

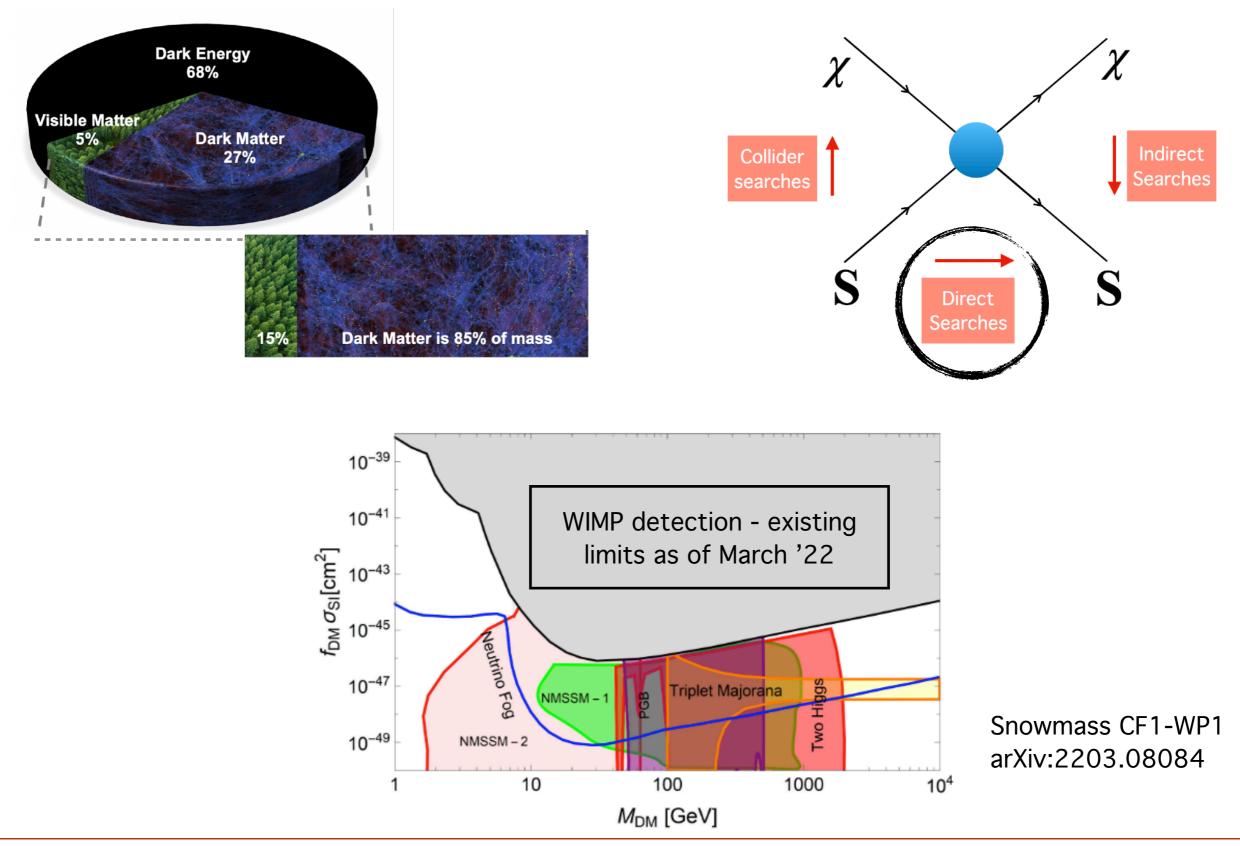
First results from the LUX-ZEPLIN (LZ) dark matter experiment

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CIPANP 2022 14th Conference on the intersections of Particle and Nuclear Physics

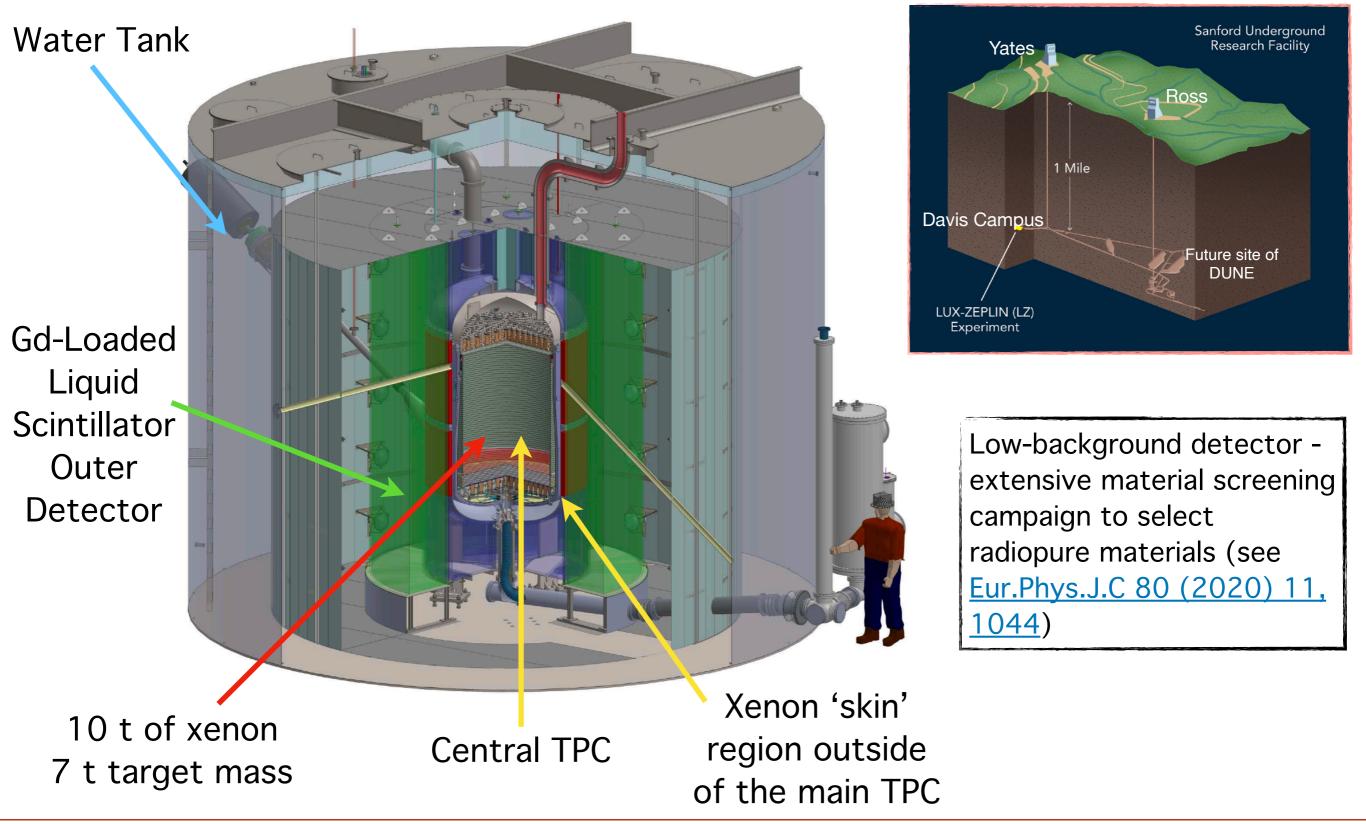


Introduction



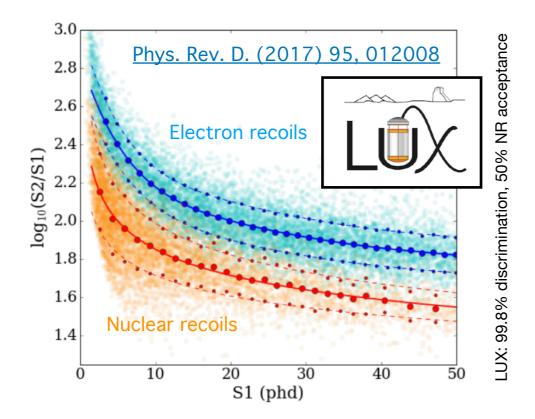
14th Conference on the Intersections of Particle and Nuclear Physics

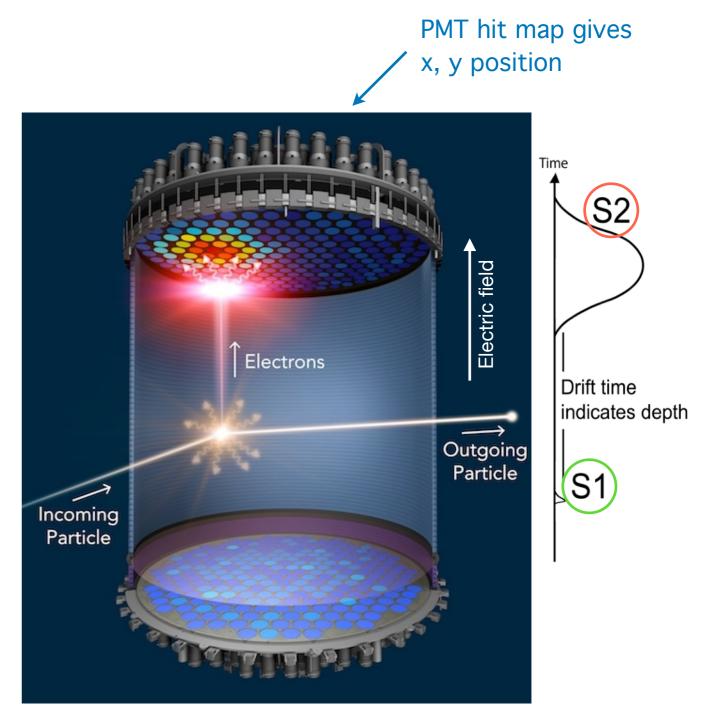
The LUX-ZEPLIN (LZ) experiment



Two-phase xenon time projection chamber

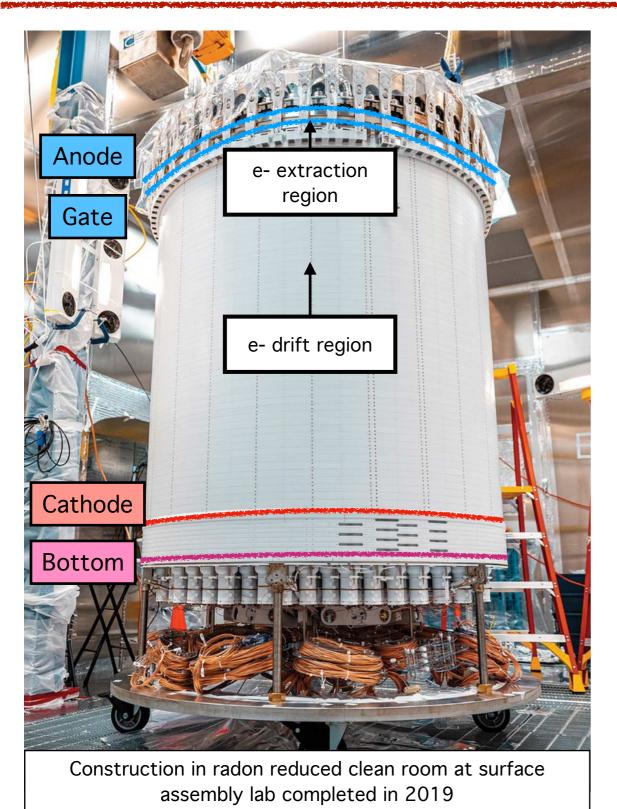
- S1 prompt scintillation
- S2 delayed scintillation after ionization electrons are drifted and extracted in gas phase
- For each event in the detector with an S1 and S2 signal, we can determine:
 - Position
 - Energy (threshold ~ few keV)
 - Recoil identification





LZ: 494 PMTs in the top and bottom arrays combined7 tonnes of liquid xenon in the 'active' region (where the electric field is applied)

The LZ TPC

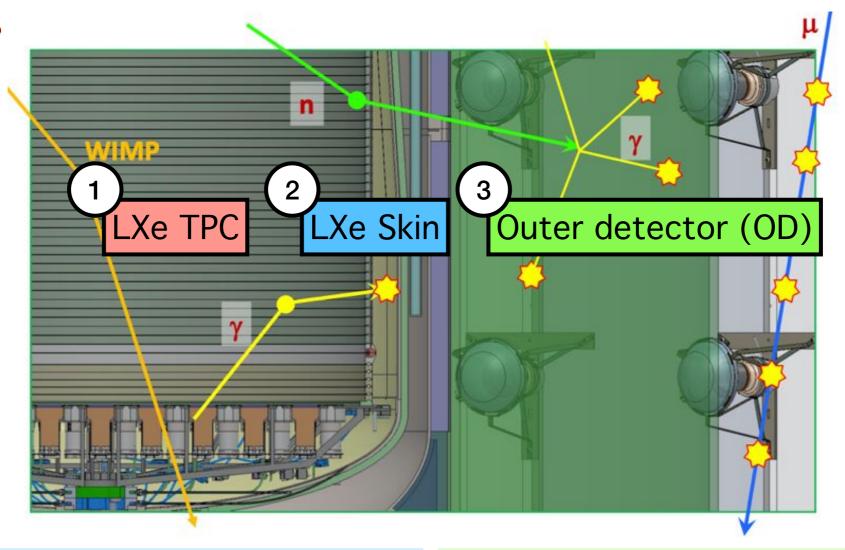


- ▶ 1.5 m diameter x 1.5 m height
- 7 t liquid xenon target
- PTFE construction for light collection
- 494 3" PMTs in two arrays on top and bottom
- 4 grids (bottom, cathode, gate, anode)
- Field cage to define TPC
- 3 spill-over weirs to define liquid surface



Bottom PMT array and cathode electrode during construction

3-in-1 in grated detector system



See talk on the LZ OD by H. Birch in session DM-2

<u>LXe Skin</u>

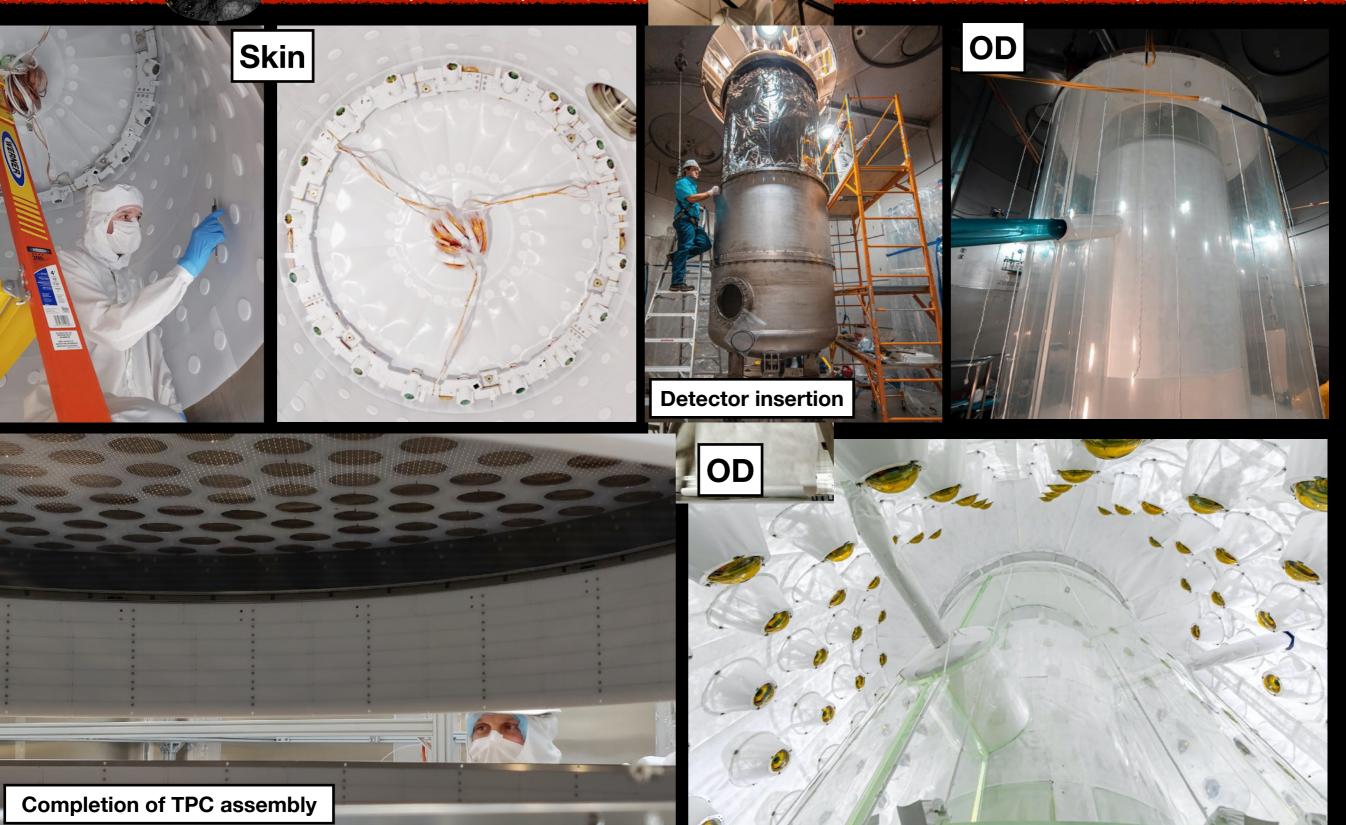
- 2 t of LXe surrounding the TPC
- 1" and 2" PMTs at top and bottom of the 'skin' region
- Lined with PTFE to maximize light collection
- Anti-coincidence detector for γ-rays

The Outer Detector (OD)

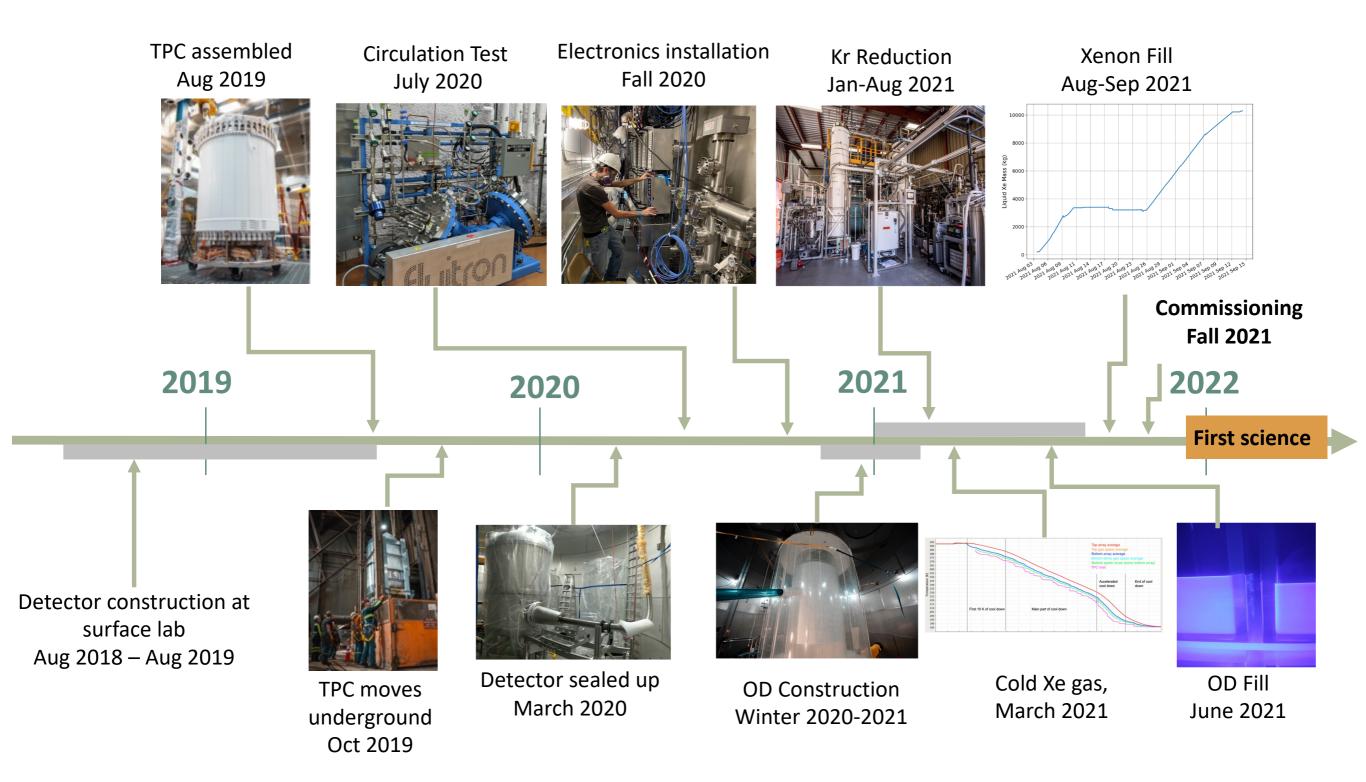
- 17 t of Gd-loaded liquid scintillator in acrylic vessels
- 120 8" PMTs mounted in the water tank
- Anti-coincidence detector for γ-rays and <u>neutrons (88% tagging efficiency</u>, measured in city with Aml i neutrons)

measured in situ with AmLi neutrons)



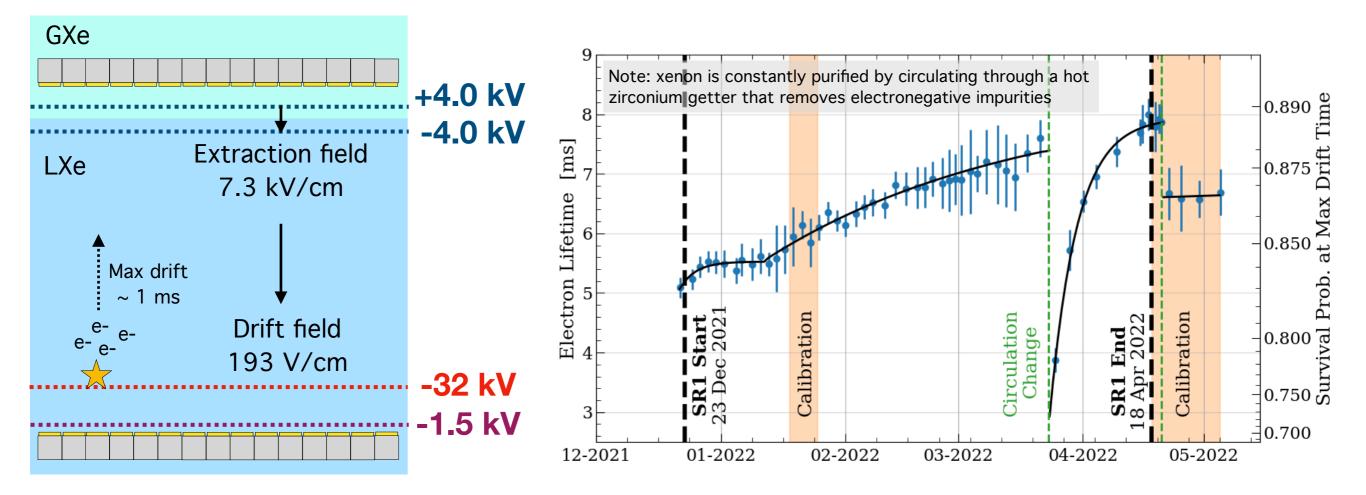


The LZ timeline



LZ Science Run 1

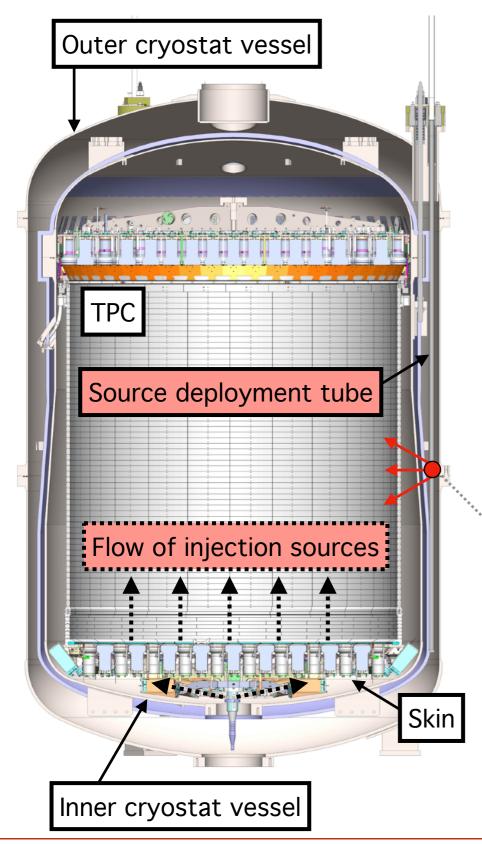
Goal: Demonstrate physics capability of the LZ detector



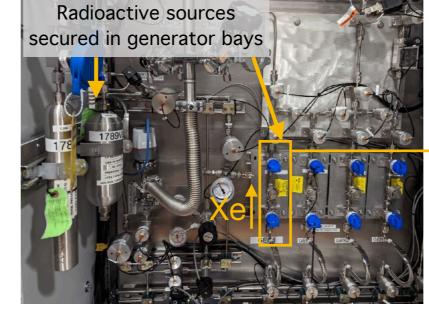
- Data from Dec 23rd '21 to May 12th '22
- Mid-run and post-run calibrations
- WIMP search live time = 60 days
- Engineering run data not blinded

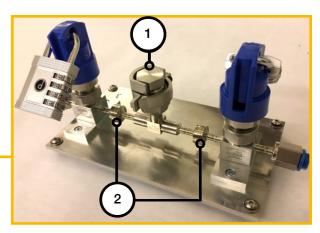
- > 97% of PMTs operational
- Liquid T = 174.1 K (0.02% variation)
- Gas P = 1.791 bar(a) (0.2% variation)
- Liquid level stable within 10 microns
- Gas Circulation ~ 3.3 t/day

Electron recoil calibrations



Injection sources (dispersed into LXe)





 Parent nuclide (producing daughter calibration isotope) enclosed in VCR cap.
 Filter elements for incoming and outgoing xenon flow

Methane tagged with <u>tritium</u>, CH₃T (β ; 18.6 keV endpoint) <u>Kr83m</u> (γ ; 32.1 keV, 9.4 keV) <u>Rn220</u> (γ , β , α ; various energies)

Sealed sources in calibration tubes

- x3 deployment tubes between inner and outer cryostat vessels
- Laser-guided deployment to specific z-positions at 5 mm precision

Threads for securing to deployment system

Typical length ~ 10 cm

Nuclear recoil calibrations

AmLi source

$$\alpha + ^7 \operatorname{Li} \rightarrow n + ^{10} \operatorname{B}$$

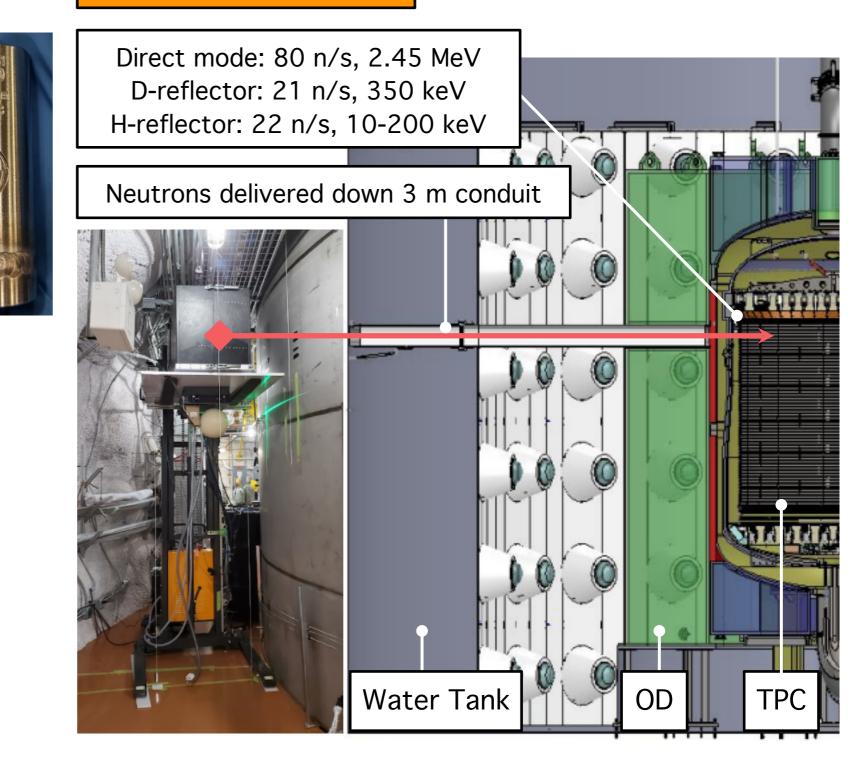
- Three AmLi sources deployed in calibration source tubes.
- Allows for a scan of different detector depths.
- Tungsten enclosure to contain low energy γ-rays.

YBe source

$$\gamma + {}^9 \mathbf{Be} \to n + {}^8 \mathbf{Be} +$$

- Photoneutron source for low energy nuclear recoil calibration at threshold.
- Deployment to top of cryostat vessel (between OD top tanks).
- Demonstrated during commissioning at different fields to the final WIMP-search.

DD Neutron Generator

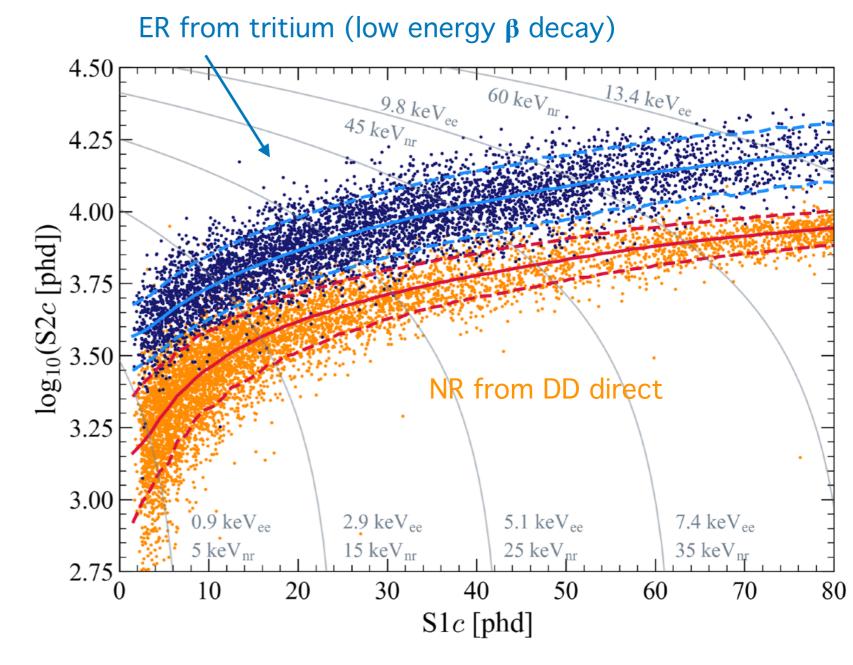


Calibrations

 NEST-based electron recoil model tuned to tritium data, then propagated to nuclear recoil model and verified with DD data.

Detector parameters

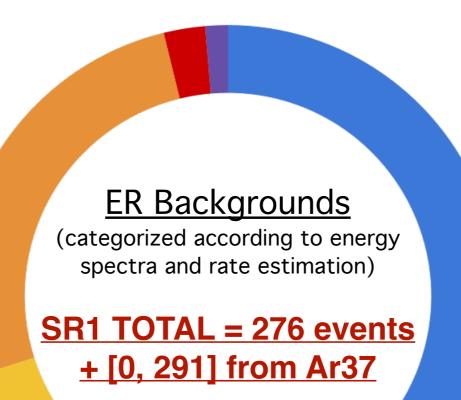
- Light gain, g1 = 0.114 ± 0.002 phd/photon
- Charge gain, g2 = 47.1 ± 1.1 phd/ photon
- Single electron size = 58.5 phd
- ER / NR discrimination = 99.75% for flat NR response



Background model

Xenon contaminants

- Pb214 (Rn222)
- Pb212 (Rn220)
- ► Kr85
- Xe136 (2νββ)
- ► Ar37
- ► Xe127
- Xe124 (double ecapture)



Solar neutrinos (ER) ▶ pp + Be7 + N13

Detector materials (ER)

γ-rays from U238,
 Th232, K40, Co60
 contamination

NR Backgrounds

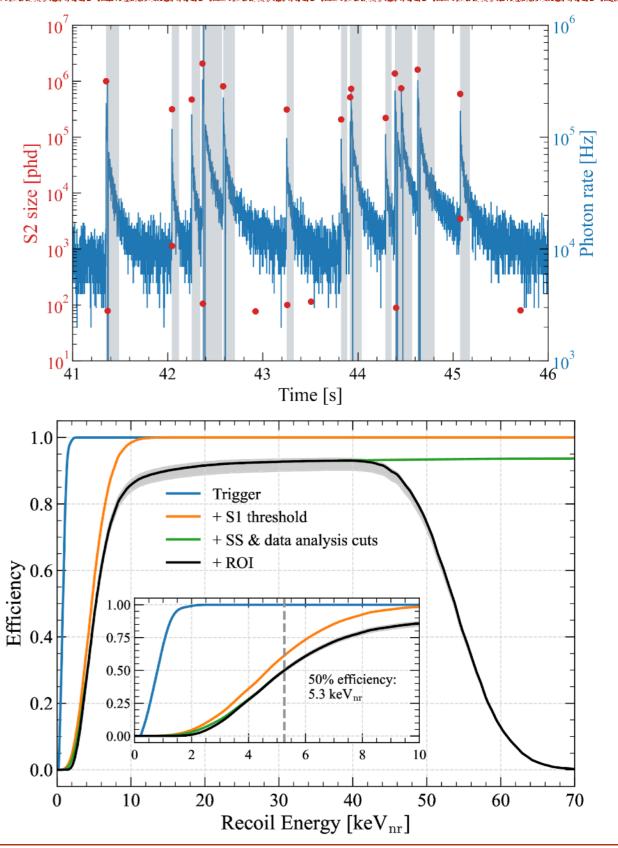
- B8 Solar Neutrinos
- Neutrons from detector materials (α, n) or spontaneous fission.
- SR1 TOTAL = 0.15 events

Accidental coincidences of isolated S1 and S2 pulses effectively eliminated after analysis selections

Data quality selections

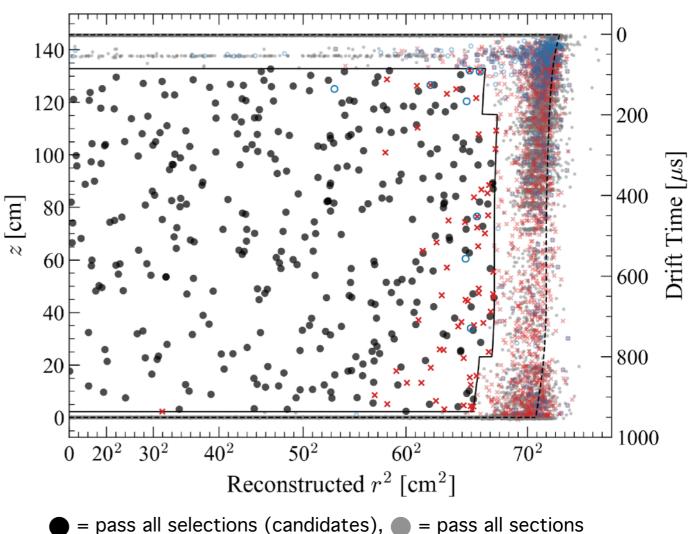
- Two broad categories of data selections allow us to remove data based on bad quality:
 - 1. Time-based:
 - Exclude periods with high rates of spurious activity (e.g. electron and photon emission)
 - 2. Pulse-based:
 - Exclude events based on outlier pulse characteristics
 - Impacts signal acceptance studied using tritium and AmLi data
 - 50% efficiency at 5.3 keV nuclear recoil energy

All cuts developed on calibration data or search data outside the WIMP search region of interest



Fiducial volume and vetoes

- S2 charge-loss close to TPC wall leads to poor position resolution at radial boundary
 - Choose a central <u>fiducial volume</u> simultaneously with S2 threshold to make wall background leakage negligible for this analysis.
 - 5.5 t fiducial mass (measured by uniformly dispersed tritium source)
- Prompt (< 0.5 μs) Skin and OD tag:</p>
 - Reduces naked L-, M-shell Xe127 background by x5 by tagging γ-ray that escapes the TPC



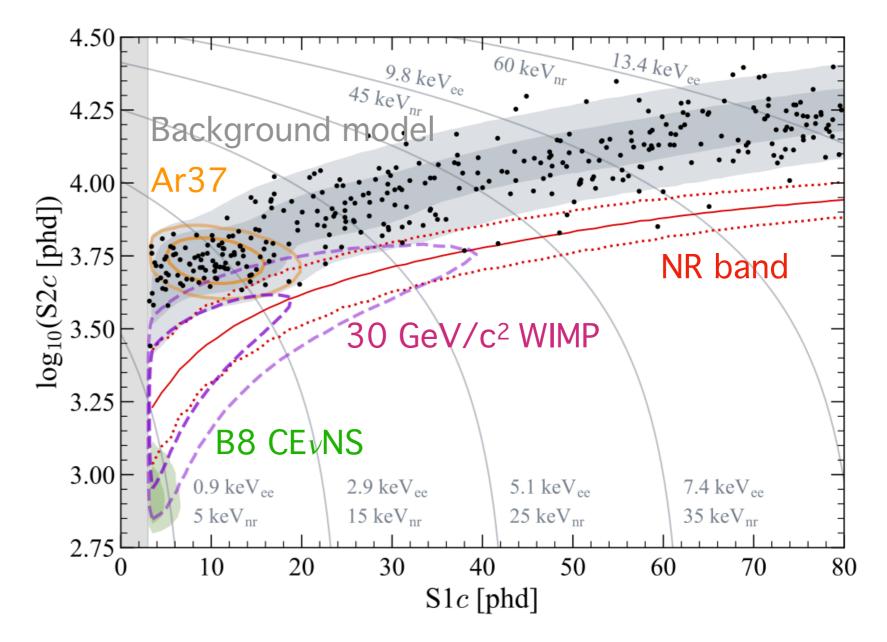
• = pass all selections (candidates), • = pass all sections outside fiducial volume, X = skin tagged, • = OD tagged

- Delayed OD (and skin) tag:
 - 1200 µs window, ~ 200 keV threshold for n-capture tag 5% false veto rate
 - Constraint on neutron background 0^{+0.2} for this analysis

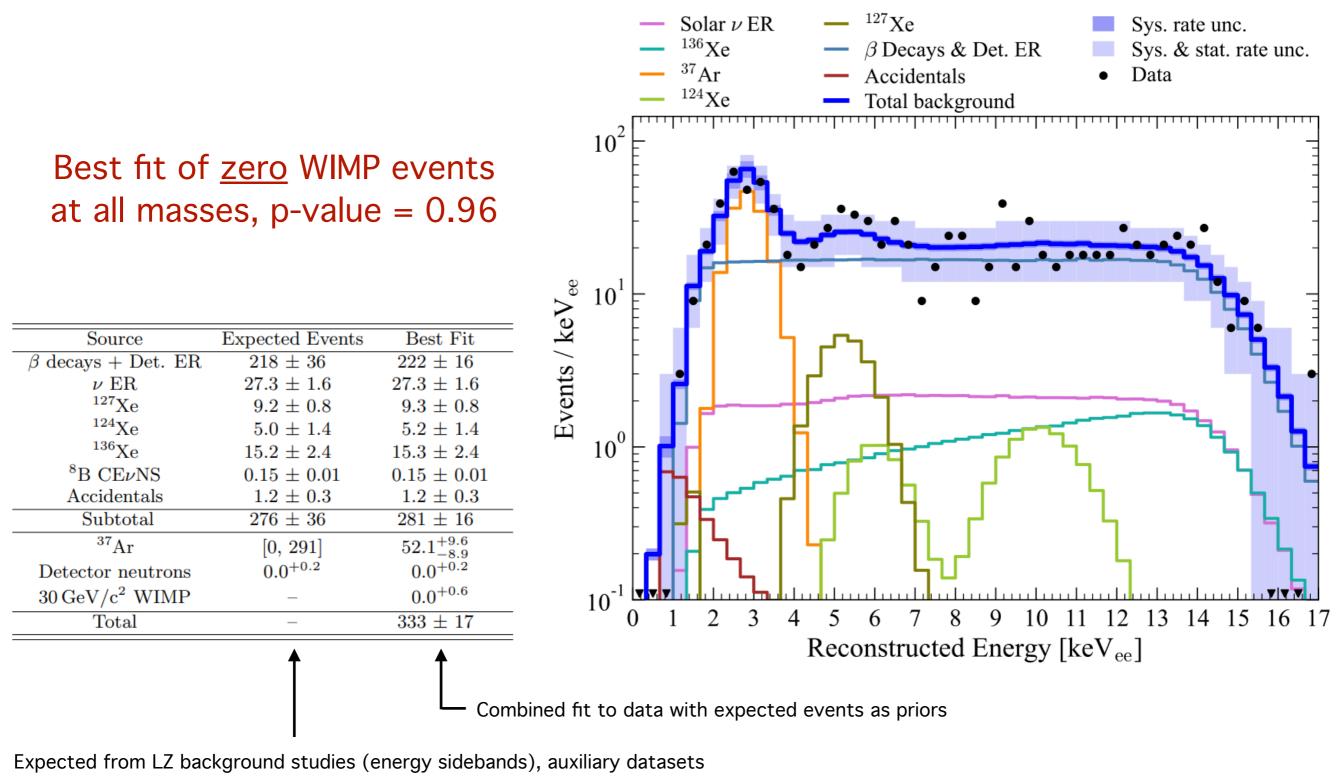
Candidates

60 live days, 5.5 t fiducial volume, 0.9 t years exposure

- ► 335 events in final dataset
- Define a WIMP search 'region-ofinterest' for a Profile Likelihood Ratio (PLR) analysis:
 - ▶ 3 phd < S1c < 80 phd
 - S2 > 600 phd (~ 10 extracted electrons)
 - ▶ S2c < 10⁵ phd



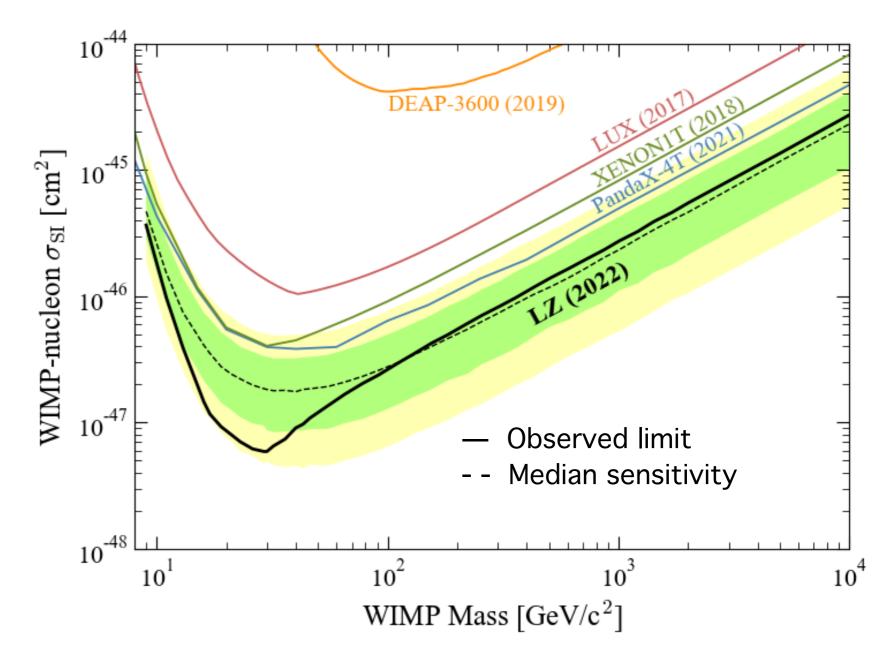
Results - best fits



(e.g. measured half lives, rate predictions from other data or simulations)

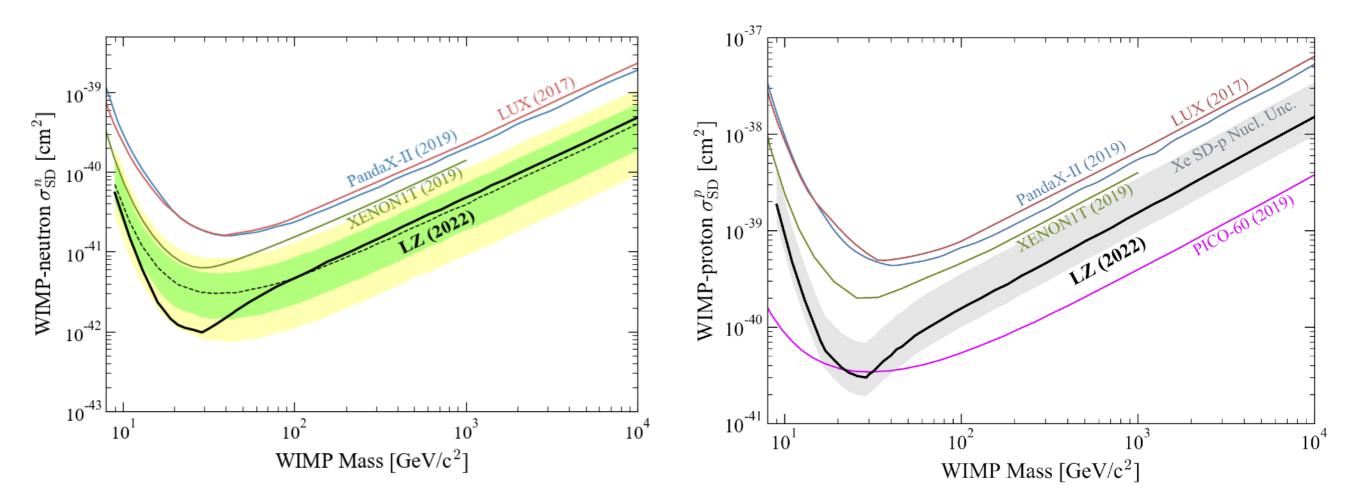
Results - spin independent interactions

- Frequentist, 2-sided PLR test statistic
- Power constrain (*) at π_{crit} =
 0.32 for discovery power
- Best limit of σ_{SI} = 5.9 x 10⁻⁴⁸ at 30 GeV/c²
- Green and yellow are the 1σ and 2σ sensitivity bands.
- Assume a spin independent (scalar) WIMP-nucleon interaction



(*) Following recommendations from community white paper: Eur. Phys. J. C 81, 907 (2021)

Results - spin dependent interactions



- Same statistical treatment as spin-independent case
- Assume a spin dependent WIMP-proton and WIMP-neutron interaction
- Xe has two isotopes with non-zero nuclear spin (both with unpaired neutrons)
 - WIMP-proton sensitivity through higher-order nuclear effects
 - Grey uncertainty band due to theoretical uncertainties on nuclear structure factors. A similar uncertainty applies for all other xenon experiments on this plot (i.e. PandaX-II, LUX, and XENON1T).

Conclusions

- LZ is up and running and taking science data
 - All detectors are performing well and backgrounds are within expectation
- With its first run, LZ has set new limits on WIMP interactions:
 - Paper: <u>arXiv:2207.03764</u> (currently under PRL review)



 LZ is expected to run for 1000 live days, and a broad physics program lies ahead - stay tuned!

Next generation - XLZD consortium

- MOU between LZ, XENON and DARWIN collaborations to work toward a G3 xenon observatory.
- First meeting 27-29 June 2022 at Karlsruhe Institute of Technology.
- See <u>https://xlzd.org</u> and white paper (<u>arXiv:2203.02309</u>).



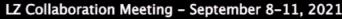
Thank you for listening

35 Institutions: 250 scientists, engineers, and technical staff

- **Black Hills State University**
- **Brandeis University**
- **Brookhaven National Laboratory**
- **Brown University**
- **Center for Underground Physics**
- **Edinburgh University** •
- Fermi National Accelerator Lab.
- Imperial College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- **Northwestern University**
- Pennsylvania State University
- **Royal Holloway University of London**
- **SLAC National Accelerator Lab.**
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
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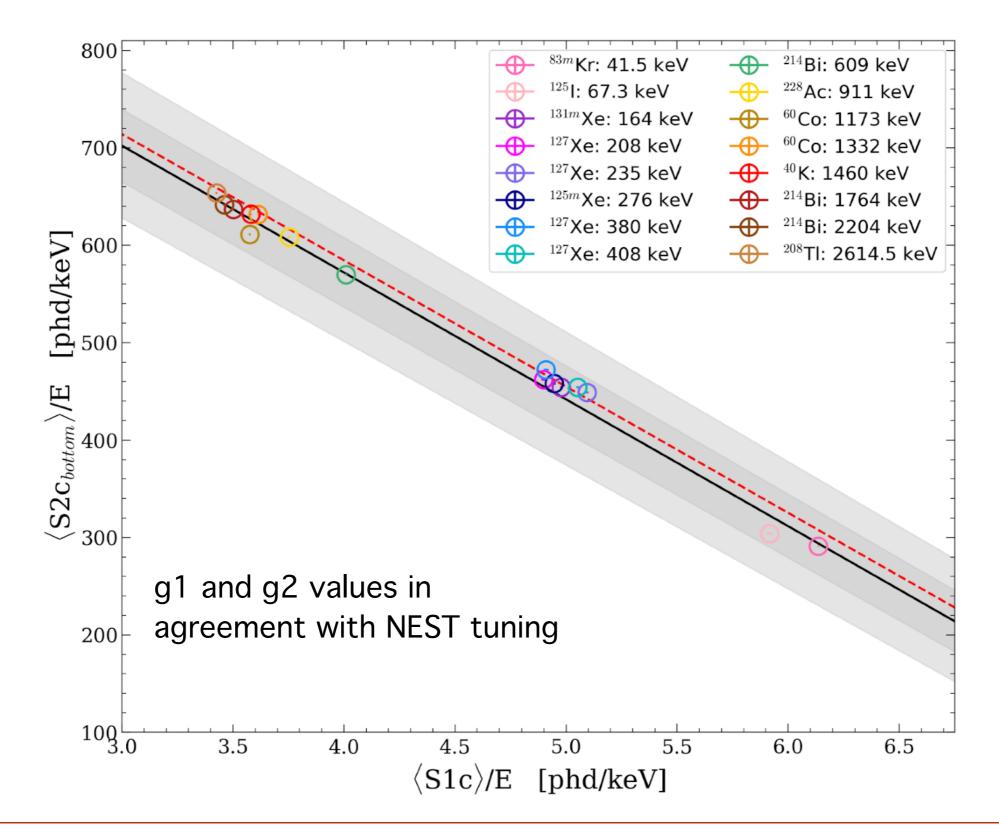




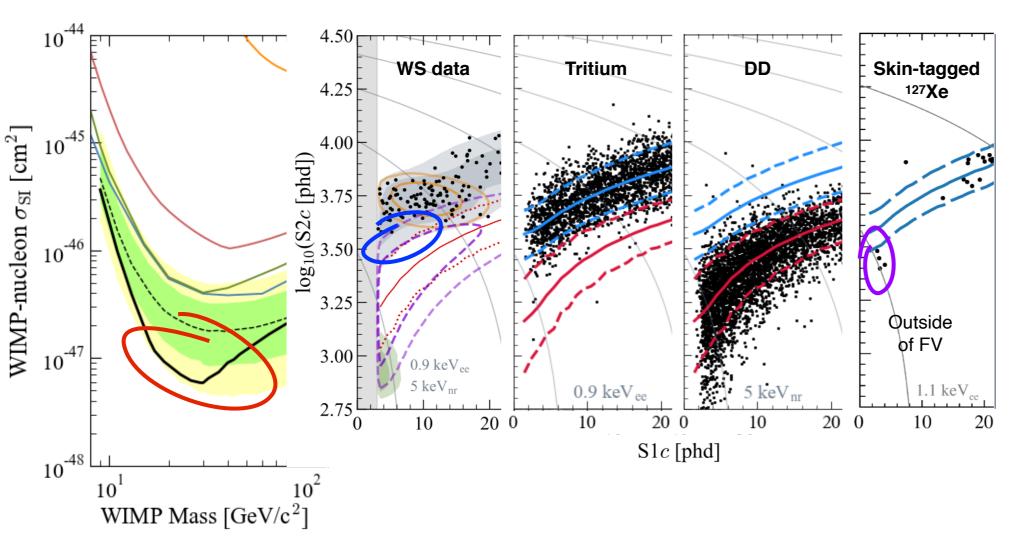


Backup slides

LZ Doke Plot



Downward fluctuation of the limit



Bare M-shell decays of 127Xe populate near deficit region. Observed rate of M-shell decays with coincident γ -ray tagged by the skin is consistent with expectation, given signal efficiencies.

Deficit appears consistent with underfluctuation of background.

↑

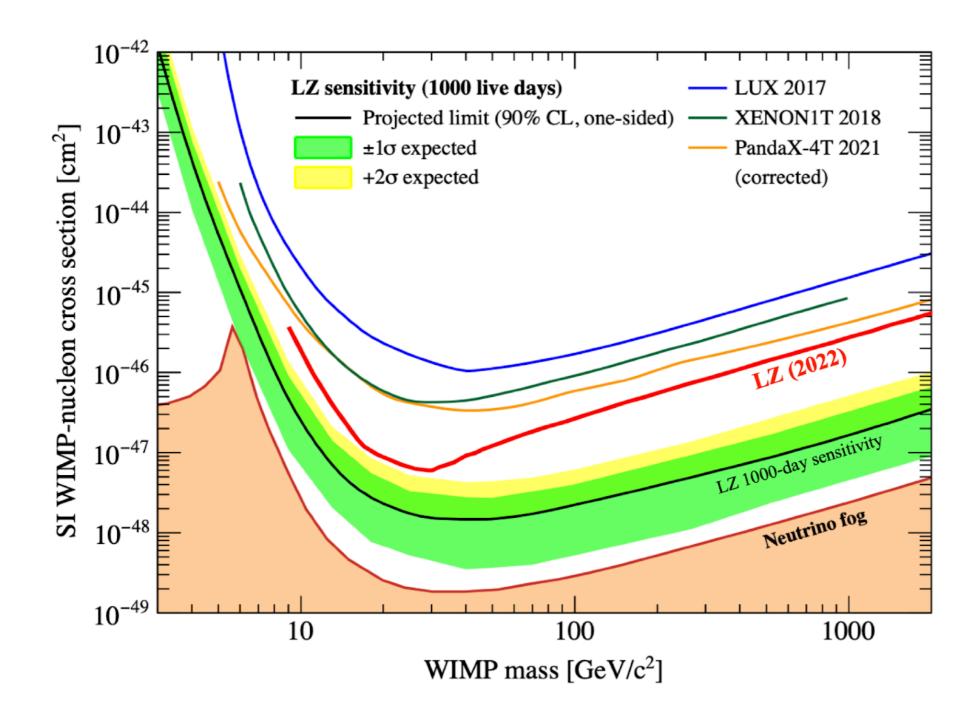
Downward fluctuation in the observed upper limit near 30 GeV/c2 is a result of the <u>deficit</u> of events under the 37Ar population.

Due to background under-fluctuation or unaccounted for signal inefficiency? Probe the latter. Tritium data analyzed identically to WS data. Deficit region is well-covered.

DD data also shows deficit region is wellcovered.

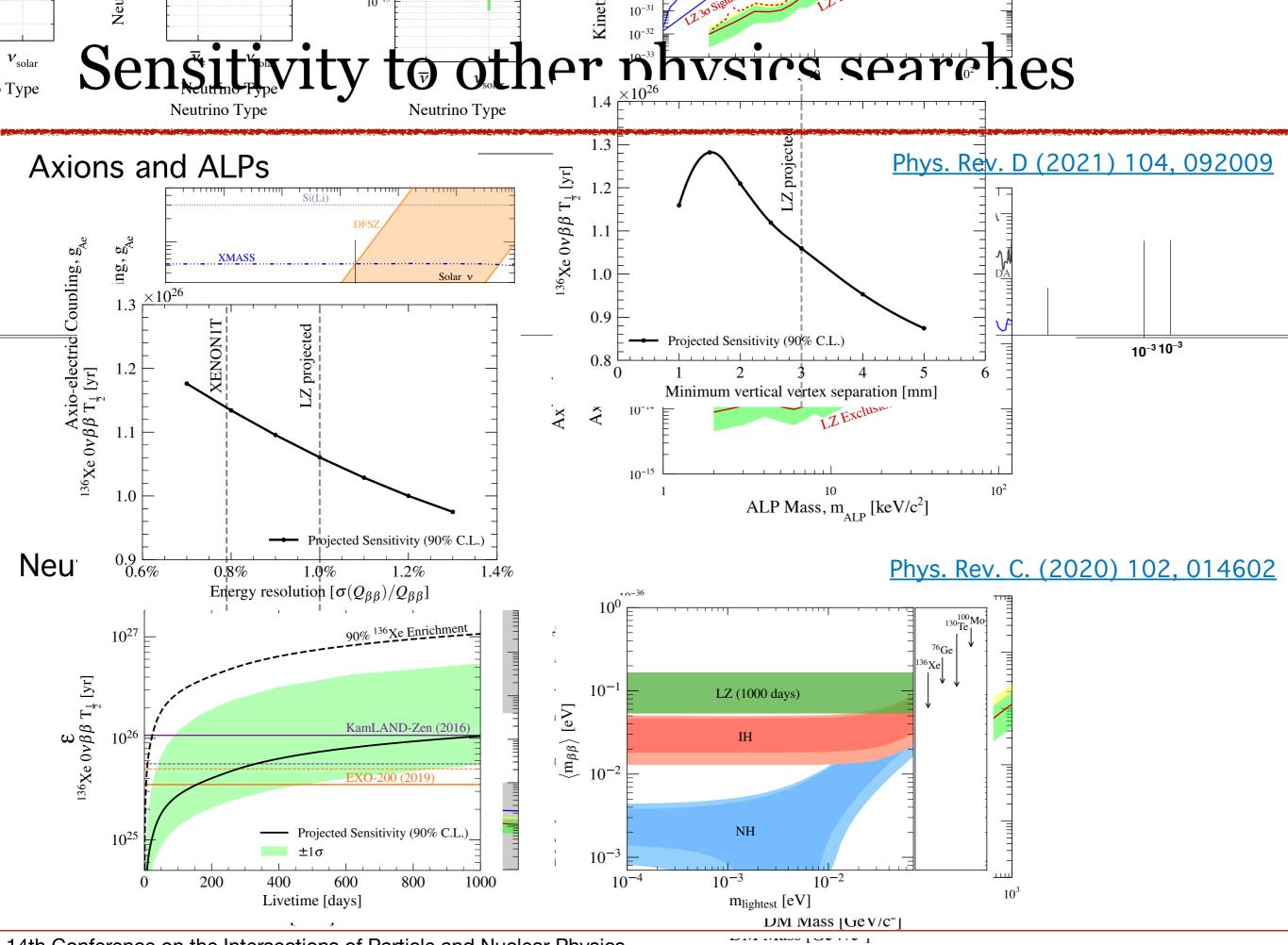
(Not shown here) AmLi neutron calibration data also shows deficit region well-covered.

LZ projected 1000 day WIMP sensitivity

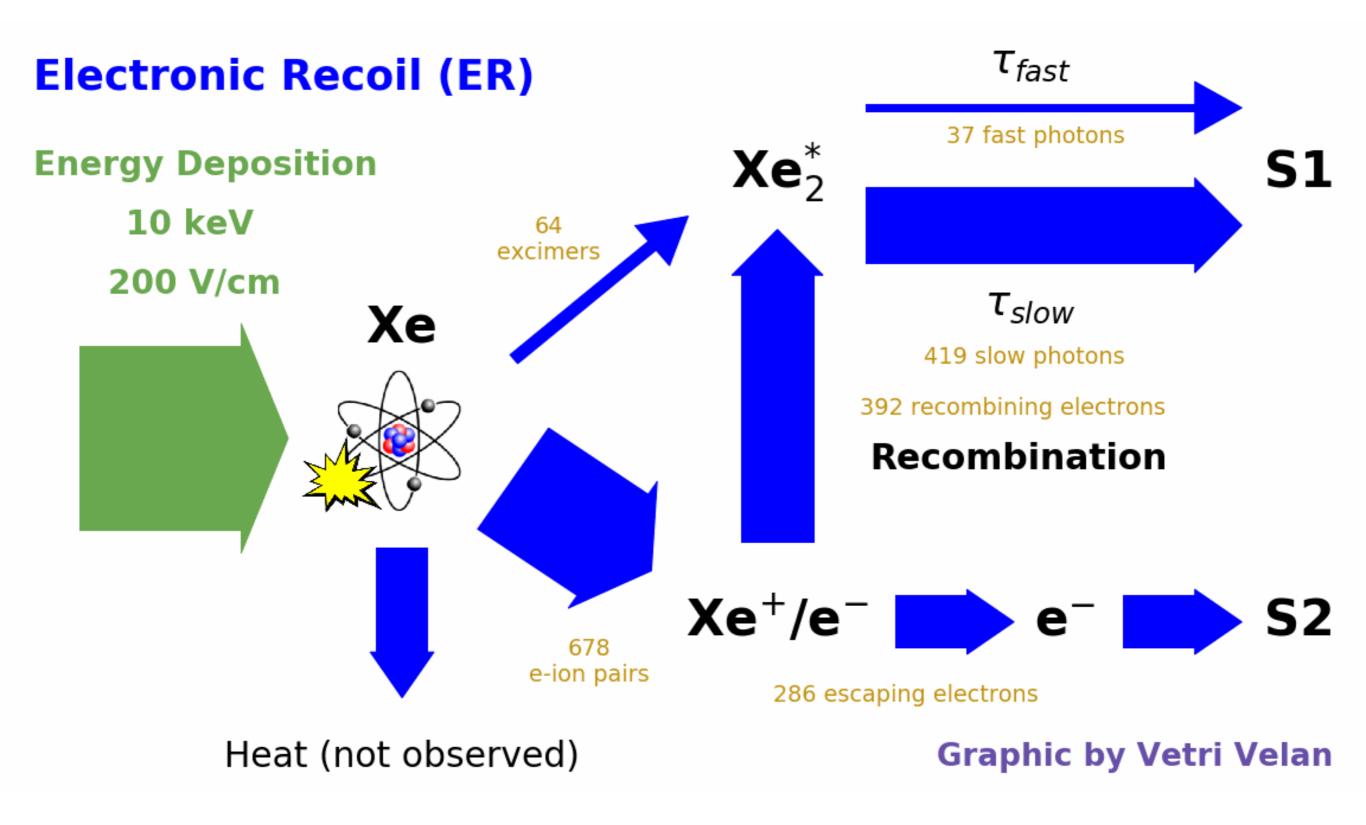


90% CL minimum (one sided) of 1.4 x 10⁻⁴⁸ cm² at 40 GeV/c² from <u>Phys. Rev. D 101</u> (2020), 052002

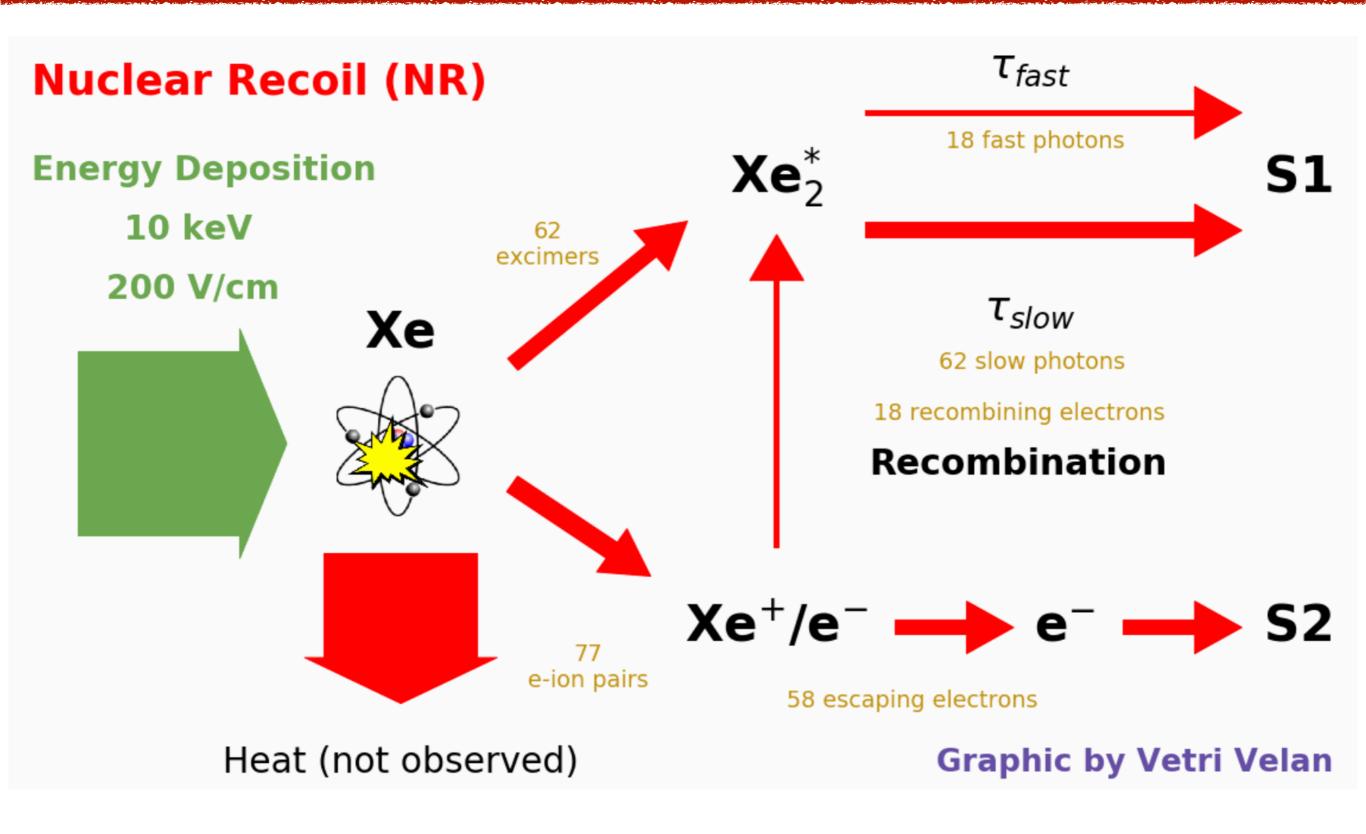
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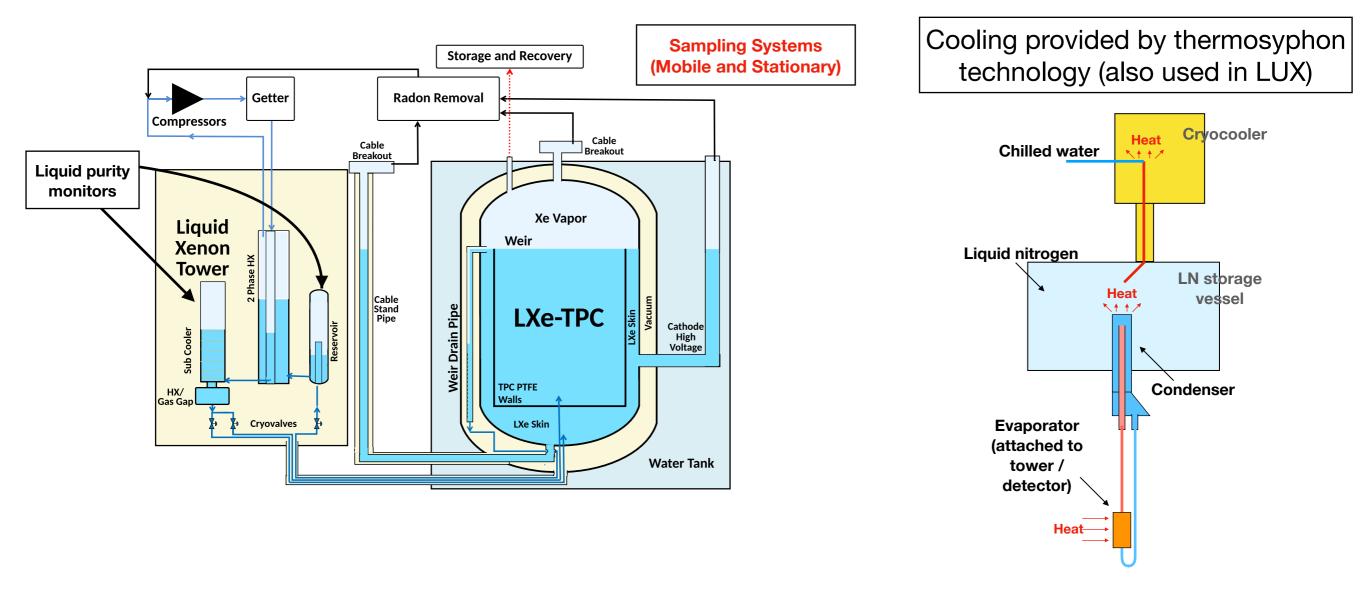
Xenon microphysics



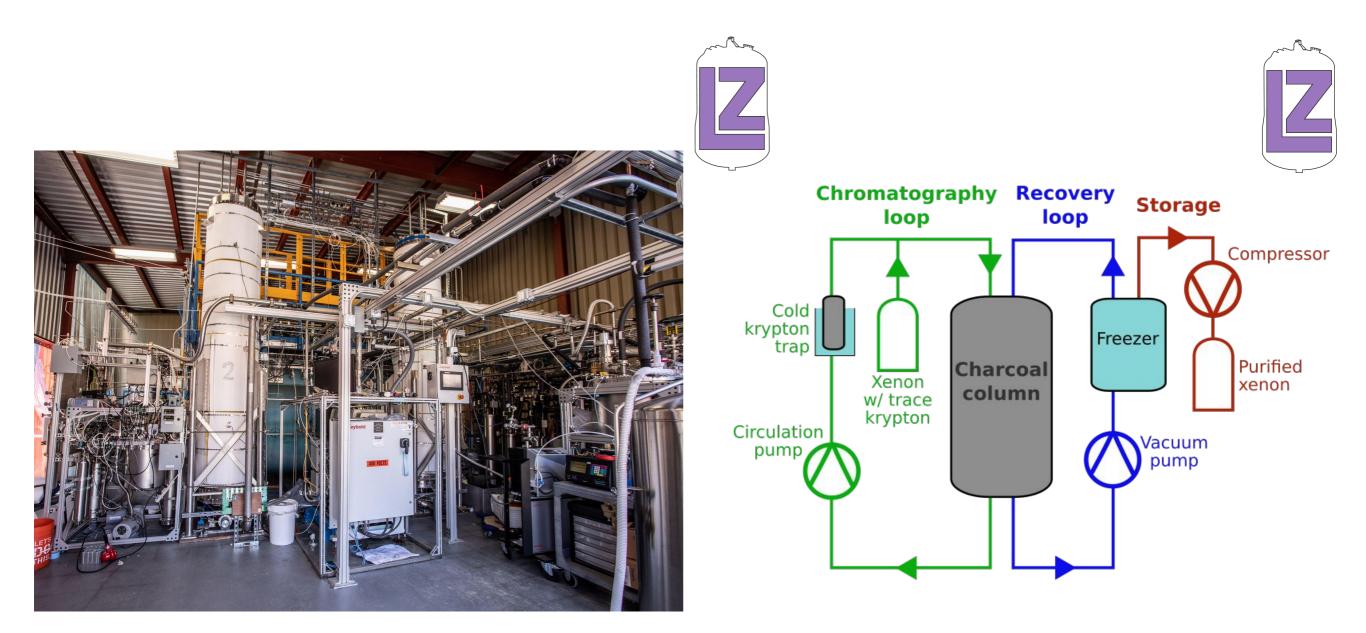
Xenon microphysics



LZ circulation system



LZ Krypton removal (gas chromatography)



- Yenon purified prior to being added to LZ.
 - Concentration reduced from 1-10 ppb (g/g) to < 300 ppq (g/g).</p>
 - Naked beta-decay Kr85 no longer a limiting background