

Searches for neutron oscillations with HIBEAM and NNBAR

ESS

Future NNbar beamline

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European Spallation
Source, Lund, Sweden

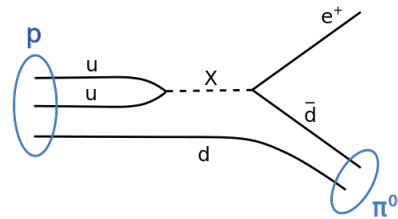
Violation of Baryon Number B

- B, L are accidental symmetries; $B - L$ conserved in the SM
- Global symmetry; BNV not forbidden
- Broadly, SM extensions to explain BAU contain BNV
 - GUT – B necessarily violated (proton decay)
 - RPV SUSY, extra dimensions
 - L-R unification theories, $n \rightarrow \bar{n}$ symbiosis with $0\nu\beta\beta$ ($\Delta L = 0$)
 - Post-sphaleron baryogenesis
 - Co-baryogenesis scenarios, mirror matter
- Not if, but at what scale is BNV?

Complementary Approaches

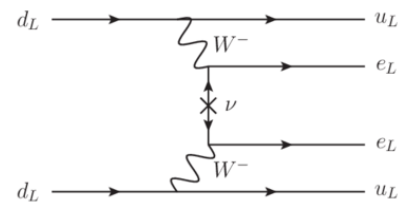
$$p \rightarrow e^+ + \pi^0$$

$$\Delta B \neq 0, \Delta L \neq 0$$



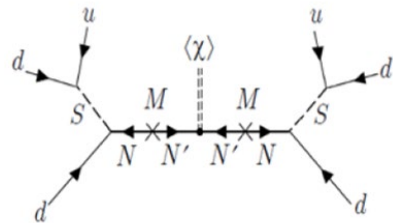
$$0\nu\beta\beta$$

$$\Delta B = 0, \Delta L \neq 0$$



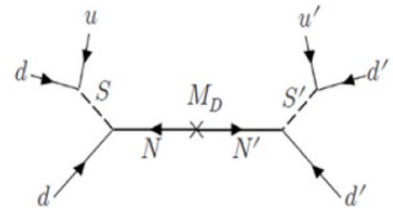
$$n \rightarrow \bar{n}$$

$$\Delta B = 2, \Delta L = 0$$



$$n \rightarrow n'$$

$$\Delta B = 1, \Delta L = 0$$



- p decay, $0\nu\beta\beta$, n conversions probe different selection rules

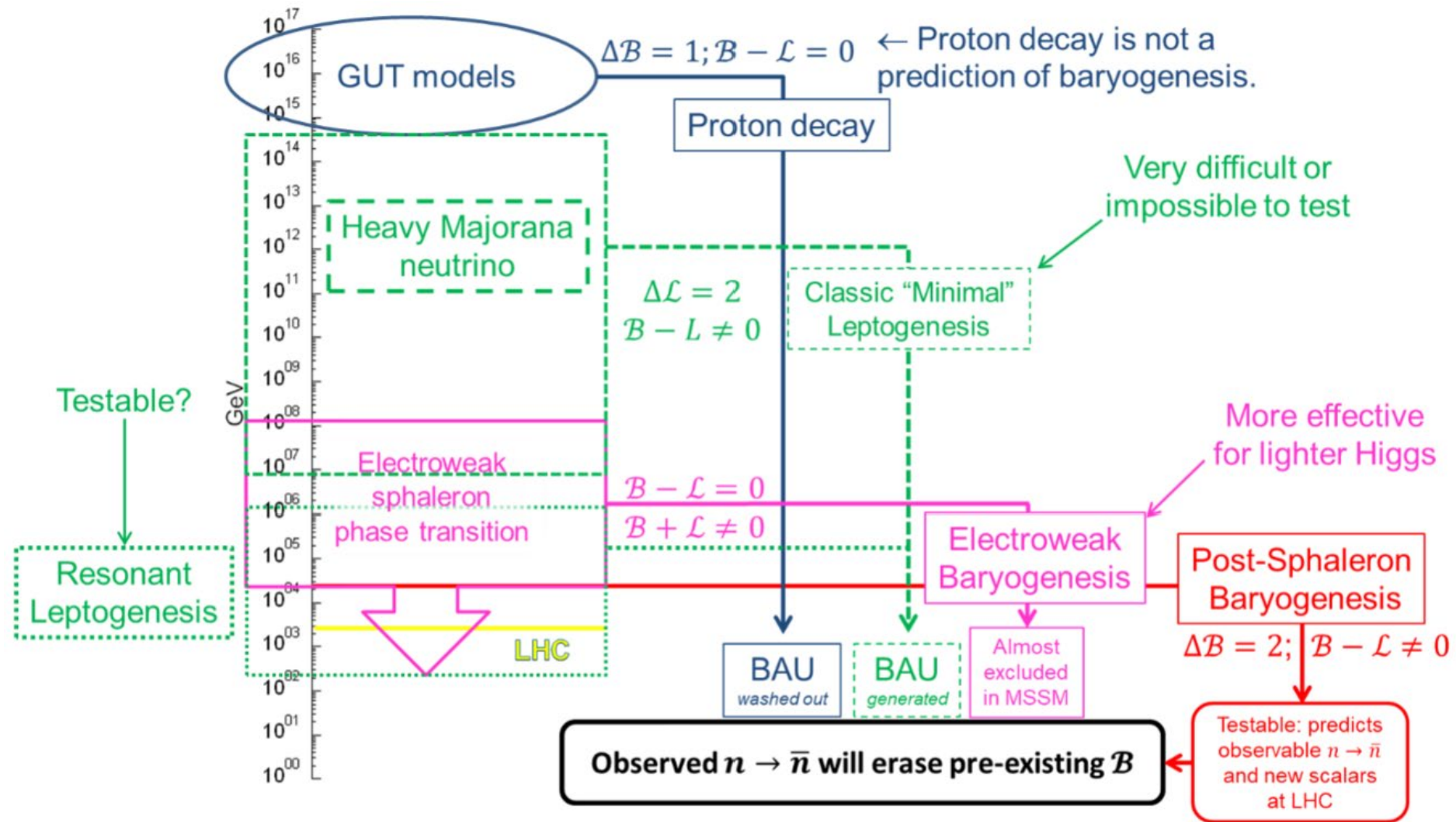
$$\Delta B \neq 0, \Delta L \neq 0, \Delta[B - L] = 0$$

$$\Delta B = 0, \Delta L \neq 0, \Delta[B - L] \neq 0$$

$$\Delta B \neq 0, \Delta L = 0, \Delta[B - L] \neq 0$$

- $n \rightarrow \bar{n}$, dinucleon decays sensitive to BNV-only
- x1000 improvement in $n \rightarrow \bar{n}$ sensitivity on horizon
- Neutral particle can mix with dark matter, $n \rightarrow n'$ and $n \rightarrow n' \rightarrow \bar{n}$

Baryogenesis models



from Y. Kamyshkov and J. Barrow

$n \rightarrow \bar{n}$ oscillations

- 2 state system with potential difference for neutron vs antineutron

- Neutron inside nucleus: ~ 100 MeV

- Free neutron in Earth's magnetic potential $\mu \cdot B$ (10^{-18} MeV)

$$\hat{H} = \begin{pmatrix} m + \vec{\mu}(\vec{B} \cdot \vec{\sigma}) & \varepsilon \\ \varepsilon & m - \vec{\mu}(\vec{B} \cdot \vec{\sigma}) \end{pmatrix} \quad \psi = \begin{pmatrix} n \\ \bar{n} \end{pmatrix}$$

- Strong limits on $\tau_{n \rightarrow \bar{n}}$: Mass splitting from nonzero $V \gg$ off-diagonal term ε

$$\Delta E(1 \text{ nT}) = E_n - E_{\bar{n}} = 2\mu B \sim 10^{-22} \text{ MeV} \quad \varepsilon < 10^{-29} \text{ MeV}$$

- Need “Quasi-free limit”: Uncertainty principle $\Delta E \Delta t \ll \hbar$

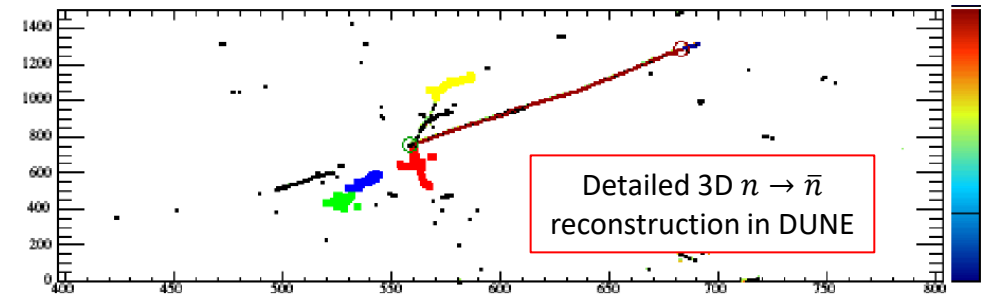
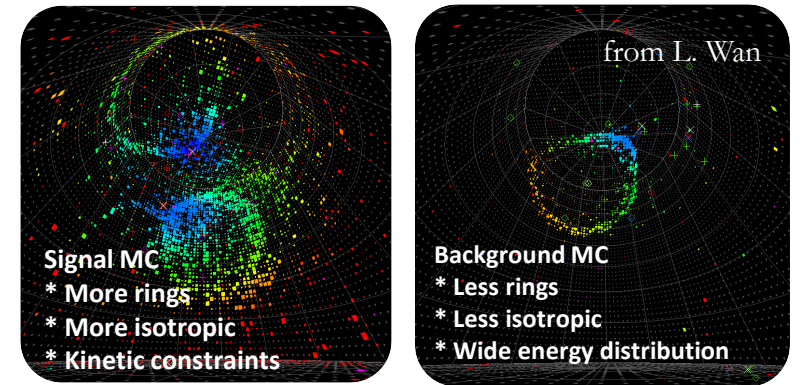
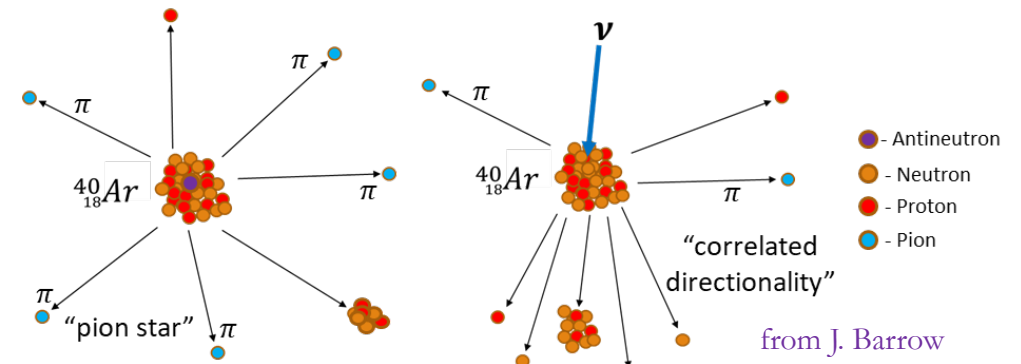
- Free neutron TOF ~ 0.1 s $\rightarrow B \sim 10$ nT [NIMA 320 \(1992\) 569](#)

- Bound neutrons “free” inside nucleons for $\Delta t \approx \hbar/E_{\text{binding}} \sim 5 \times 10^{-22}$ s (nuclear suppression factor)

- In this limit $P_{n \rightarrow \bar{n}}(t) = \left(\frac{t_{\text{free}}}{\tau_{n \rightarrow \bar{n}}}\right)^2 \frac{1}{\tau_{n \rightarrow \bar{n}}} = \varepsilon$ (Figure of Merit)

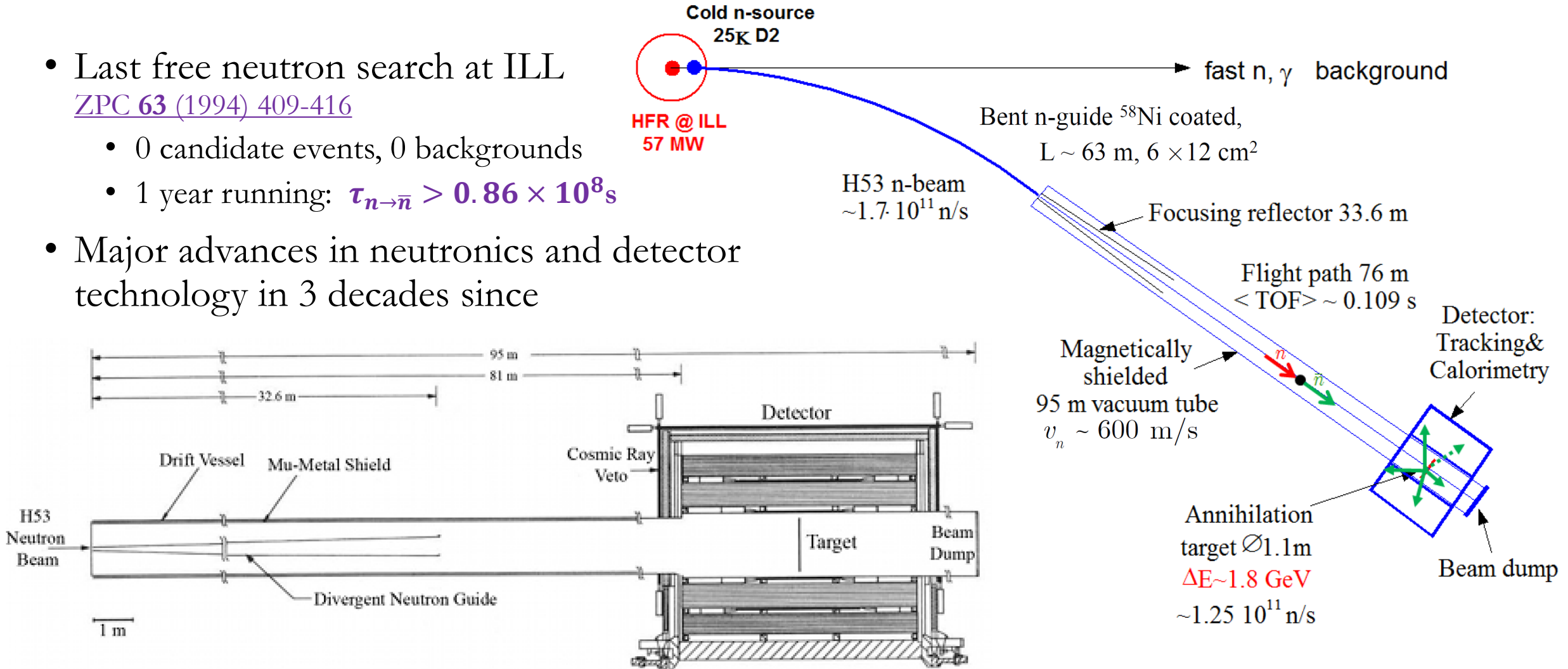
Searches for $n \rightarrow \bar{n}$ in nuclei

- Task: distinguish characteristic \bar{n} signature (“pion star”) from backgrounds such as ν_{atm} background
 - Constraints: topology, 1.88 GeV, total momentum 0, average 5 pions
- Super-K Runs I/II/III/IV [PRD 103 \(2021\) 012008](#)
 - 11 candidate events, 9.3 expected backgrounds
 - 3.6×10^{32} years at 90% C.L, $\tau > 4.7 \times 10^8$ s
- Sensitive searches potentially possible in Hyper-K, NOvA, MicroBooNE
- DUNE: improved background rejection [arXiv:2002.03005](#)
 - Offer lower KE threshold (e.g. protons), higher resolution, bubble chamber-like images, PID & dE/dx
 - Expected reach: $\tau_{n \rightarrow \bar{n}} > 5.53 \times 10^8$ s



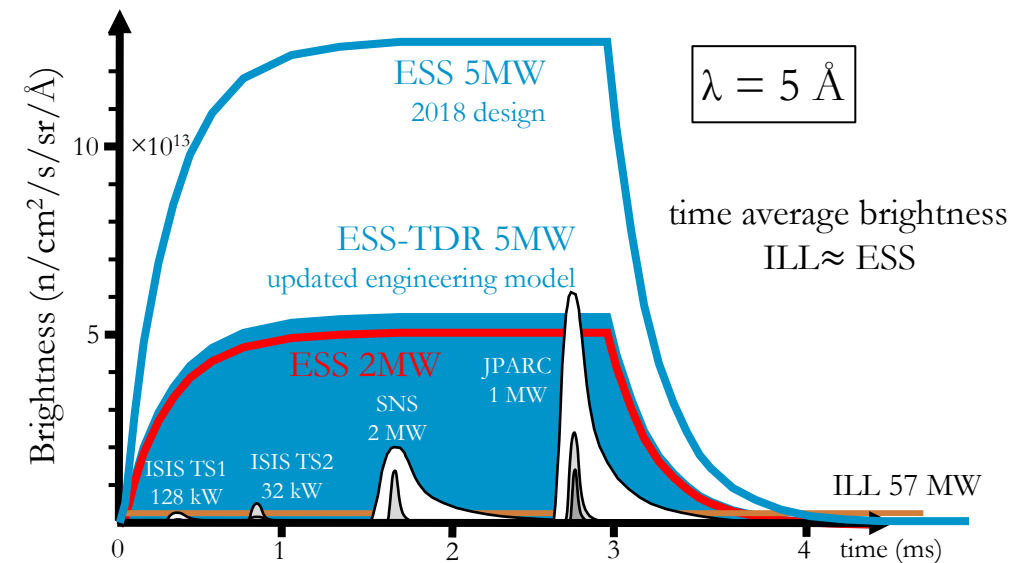
Free $n \rightarrow \bar{n}$ search at ILL

- Last free neutron search at ILL
[ZPC 63 \(1994\) 409-416](#)
 - 0 candidate events, 0 backgrounds
 - 1 year running: $\tau_{n \rightarrow \bar{n}} > 0.86 \times 10^8 \text{ s}$
- Major advances in neutronics and detector technology in 3 decades since



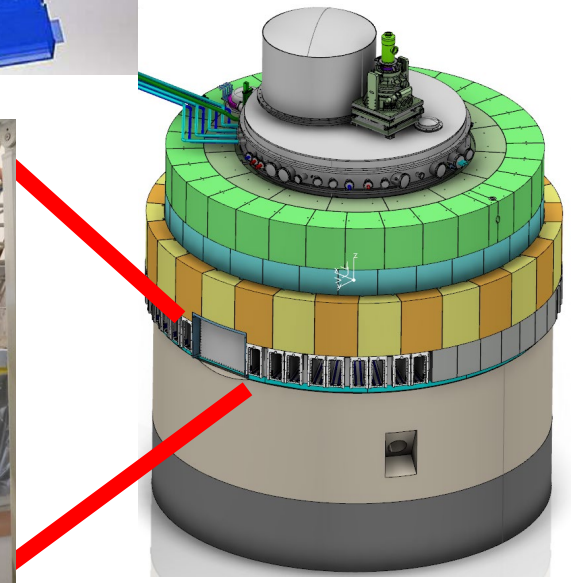
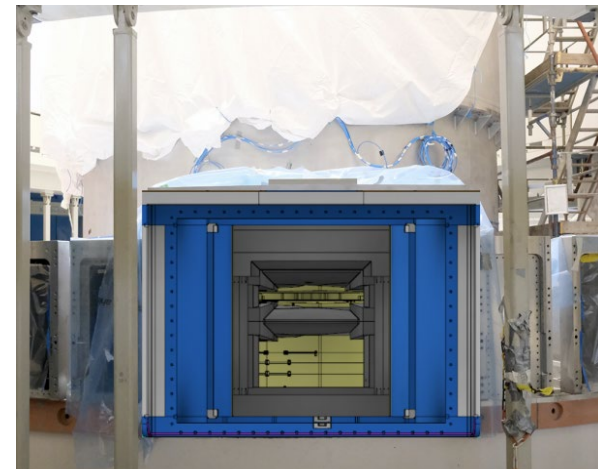
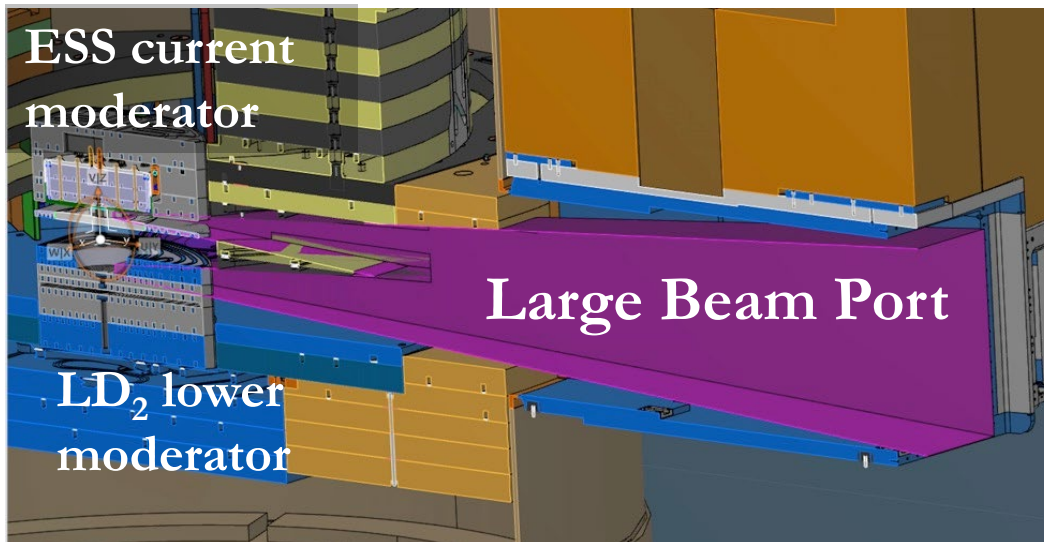
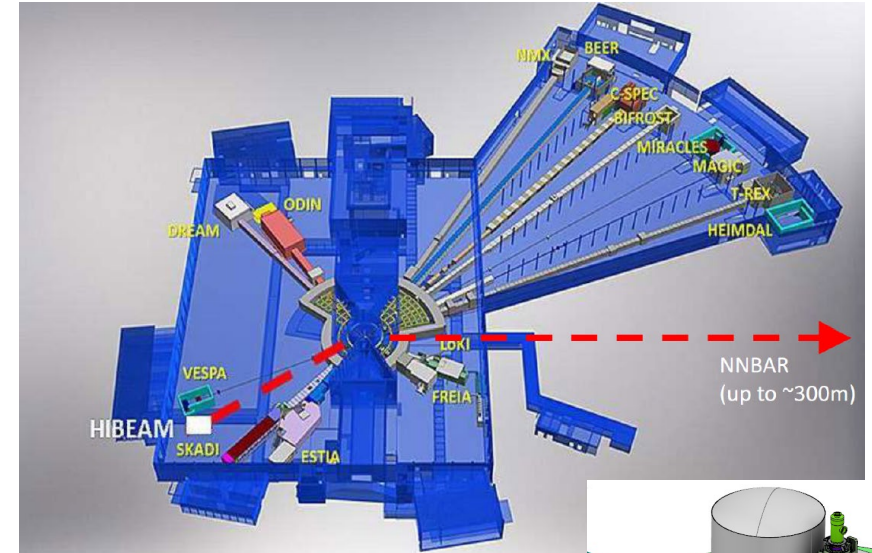
The European Spallation Source

- World's most powerful neutron spallation source, primarily serves neutron scattering community
- 2 GeV protons (3ms long pulse, 14 Hz) hit rotating tungsten target
- 2 MW beam on target ~ 2027, upper moderator, first suite of 15 instruments in user program
- 22 instruments, lower moderator > 2027
- Possibility to upgrade to 5 MW in future project
- [2018 Capability Gap Analysis](#): inclusion of particle physics given highest priority (with 1 other) => particle physics beamline in TDR of 22 instruments



Golden opportunity for $n \rightarrow \bar{n}$ at ESS

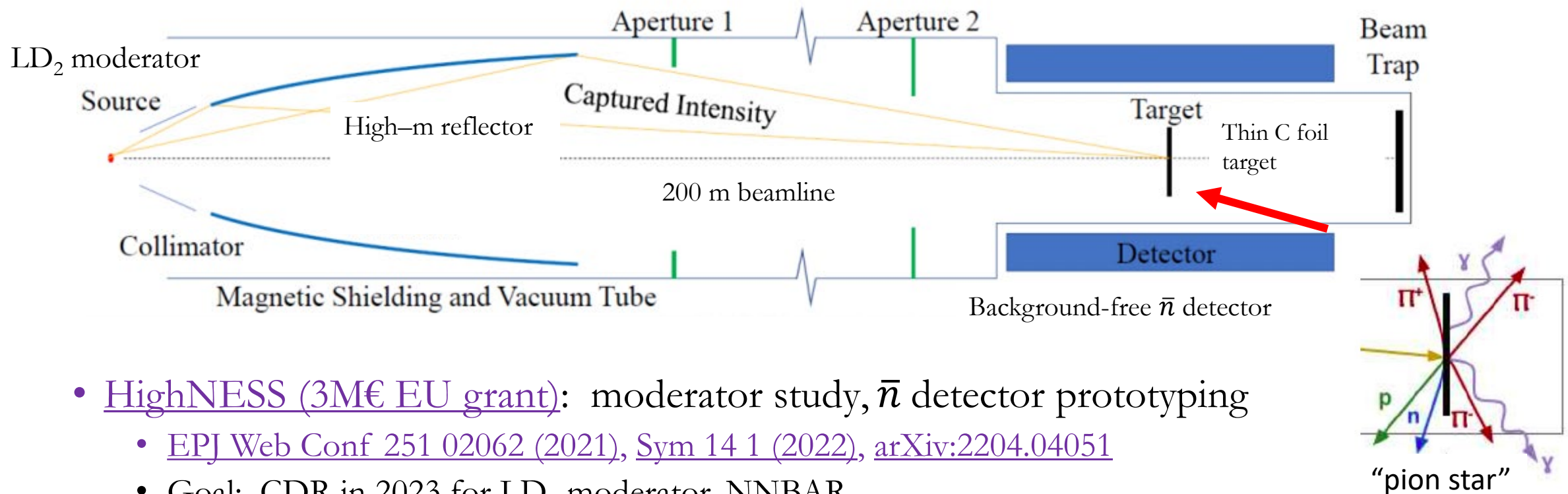
- Substantial investment from ESS with $n \rightarrow \bar{n}$ in mind, to maximize FOM: $N\langle t^2 \rangle$
 - “Large Beam Port” now constructed
 - Up to 300 m beamline
- Fundamental Physics leading design of LD₂ lower moderator [J Phys G 48 070501 \(2021\)](#)



The NNBAR Experiment



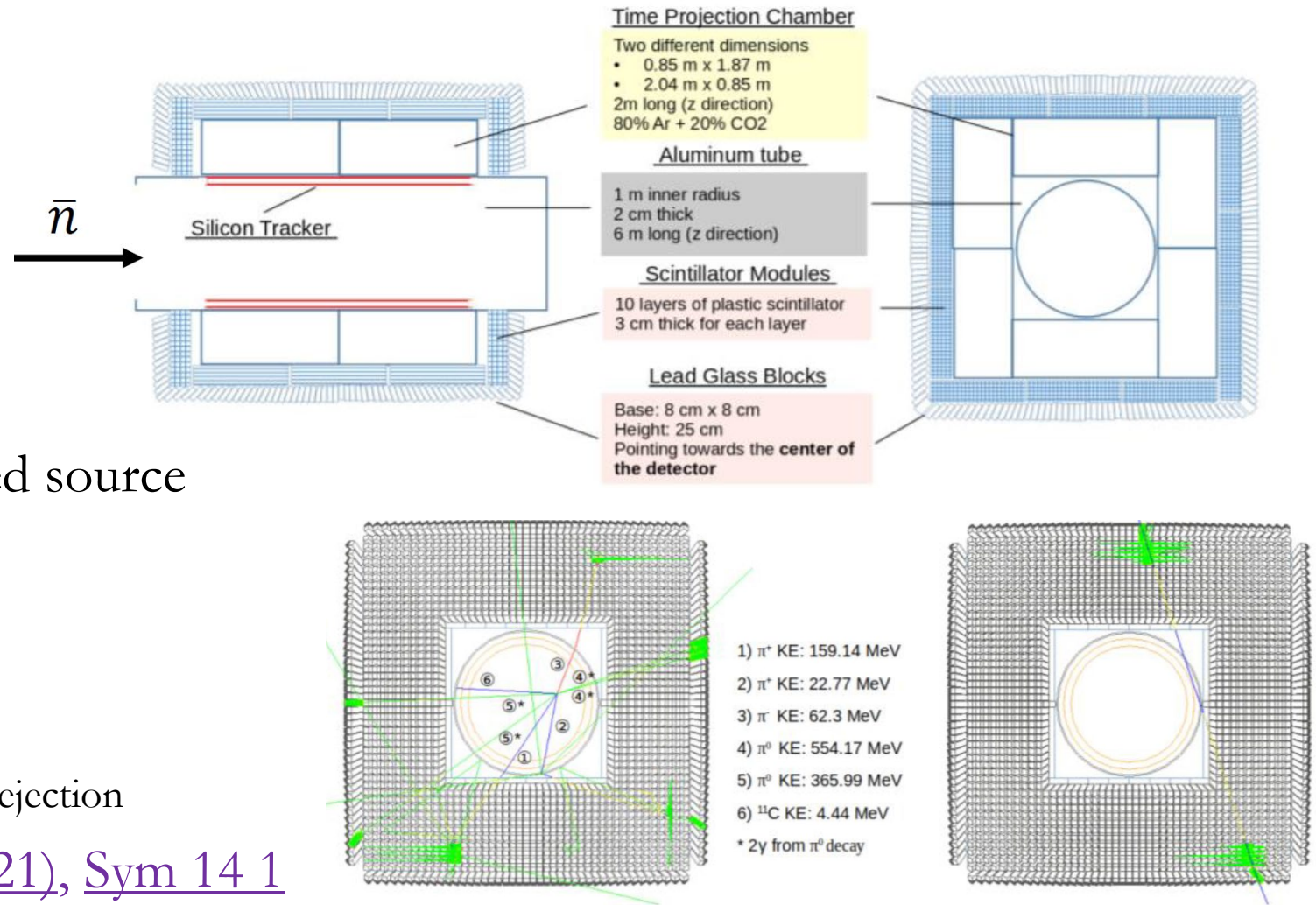
- NNBAR at ESS: Leverage 3 decades of advances: moderator design, neutronics, detection, reconstruction techniques $\times 1000$ sensitivity of ILL [J Phys G 48 070501 \(2021\)](#)



- [HighNESS \(3M€ EU grant\)](#): moderator study, \bar{n} detector prototyping
 - [EPJ Web Conf 251 02062 \(2021\)](#), [Sym 14 1 \(2022\)](#), [arXiv:2204.04051](#)
 - Goal: CDR in 2023 for LD₂ moderator, NNBAR
- Staged R&D program ORNL – HIBEAM – NNBAR

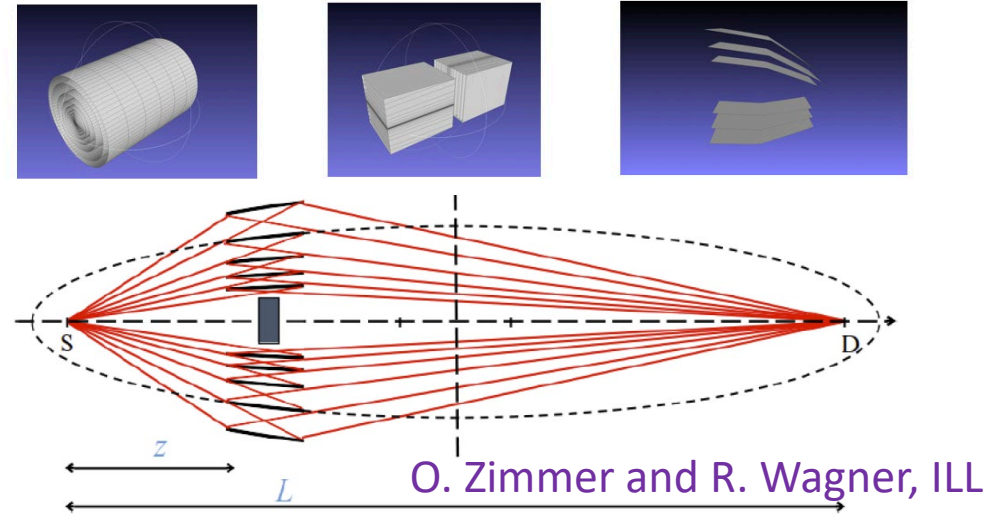
NNBAR Detector

- Detector concept:
 - Silicon inner tracker
 - TPC
 - Scintillator range detector
 - Lead-glass calorimeter
- Strong constraints: 1-2 GeV, 4-7 pions, precise vertex, pulsed source
- Requirements:
 - Reconstruct multi-pion final state
 - Invariant mass reconstruction
 - Particle ID
 - Timing sensitivity for background rejection
- [EPJ Web Conf 251 02062 \(2021\)](#), [Sym 14 1 \(2022\)](#), [arXiv:2107.02147](#)



Factors improvement in NNBAR over ILL

Brightness		≥ 1
Moderator Temperature	Colder neutron <TOF>, quadratic sensitivity	≥ 1
Moderator Area	Large aperture required	2
Angular Acceptance	2D = quadratic sensitivity	40
Length	\propto time, quadratic sensitivity	5
Run Time	ILL run = 1 year	3
Total gain vs ILL	Figure of merit: Nt^2	≥ 1000



- Assumes ESS @ 5 MW, optimized moderator + 2 MW may be similar; 200 m baseline, 50% detection efficiency (as in ILL), no bounces
- Large gains from neutron reflector: supermirror reflectors now commercially available ($m=6$); advanced concepts like nested reflectors [JNR 30 \(2018\) 91](#)
- $\times 1000$ increase in sensitivity, reaching $\tau \sim 2-3 \times 10^9$ s
- Multiple reflections? [PLB 795 \(2019\) 362-365](#); [PRL 122 \(2019\) 221802](#); [PLB 808 \(2020\) 135636](#)

Road to NNBAR

- R&D, \bar{n} detector prototype
- CDR for LD₂ moderator, HIBEAM/NNBAR
- ORNL $n \rightarrow n'$ program

2022

2024

2028

- HIBEAM high precision $n \rightarrow n'$ program
- Low sensitivity $n \rightarrow \bar{n}$

>2028

>2030

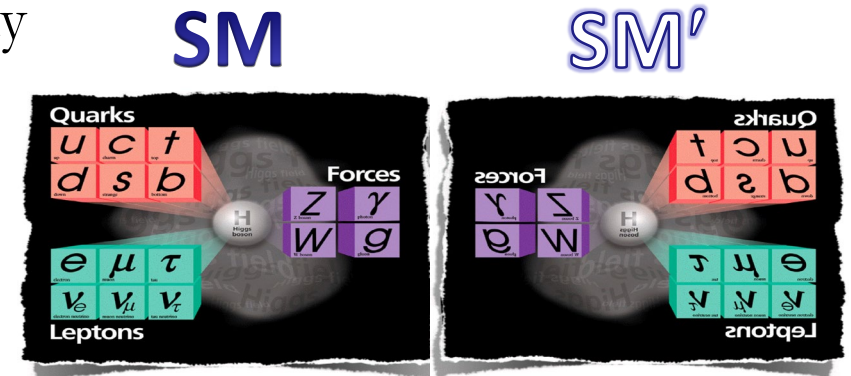
- TDR
- Small scale experiments at ESS Test Beamline

- NNBAR high sensitivity $n \rightarrow \bar{n}$ ($\times 1000$ ILL)

Sterile neutrons

- Mirror matter: identical copy of SM with opposite parity
[[Phys.Usp. 50 \(2007\) 380-389](#), [From Fields to Strings 3 \(2015\) 2147](#)]

- Mirror sector was proposed to restore L-R symmetry
[[Phys.Rev. 104 \(1956\) 254-258](#)]
- No new parameters. Z_2 symmetry
- MM and SM don't interact via known SM interactions except gravity [[Sov.J.Nucl.Phys. 3 \(1966\) 6](#)]

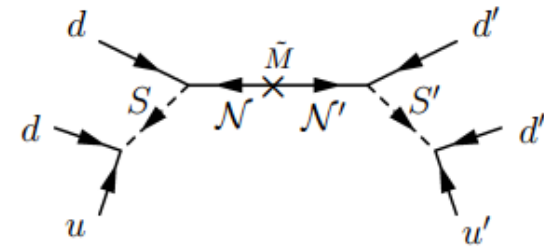


- MM a viable DM candidate [[PLB 503 \(2001\) 362](#), [IJMPA 29 \(2014\) 1430013](#)]
 - Possibly related to [sterile neutrino anomaly](#), [GZK limit](#)
- Predictions of nn' mixing in Mirror Matter models [PRL 96 081801 \(2006\)](#)
 - Apparent BNV: Global $\mathcal{B} = \mathcal{B} + \mathcal{B}'$?
- Recent activity: searches using PSI nEDM apparatus [Sym 14 503 \(2022\)](#), STEREO hidden neutron search [PRL 128 061801 \(2022\)](#), UCN-based search at ILL [Saenz, Moriond-EW \(2022\)](#), SNS regeneration search [PRL 128 212503 \(2022\)](#)

$n \rightarrow n'$ Oscillations

- Similar concept to $n - \bar{n}$: mixing of neutron with sterile twin
- Small mirror magnetic field \mathbf{B}' possible from MM captured by earth

$$\mathcal{H}_{int} = \begin{pmatrix} m + \mu\boldsymbol{\sigma} \cdot \mathbf{B} & \epsilon \\ \epsilon & m' + \mu'\boldsymbol{\sigma} \cdot \mathbf{B}' \end{pmatrix} \quad \text{oscillation time } \tau_{nn'} = \frac{1}{\epsilon}$$



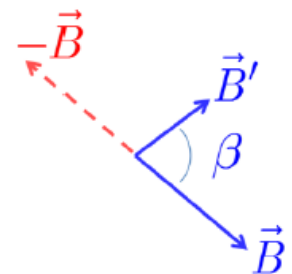
- $n - n'$ mass splitting (10^{-24} GeV) or magnetic field (mG) can strongly suppress oscillation:

- Not sensitive to large $\Delta m_{nn'}$ in laboratory, control \vec{B} for resonance in probability:

$$P(n \rightarrow n') = \frac{\sin^2[(\omega - \omega')t]}{[(\omega - \omega')]^2 2\tau^2} + \frac{\sin^2[(\omega + \omega')t]}{(\omega + \omega')^2 2\tau^2} + \cos \beta \left[\frac{\sin^2[(\omega - \omega')t]}{(\omega - \omega')^2 2\tau^2} - \frac{\sin^2[(\omega + \omega')t]}{(\omega + \omega')^2 2\tau^2} \right]$$

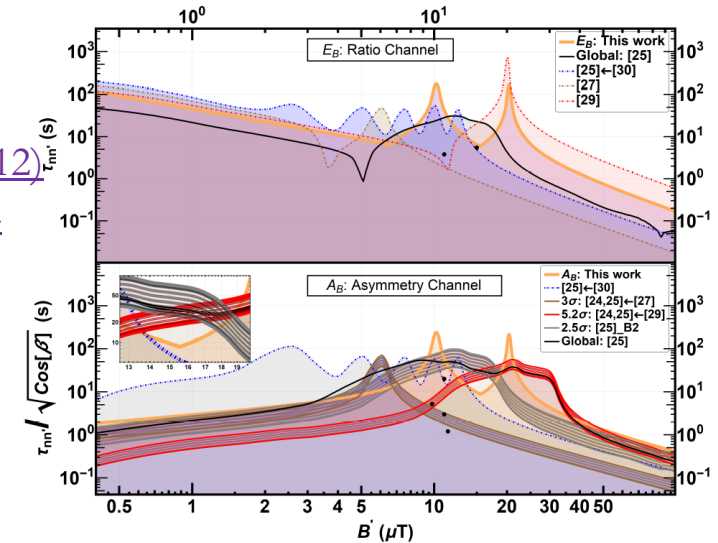
$$\omega = \frac{1}{2} |\mu B|, \quad \omega' = \frac{1}{2} |\mu' B'|, \quad \mu = \mu' \text{ and } \tau = \frac{1}{\epsilon}$$

- Near $B \approx B'$ resonance: $P(n \rightarrow n') \propto \left(\frac{t_{free}}{\tau_{n \rightarrow n'}} \right)^2$. Signal maximum when $\cos \beta = 1$



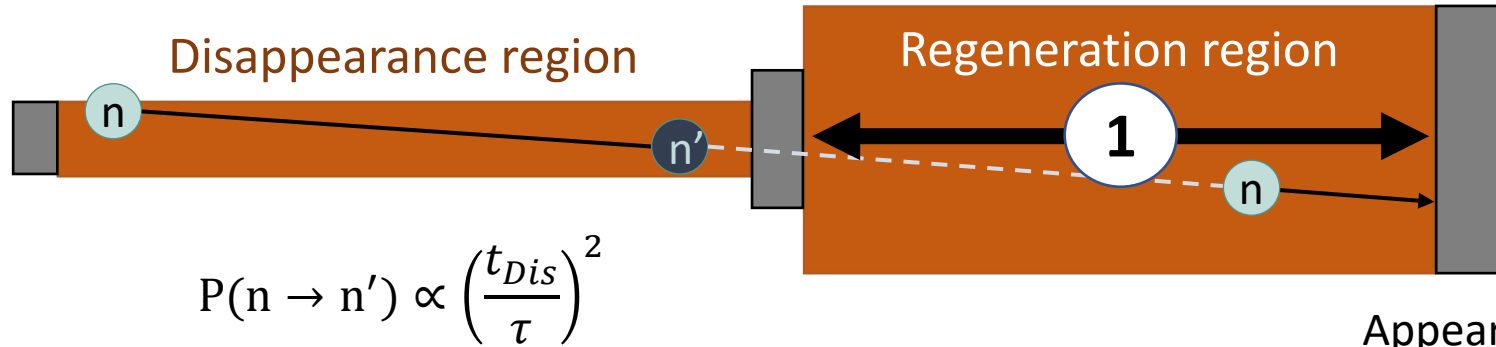
Complementary approaches: UCN vs CN

- Strongest limits from UCN disappearance searches, but some anomalies observed [NIMA 611 \(2008\) 137-140](#), [EPJC 72 \(2012\) 1974](#), [EPJC 78 \(2018\) 717](#), [PRD 80 \(2009\) 032003](#), [PLB 812 \(2021\) 135993](#), [arXiv:2111.02794](#)
 - P Mohanmurthy talk Thursday
- CN intensities much higher; UCNs=1000's bounces
- CN experiments need long, large area beam tubes; UCN bottles much smaller
- CN robust alternative using “regeneration” technique [PRD 96 \(2017\) 035039](#)
- First regeneration search attempt with 6m flight path: $\tau_{nn'} > 2.7$ s [Schmidt 2007](#)

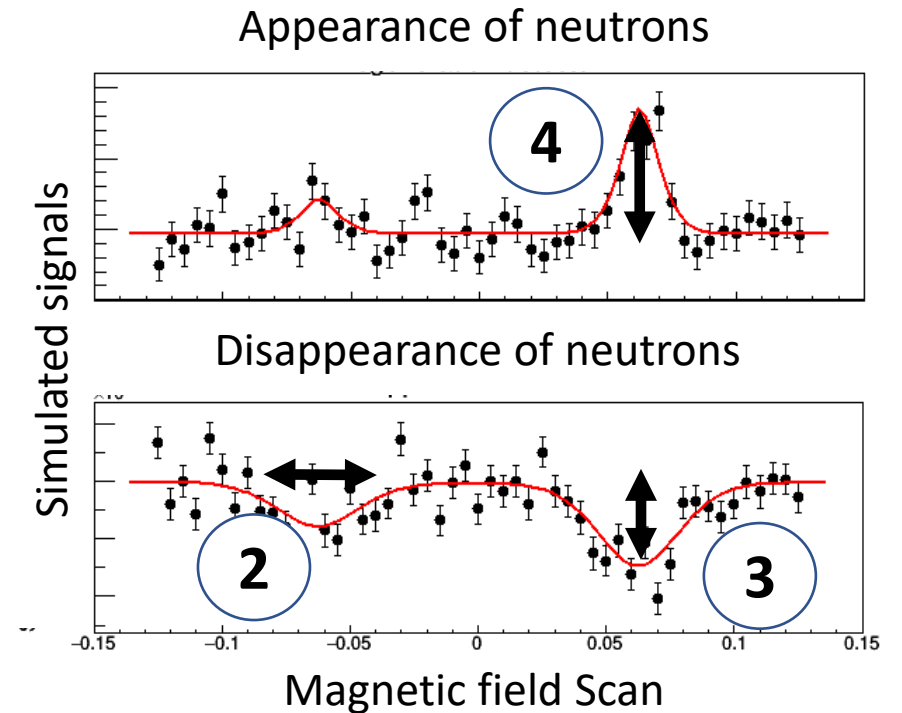


$n \rightarrow n' \rightarrow n$ Disappearance and Regeneration

$$P(n \rightarrow n' \rightarrow n) \propto \left(\frac{t_{Dis}}{\tau}\right)^2 \left(\frac{t_{Reg}}{\tau}\right)^2$$

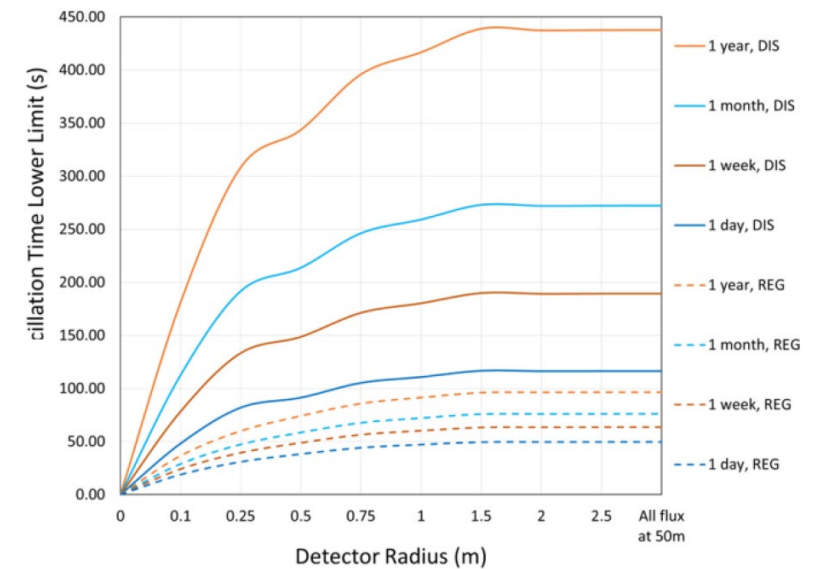
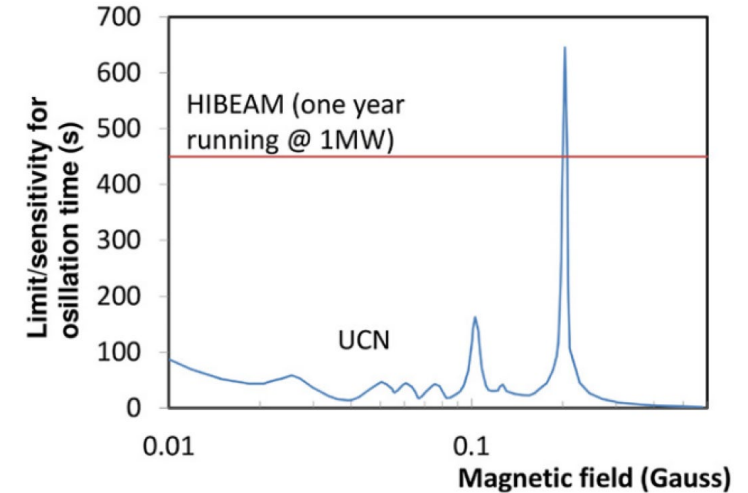
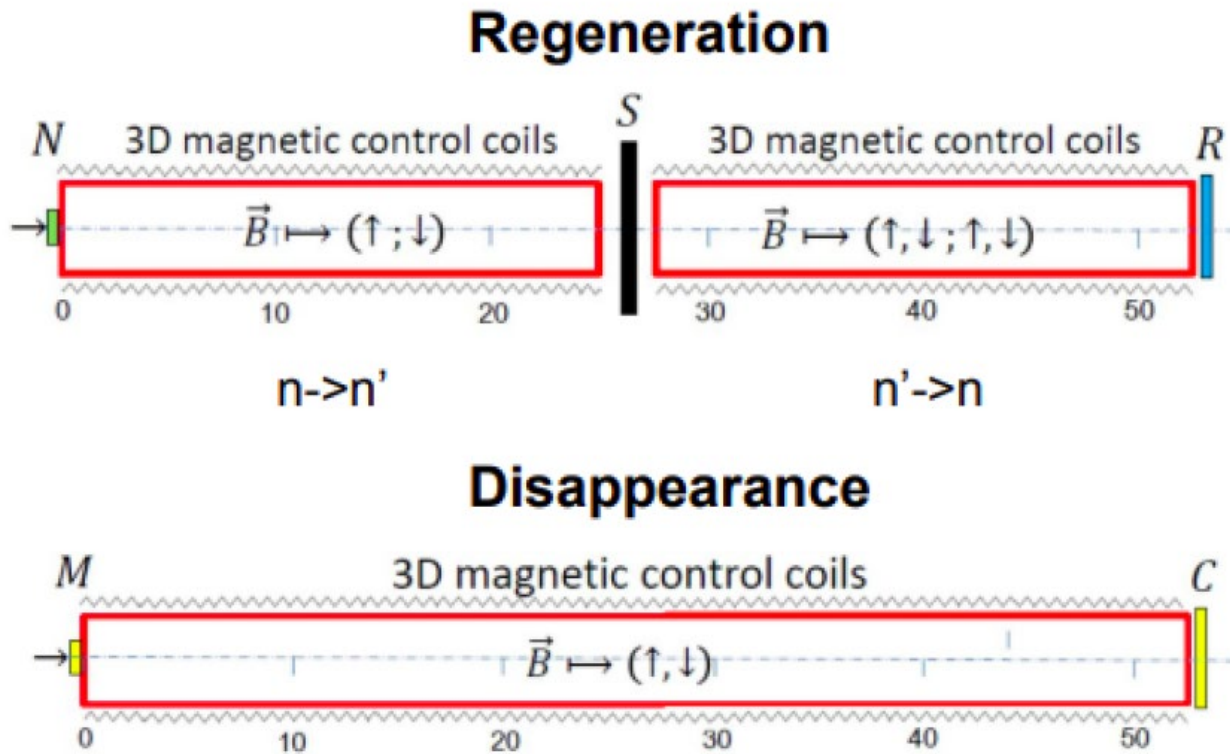


1. High neutron flux + long, large area guides
2. Magnetic field uniformity and control (\sim mG)
3. Precise monitoring of changes in transmission
4. Regeneration: large area, low bkgd detector



HIBEAM sensitivity to $n \rightarrow n'$

- Assume ESS@1MW, 50m, 1 year operation, 1 n/s bkgd



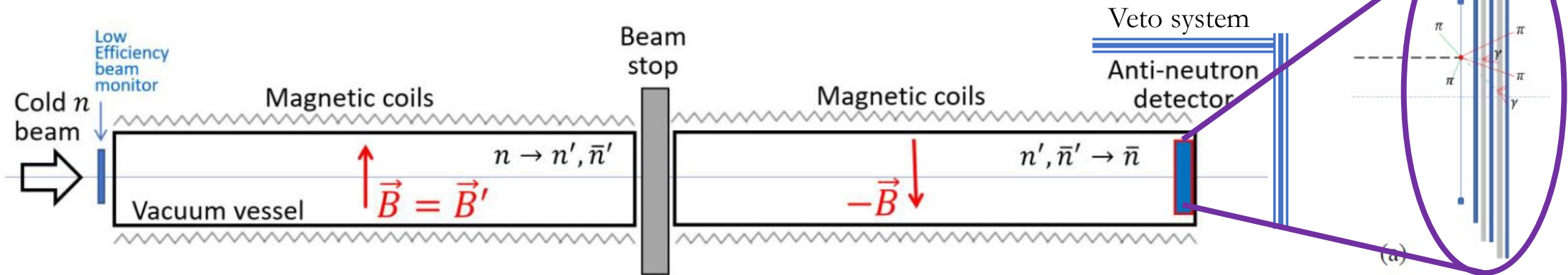
Mirror neutrons and antineutrons

- Straightforward extension of formalism to consider $n \rightarrow \bar{n}, n', \bar{n}'$
- Mirror neutrons provide shortcut for neutron-antineutron oscillations

[EPJC 81 \(2021\) 33](#)

- Connection to cobaryogenesis
[IJMP 33 1844034 \(2018\)](#)
- Shortcut via new term $\delta_{n\bar{n}'}$

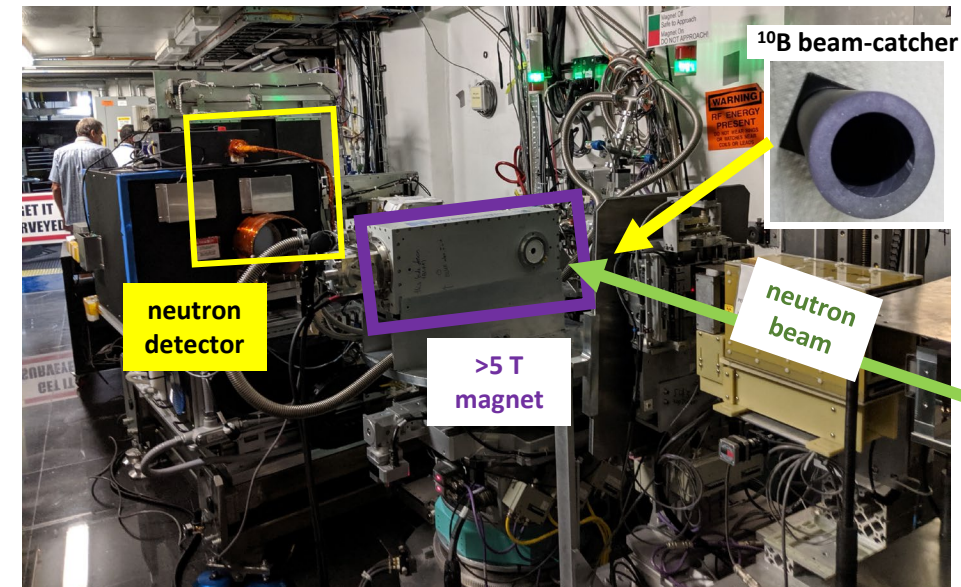
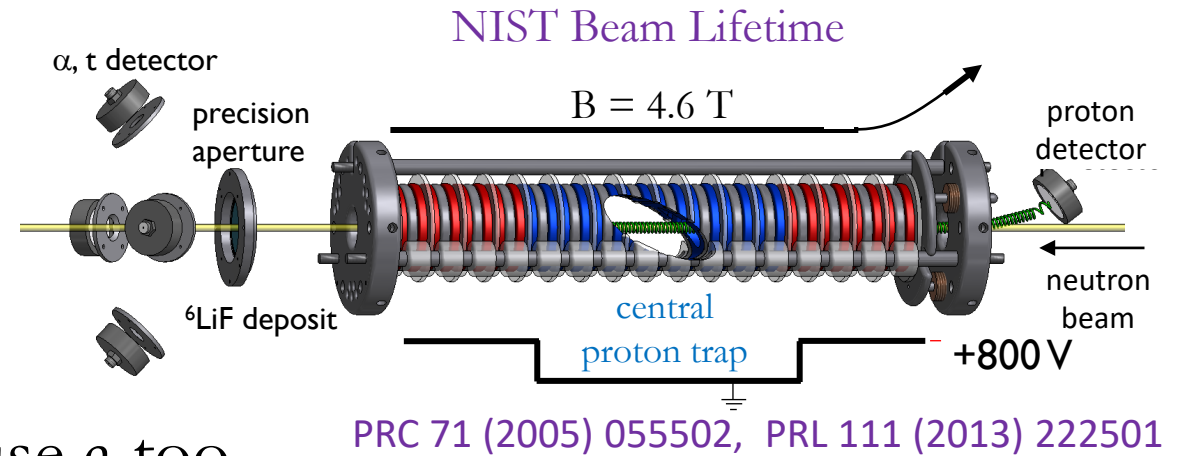
$$\mathcal{H}_{int} = \begin{pmatrix} m + \mu\sigma \cdot \mathbf{B} & \epsilon_{n\bar{n}} & \alpha_{nn'} & \delta_{n\bar{n}'} \\ \epsilon_{n\bar{n}} & m - \mu\sigma \cdot \mathbf{B} & \delta_{n\bar{n}'} & \alpha_{nn'} \\ \alpha_{nn'} & \delta_{n\bar{n}'} & m' + \mu'\sigma \cdot \mathbf{B}' & \epsilon_{n\bar{n}} \\ \delta_{n\bar{n}'} & \alpha_{nn'} & \epsilon_{n\bar{n}} & m' - \mu'\sigma \cdot \mathbf{B}' \end{pmatrix}$$



- Staged approach for R&D for new $n \rightarrow \bar{n}$ experiment, similar sensitivity as ILL

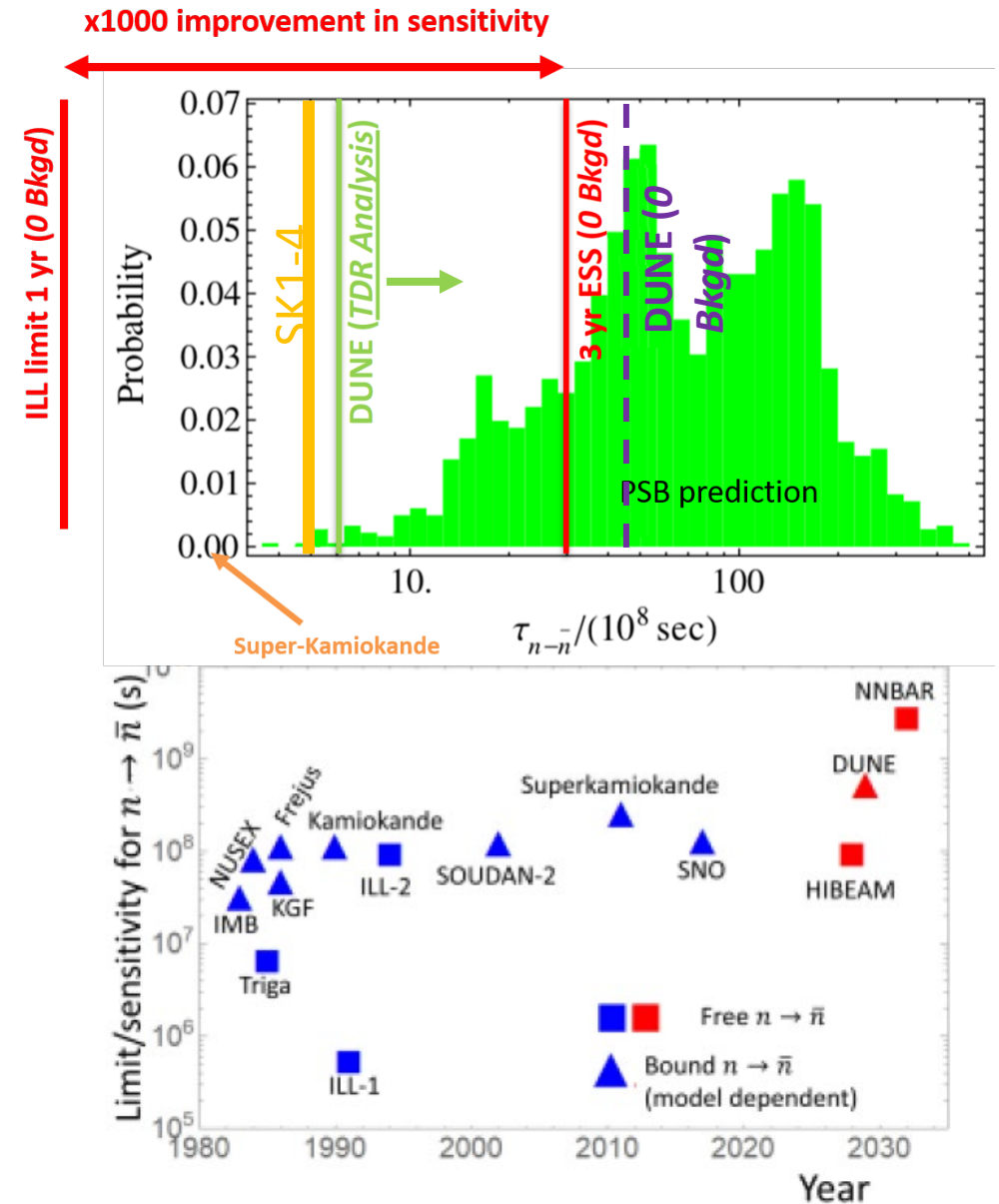
$n \rightarrow n'$ at ORNL

- Neutron Lifetime precision improves (see R Pattie talk), but UCN/CN discrepancy unresolved
- Small mass splitting in n, n' could cause a too-high NIST Beam Lifetime result [EPJC 79 \(2019\) 484](#)
- Search conducted at ORNL SNS using neutron scattering instrument MagRef
- Probed full parameter space to confirm/refute explanation [PRL 128 \(2022\) 212503](#) more sensitive searches in progress (see F Gonzalez talk) [arXiv:1710.00767](#), [EPJWebConf 219 \(2019\) 07002](#)



Outlook for $n \rightarrow \bar{n}$

- Green: range predicted by PSB model [PRD 87 115019 \(2013\)](#)
- Recent Super-K result [PRD 103 \(2021\) 012008](#) approaching that of lower DUNE limit [arXiv:2002.03005](#)
- NNBAR free neutron search complementary, with discovery potential



Summary

- Searches for Baryon Number Violation is strongly motivated
 - Neutrons are under-explored territory in worldwide BLV program
- Fundamental physics program enabled by investments in Large Beam Port and LD₂ lower moderator optimization
- Searches for $n \rightarrow n'$ and $n \rightarrow n' \rightarrow \bar{n}$ in HIBEAM address BNV and dark matter, enable unique opportunity for early R&D for future high sensitivity search for $n \rightarrow \bar{n}$
- NNBAR can produce the world leading limit on $n \rightarrow \bar{n}$ oscillations
- Strong interdisciplinary community in nuclear physics, particle physics and neutronics—exciting opportunities ahead!

HIBEAM/NNBAR Collaboration

- 26 institutions in 8 countries signed NNBAR LOI (2015)
- HIBEAM-NNBAR white-paper (J. Phys. G: Nucl. Part. Phys. 48 (2021)) with over 100 authors.
- Conceptual Design Report to be delivered by end of 2023.
- Currently funded by EU grant (EUR 3M), Swedish Research council (total of ~ EUR 1.3M)
- Current focus of the work are simulations and the detector prototype
- Potential for institutions to getting involved (only “in-kind” contributions expected)

New high-sensitivity searches for neutrons converting into antineutrons and/or sterile neutrons at the European Spallation Source

A. Addazi^{h,at}, K. Anderson^{aq}, S. Ansell^{bm}, K. S. Babu^{az}, J. Barrow^w, D. V. Baxter^{d,e,f}, P. M. Bentley^{ac}, Z. Berezhiani^{b,l}, R. Bevilacqua^{ac}, R. Biondi^b, C. Boehm^{ba}, G. Brooijmans^{an}, L. J. Broussard^{aq}, B. Dev^{ay}, C. Crawford^d, A. D. Dolgov^{ai,ao}, K. Dunne^{ba}, P. Fierlinger^o, M. R. Fitzsimmons^w, A. Fominⁿ, M. Frost^{aq}, S. Gardiner^e, S. Gardner^z, A. Galindo-Uribarri^{aq}, P. Geltenbort^p, S. Girmohanta^{bb}, E. Golubeva^{ah}, G. L. Greene^w, T. Greenshaw^{aa}, V. Gudkov^k, R. Hall-Wilton^{ac}, L. Heilbronn^x, J. Herrero-Garcia^{be}, G. Ichikawa^{bf}, T. M. Ito^{ab}, E. Iverson^{aq}, T. Johansson^{bg}, L. Jönsson^{ad}, Y.-J. Jwa^{an}, Y. Kamyshev^w, K. Kanaki^{ac}, E. Kearns^g, B. Kerbikov^{al,aj,ak}, M. Kitaguchi^{ip}, T. Kittelmann^{ac}, E. Klinkby^{ae}, A. Kobakhidze^{bl}, L. W. Koerner^s, B. Kopeliovich^{bi}, A. Kozela^y, V. Kudryavtsev^{ax}, A. Kupsc^{bg}, Y. Lee^{ac}, M. Lindroos^{ac}, J. Makkinje^{an}, J. I. Marquez^{ac}, B. Meirose^{ba,ad}, T. M. Miller^{ac}, D. Milstead^{ba,*}, R. N. Mohapatra^l, T. Morishima^{ap}, G. Muhrer^{ac}, H. P. Mumm^m, K. Nagamoto^{ap}, F. Nesti^l, V. V. Nesvizhevsky^p, T. Nilsson^f, A. Oskarsson^{ad}, E. Paryev^{ah}, R. W. Pattie, Jr.^l, S. Penttilä^{aq}, Y. N. Pokotilovski^{am}, I. Potashnikova^{bi}, C. Redding^x, J.-M. Richard^{bj}, D. Ries^{af}, E. Rinaldi^{au,bc}, N. Rossi^b, A. Ruggles^x, B. Rybolt^l, V. Santoro^{ac}, U. Sarkar^v, A. Saunders^{ab}, G. Senjanovic^{bd,bn}, A. P. Serebrovⁿ, H. M. Shimizu^{ap}, R. Shrock^{bb}, S. Silverstein^{ba}, D. Silvermyr^{ad}, W. M. Snow^{d,e,f}, A. Takibayev^{ac}, I. Tkachev^{ah}, L. Townsend^x, A. Tureanu^q, L. Varriano^l, A. Vainshtein^{ag,av}, J. de Vries^{ah,bh}, R. Woracek^{ac}, Y. Yamagata^{bk}, A. R. Young^{as}, L. Zanini^{ac}, Z. Zhang^{af}, O. Zimmer^p

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