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### **SBND Detector Status and First Results**

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Search for sterile neutrino with  $\sim$  eV mass scale

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- Is the low energy access coming from electrons or photons?  $\rightarrow$  Improved detector technology!!!



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SBND BINE SOUICE		Active Mass (ton)	Baseline (m)	Operation
Active mass: 112 t Distance: 110 m Operation: 2024- MicroBooNE	SBND	112	110	2024 -
Active mass: 89 t Distance: 470 m Operation: 2015-2021	MicroBooNE	89	470	2015 - 2021
Active mass: 476 t Distance: 600 m Operation: 2021-	ICARUS	476	600	2021 -

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- Liquid argon time projection chamber (LArTPC) detector technology
- Three LArTPCs at different baselines
  - Combined analyses to maximize sensitivity for eV-scale sterile neutrino: also  $\nu_{\mu}$  disappearance
- A wide physics program includes precise  $\nu$ -Ar interaction studies and BSM searches (dark matter, HNLs)



#### Short Baseline Neutrino Detector (SBND)

Proximity: 110 m from the BNB target

- High neutrino flux  $\rightarrow$  high neutrino event rate  $\rightarrow$  will record the world's largest  $\nu$ -Ar scattering data
  - ~3 years of operation  $\rightarrow$  ~6 million neutrino events





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NO*v*A: *Phys.Rev.D* 106 (2022) 3, 032004 T2K: Phys.Rev.D 108 (2023) 7, 072011





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- NO*v*A: *Phys.Rev.D* 106 (2022) 3, 032004 T2K: Phys.Rev.D 108 (2023) 7, 072011
- Excellent sensitivity to BSM particles as a beam dump experiment





#### SBND: (1) Subsystems: (a) Cryostat and Cryogenics

Under 1 atm, argon boiling point is 87.3 K and freezing point is 83.8 K  $\,$ 

• Utilize membrane cryostat which is an LNG industry technology: same with DUNE!







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- A very successful collaboration between CERN and Fermilab engineers and physicists







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- A field cage for uniform 500 V/cm E-field
- Three readout wire planes for each TPC: one collection (vertical) and two induction ( $\pm 60^{\circ}$ ) with 3 mm pitch
- TPC cold electronics
  - Shaping, amplification and digitization in LAr (~87 K) for 11,264 wires



#### SBND: (1) Subsystems: (c) Photon detection system (PDS)

PDS)

120 PMTs (Hamamatsu R5912-mod 8") and 192 X-ARAPUCA devices

- Mixture of coated and uncoated devices
   \*Coated: for wavelength shifting of scintillating photons (λ ~ 127 nm) to photocathode sensible regions (λ ~ 440 nm)
   Reflective foil panels coated with wavelength shifter embedded into the cathode plane
- Coated ones could collect direct VUV ( $\lambda$  ~ 127 nm), uncoated ones collect only wavelength shifted photons

Timing resolution is  $\sim O(1 \text{ ns})$ : essential for the trigger system











Cosmic ray tagger (CRT)

• Surrounds the SBND cryostat on all 6 sides

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Cosmic ray tagger (CRT)

- Surrounds the SBND cryostat on all 6 sides
- Built from 11.2 cm wide scintillator strips
  - Wavelength shift (WLS) fibers
  - Silicon photomultipliers (SiPMs) for light collection





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December 2022

• Detector was fully assembled and moved to the detector hall (SBN-ND)







Dec. 2022 Fully assembled Moved to the detector hall



April 2023

Moved the detector into the cryostat



IIDec. 2022April. 2023Fully assembledDetectorMoved to the detector hallinside cryostat





July 2023

Cathode HV feedthrough installed











"Warm" commissioning

- Install cables
- TPC readout: front-end mother board to disk
- Install purity monitors
- Verify detector control systems .....









March 2024

• Filled with liquid argon











SBND DETECTO

July 2024

Cathode voltage reached to the design target 100 kV



April 2023 July 2023 Detector inside cryostat Feedthrough

"Warm" Mar. 2024 Commissioning Filling LAr

July 2024 Cathode HV @ 100 kV

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September 2024

Dec. 2022

Fully assembled

Moved to the detector hall

Installation of top CRT layers finished

"Warm"

Commissioning

Mar. 2024

Filling LAr

**July 2024** 

Cathode HV

@ 100 kV

**July 2023** 

Cathode HV

Feedthrough

• The last pice of the SBND!

**April 2023** 

Detector

inside cryostat





Sep. 2024

Top CRT layers

installed

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#### SBND: (2) Toward Operation

February 2025

- Trigger menu for the Run1 is fixed
  - Includes trigger for BNB gate + light activity inside the TPC
  - BNB spill length (~ 1.6  $\mu s$ ) appears in PMT and CRT data



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#### SBND: (3) DAQ Performance

Collected data and DAQ efficiency

- From 2024/12/20 to 2025/02/13, SBND had recorded zero bias data using BNB gates
- Since Feb. 13th 2025, SBND is recording data with light-based trigger and BNB gates
- Data recording efficiency is very high: 98.4%
- Collected  $2.57 \times 10^{20}$  POT as of 2025/05/20  $\rightarrow$  already more than 1M neutrino events!!!





Noise

- For the longest (most noisy) wires, noise RMS is less than 3 ADCs in general
- Excellent signal to noise ratio in raw data: 32.5 for the collection plane



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- Impurities inside liquid argon absorb drifting electrons: design electron lifetime goal is 3 ms
- Observed electron lifetime is longer than 10 ms



#### SBND: (4) TPC Performance

Noise

- For the longest (most noisy) wires, noise RMS is less than 3 ADCs in general
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Electron lifetime

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Calorimetry

 Measured energy deposition per unit length (dE/dx) for muon and proton candidate tracks show good agreement with the most probable value (MPV) curve expected by the Landau-Vavilov theory





## SBND: (4) Timing Performance

Reproduction of the BNB spill structure

- For a spill (~1.6  $\mu \rm s$ ), there are 81 proton bunches
  - Bunch spacing ( $\Delta T_{bunch}$ ): 18.936 ns, bunch width ( $\sigma_{bunch}$ ): 1.308 ns
- Using BNB gate plus light trigger and CRT timing information
  - Reproduced  $\Delta T_{bunch}$  very well with 18.931 ns and all 81 bunches
  - Measured bunch width of 3.510 ns indicates that CRT has about 3.3 ns timing resolution





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The SBN and SBND offer world class physics program

- Search for sterile neutrino with  $\sim$  eV scale mass
- Precise  $\nu$ -Ar interaction studies
- BSM searches: dark matter, heavy neutral leptons, ...

SBND finished its commissioning at Summer 2024

- DAQ is stable: physics run since December 2024
- TPC: electron lifetime longer than the goal and good calorimetry performance with low noise level
- TPC, PDS and CRT are well integrated as a single detector: SBND!!!

First physics run of SBND is ongoing

- Since the last December, already more than 1 M neutrino events are recorded as of May 20th
- Physics results are coming soon. Stay tuned!





#### Thank you!!

**The SBND Collaboration** 



updated May 2025

Including both scientific and technical personnel Argonne National Lab, USA: Z. Djurcic, M. Goodman, A. Papadopoulou \*Spo kespeo ple University of Bern, Switzerland: S. Mulleriababu, M. Weber Brookhaven National Lab, USA: M. Bishai, M. Carneiro, H. Chen, J. Farrell, J. Fried, S. Gao, X. Qian, V. Radeka, E. Raguzin, J. Smith, C. Thorn, E. Worcester, M. Worcester, B. Yu, H. Yu, C. Zhang, M. Zhao University of California Santa Barbara, USA: A. Antonakis, S. Brickner, D. Caratelli, X. Luo University of Campinas, Brazil: P. Holanda, A. Machado, O. Peres, V. do Lago Pimentel, E. Segreto CERN: O. Beltramello, J. Bremer, M. Chalifour, A. de Roeck, L. Di Giulio, C. Fabre, J. Hrivnak, U. Kose, B. Lacarelle, D. Mladenov, M. Nessi, S. Palestini, F. Pietropaolo, X. Pons, F. Resnati, A. Rigamonti, E. Seletskaya, S. Tuflani, A. Zani University of Chicago, USA: A. Bhat, A. Ereditato, B. Fleming, D. Franco, L. Hagaman, M. Jung, M. King, N. Rowe, D. Schmitz\*, L. Tung, T. Wester, A. White CIEMAT, Spain: R. Alvarez-Garrote, J. Crespo-Anadón, C. Cuesta, S. Dominguez-Vidales, I. Gil-Botella, J. Romeo-Araujo Colorado State University, USA: D. Carber, L. Kashur, R. Lazur, M. Mooney, D. Totani Columbia University: L. Arnold, L. Camilleri, C. Chi, S. Chung, D. Kalra, G. Karagiorgi, N. Oza, M. Ross-Lonergan, M.H. Shaevitz, B. Sippach University of Edinburgh: A. Hamer, L. Kotsiopoulou, M. Nebot-Guinot, H. Parkinson, M. Reggiani-Guzzo, A. Szelc Federal University of ABC, Brazil: C. Moura Federal University of Alfenas, Brazil: G. Valdiviesso Instituto Tecnológico de Aeronáutica, Brazil: F. Marinho, L. Paulucci Fermilab: R. Acciarri, W. Badgett, L. Bagby, V. Basque, M. Betancourt, D. Caratelli, F. Cavanna, O. Dalager, M. Del Tutto, V. Di Benedetto, S. Dixon, S. Dytman, S. Gardiner, M. Geynisman, H. Greenlee, S. M. Kancharla, C. James, T. Junk, W. Ketchum, M.J. Kim, J.Y. Li, L. Liu, P. Machado, M. Micheli, D. Montanari, J. Mueller, T. Nichols, B. Norris, M. Nunes, S. Oh, O. Palamara\*, J. Paton, V. Pandey, Z. Pavlovic, D. Pushka, G. Putnam, A. Schukraft, S. Shetty, M. Stancari, A. Stefanik, T. Strauss, D. Torretta, M. Toups, L. Wan, P. Wilson, L. Yates, J. Zennamo University of Florida: B. Carlson, C. Fan, I. Furic, H. Ray University of Granada: D. Garcia Gamez, L. Pelegrina-Gutierrez, A. Sanchez-Castillo, P. Sanchez-Lucas, A. Vazquez-Ramos, B. Zamorano Illinois Institute of Technology: D. Andrade-Aldana, M. Hernandez-Morquecho, B. Littlejohn, J. Mclaughlin Imperial College London: A. Navrer-Agasson, P. Hamilton, S. Söldner-Rembold University of Kansas: M. Andriamirado, M.B. Brunetti Lancaster University: A. Blake, R. Coacklev, D. Brailsford, B. McCusker, J. Nowak, P.N. Ratoff University of Liverpool: C. Andreopoulos, K. Mavrokoridis, D. Payne, J. Plows, M. Roda, B.A. Slater, C. Touramanis Los Alamos National Lab: W. Foreman, S. Gollapinni, W.C. Louis, A. Schneider, R.G. Van de Water, E. Yandel Louisiana State University: E. Belchior, P. Singh, H. Wei University of Manchester: J. Bateman, J. Evans, N. Lane, C. Thorpe University of Michigan: B. Bogart, J. Spitz University of Minnesota: A. Furmanski, N. Pallat Mount Holyoke College: S. Balasubramanian University of Oxford: A. Barnard, K. Duffy, P. Green University of Pennsylvania: M. Dubnowski, J. Klein, T. Kroupova, J. Sensenig Queen Mary University, UK: Y. Dabburi, N. McConkey Rutgers University: K. Lin, A. Mastbaum University of Sheffield: A. Beever, A. Ezeribe, R. Jones, V.A. Kudryavtsev, H. Lav, A. Moor, L. Nguyen, H. Scott University of Sussex: R. Darby, C. Griffith, S. Kr Das Syracuse University: A. Filkins, R. Rajagopalan, M. Soderberg Texas A&M University: K. Kelly University of Texas, Arlington: L. Aliaga-Soplin, A. Brandt, R. Castillo, M. Dall'Olio, F. J. Nicolas-Arnaldos, S. Yadav, J. Yu Tufts University: O. Alterkait, Z. Imani, T. Wongjirad Virginia Tech: C. Mariani, G. Moreno-Granados, P. Roy University of Warwick: A. Chappell, X. Lu, J. Marshall, A. Wilkinson

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# **‡**Fermilab





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• Both detectors are filled with mineral oil as scintillating material, and ~ 1k PMTs for light detection







There could be

-  $\nu_{\rm e}$  (dis)appearance,  $\nu_{\mu}$  disappearance

$$U_{3+1} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{bmatrix}$$

$$\begin{split} \sin^2(2\theta_{\mu\mu}) &= 4(1 - |U_{\mu4}|^2)|U_{\mu4}|^2,\\ \sin^2(2\theta_{ee}) &= 4(1 - |U_{e4}|^2)|U_{e4}|^2,\\ \sin^2(2\theta_{e\mu}) &= 4|U_{e4}|^2|U_{\mu4}|^2, \end{split}$$

$$P(\nu_{\mu} \rightarrow \nu_{e}) = \sin^{2}2\theta_{\mu e}\sin^{2}(\Delta m_{41}^{2}L/E),$$
  

$$P(\nu_{e} \rightarrow \nu_{e}) = 1 - \sin^{2}2\theta_{ee}\sin^{2}(\Delta m_{41}^{2}L/E),$$
  

$$P(\nu_{\mu} \rightarrow \nu_{\mu}) = 1 - \sin^{2}2\theta_{\mu\mu}\sin^{2}(\Delta m_{41}^{2}L/E),$$





Spills with 5 Hz rate

- A spill is about 1.6  $\mu$ s long
- A spill consists of 81 bunches (protons packet)
  - A bunch consists of ~  $5 \times 10^{10}$  protons
  - With 18.936  $\pm$  0.001 ns spacing: 81 x 18.936 = 1533.816 ns ~ 1.53  $\mu$ s
  - With 1.308  $\pm$  0.001 ns width
- Injection energy: 400 MeV
- Extraction energy:8 GeV



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7 pulses of 15 Hz in a raw delivered at 0.769 Hz  $\rightarrow$  (Rate)= 5.3 Hz

