

Parity-Odd Modes in the 4PCF of BOSS DR12 CMASS and LOWZ

J. Hou
Z. Slepian
R.N. Cahn

<https://arxiv.org/abs/2110.12004>

<https://arxiv.org/abs/2206.03625>



CIPANP: August 31, 2022

The Renaissance was known for self-
portraits



How did they do it?



Mirrors

"Art is the mirror of nature"

Borrow from Renaissance portraiture to
study nature, and ask:

Is nature a mirror of itself?

Parity

CPT Theorem: Luders 1940—any Lorentz-invariant theory

Yet we have 1 billion baryons for every anti-baryon—fortunately

Requires breaking of CP symmetry

Perhaps just P symmetry

Inflation

Vanilla inflation predicts no parity-violation

But if there is a new force (or coupling, if you prefer), there could be

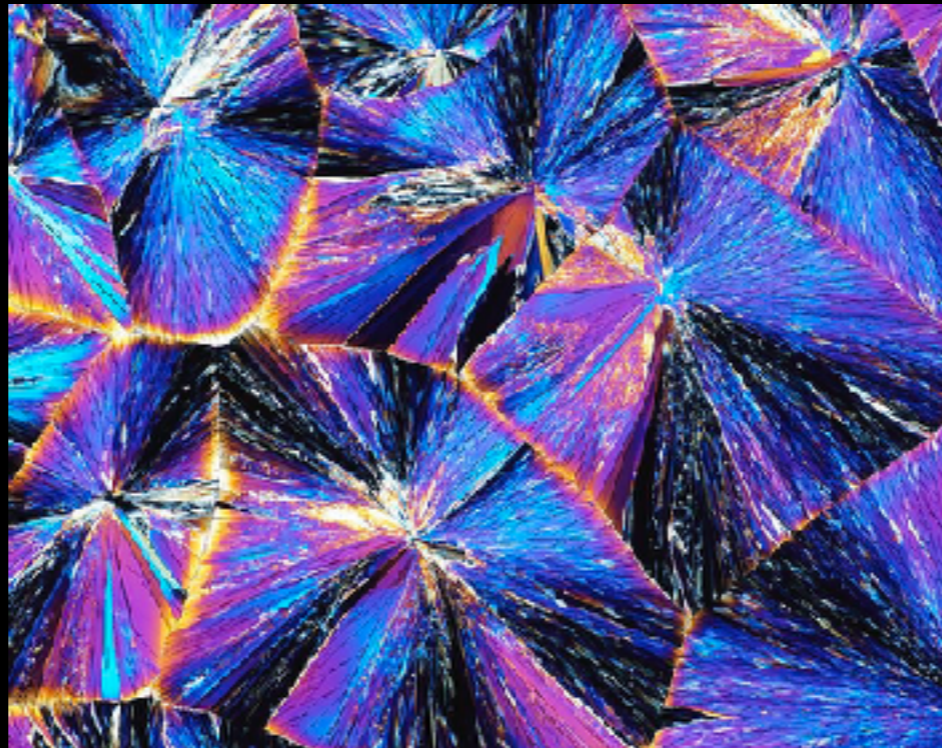
e.g. a term in the Lagrangian such as Maxwell tensor \times Maxwell tensor dual

Or a term in the Riemann tensor and its dual

Back to history

Pasteur: 1848—noticed that synthetically-produced tartaric acid produced no polarization of light when the light passed through it

But naturally-occurring tartaric acid did

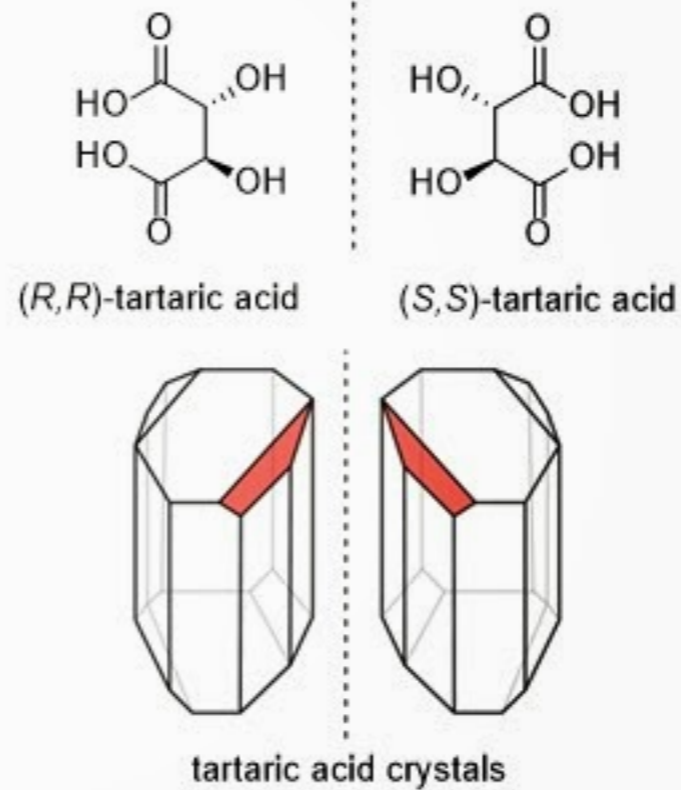


Another history lesson

Also known as "stereoactive isomers" or
"enantiomers"



Louis Pasteur



So let us search for such "crystals"

Need to be 3D

Point = no

Line = no

Triangle = no (sadly for the 3PCF)

Tetrahedron = YES



The Galaxy 4PCF

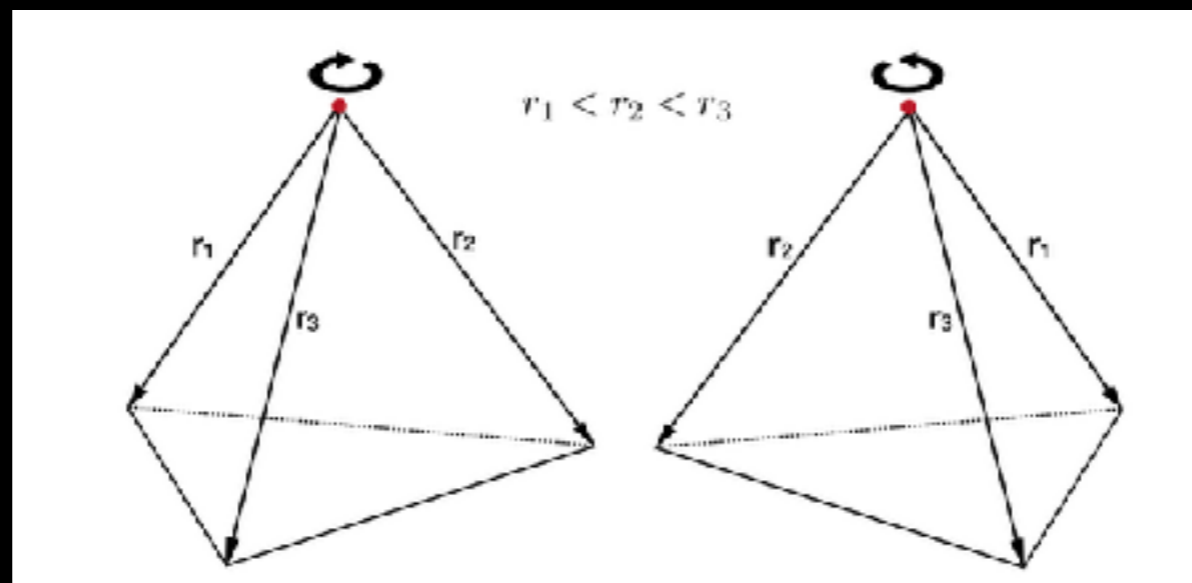
Fun part's over—or maybe just starting!

-> From human history, back to
cosmology

4PCF = excess clustering of tetrahedra
over and above what a spatially random
distribution of points would have

4PCF Basis

Parametrize 4PCF by six variables:
3 side lengths from a given vertex,
3 angular momenta for dependence on angles
around it



Angular basis functions = "isotropic functions" —
total-angular-momentum-zero products of 3
spherical harmonics ("primary" galaxy is at the
origin, so no harmonic needed for it)

The Galaxy 4PCF

Seek radial coefficients giving projection of 4PCF onto this angular basis:

$$4\text{PCF} = \sum_{l_i} \zeta_{l_1 l_2 l_3}(r_1, r_2, r_3) \mathcal{P}_{l_1 l_2 l_3}(\hat{r}_1, \hat{r}_2, \hat{r}_3)$$

$$\mathcal{P}_{l_1 l_2 l_3}(\hat{r}_1, \hat{r}_2, \hat{r}_3) = \sum_{m_i} \begin{pmatrix} l_1 & l_2 & l_3 \\ m_1 & m_2 & m_3 \end{pmatrix} Y_{l_1 m_1}(\hat{r}_1) Y_{l_2 m_2}(\hat{r}_2) Y_{l_3 m_3}(\hat{r}_3)$$

Even/Odd Split

Both even and odd-parity basis functions

In standard picture, projection of 4PCF onto odd basis functions should be zero (up to fluctuations)

$$\begin{aligned}\mathcal{P}_{111}(\hat{\mathbf{r}}_1, \hat{\mathbf{r}}_2, \hat{\mathbf{r}}_3) &= -i \frac{3}{\sqrt{2}} (4\pi)^{-3/2} \hat{\mathbf{r}}_1 \cdot (\hat{\mathbf{r}}_2 \times \hat{\mathbf{r}}_3), \\ \mathcal{P}_{122}(\hat{\mathbf{r}}_1, \hat{\mathbf{r}}_2, \hat{\mathbf{r}}_3) &= i \sqrt{\frac{45}{2}} (4\pi)^{-3/2} \hat{\mathbf{r}}_1 \cdot (\hat{\mathbf{r}}_2 \times \hat{\mathbf{r}}_3) (\hat{\mathbf{r}}_2 \cdot \hat{\mathbf{r}}_3), \\ \mathcal{P}_{133}(\hat{\mathbf{r}}_1, \hat{\mathbf{r}}_2, \hat{\mathbf{r}}_3) &= -i \frac{15}{4} \sqrt{7} (4\pi)^{-3/2} \hat{\mathbf{r}}_1 \cdot (\hat{\mathbf{r}}_2 \times \hat{\mathbf{r}}_3) \left[(\hat{\mathbf{r}}_2 \cdot \hat{\mathbf{r}}_3)^2 - \frac{1}{5} \right], \\ \mathcal{P}_{144}(\hat{\mathbf{r}}_1, \hat{\mathbf{r}}_2, \hat{\mathbf{r}}_3) &= +i \frac{21\sqrt{15}}{4} (4\pi)^{-3/2} \hat{\mathbf{r}}_1 \cdot (\hat{\mathbf{r}}_2 \times \hat{\mathbf{r}}_3) \left[(\hat{\mathbf{r}}_2 \cdot \hat{\mathbf{r}}_3)^3 - \frac{3}{7} (\hat{\mathbf{r}}_2 \cdot \hat{\mathbf{r}}_3) \right], \\ \mathcal{P}_{223}(\hat{\mathbf{r}}_1, \hat{\mathbf{r}}_2, \hat{\mathbf{r}}_3) &= -i 15 \sqrt{\frac{5}{8}} (4\pi)^{-3/2} \hat{\mathbf{r}}_1 \cdot (\hat{\mathbf{r}}_2 \times \hat{\mathbf{r}}_3) \left[(\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{r}}_3) (\hat{\mathbf{r}}_2 \cdot \hat{\mathbf{r}}_3) - \frac{1}{5} \hat{\mathbf{r}}_1 \cdot \hat{\mathbf{r}}_2 \right],\end{aligned}$$

We have our basis

What about an *algorithm*?

How do you measure the 4PCF fast?

Naive combinatorics is explosive!

Computing Summary



Lots of math: use GPUs

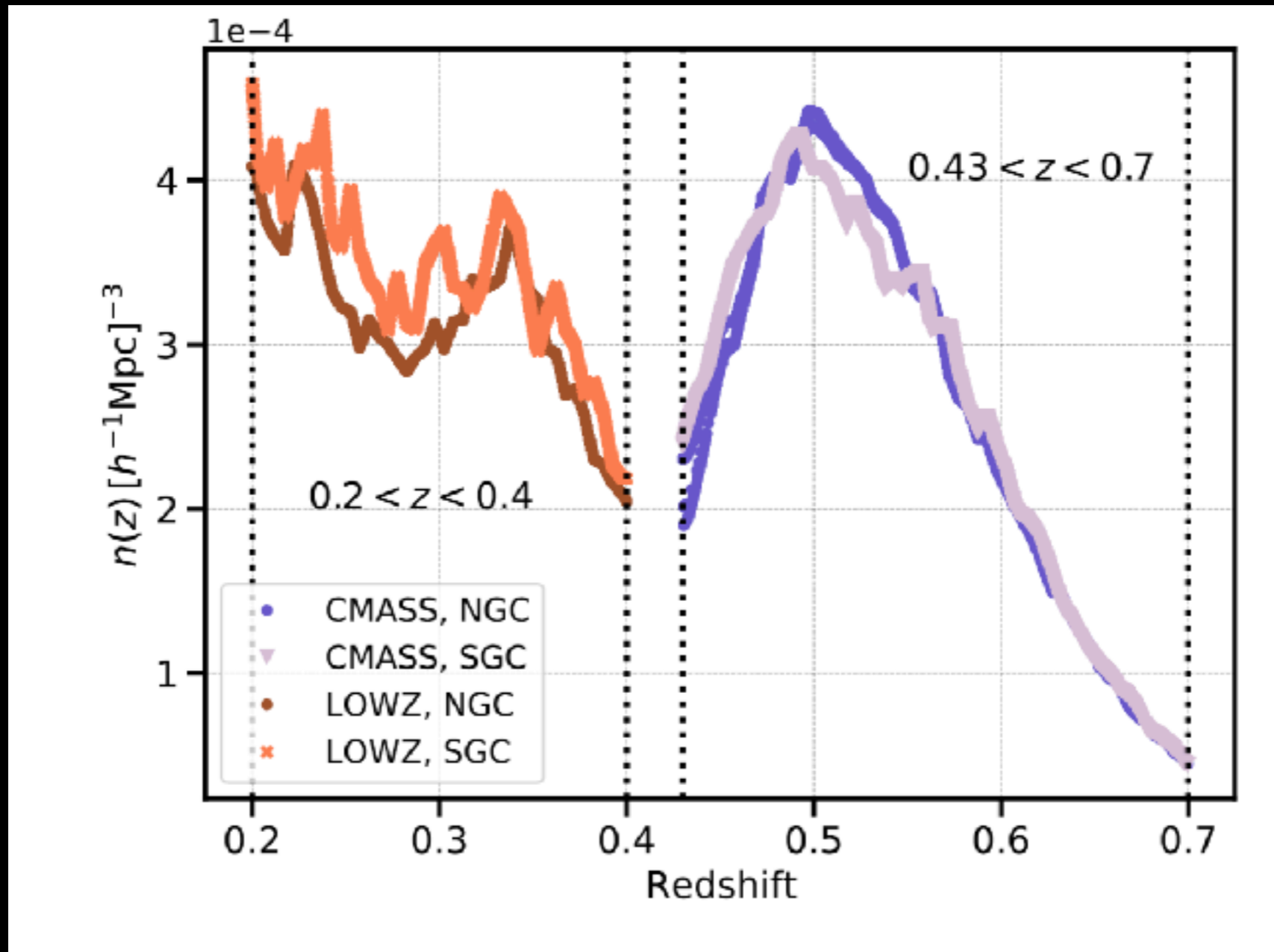
Code for 2-6PCF optimized for NVIDIA A100s



7.5M CPU hours = 20 months day and night on a 500-core cluster

1 month on 69 NVIDIA A100 GPUs

Data



SDSS DR12 BOSS CMASS and LOWZ samples

Covariance Matrix

Our fiducial analysis had $\sim 9,000$ degrees of freedom, \gg number of mocks

Driven by fine radial binning needed to avoid incoherent co-adding of parity-odd signal if one sums over vertices of a tetrahedron

-> Cannot derive invertible covariance matrix from mocks

-> Other solutions

Covariance Matrix

3 Methods

- 1) Analytic: use GRF to compute analytic covariance matrix
- 2) Compressed: diagonalize analytic covariance, select a sub-space of "best" e-vectors, measure covariance directly from mocks in that smaller space
- 3) Direct: restrict angular momentum and/or side-length range to make d.o.f. small enough to determine C directly from mocks

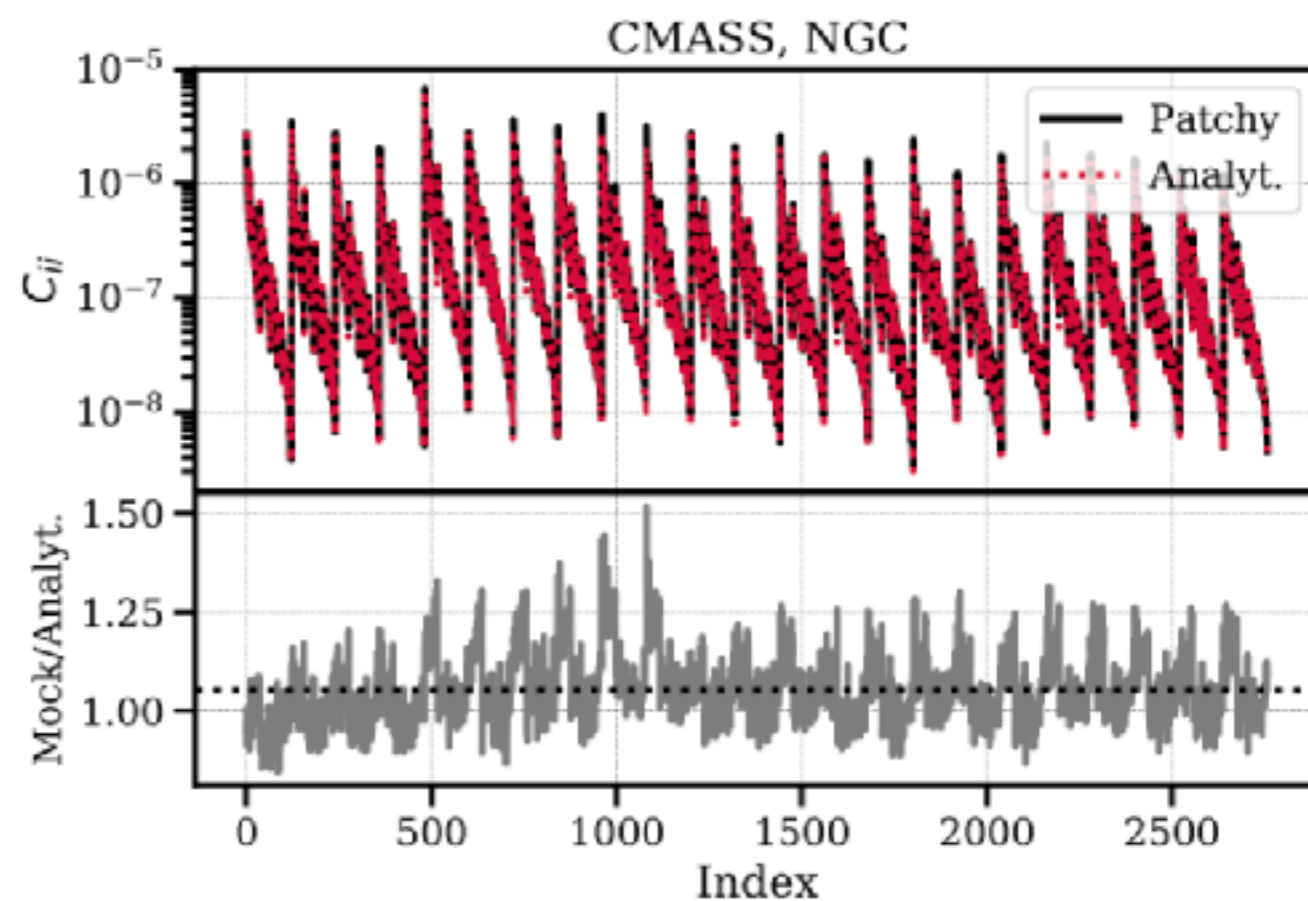
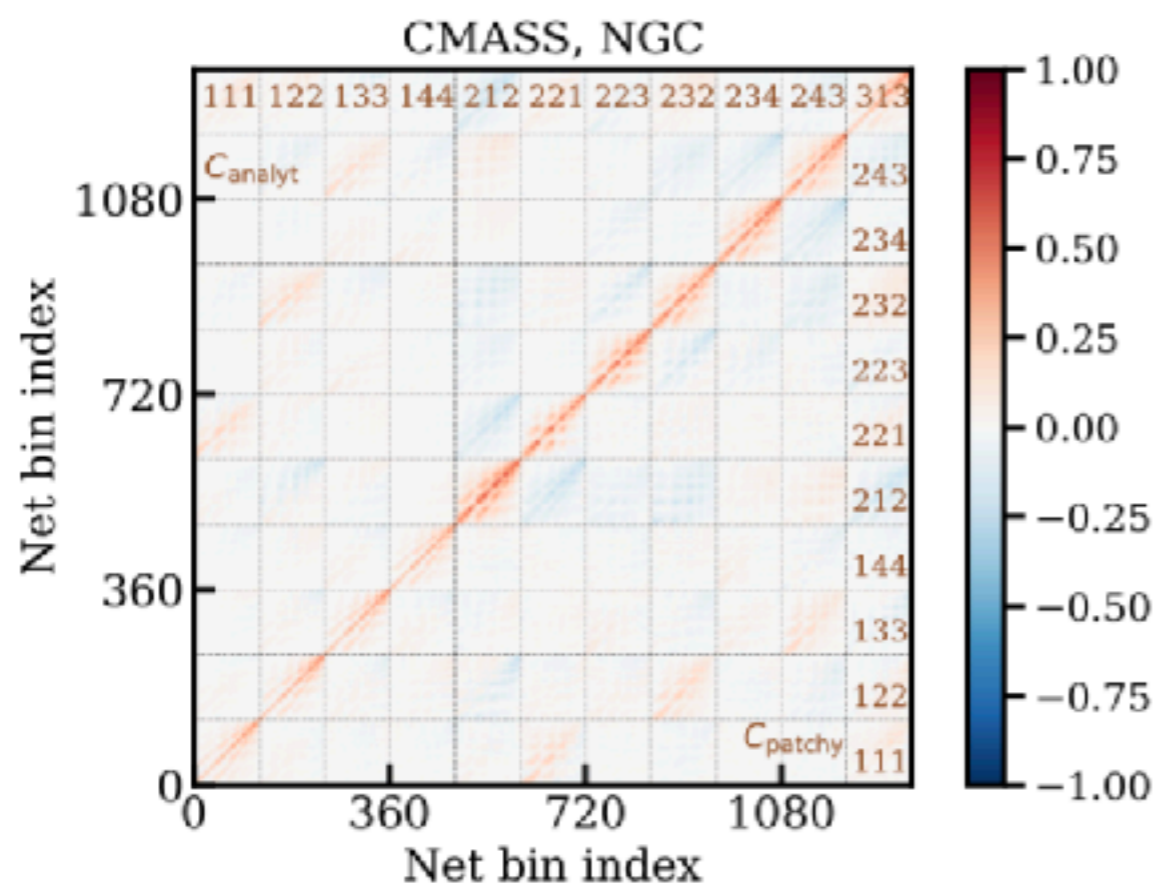
The "Serious" Slide

$$\text{COV}_{\Lambda, \Lambda'}(r_1, r_2, r_3; r'_1, r'_2, r'_3) = \text{COV}_{\Lambda, \Lambda'}^{\text{I}}(r_1, r_2, r_3; r'_1, r'_2, r'_3) + \text{COV}_{\Lambda, \Lambda'}^{\text{II}}(r_1, r_2, r_3; r'_1, r'_2, r'_3).$$

$$\begin{aligned} \text{COV}_{\Lambda, \Lambda'}^{\text{I}}(r_1, r_2, r_3; r'_1, r'_2, r'_3) &= (4\pi)^4 \sum_G (-1)^{\Sigma(\Lambda)(1-\varepsilon_G)/2} \sum_{L_1 L_2 L_3} \mathcal{D}_{L_1 L_2 L_3}^{\text{P}} \mathcal{C}_{000}^{L_1 L_2 L_3} \begin{Bmatrix} \ell_{G1} & \ell_{G2} & \ell_{G3} \\ \ell'_1 & \ell'_2 & \ell'_3 \\ L_1 & L_2 & L_3 \end{Bmatrix} \\ &\times \int \frac{s^2 ds}{V} \prod_{i=1}^3 \left[(-1)^{(-\ell_{Gi} - \ell'_i + L_i)/2} \mathcal{D}_{\ell_{Gi} \ell'_i L_i}^{\text{P}} \mathcal{C}_{000}^{\ell_{Gi} \ell'_i L_i} \xi(s) f_{\ell_{Gi} \ell'_i L_i}(r_{Gi}, r'_i, s) \right], \end{aligned}$$

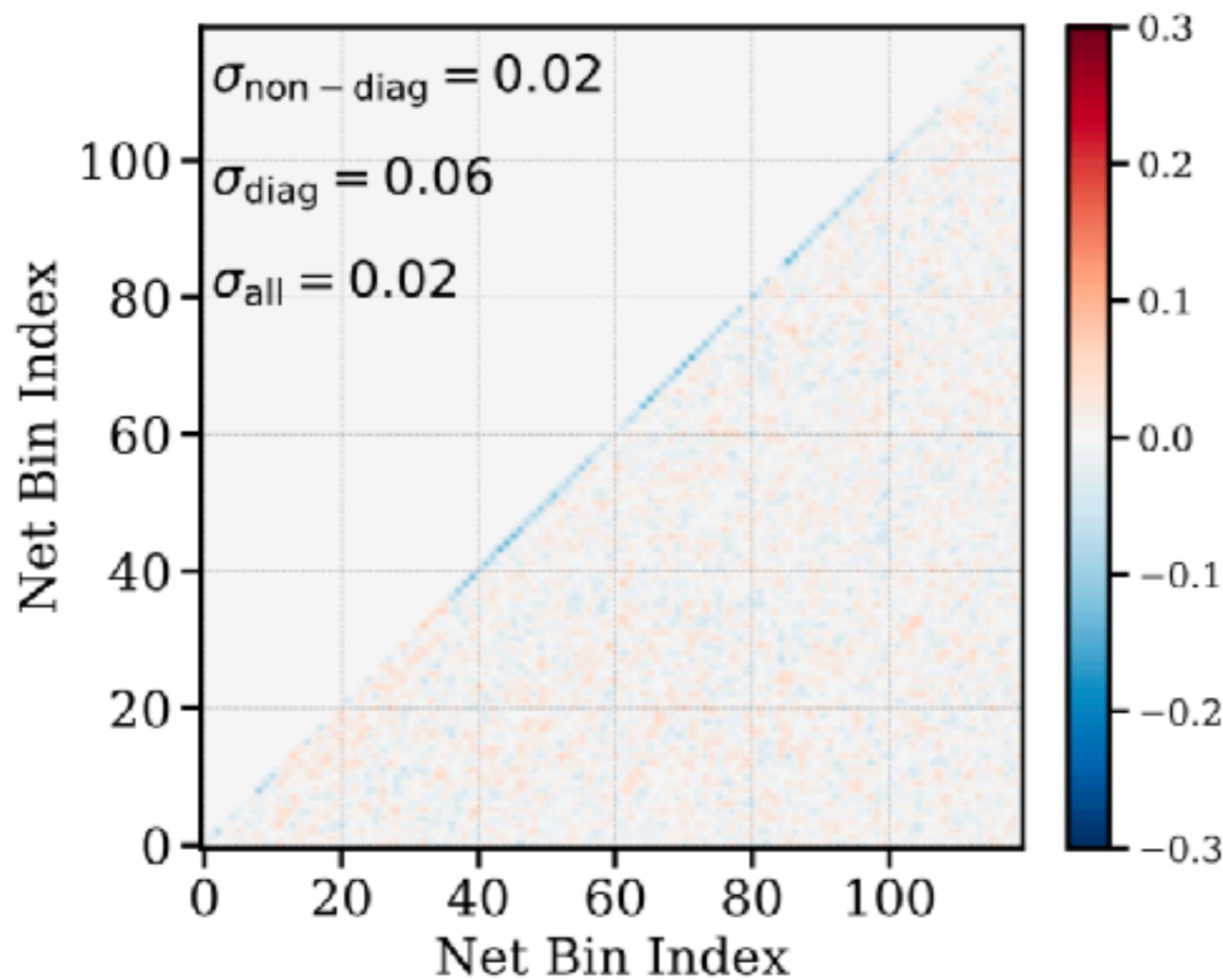
$$\begin{aligned} \text{COV}_{\Lambda, \Lambda'}^{\text{II}}(r_1, r_2, r_3; r'_1, r'_2, r'_3) &= (4\pi)^4 \sum_{G, H} (-1)^{\Sigma(\Lambda')(1-\varepsilon_H)/2} \sum_{L_1 L_2 L_3} \mathcal{D}_{L_1 L_2 L_3}^{\text{P}} \mathcal{C}_{000}^{L_1 L_2 L_3} \begin{Bmatrix} \ell_{G1} & \ell_{G2} & \ell_{G3} \\ \ell'_{H1} & \ell'_{H2} & \ell'_{H3} \\ L_1 & L_2 & L_3 \end{Bmatrix} \\ &\times \int \frac{s^2 ds}{V} \prod_{i=1}^3 \left[(-1)^{(-\ell_{Gi} - \ell'_{Hi} + L_i)/2} \mathcal{D}_{\ell_{Gi} \ell'_{Hi} L_i}^{\text{P}} \mathcal{C}_{000}^{\ell_{Gi} \ell'_{Hi} L_i} \right] \\ &\times f_{\ell_{G1} 0 \ell_{G1}}(r_{G1}, 0, s) f_{0 \ell'_{H1} \ell'_{H1}}(0, r'_{H1}, s) f_{\ell_{G2} \ell'_{H2} L_2}(r_{G2}, r'_{H2}, s) f_{\ell_{G3} \ell'_{H3} L_3}(r_{G3}, r'_{H3}, s). \end{aligned}$$

Comparison of Mock and Analytic Covariance

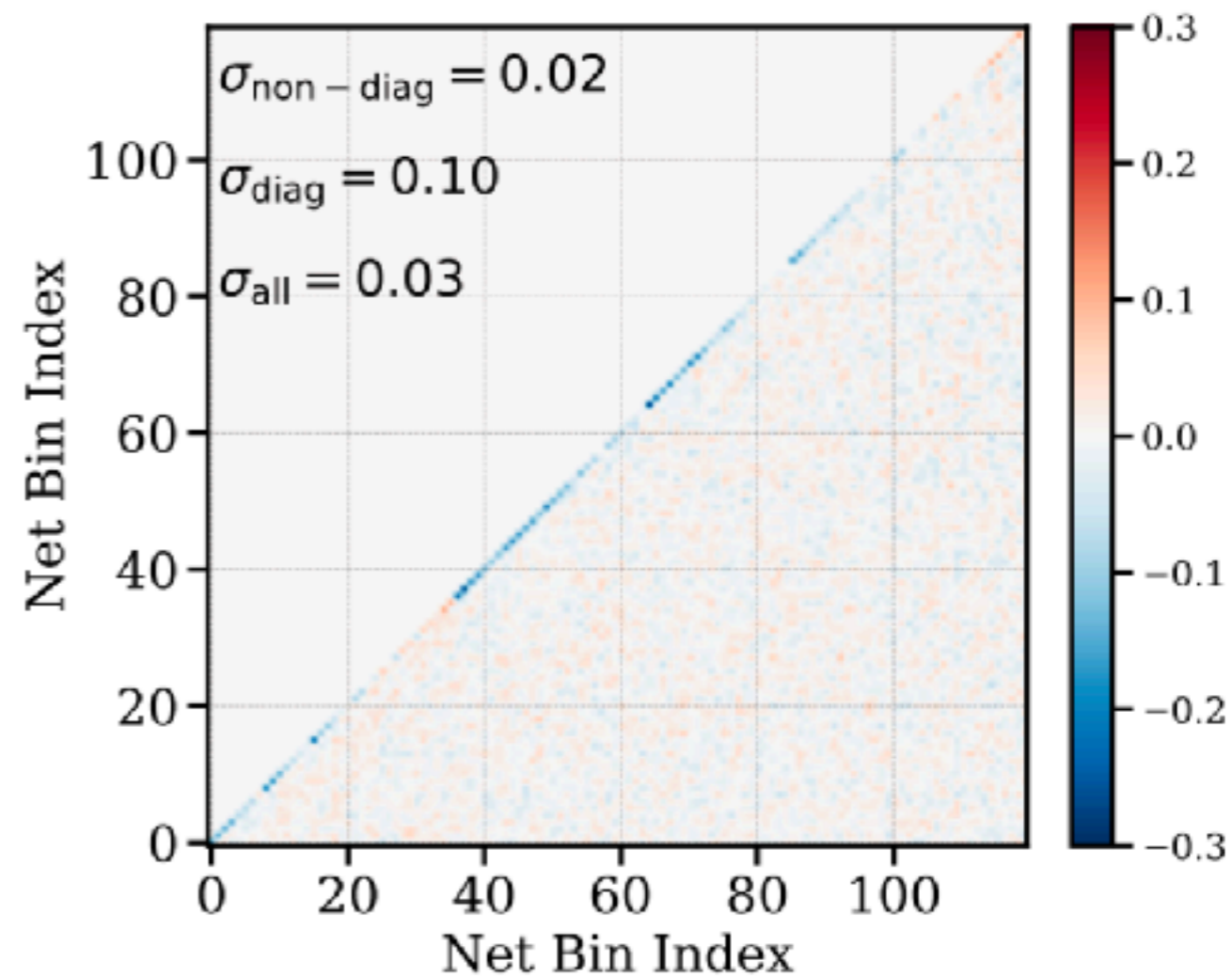


Half-Inverse Test

$$\mathbf{C}_{\text{analyt}}^{-1/2} \mathbf{C}_{\text{Patchy}}^{\text{CMASS}} \mathbf{C}_{\text{analyt}}^{-1/2} - \mathbf{1}$$



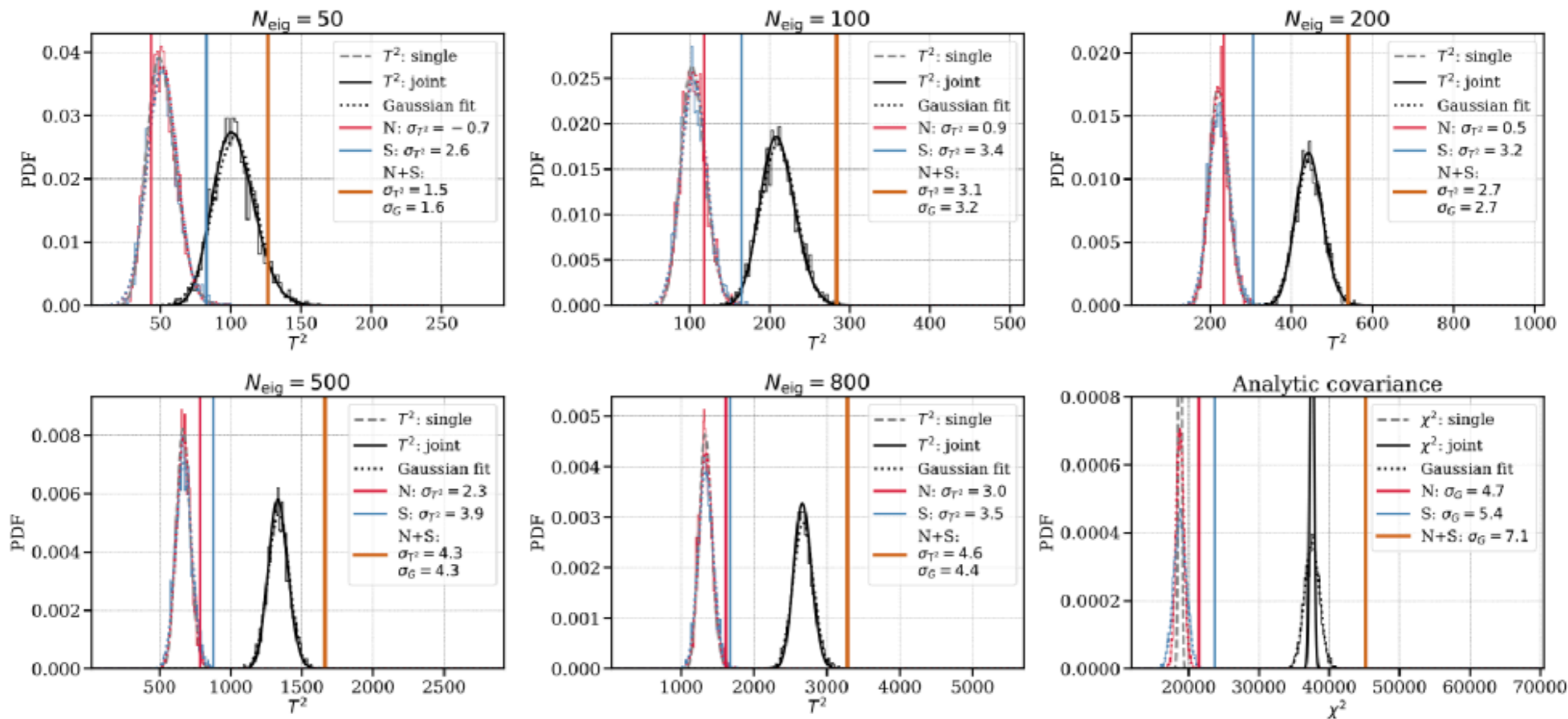
$$\mathbf{C}_{\text{analyt}}^{-1/2} \mathbf{C}_{\text{Patchy}}^{\text{LOWZ}} \mathbf{C}_{\text{analyt}}^{-1/2} - \mathbf{1}$$



BOSS Results

CMASS

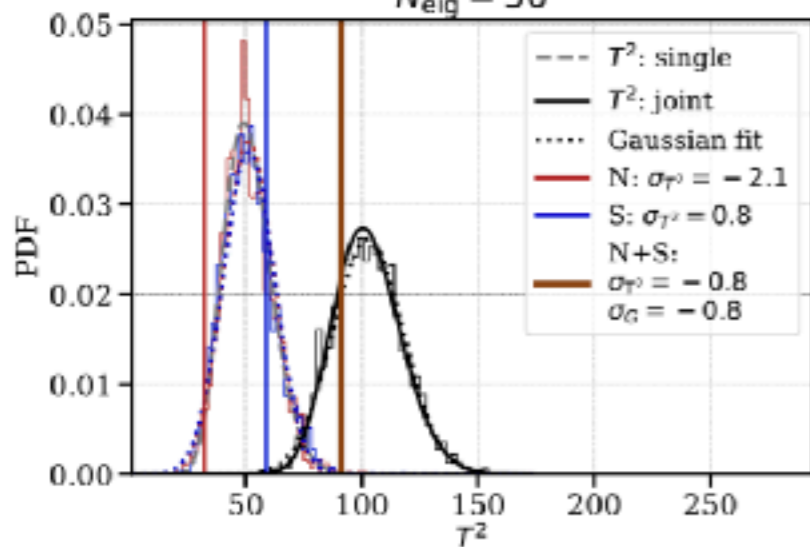
CMASS, 18 bins, $\ell_{\max} = 4$



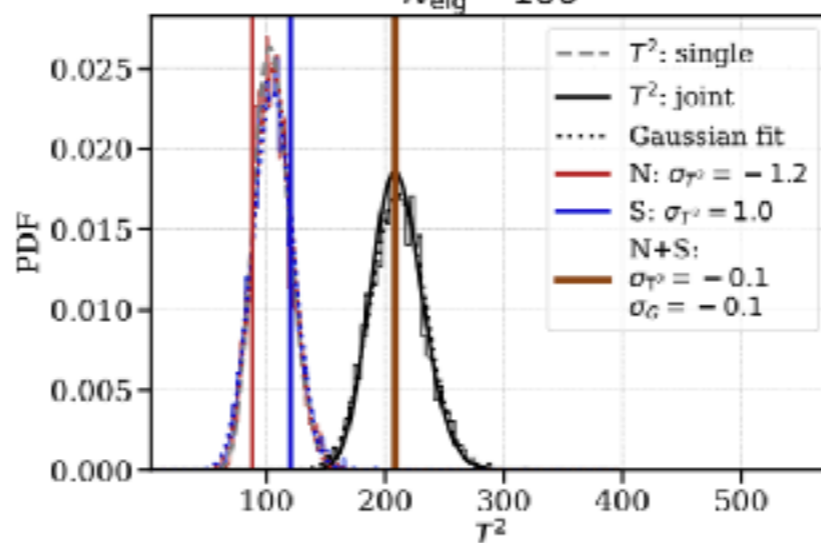
LOWZ

LOWZ, 18 bins, $l_{\max} = 4$

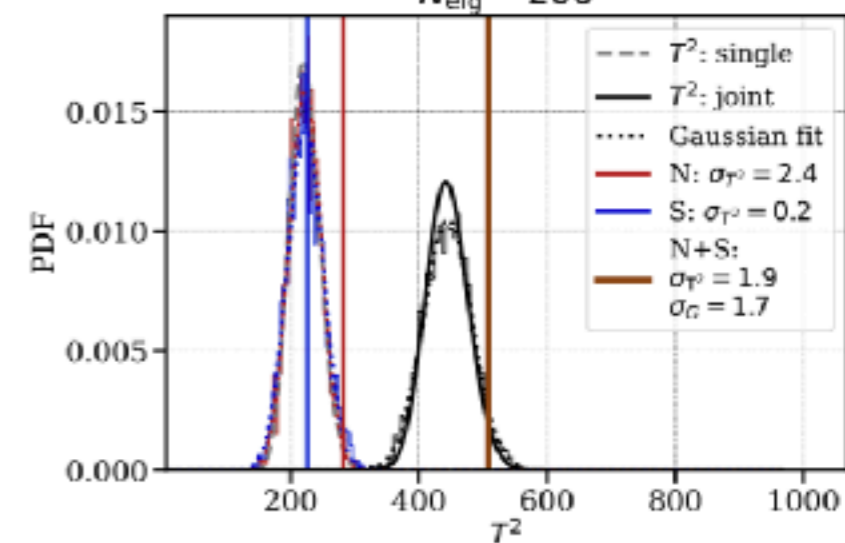
$N_{\text{eig}} = 50$



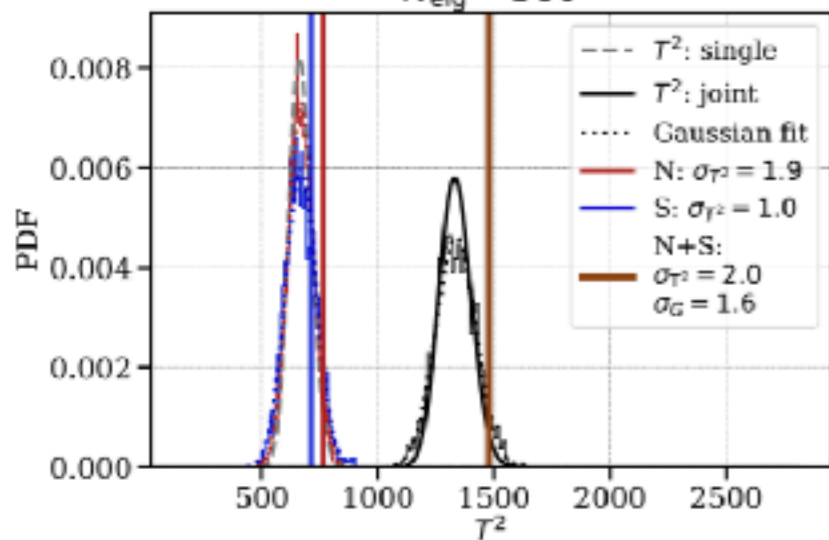
$N_{\text{eig}} = 100$



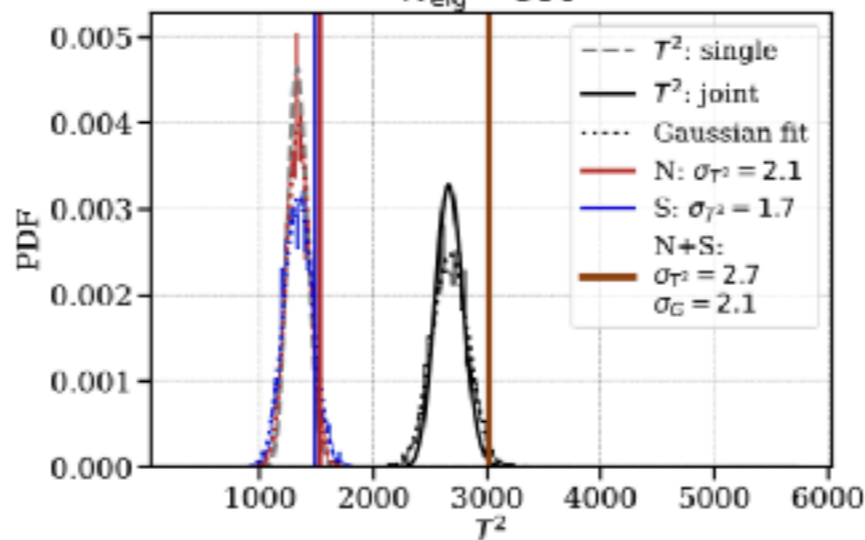
$N_{\text{eig}} = 200$



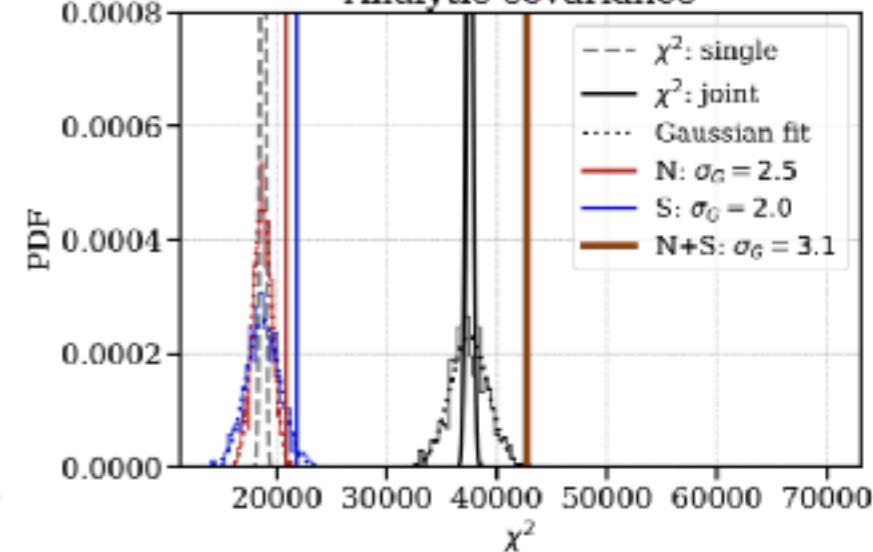
$N_{\text{eig}} = 500$



$N_{\text{eig}} = 800$



Analytic covariance



Summary Table

| Bins | σ_{eff} , Detection Significance | CMASS | | LOWZ | |
|------|---|-------|-----|------|-----|
| | | NGC | SGC | NGC | SGC |
| 18 | split | 4.7 | 5.4 | 2.5 | 2.0 |
| | joint | 7.1 | | 3.1 | |
| 10 | split | 3.3 | 2.6 | 1.8 | 1.9 |
| | joint | 4.0 | | 2.7 | |
| 6 | split | -0.7 | 1.2 | - | - |
| | joint | 0.4 | | - | |

Self-Calibration Using Even-Parity Sector

Demand consistency in even-parity modes ->
control covariance to prevent mismatch in true
C of data vs. **C** of mocks

| Consistency in Even-Parity Sector | CMASS 18 bins | Rescaling Factor | Rescaled Odd Detection Significance |
|--------------------------------------|------------------------|---------------------|--|
| 1 standard deviation | Analytic Covariance | 0.88 | 2.0σ |
| | $N_{\text{eig}} = 800$ | 0.98 | 4.0σ |
| 3 standard deviation | Analytic Covariance | 0.94 | 4.6σ |
| | $N_{\text{eig}} = 800$ | - | 4.4σ |

Systematics

Explored a large number

Few of the most interesting/salient ones

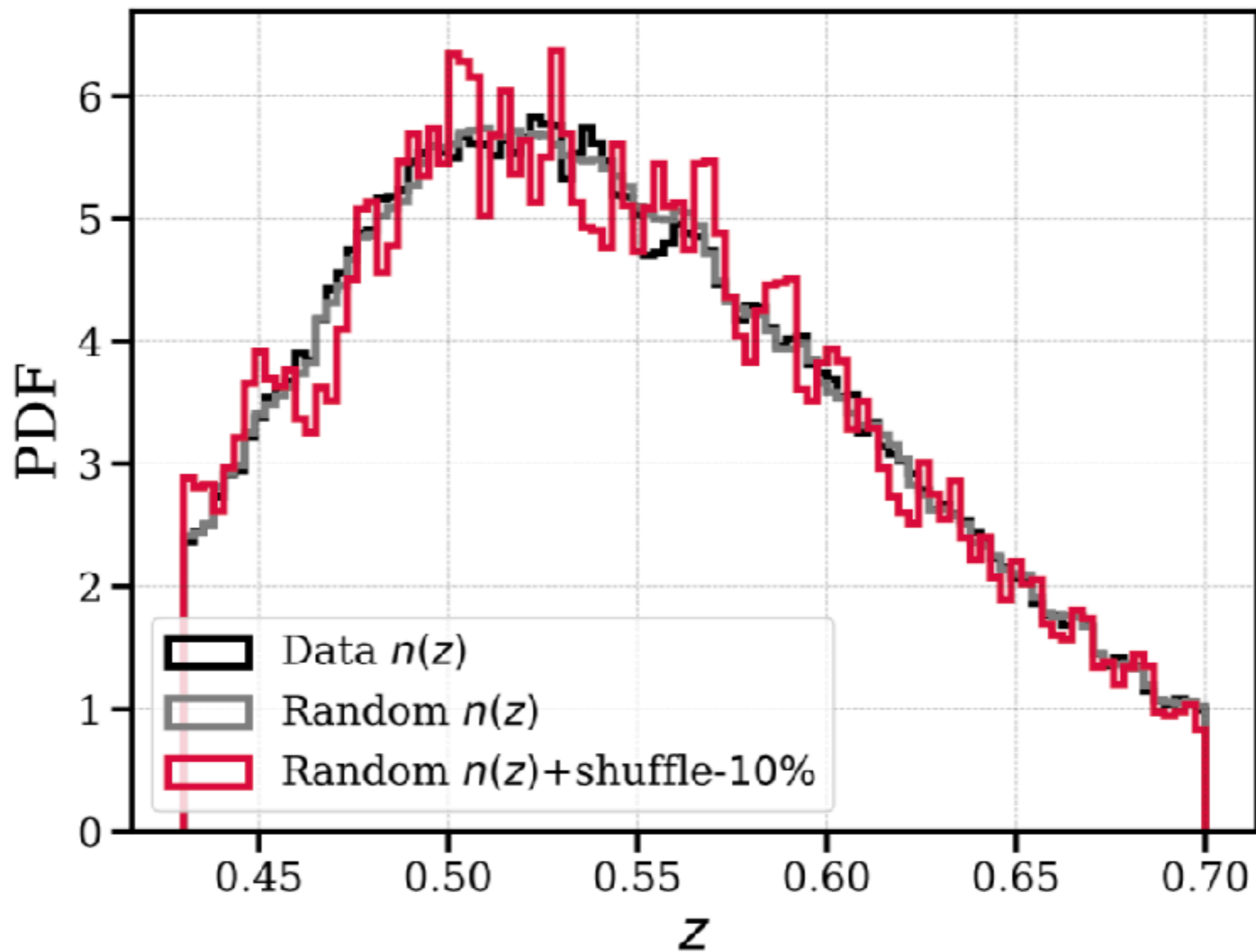
Redshift Errors

Other Weights Being Wrong

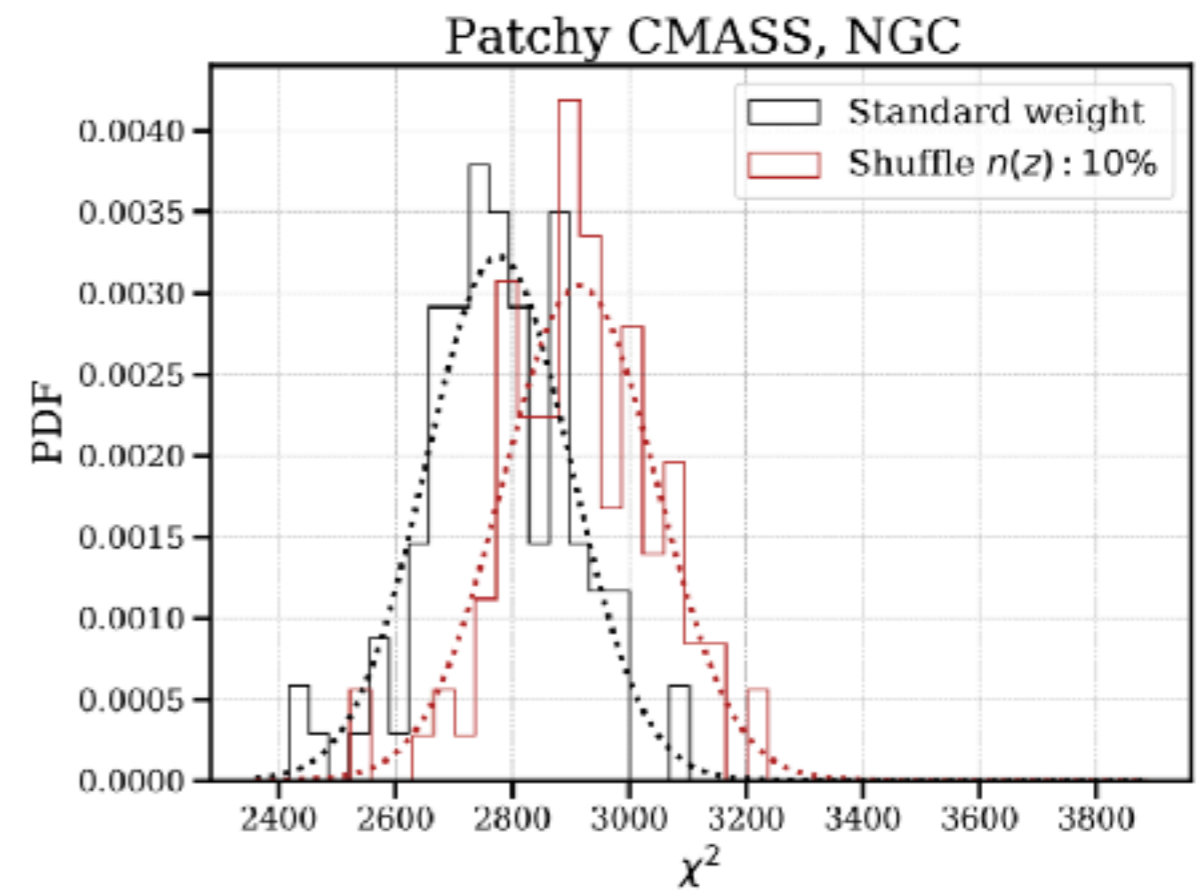
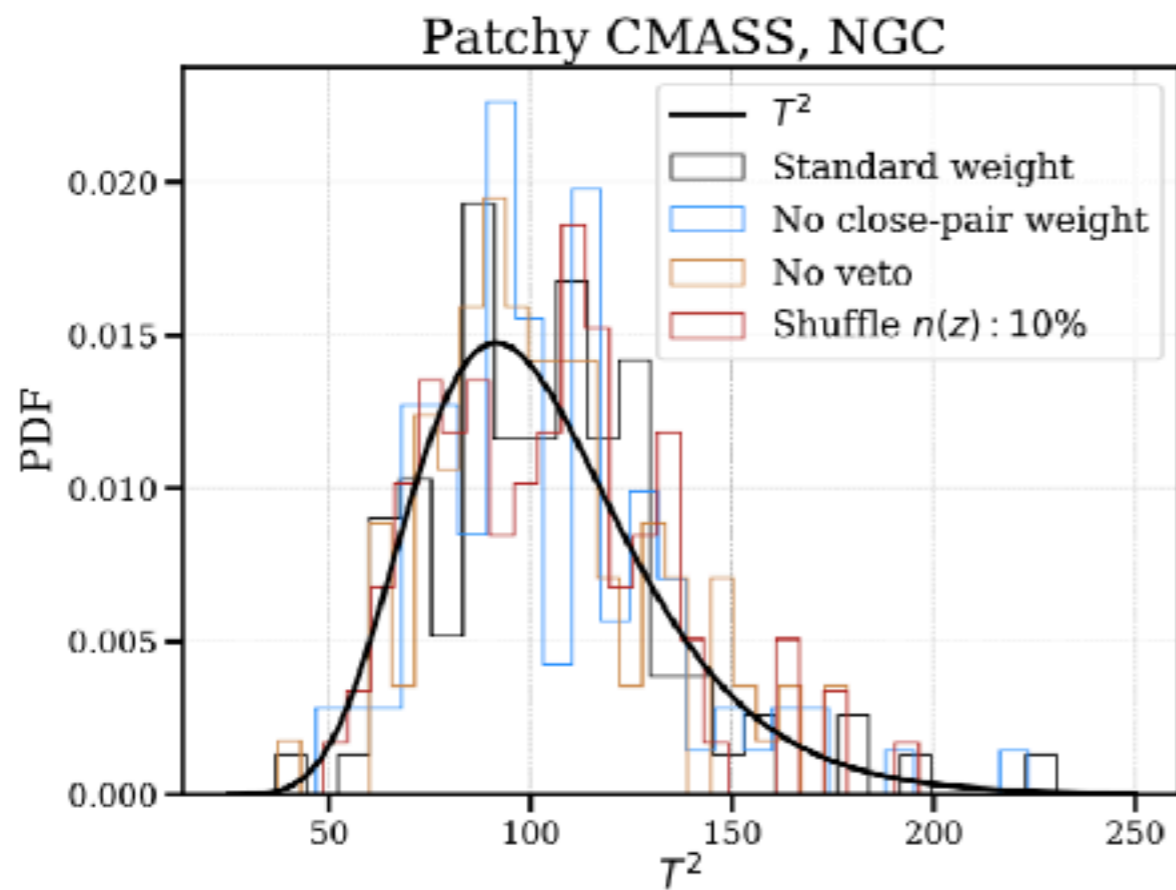
Split Sky

Fiber-magnitude and plate-location
dependent redshift failures

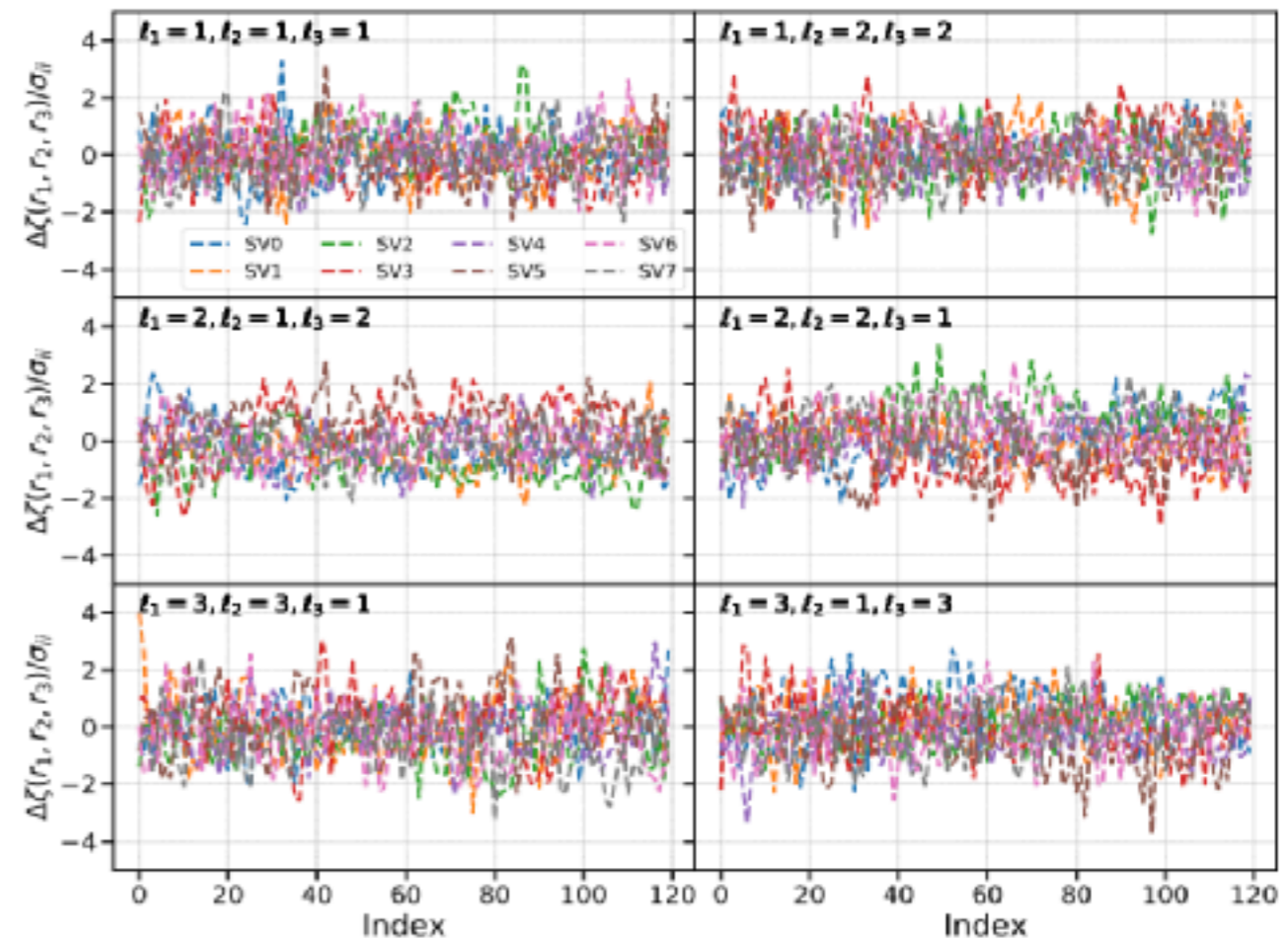
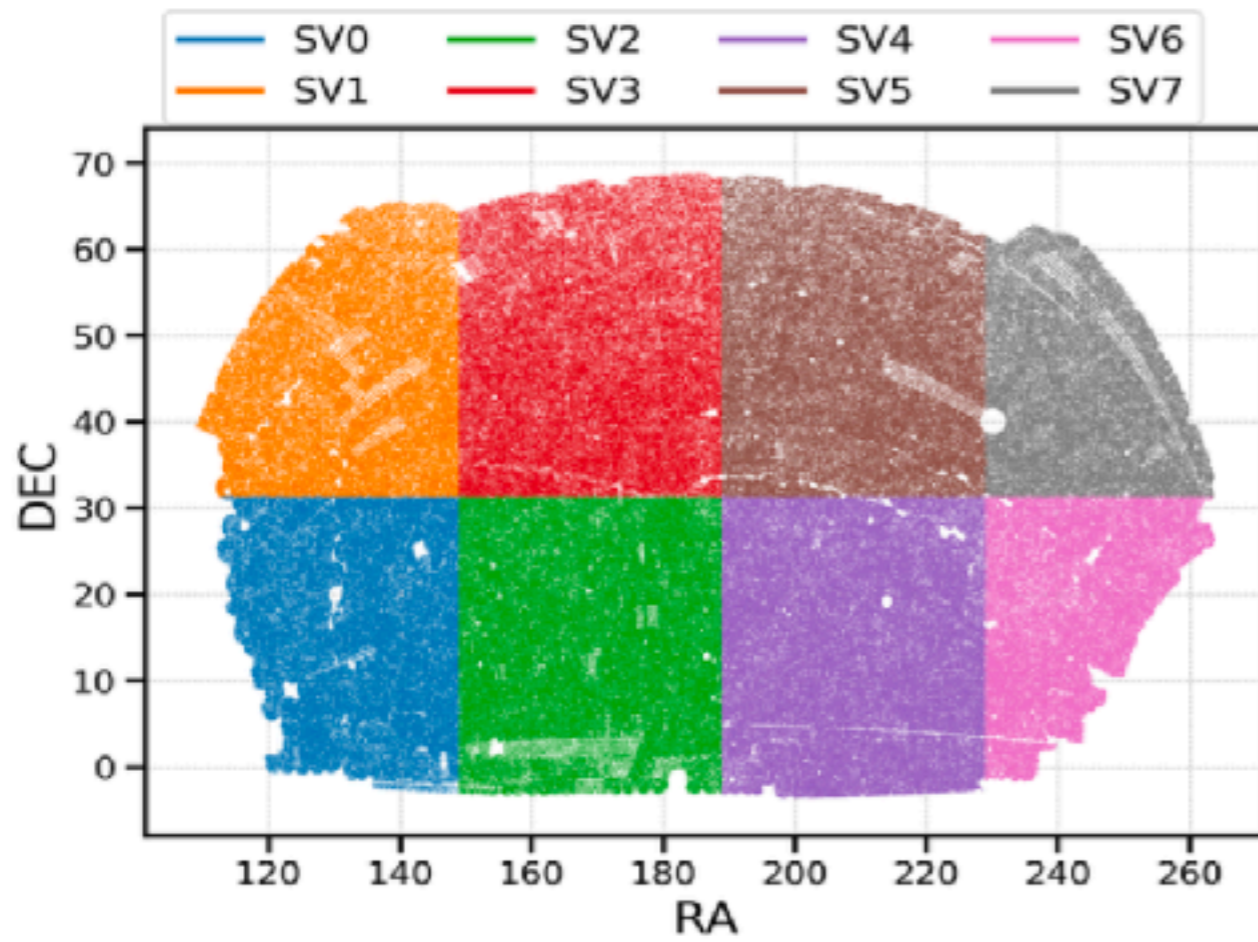
Redshift Errors



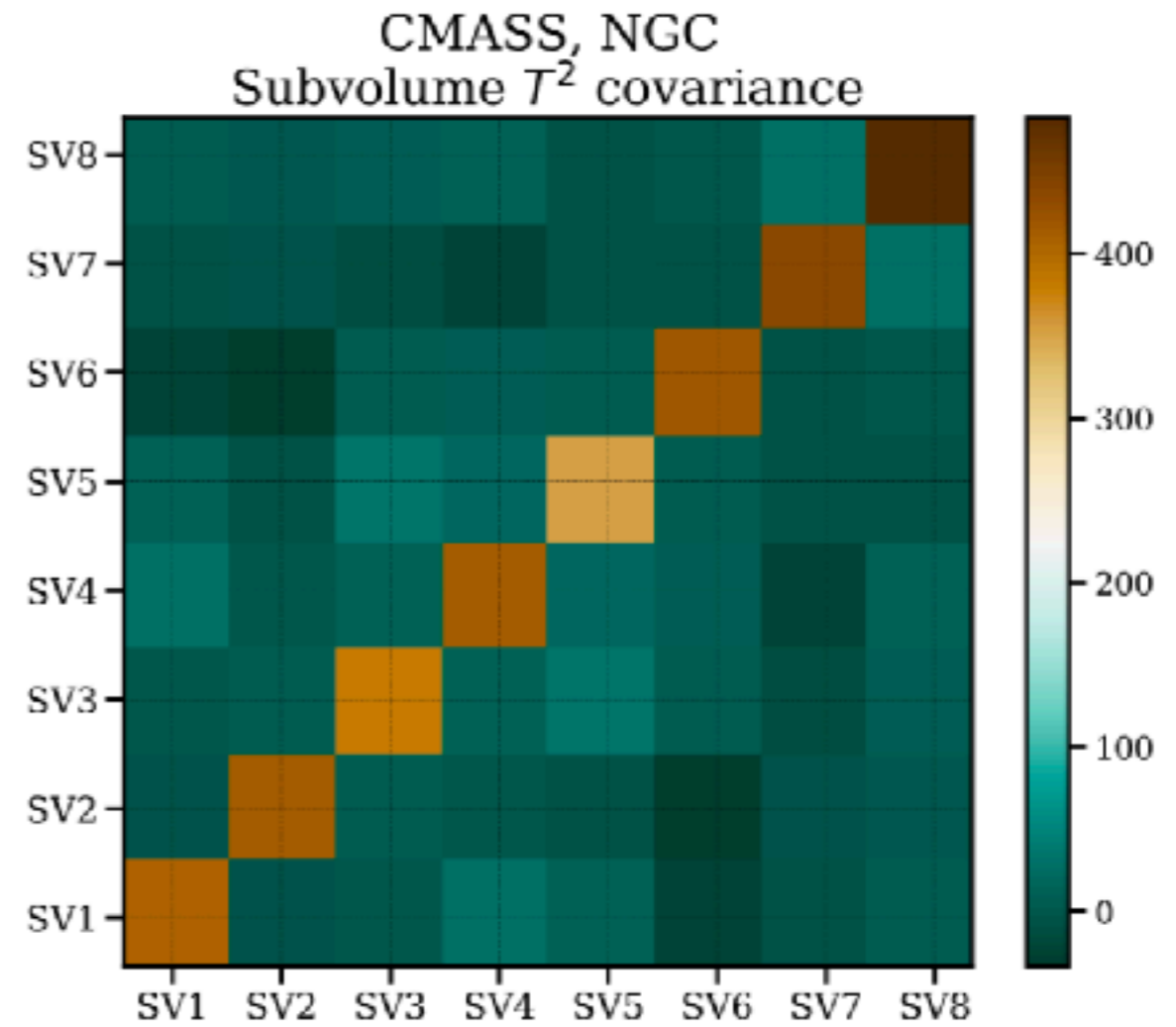
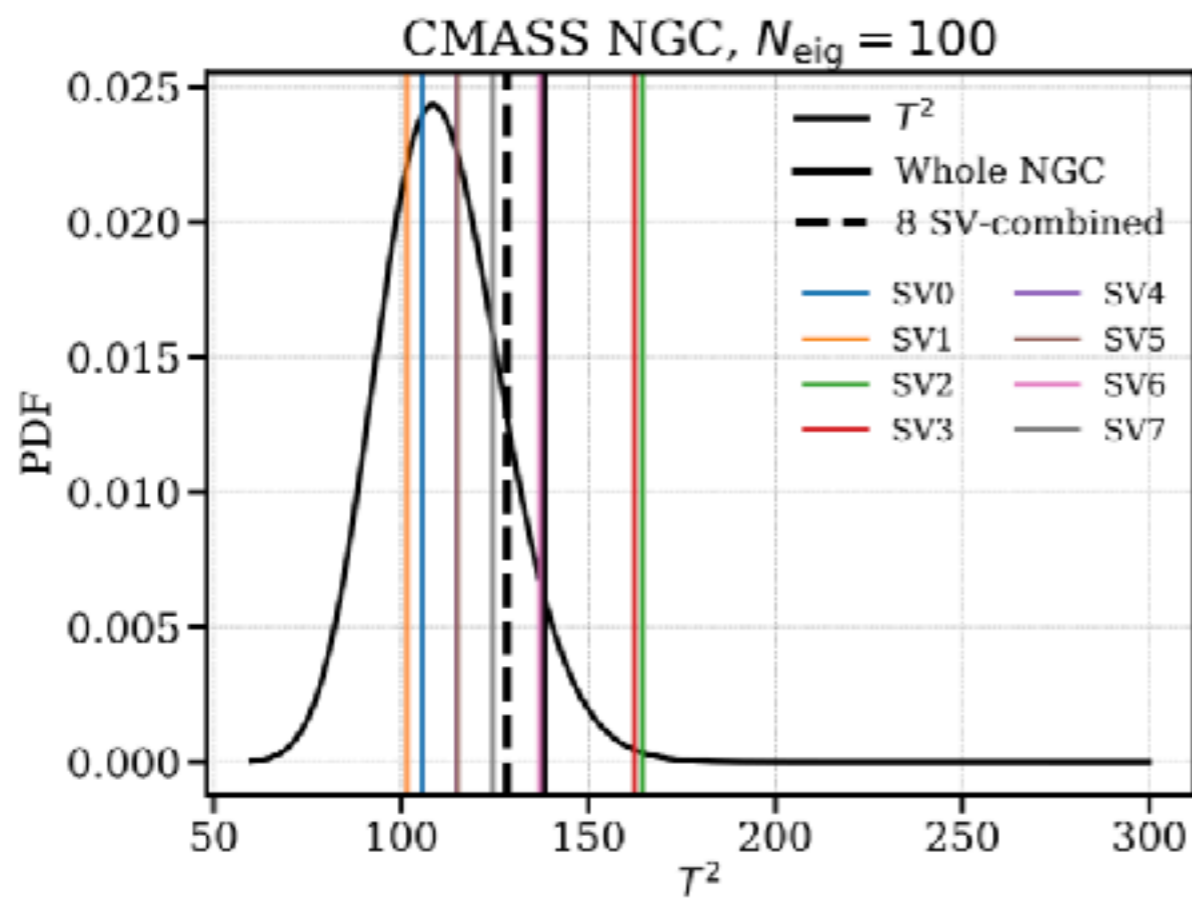
Redshift Errors & Other Weights



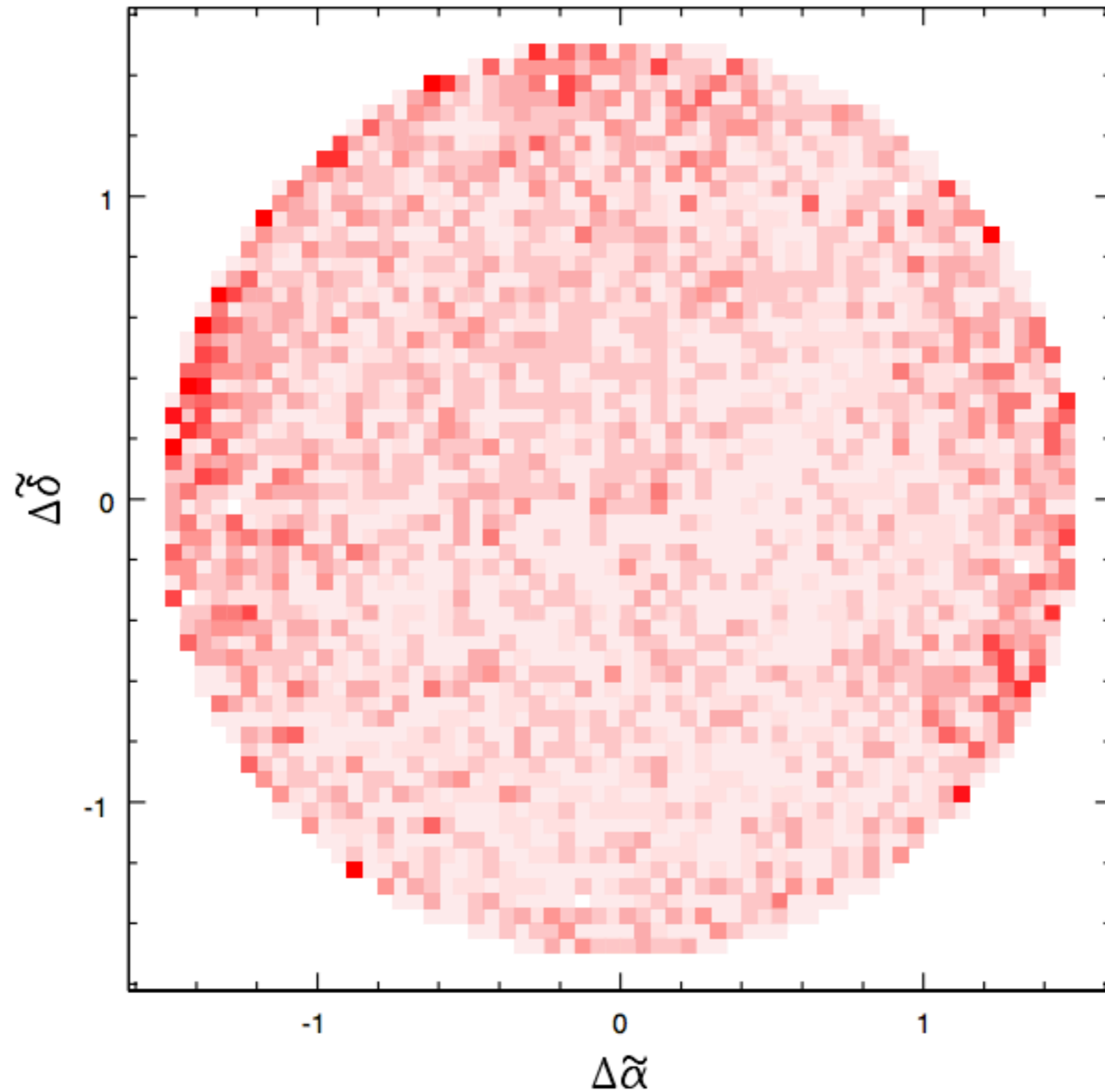
Split Sky



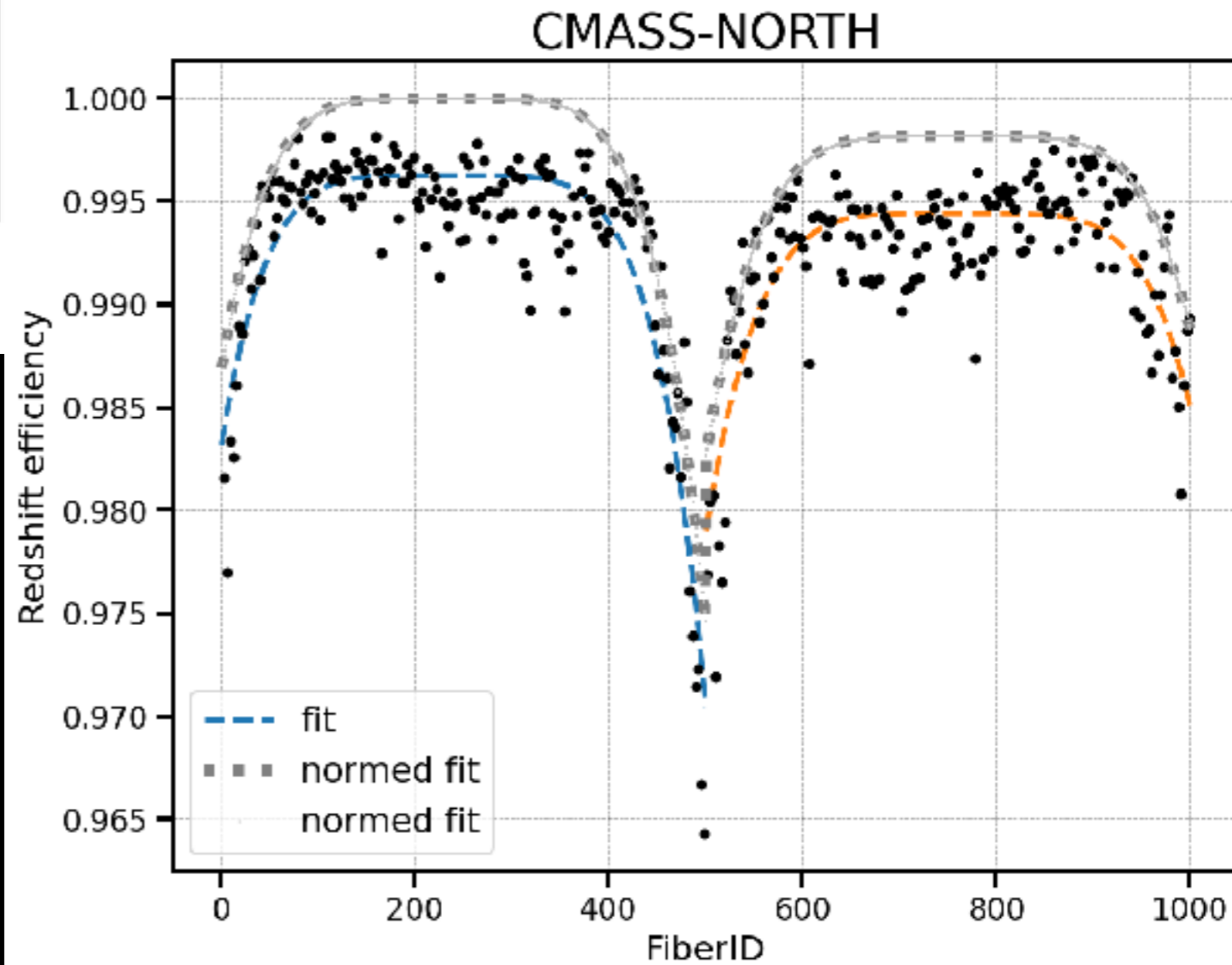
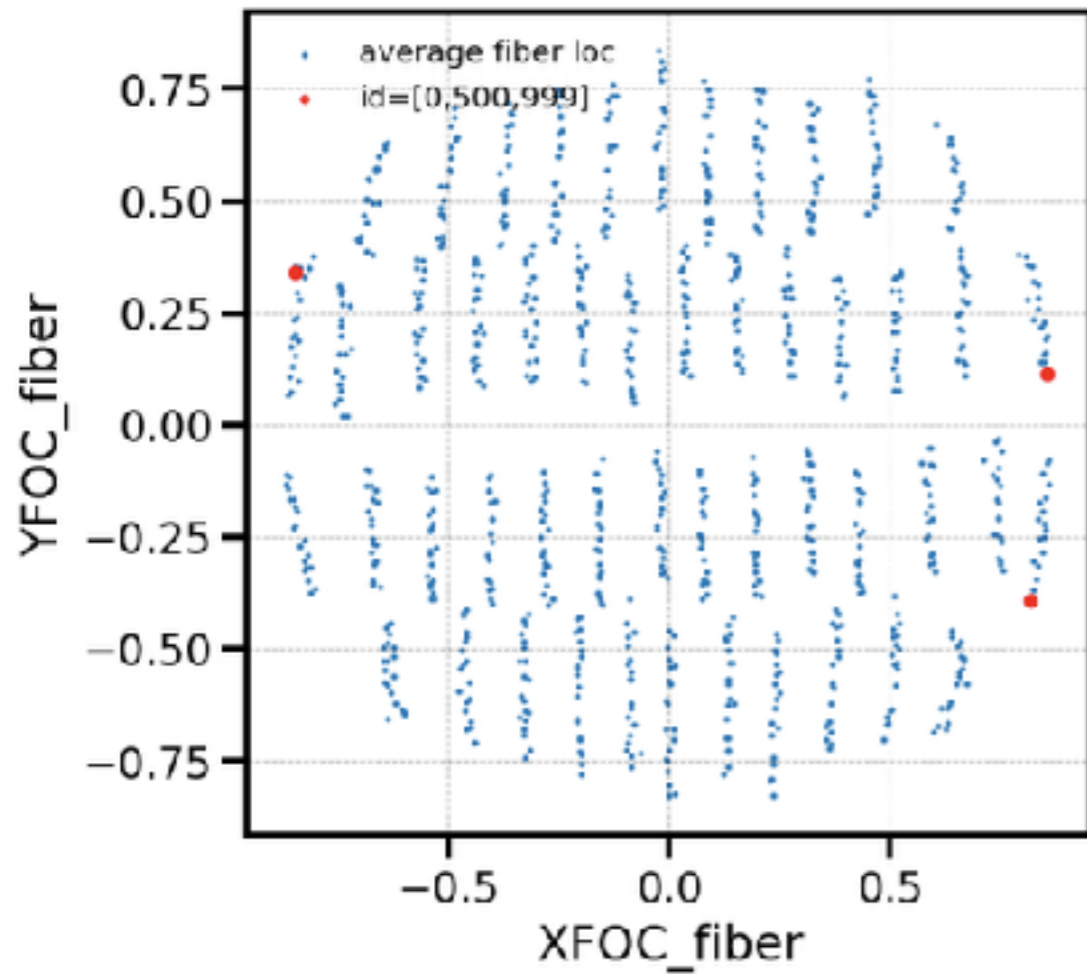
Split Sky



Redshift Failures



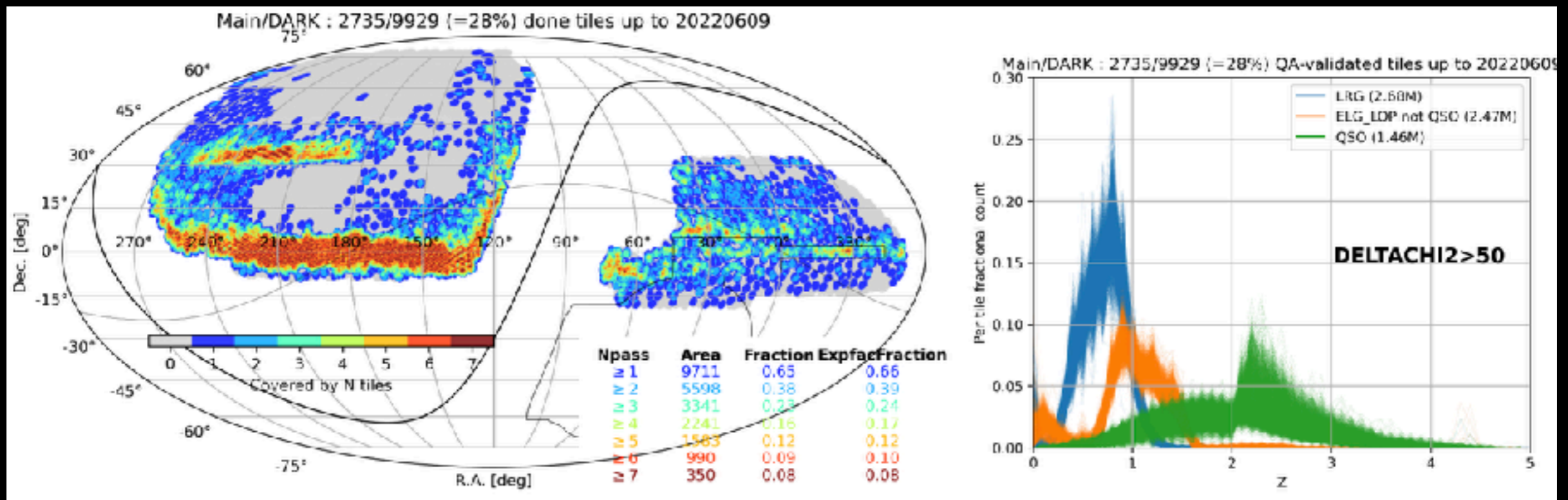
Redshift Failures



Systematics Summary

| Parity | Number of Bins | Redshift Failure | Fiducial Cosmology | Distortion in Selection Function |
|--------|----------------|------------------|--------------------|----------------------------------|
| Odd | 10 | 1.5 | ~ 1.5 | ~ 1.0 |
| | 18 | 2.5 | ~ 2.5 | ~ 1.7 |
| Even | 10 | 0.9 | ~ 0.7 | ~ 0.6 |
| | 18 | 1.8 | ~ 1.9 | ~ 2.4 |

Outlook

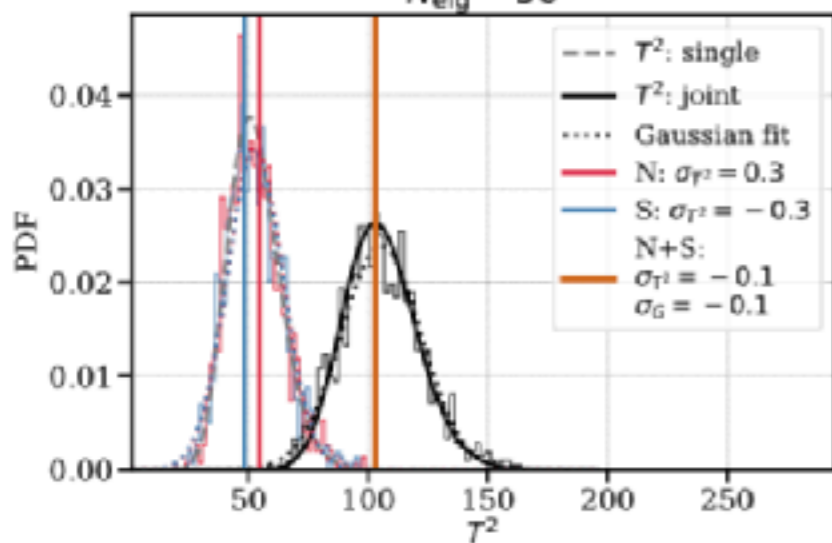


Thanks!

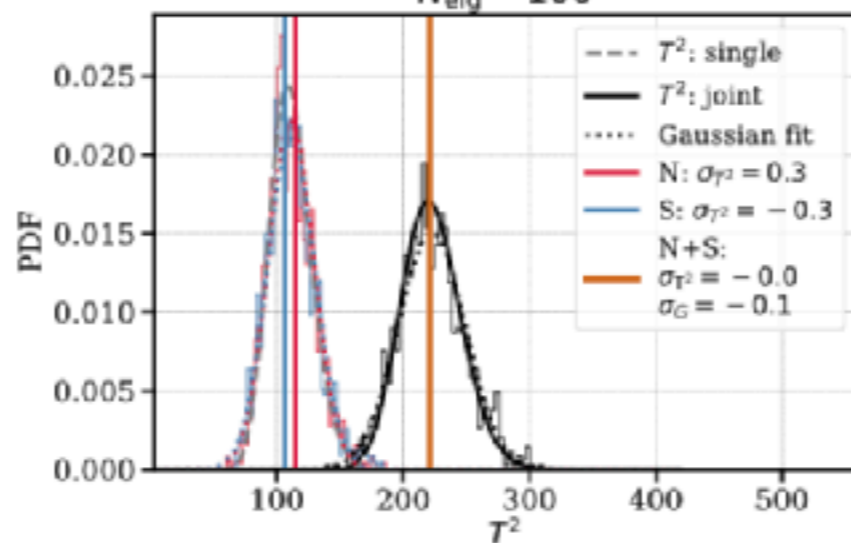


CMASS, 6 bins, $\ell_{\max} = 4$

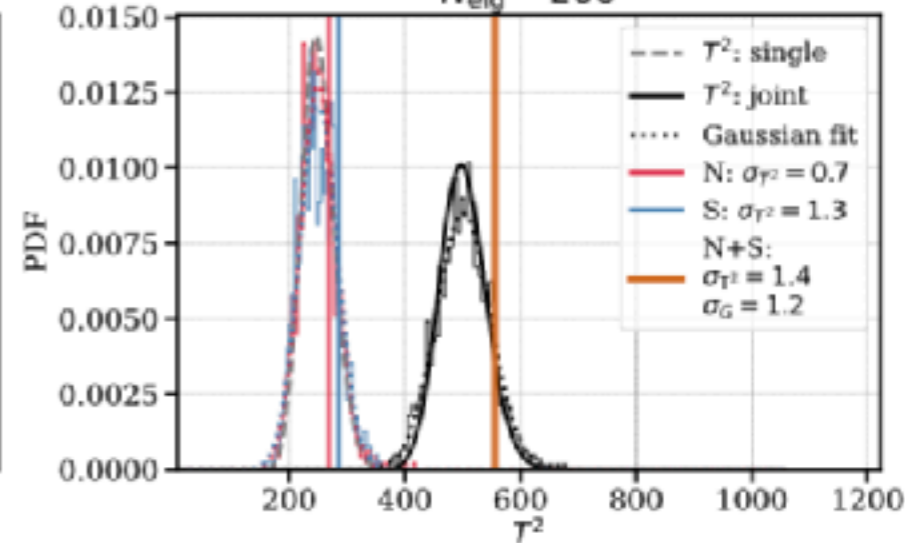
$N_{\text{eig}} = 50$



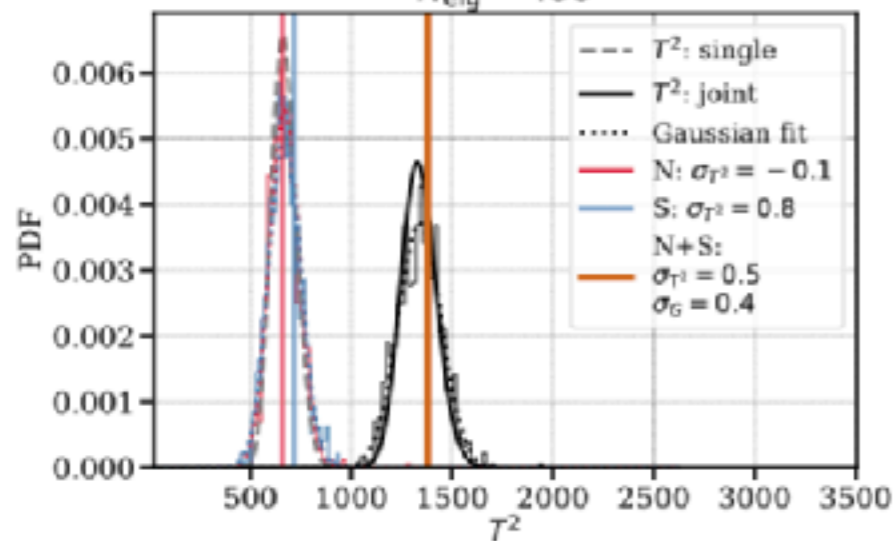
$N_{\text{eig}} = 100$



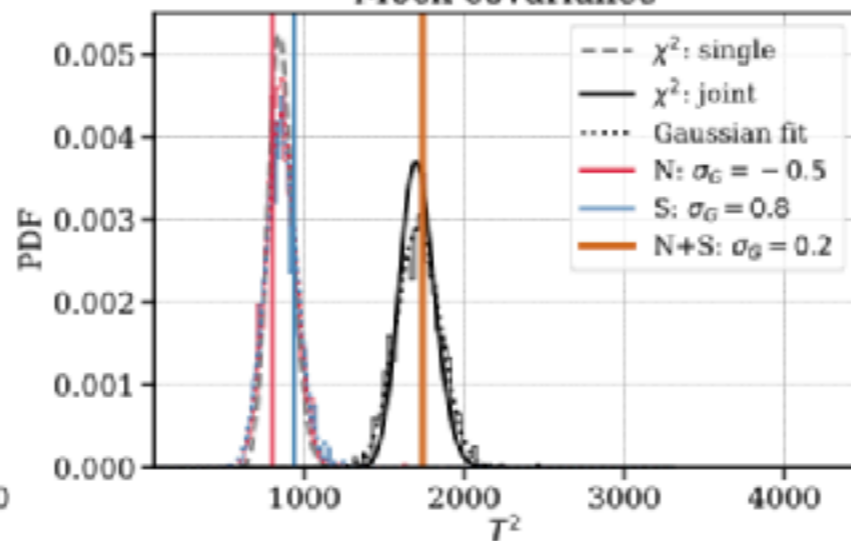
$N_{\text{eig}} = 200$



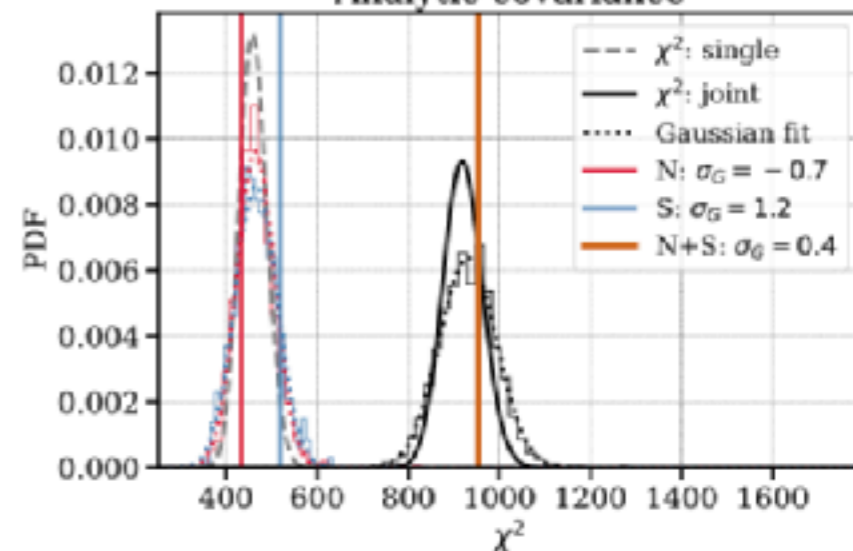
$N_{\text{eig}} = 400$



Mock covariance



Analytic covariance



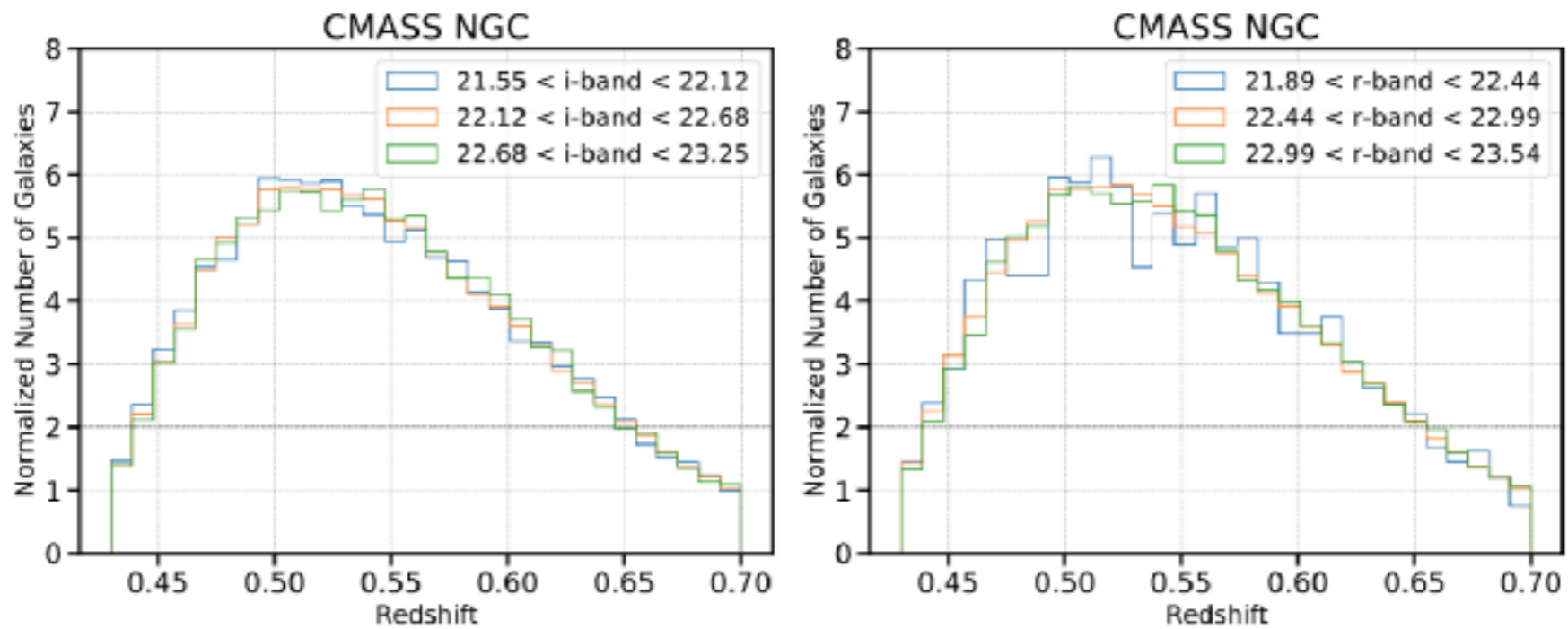


Figure D1. Redshift dependence on imaging depth for CMASS NGC. *Left:* The normalized galaxy number counts as a function of redshift for the *i*-band, where we split the sample into three bins in imaging depth. *Right:* Same as the left but for the *r*-band. Here we can see that the three bins in *i*- and *r*-band have very similar $n(z)$, which implies that the imaging depth is unlikely to impact our analysis.

