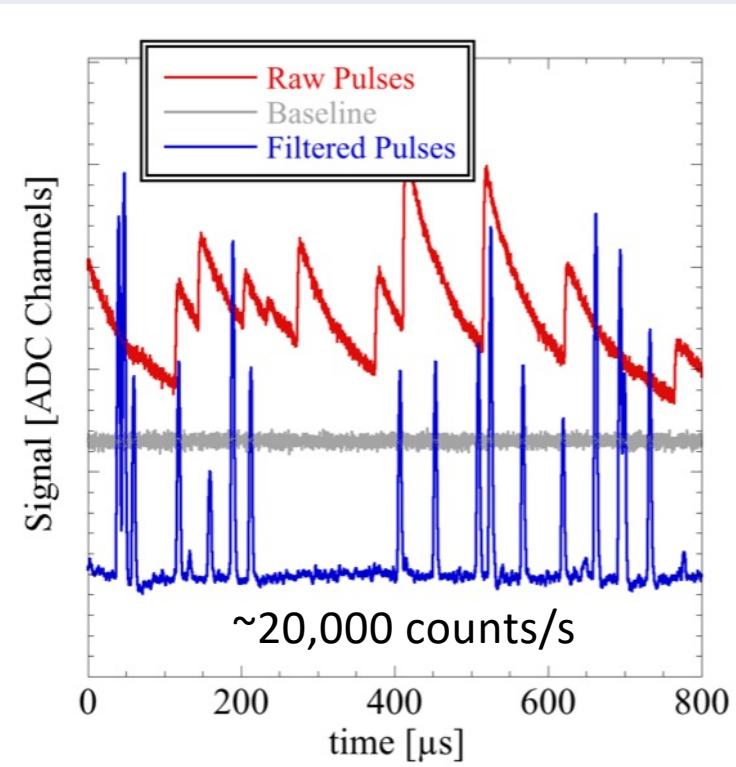
Specialized preamplifiers for STJs are commercially available.

STJ Characterization at High Speed

Nb-STJs ($\tau_{\text{decay}} \approx \text{few } \mu\text{s}$)



Ta-STJs ($\tau_{\text{decay}} \approx \text{few } 10 \mu\text{s}$)



STJ Signal:

- Single exponential decay constant

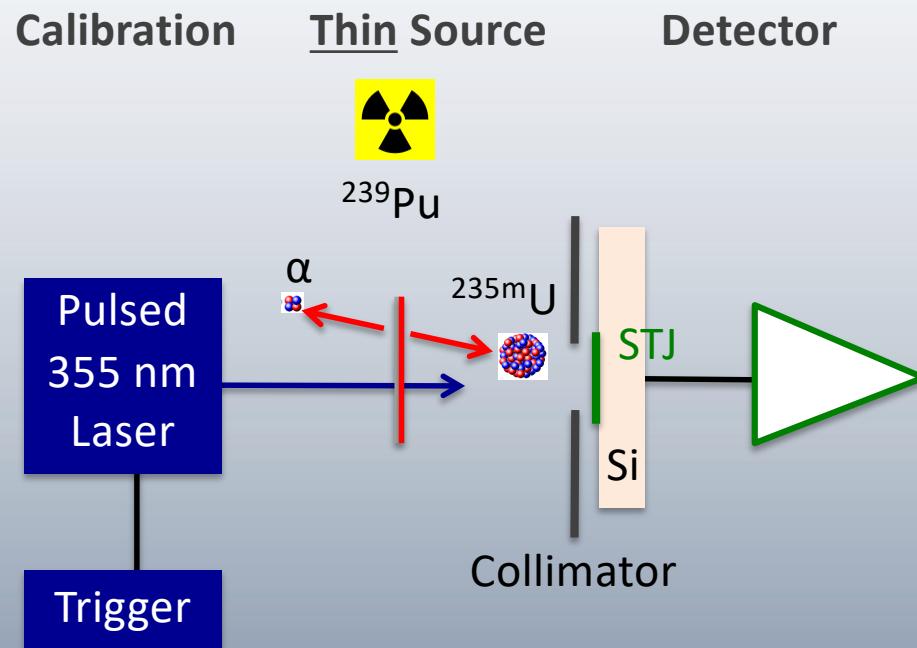
DSP pulse processing:

- Trapezoidal filter
- Pile-up rejection

STJs detectors can be operated at rates well above 1000 counts/s per pixel.

Nice, but only at low energies.

Low-Energy Nuclear Spectroscopy of ^{235m}U



- Absorb ^{235m}U from ^{239}Pu decay in STJ
- Measure ^{235m}U decay with 2 eV resolution
- Result: ^{235m}U energy: 76.737 ± 0.018 eV

