Neutral Weak Form Factor Measurements from the PREX-II and CREX Experiments

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PVES and Nuclear Structure

Parity Violating Electron Scattering

• Measure scattering asymmetry from left & right handed polarized electrons:

$$A_{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L} = \frac{\mathcal{M}^*_{\gamma} \mathcal{M}_W}{\mathcal{M}^2_{\gamma}} \quad (1)$$

- First measurement: E122 at SLAC (1978)
- PVES experiments have broad scope
- Also useful in pioneering technology and methods for ensuring beam stability
- Progressively lower asymmetries measured over generations



Weak Charge & Neutron Skin



- Theory predicts "neutron skin"
- Excess neutrons pushed radially outwards against surface tension due to pressure of nuclear matter
- Asymmetry correlated with EM & weak form factors:

$$A_{PV} \approx \frac{G_F q^2}{4\pi\alpha\sqrt{2}} \frac{F_W}{F_{ch}} \qquad (2)$$

	Proton	Neutron
EM Charge	1	0
Weak Charge	0.08	1



Current theory models predicting neutron radius and asymmetry highly correlated, but not constrained well by neutron radius measurements

X. Roca-Maza et al 2011 Phys. Rev. Lett. 106 252501

PREX-II & CREX

	PREX-II	CREX
Target	²⁰⁸ Pb	⁴⁸ Ca
Target Thick.	0.5 mm	6 mm
Beam Energy	1 GeV	2.2 GeV
$\langle Q^2 \rangle$	0.00616 GeV ²	0.0297 GeV ²
$\langle \theta \rangle$	5°	5°

- Two PVES experiments, run at JLab Hall A in 2019-20
- Neutron radius from A_{PV} measurements
- Polarized electron beam on unpolarized high-Z targets
- ²⁰⁸Pb radius: test of uniform nuclear matter
- ⁴⁸Ca radius: test of different nuclear structure models
- Overall goal: $\approx 3\%$ error on A_{PV}^{208} and A_{PV}^{48}

Predicted asymmetries on the order of 1 ppm!





Experimental Apparatus

Polarized Source





Image Credit: C. Palatchi (2019)

Left: RTP pockels cell in its mount. Right: An example of helicity patterns like those used for PREX-II.

- Pockels cell sends circularly polarized laser onto GaAs photocathode
 - Pockels cell controls helicity-flipping
- Insertable Half-Wave Plate independently flips helicity sign
- EM Wien filter also flips helicity (was changed every few weeks)
- Mott polarimeter checks polarization of beam out of injector



Targets



• Multiple ²⁰⁸Pb targets to account for cumulative

One target ladder for both experiments

• 10 ²⁰⁸Pb targets, 1 ⁴⁸Ca target

Above Left: Detector hit distribution from a new ²⁰⁸ Pb target **Above Right:** Detector hit distribution from a damaged ²⁰⁸ Pb target

Right: The target ladder used during PREX-II and CREX, with target positions labeled



Detectors

Integrating Detectors:

- Fused silica quartz Čerenkov detectors, two in each spectrometer arm
- \bullet Built for beam currents up to 150 μA
- Takes asymmetry data

Counting Detectors:

- VDCs, GEMs and scintillators
- Can only take low current data
- Used for optics, Q² measurements, detector alignment, other systematic checks

Right: HRS optics with detectors outlined. Inelastics are steered out of the path of the main detector by the dipole but can still be measured in the counting detectors.





PREX-II/CREX detectors deployed inside the HRS huts

Systematics Controls

Beam Corrections



Above: in **black**, the asymmetry distribution without trajectory corrections. In red, the asymmetry distribution with trajectory corrections. (For PREX-II)

- Uncorrected, the systematic uncertainty from beam backgrounds would be large compared to counting statistics
- Asymmetry correction is:

$$A_{PV} = \boxed{A_{raw} - A_Q} - \sum_i \alpha_i \Delta x_i - \alpha_E A_E$$
(3)

- Beam modulation & regression both methods to extract α_i's
- Both modulation and regression data monitored consistently throughout experiments' running

Beam Corrections

Regression

ΔE

-1.6±5.6





Lagrangian Multipliers



- Position, angle and energy information taken from position monitors
- Trajectories correlated with changes in measured A_{raw}
- Magnetic coils alter beam trajectory periodically, in sequence
- Correlation slopes extracted from beam changes
- Create set of Lagrangian multipliers to minimize regression correction
- Constraints come from modulation slopes

Beam Corrections (PREX-II)



Above: PREX-II regression results. Below: PREX-II

regression	uncertainty	table.
------------	-------------	--------

Source	Mean	3% Err
	[ppb]	[ppb]
E	-70.44	2.11
X1	-22.33	0.67
X2	9.70	0.29
Y1	22.50	0.68
Y2	-2.84	0.06
Others	3.04	0.09
Total	-60.38	2.5

PREX-II:

- Energy regression is highest source of uncertainty
 - Energy slopes have same sign in L-R HRS detectors and don't cancel out
- Average energy slope: ${\sim}36$ ppm/ μ m
- Average difference between Lagrangian and regression slopes: ${\sim}0.14~\text{ppm/um}$
- Determined \approx 3% uncertainty
 - Determination from difference between regression corrections and Lagrange multiplier corrections

CREX Compton Polarimetry





- Polarimeter consists of:
 - Magnetic chicane to steer beam
 - Fabry-Perot cavity on laser table
 - Photon calorimeter
 - High-speed DAQ system
- Laser λ =532 nm

- $\approx\!30$ MeV Compton edge for PREX-II; $\approx\!150$ MeV Compton edge for CREX
- Laser polarization measured on table
- Runs concurrently with main experiment

CREX Compton Polarimetry



CREX Polarization Measurements (Compton & Moller)

Above: Møller and Compton polarimetry data for CREX. All uncertainties plotted are statistical only. Møller data courtesy of E. King.



- Time-dependent polarization
- Laser polarization is dominant systematic, followed by photon-cone pointing direction
- Most accurate Compton beam polarimetry measurement

Inelastics Rejection (CREX)

- Hall A HRS's separate inelastics by dipole steering
- Non-tracking "quartz" detectors measure *A_{PV}*
- Other tracking detectors measure inelastic peak positions







- Frequent detector realignment keeps quartz edge off inelastic peaks
- CREX overall systematic contribution: 0.89%

Left: Elastic and inelastic peak locations relative to quartz detector edge

Asymmetry Results

PREX-II Results

PREX-II Asymmetry

$$\mathcal{A}_{PV}^{208} = 550 \pm 16 (ext{stats}) \pm 8 (ext{syst})$$
 parts per billion

Neutron Skin Thickness

$$R_n^{208} - R_p^{208} = 0.283 \pm 0.071 \text{ fm}$$

Uncert. Source	A _{PV} uncert. contribution
Polarization	0.95%
Acceptance	0.83%
normalization	
Beam correction	0.54%
Detector nonlin-	0.49%
earity	
Carbon contami-	0.26%
nation	
Charge Correction	0.04%
Inelastic Contami-	0.02%
nation	
Total	1.48%



PREX-II paper published in PRL! Phys.Rev.Lett. 126 (2021) 17, 172502

Above: New baryon density curves calculated from PREX-II results.

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PREX-II in the News



PREX-II in the News



CREX Results

CREX Asymmetry

 $\mathcal{A}_{PV}^{48} = 2668 \pm 106 (ext{stats}) \pm 39 (ext{syst})$ parts per billion

Neutron Skin Thickness

$$R_n^{48} - R_p^{48} = 0.121 \pm 0.035 \; {
m fm}$$

Uncert. Source	A_{PV} uncert. contribution	0.06 0.05
Acceptance normaliza-	0.90%	
tion		0.04
Inelastic Contamination	0.82%	
Transverse asymmetry	0.49%	
Polarization	0.39%	0.02
Radiative Corrections	0.37%	0.01 -
Beam Correction	0.27%	
Nonlinearity	0.26%	0
Isotopic purity	0.11%	-0.01000.5111.522.5
Total	1.49%	g (fm ⁻¹)

Published in PRL! D. Adhikari *et al* Phys. Rev. Lett. **129** 042501

Above: CREX measured $F_{ch} - F_W$ compared to one family of theory models with different weak radii.

Summary



Image Credit: M. Petrusky (2019)

- Two experiments of current PVES generation
 - Reducing and controlling systematic uncertainties from beam
 - Accurate beam polarimetry
 - Specialized "integrating" detectors
- PREX-II ran in summer 2019
- Measured neutron radius of ²⁰⁸Pb
- CREX ran on ⁴⁸Ca target in spring/summer 2020
- PREX-II and CREX both published in PRL

THANK YOU!

Accurate Determination of the Neutron Skin Thickness of ²⁰⁸Pb through Parity-Violation in Electron Scattering (The PREX Collaboration)

D. Adhkadi, H. Albatalarki, P. D. Anforde, Y. K. Malok, P. S. Armstrong, T. Xwertl, Y. C. Ayerbe Gayon, J. S. Barcesk, Y. Bellin, T. B. Baninovith, S. J. Bonesko, H. Bhatty, D. Blattar, P. Blattar, H. Mattar, D. Blattar, P. Mattar, J. S. Bonesko, P. S. Data, A. Camosom, G. D. Catsu, Y. Y. Chen, Y. C. Chark, Y. J. C. Garla, Y. J. C. Gandall, Y. Gandi, T. Gartan, M. M. Gardzin, W. J. Cohney, C. Ganda, Y. D. Unita, W. F. Heinhord, E. Fachey, H. C. Gall, ^{11,11}, D. Gashell, Y. Gandar, K. K. Kang, Y. Ka

Precision Determination of the Neutral Weak Form Factor of ⁴⁸Ca

(The CREX Collaboration)

D. Athibari, H. Albataineh, D. Andorek, P. K. Anizd, D. S. Amstoragi, T. Avereti, C. Ayerle Gayon, S. K. Barcze, Y. Deflari, T. S. Bominostarka, J. F. Bowech, H. Batt, T. Bhatta Pathak, P. B. Bherwark, B. Bherwark, B. Bhatta Pathak, P. B. Bherwark, B. Bherwark, B. K. Bartz, P. J. Bellerwark, J. C. Cornejo, M. M. Dalow, F. P. Darta, A. D. Gayano, H. S. Botta, P. J. Charne, J. J. C. Ferder, J. J. C. Gardon, J. M. Charne, T. C. Tarta, A. D. Barjanov, P. C. Bartz, P. J. Bellerwark, B. C. Gardon, J. M. Dalow, F. D. Karta, A. D. Barjanov, P. C. Hand, P. J. C. Ferder, J. J. C. Gardon, J. K. Berker, J. C. Gardon, J. M. Bartz, J. A. Bartz, J. C. Bartz, J. L. Bartz, J. J. Bartz, J. C. Bartz, J. K. Bartz, J. S. Bartz, J. Bartz,

Thank you to our 100+ collaborators!

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C. Korpal,⁴ P.M. King,² D.D. King,² M. Kunsas,²⁴ K.S. Kunsas,¹⁷ T. Kung,¹⁸ N. Lashig-Cohltra;¹⁴ K. G. Kevrick,¹⁴ H. Lati,¹⁷ N. Lashig,¹⁵ Cohltra;¹⁵ M. Mahord, ¹⁴ M. Mark,¹⁶ M. Markashi,¹⁶ D. McNuhy,¹⁰ M. M. Radhall,¹⁴ V. Owa,¹⁷ C. Patrich,¹¹ D. P. Pandoy,¹⁶ S. Patric,¹⁴ K.D. Pashish,¹¹ C.J. M. K. Bandar,¹¹ J. M. Pandor,¹¹ S. Patric,¹⁴ K.D. Pashish,¹¹ C.J. M. K. Bandar,¹¹ D. T. Jang,¹⁶ J. R. Rimer,¹⁷ R. Richten,¹² S. Raiman,¹⁰ A. Raiharya,²⁰ P. Souhay,²¹ D. Soeks,¹⁴ A. Shaharya,²⁰ P. Souhay,²¹ D. Soeks,¹⁴ A. Shaharya,²⁰ P. Souhay,²¹ D. Yang,¹⁶ J. Yu, Man,¹⁴ S. Zawa,¹¹ L. W. Kunz,¹⁴ J. Zamar,¹¹ L. Sauda,¹⁶ S. Buchar,¹¹ K. Jandar,¹⁶ S. Markan,¹¹ C. M. Kunz,¹⁴ J. Zamar,¹¹ L. Sauda,¹⁶ S. Markan,¹⁴ J. Mahinya,²⁶ P. Souhay,²⁴ D. Yang,¹⁶ J. Yu, Mann,¹⁵ J. Zamar,¹¹ L. Sauda,¹⁶ S. Markan,¹⁶ J. M. Kunz,¹⁶ J. Jang,¹⁶ M. M. Kunz,¹⁶ J. Jang,¹¹ L. M. Kunz,¹⁶ J. Zamar,¹¹ L. Sauda,¹⁶ S. Markan,¹⁶ J. Markan,¹⁶

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Backup Slides

²⁰⁸Pb

- ²⁰⁸Pb: neutron rich double-magic nucleus
- Can be modeled as uniform nuclear matter
 - Becomes an terrestrial laboratory to test neutron star structure
- Δ*r_{np}* highly correlated to symmetry energy





- Neutron star radius 18 orders of magnitude greater than ²⁰⁸Pb radius
- ... but they have the same EOS!

- Changed cell crystal to RTP for lower $T_{sett/e}$ and higher flip rate
- Handled intensity asymmetries with aggressive feedback program
 - Also with careful alignment of crystal
- Separate feedback system for handling position differences off the injector
- Avoid spot-size asymmetries with crystal alignment



(Rubidium Titanyle Phosphate)



PREX-II Møller Polarimetry



Images courtesy of E. King (2019)

PREX Polarimetry Results



Polarization [pct] vs Escargatoire

PREX-II Compton results with Møller polarimeter results overlaid in colored bars (stat error only)

- Systematic corrections applied
- Good consistency with Møller (within 0.8% systematic window)
- Compton raw statistics: 0.2%-0.5%

- χ^2 uneven across run periods
- Non-statistical behavior between run periods
- Percent-level polarization changes not yet explained

Inelastics Rejection (PREX-II)

- Additional excited states had contribution to PREX-II
- Inelastic peak detector location further from quartz edge
- Uncertainty introduced from PREX-II elastics is negligible



Above: Inelastic peak spectrum for ²⁰⁸Pb. Image courtesy of D. Adhikari (2021)

Q^2 & Pointing Angle



Variable	Amount
$\frac{A_{corr}}{P_{e}}$ [ppb]	549.34
A _c [ppb]	539.36
δA_c [ppb]	21.574
$\frac{\delta A_c}{A_c}$ [%]	4.00
f _c	6.29E-02
δf_c	4.63E-03
Rel. Error	0.01
from <i>f_c</i> [%]	
Rel. Error from A _c [%]	0.26

• Correction for a false asymmetry source can be applied by

$$A_{PV} = \frac{\frac{A_{corr}}{P_e} - \sum_i A_i f_i}{1 - \sum_i f_i}$$
(6)

- Where A_i every source of false asymmetry, f_i is the background fraction, and P_e is the incoming beam polarization
- Theoretical C-asymmetry calculation provided by optical simulation and rate informs correction
- In limit that $\frac{A_{corr}}{P_e} \simeq A_c$ then correction factor approaches 1



Above: Asymmetry plotted by combined IHWP IN/OUT periods.

• Search for systematic effects by measuring

$$A_{null} = \frac{A_{PV}^{lN} - A_{PV}^{OUT}}{2} \quad (7)$$

- Which should be identically zero without uncorrected systematics.
- Asymmetries should cancel out as IHWP change is just a change of sign
- Average A_{null} zero over PREX-II run

PREX-II and Neutron Stars

- PREX-II has significant implications for the deformability of neutron stars
- NICER experiment bounds the radii of neutron stars
- PREX-II result has good agreement with NICER bounds

Right: the overlap between the PREX-II and NICER error bars is marked in blue in the "Allowed" region

Image Credit: B. Reed et. al.

Phys.Rev.Lett. 126 (2021) 17, 172503



What's Next for PVES?

MOLLER



Above: Target, magnet, and detector concept for MOLLER experiment. Image courtesy of MOLLER collaboration.

- High precision measurement of weak mixing angle at low Q^2
- Expected $A_{PV} \approx 35.6$ ppb



SoLID

Above: CAD model of SoLID detector in its PVDIS configuration.

- High-luminosity, large acceptance detector
- $\bullet\,$ SIDIS, PVDIS, and J/ $\psi\,$ programs

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What's Next for PVES?

- Mainzer Microtron (MAMI) facility in Germany
- Two experiments:
 - **P2:** Ultra-precise measurement of $\sin^2 \theta_w$
 - MREX: Ultra-precise measurement of $R_n R_p$ for ²⁰⁸Pb
- MREX will improve on the precision of the asymmetry measured in PREX-II
- Both experiments to run after the completion of MESA



Above: Projected precision for PREX-II and MREX experiments compared to the range of neutron radii predicted by DFT models. Reproduced from Becker et al., Eur. Phys. J. A, 54 (2018), 11

Compton Backup

Compton Polarimetry

Images Credit: D. Gaskell (2019)





Optics table device layout

• Polarimeter consists of:

- Magnetic chicane to steer beam
- Fabry-Perot cavity on laser table
- Photon calorimeter
- High-speed DAQ system
- Laser/Amp outputs at λ =1064 nm, but is doubled to λ =532 nm
- Laser polarization measured on table

Polarimeter Laser



- PD
- Laser polarization achieved with combination of half waveplates and quarter waveplates
- Retro-reflected photodiode (RRPD) power anti-correlated with degree of circular polarization (DOCP)
- Laser polarization highly controlled and measurable
- Laser input state can become arbitrary

Laser Cycling

- To handle shifts in background, we periodically flip off the laser
- Backgrounds calculated on cycle-to-cycle basis
- 1 cycle = a laser-on period, sandwiched by two laser off periods

Acc0/NAcc0:mpsCoda {mpsCoda>=524000 && mpsCoda<566377}





Top: Plot of helicity correlated differences vs time for one PREX-II run Bottom: Plot of sums for same time period

In each plot, low variation of the integrated signal is likely indicative of healthy data. In all plots, blue represents laser-on periods, red represents laser-off.

Data shown here was taken over a \approx 90 minute period.

Laser Polarization Model



Before run:

- NPBS used to used to characterize DOCP in-cavity
- Complicated by birefringence of cavity mirrors
- Entrance scans, exit scans, cavity scans. . .

During run:

- Additional verification: running off 100%
- QWP/HWP changed for multiple snails to alter DOCP
- Saw polarization magnitude decrease
- Laser systematic still being calculated!
 - Study on birefringence parameters still pending



Compton Photon Detector



Images Credit: J. C. Cornejo (2019)



- Detector Components:
 - Pb Collimator
 - Pb Sync Shield
 - GSO scintillator
 - PMT and DAQ readout
- Signals read out per rapidly-flipping helicity state
- Measure helicity-correlated asymmetry
- LED's allow for *in-situ* detector tests

ADC Response IBAU 3600 3500

3700



Single Compton-Edge PREX Pulse

Example photon pulse with energy matching the PREX Compton-edge. The CREX Compton

edge photons had about 4x greater energy 2022-09-02

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Polarimetry Measurement: Integrating Method

How to measure a Compton Asymmetry: Integrate the signal over pedestal per helicity state. Measure signal *S*, for each laser state ON, OFF and helicity state +, -. Helicity pattern difference (Δ), sum (*Y*), and asymmetry (A) distributions are calculated:

$$\mathcal{A}_{exp} = \langle \mathcal{A}_{ON}
angle - \langle \mathcal{A}_{OFF}
angle = \mathcal{P}_{e} \mathcal{P}_{\gamma} \langle \mathcal{A}_{I}
angle$$

With laser DOCP \mathcal{P}_{γ} , energy-weighted average analyzing power $\langle \mathcal{A}_l \rangle$, and beam polarization \mathcal{P}_e .

$$\begin{split} \Delta_{ON} &= S^+_{ON} - S^-_{ON} \\ \Delta_{OFF} &= S^+_{OFF} - S^-_{OFF} \\ Y_{ON} &= S^+_{ON} + S^-_{ON} \\ Y_{OFF} &= S^+_{OFF} + S^-_{OFF} \\ \mathcal{A}_{ON} &= \frac{\Delta_{ON}}{Y_{ON} - \langle Y_{OFF} \rangle} \\ \mathcal{A}_{OFF} &= \frac{\Delta_{OFF}}{\langle Y_{ON} \rangle - \langle Y_{OFF} \rangle} \end{split}$$





Compton Spectra



Detector Corrections

Nonlinearity

- 1 kHz pulser system w/ load = PREX signal
 - Track $Yield(var + \Delta) Yield(var)$
- Nonlinearity out to 2*CE
- Very small analyzing power correction ($\approx 0.08\%$ for PREX)



Gain Shift

- Evidence of small change in pulse size with background signal size
- Nonzero shift necessitates dynamic correction
- Bench tests of gain shift done
 - Analysis yielded correction factor for \mathcal{A}_{exp}

Left: Evidence of a gain shift from a linearity run taken during beam operations

Differential Non Linearity

