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QUANTUM TOMOGRAPHY
OF CHIRAL MEDIA:
FROM QUARK-GLUON
PLASMA TO WEYL
SEMIMETALS TO
AXIONS

CIPANP 2025, Madison, WI

9 June 2025

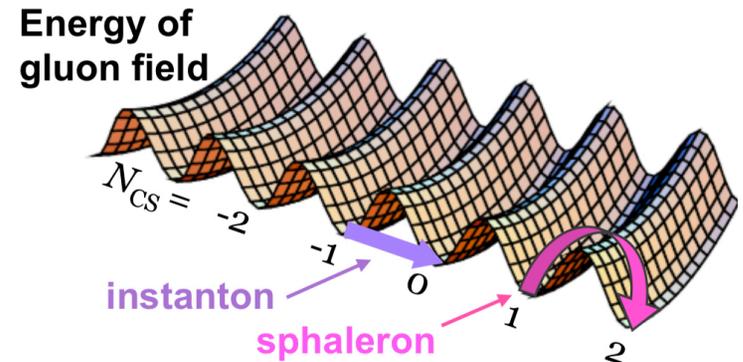
CHIRAL MAGNETIC EFFECT

Sphaleron transitions in hot nuclear matter \Rightarrow
chiral imbalance \Rightarrow Chiral magnetic effect:

$$\mathbf{j} = b_0 \mathbf{B} \quad b_0 = \frac{e^2 \mu_5}{2\pi^2}$$

Kharzeev, McLerran Warringa (2008), Fukushima, Kharzeev, Warringa (2008)

$$b_0 \propto \mu_5 = \text{chiral chemical potential}$$



How does the chiral magnetic current impact jets in QGP?

Closely related problems:

- Fast charges in Weyl semimetals
- Cosmic rays in the axion field

GLUON/PHOTON DISPERSION RELATION

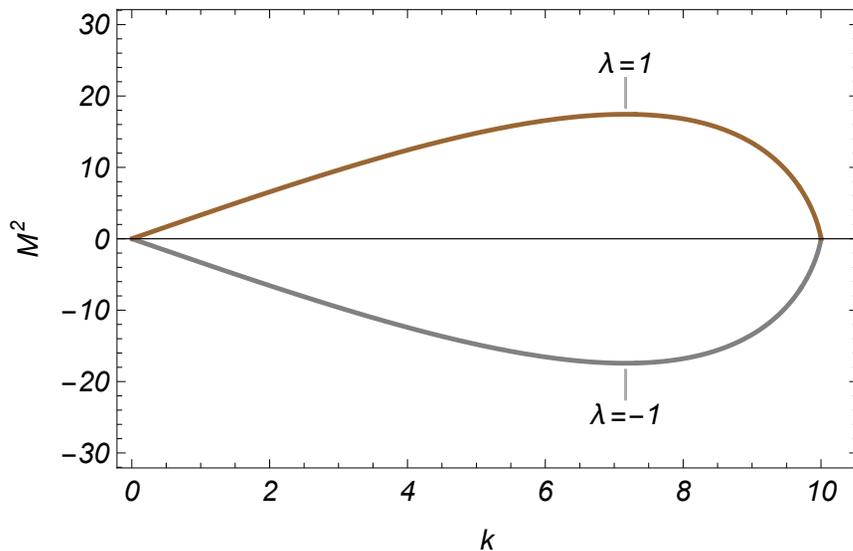
- It modifies Amper's law: $\nabla \times \mathbf{B} = \dot{\mathbf{E}} + \mathbf{j} + b_0 \mathbf{B}$

such that the wave equation becomes: $\nabla^2 \mathbf{A} = \partial_t^2 \mathbf{A} - b_0 \nabla \times \mathbf{A}$

- The plane-wave solutions: $\mathbf{A} = \epsilon_\lambda e^{-i\omega t + i\mathbf{k} \cdot \mathbf{x}}$ $\lambda = \pm 1$ is Right or Left polarization.

where $\omega^2 - \mathbf{k}^2 = -\lambda b_0 |\mathbf{k}|$

- Thus, the dynamical gluon or photon mass can be real or imaginary.

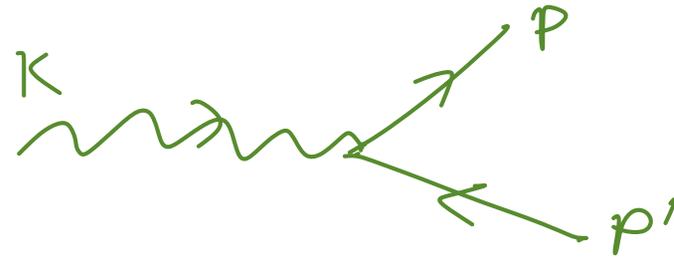
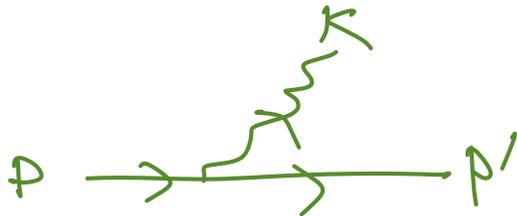


The effective photon mass M squared at finite chiral chemical potential in EM plasma.

Akamatsu, Yamamoto (2013)

NEW PROCESSES INDUCED BY CME

KT (2018)



Energy-momentum conservation: $k^2 = (p \pm p')^2 = 2m(m \pm E)$

Photon dispersion: $\omega^2 - \mathbf{k}^2 = -\lambda b_0 |\mathbf{k}|$ $b_0 \propto \mu_5 = \text{chiral chemical potential}$

\Rightarrow $1 \rightarrow 2$ processes are forbidden if $b_0 = 0$

- Pair production is allowed if $\lambda b_0 < 0$
- Gauge boson radiation is allowed if $\lambda b_0 > 0$ (*Chiral Cherenkov radiation*)

NEW CHANNELS IN QED

1. Photon production

KT (2018)

$$\frac{dW^{e \rightarrow \gamma^+ e}}{dx} = \frac{\alpha b_0}{4} \left(\frac{(1-x)^2 + 1}{x} - \frac{2m^2}{b_0 E} \right) \quad x \leq \frac{1}{1 + \frac{m^2}{b_0 E}}$$

Only right-handed photons are produced with transverse momentum

$$k_{\perp}^2 = -x^2 m^2 + (1-x)x\lambda b_0 E$$

2. Pair production

$$\frac{dW^{\gamma^- \rightarrow e\bar{e}}}{dx} = \frac{\alpha b_0}{4} \left(x^2 + (1-x)^2 + \frac{2m^2}{\omega b_0} \right) \quad x_1 \leq x \leq x_2$$
$$x_{1,2} = \frac{1}{2} \left(1 \mp \sqrt{1 - \frac{4m^2}{b_0 \omega}} \right)$$

Only left-handed photons decay.

- These processes induce the overall polarization of a jet and contribute to the energy loss.

$$-\frac{dE}{dz} = \int_0^1 \frac{dW^{e \rightarrow \gamma^+ e}}{dx} x E dx$$

NEW CHANNELS IN QCD

Hansen, KT (2024)

A free gluon is has the same dispersion is a free photon.

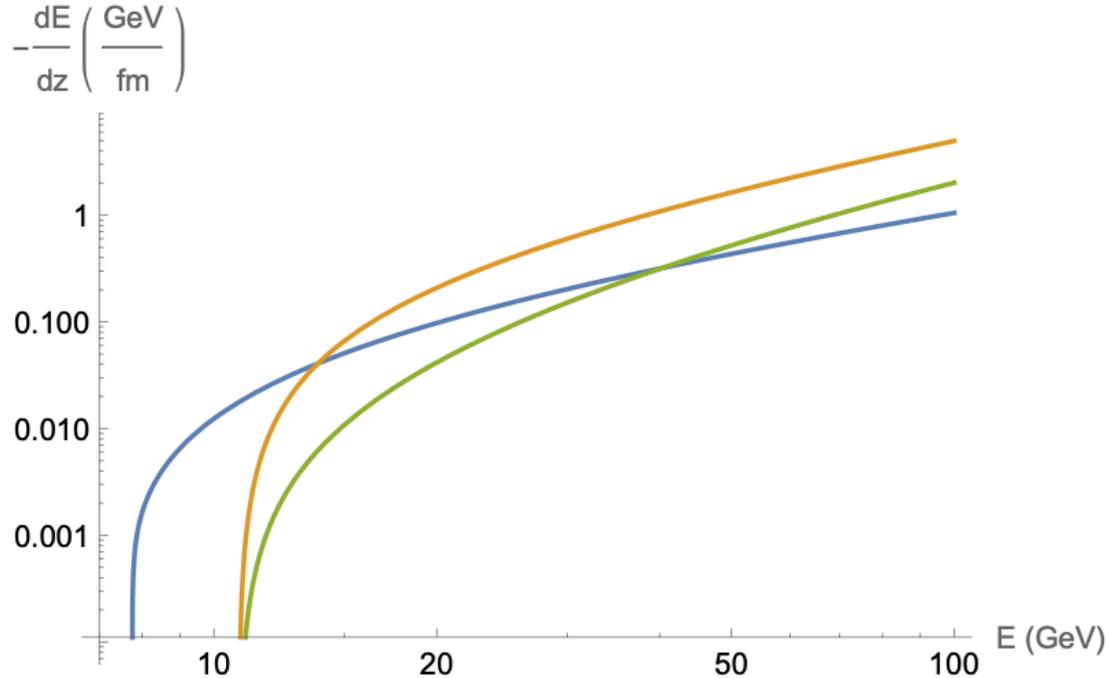
New non-Abelian term in the Lagrangian: $b_0 \epsilon^{ijk} \frac{g}{6} f^{abc} A_i^a A_j^b A_k^c$

$$\begin{aligned} i\mathcal{M}_{g_R \rightarrow g_R g_R}^A &= \frac{k_\perp}{x(1-x)}, \\ iM_{g_L \rightarrow g_R g_L}^A &= \frac{(1-x)k_\perp}{x}, \quad +\mathcal{O}\left(\frac{b_0, m}{E}\right) \\ iM_{g_L \rightarrow g_L g_R}^A &= \frac{xk_\perp}{(1-x)}, \\ iM_{g_L \rightarrow g_R g_R}^A &= \mathcal{O}(k_\perp/E, b_0/E), \\ iM_{g \rightarrow gg}^B &= \frac{-b_0 k_\perp (\lambda(1-x) + \lambda'x + \lambda_0)}{x(1-x)E} \quad (\text{small}) \end{aligned}$$

- Both right and left handed gluons can be produced and decay, but with different amplitudes.

ENERGY LOSS IN QGP DUE TO CHIRAL CHERENKOV

Hansen, KT (2024)



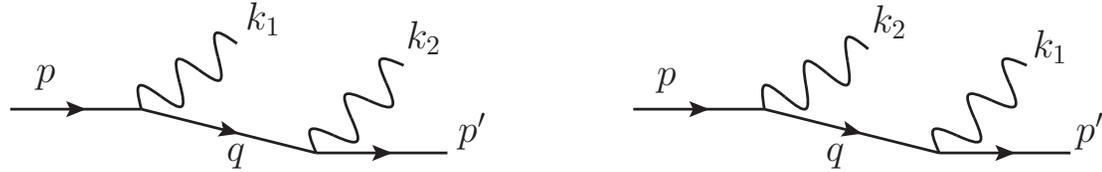
$$\begin{aligned}
 -\frac{dE_{q \rightarrow qq}}{dz} &= \frac{4\alpha_s g^2 b_0 E}{9}, \\
 -\frac{dE_{g_R \rightarrow gg}}{dz} &= \frac{3\alpha_s g^2 b_0 E}{4} \left(\ln \frac{b_0 E}{\omega_p^2} - 1 \right), \\
 -\frac{dE_{g_L \rightarrow gg}}{dz} &= \frac{3\alpha_s g^2 b_0 E}{4} \left(\ln \frac{b_0 E}{\omega_p^2} - \frac{17}{6} \right).
 \end{aligned}$$

Fig. 3. The rate of energy loss due to the Color Chiral Cherenkov radiation for $q \rightarrow qq$ (blue), $g_R \rightarrow gg$ (orange), and $g_L \rightarrow gg$ (green). Parameters: $g = 2$, $T = 300$ MeV, $b_0 = 50$ MeV.

FACTORIZATION OF SUCCESSIVE EMISSIONS

Hansen, KT (2025)

Double photon emission:



$$W^{e \rightarrow e\gamma\gamma} = \frac{1}{8E} \sum_{\text{spins}} \sum_{\lambda_1, \lambda_2} \int \frac{d^3 p'}{2E'(2\pi)^3} \int \frac{d^3 k_1}{2\omega_1(2\pi)^3} \int \frac{d^3 k_2}{2\omega_2(2\pi)^3} (2\pi)^4 \delta(p - k_1 - k_2 - p')$$

$$\times \left| \frac{e^2 \bar{u}' \not{\epsilon}_2^* (\not{p} - \not{k}_1 + m) \not{\epsilon}_1^* u}{(p - k_1)^2 - m^2 + 2iE_q/\tau} + \frac{e^2 \bar{u}' \not{\epsilon}_1^* (\not{p} - \not{k}_2 + m) \not{\epsilon}_2^* u}{(p - k_2)^2 - m^2 + 2iE_q/\tau} \right|^2,$$

The intermediate propagator has a resonance corresponding to χ Cherenkov:

$$\frac{1}{-2p \cdot k_1 + k_1^2 + 2iE_q/\tau} = \frac{1}{\omega_1 E \left(\frac{m^2}{E^2 \omega_1} \frac{\omega_1 - x_0 E}{1 - x_0} + \vartheta^2 \right) + 2iE_q/\tau}$$

If the resonance is dominant then: $\left| \frac{1}{(p - k_1)^2 - m^2 + 2iE_q/\tau} \right|^2 \approx \frac{\pi\tau}{2E_q} \delta((p - k_1)^2 - m^2)$

Factorization:
$$W^{e \rightarrow e\gamma\gamma} = \tau \int_0^E d\omega_1 \frac{dW^{e(p) \rightarrow e(q)\gamma(k_1)}}{d\omega_1} \int_0^{E_q} d\omega_2 \frac{dW^{e(q) \rightarrow e(p')\gamma(k_2)}}{d\omega_2}$$

EVOLUTION EQUATIONS OF CHIRAL CASCADE IN QED

Hansen, KT (2025)

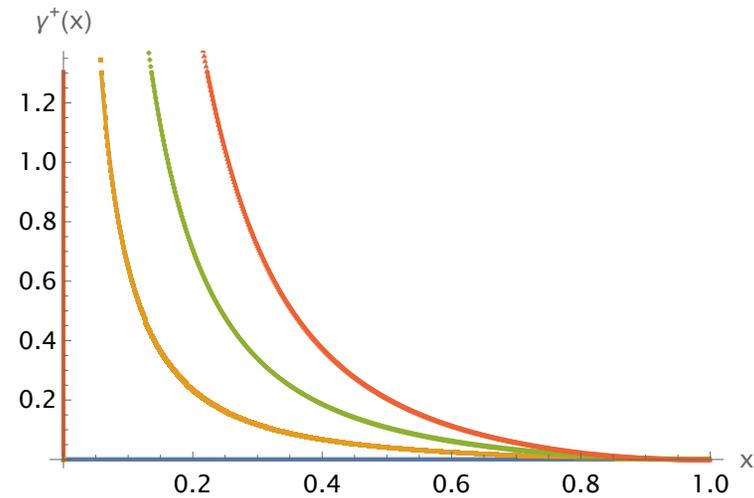
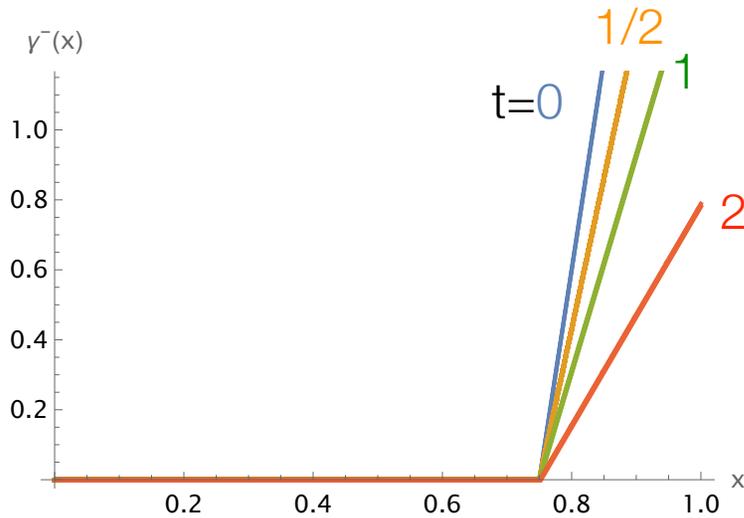
$$\begin{aligned}\frac{de(y, t)}{dt} &= \frac{1}{\tau} \int_y^1 \frac{dx}{x} \left\{ P_{ee}(x) e\left(\frac{y}{x}, t\right) + P_{e\gamma^-}(x) \gamma^-\left(\frac{y}{x}, t\right) \right\}, \\ \frac{d\bar{e}(y, t)}{dt} &= \frac{1}{\tau} \int_y^1 \frac{dx}{x} \left\{ P_{ee}(x) \bar{e}\left(\frac{y}{x}, t\right) + P_{e\gamma^-}(x) \gamma^-\left(\frac{y}{x}, t\right) \right\}, & \frac{1}{\tau} &= \frac{\alpha b_0}{4} \\ \frac{d\gamma^+(y, t)}{dt} &= \frac{1}{\tau} \int_y^1 \frac{dx}{x} \left\{ P_{\gamma^+e}(x) e\left(\frac{y}{x}, t\right) + P_{\gamma^+e}(x) \bar{e}\left(\frac{y}{x}, t\right) \right\}, \\ \frac{d\gamma^-(y, t)}{dt} &= \frac{1}{\tau} \int_y^1 \frac{dx}{x} P_{\gamma^-\gamma^-}(x) \gamma^-\left(\frac{y}{x}, t\right)\end{aligned}$$

- These are NOT standard evolution equations - they describe evolution due to anomaly and contain no conventional terms.
- The standard splitting functions are due to $1 \rightarrow 2$ processes.
- There is clear asymmetry between right and left photons.

EXAMPLE OF A CHIRAL CASCADE IN QED

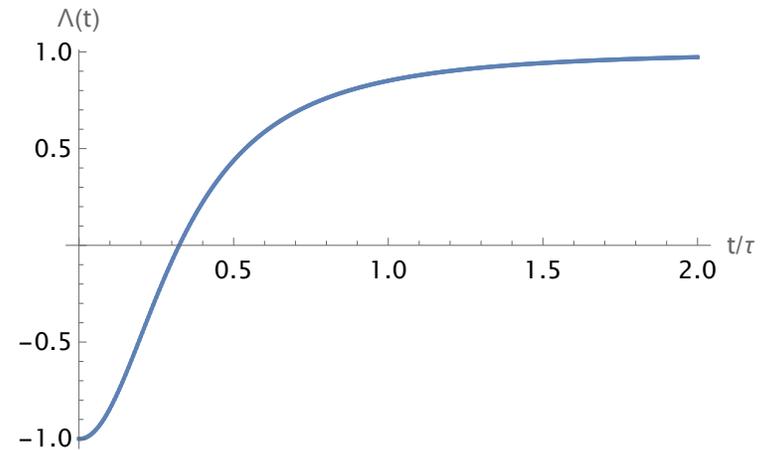
Hansen, KT (2025)

The initial condition: only left-handed photons. At late times: only right-handed ones.



The degree of polarization:

$$\Lambda(t) = \frac{\int_0^1 (\gamma^+(x, t) - \gamma^-(x, t)) dx}{\int_0^1 (\gamma^+(x, t) + \gamma^-(x, t)) dx}$$



EVOLUTION EQUATIONS OF CHIRAL CASCADE IN QCD

$$\begin{aligned}\frac{dq(y, t)}{dt} &= \frac{1}{\tau} \int_y^1 \frac{dx}{x} \left\{ P_{qq}(x) q\left(\frac{y}{x}, t\right) + P_{qg^-}(x) g^-\left(\frac{y}{x}, t\right) \right\}, \\ \frac{dg^+(y, t)}{dt} &= \frac{1}{\tau} \int_y^1 \frac{dx}{x} \left\{ N_f P_{g^+q}(x) q\left(\frac{y}{x}, t\right) + N_f P_{g^+q}(x) \bar{q}\left(\frac{y}{x}, t\right) \right. \\ &\quad \left. + P_{g^+g^+}(x) g^+\left(\frac{y}{x}, t\right) + P_{g^+g^-}(x) g^-\left(\frac{y}{x}, t\right) \right\}, \\ \frac{dg^-(y, t)}{dt} &= \frac{1}{\tau} \int_y^1 \frac{dx}{x} P_{g^-g^-}(x) g^-\left(\frac{y}{x}, t\right)\end{aligned}$$

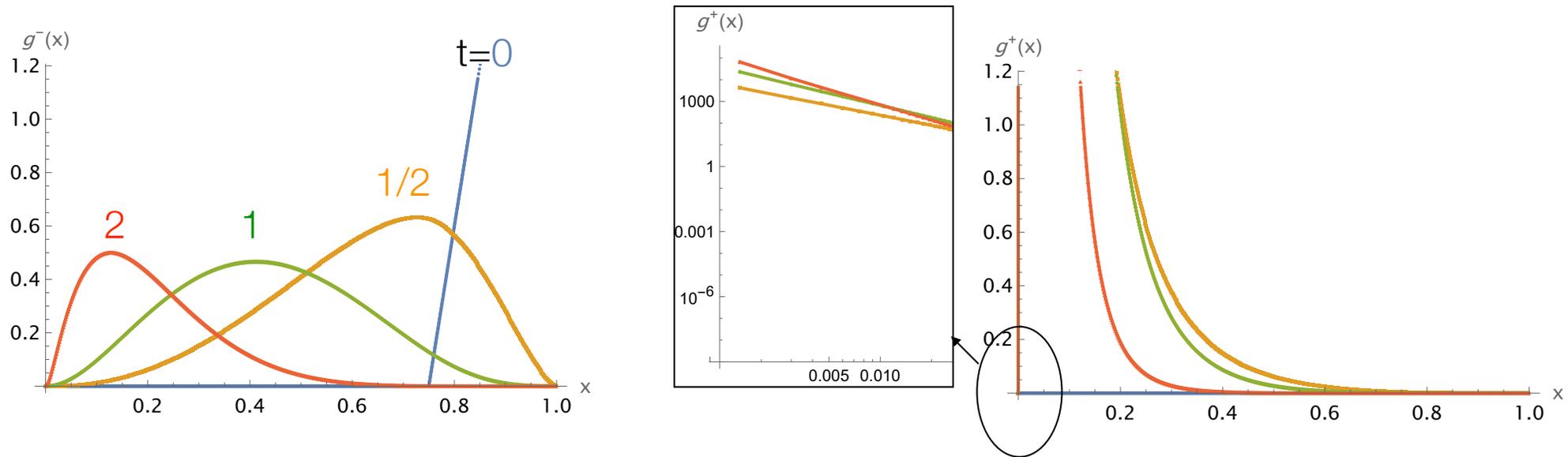
$$\begin{aligned}P_{g^+g^+} &= \frac{N_c}{x} \left[\frac{1}{(1-x)_+} + C^+ \delta(1-x) \right], \\ P_{g^-g^-} &= N_c \left[\frac{x^3}{(1-x)_+} + C^- \delta(1-x) \right]\end{aligned}$$

Charge and energy conservation: $C^+ = 0, \quad N_c C^- = \frac{11N_c}{6} - \frac{N_f}{3}$

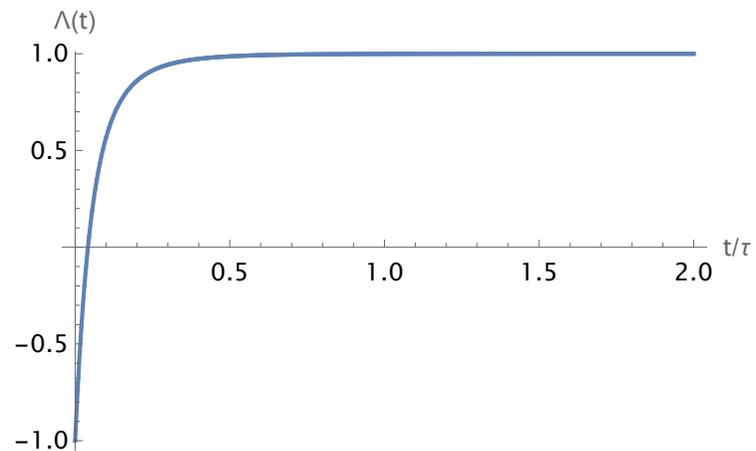
EXAMPLE OF A CHIRAL CASCADE IN QCD

Hansen, KT (2025)

The initial condition: only left-handed gluons.

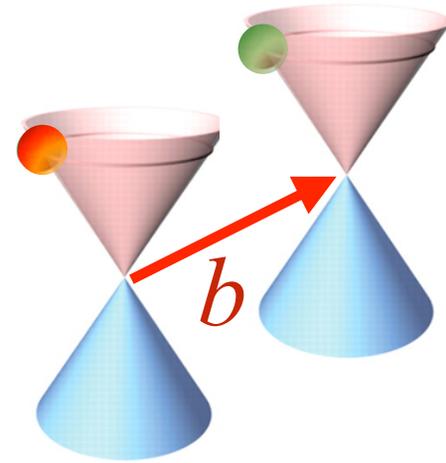
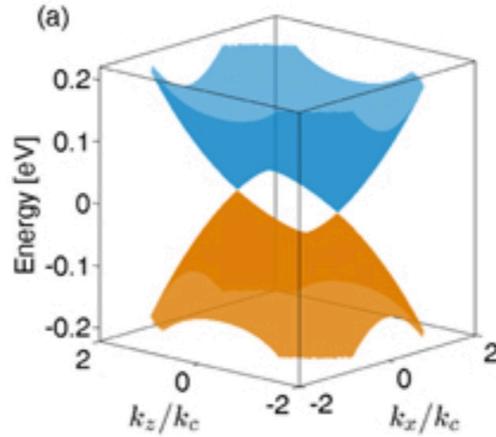


The degree of polarization:



CHIRAL CHERENKOV IN WEYL SEMIMETALS

Emergent chiral domains in Weyl semimetals:



TaAs
NbAs
NbP
TaP

$$\left. \begin{aligned} \nabla \cdot \mathbf{B} &= 0, \\ \nabla \cdot \mathbf{E} &= -\mathbf{b} \cdot \mathbf{B}, \\ \nabla \times \mathbf{E} &= -\partial_t \mathbf{B}, \\ \nabla \times \mathbf{B} &= \partial_t \mathbf{E} + \mathbf{b} \times \mathbf{E}. \end{aligned} \right\} \Rightarrow \left\{ \begin{aligned} \nabla \cdot \mathbf{B} &= 0, & \nabla \cdot \mathbf{D} &= 0 \\ \nabla \times \mathbf{E} &= i\omega \mathbf{B}, & \nabla \times \mathbf{B} &= -i\omega \mathbf{D} \end{aligned} \right.$$

where $\mathbf{D} = \mathbf{E} + \frac{i}{\omega} \mathbf{b} \times \mathbf{E} = 0$.

Dielectric tensor

$$\varepsilon = \begin{pmatrix} 1 & -ib/\omega_{\mathbf{k}\lambda} & 0 \\ ib/\omega_{\mathbf{k}\lambda} & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Qui, Cao, Huang (2017)

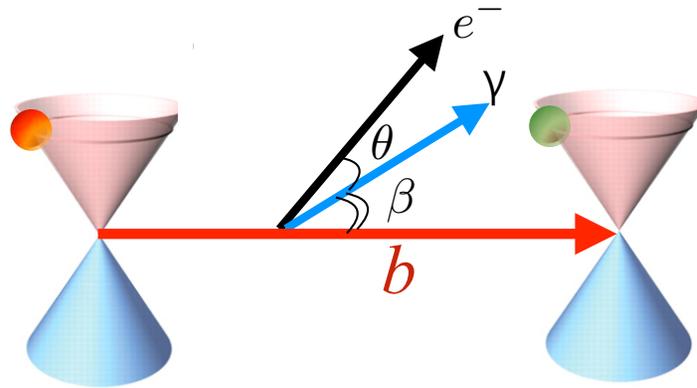
- Dispersion: $|k_i k_j - k^2 \delta_{ij} + \omega_{\mathbf{k}\lambda}^2 \varepsilon_{ij}(\omega_{\mathbf{k}\lambda})| = 0$

$$\omega^2 - \mathbf{k}^2 = \frac{1}{2} b^2 - \lambda \operatorname{sgn}(\mathbf{k} \cdot \mathbf{b}) \sqrt{(\mathbf{k} \cdot \mathbf{b})^2 + \frac{1}{4} b^4}$$

- Now the photon “mass” depends on the photon direction w.r.t. \mathbf{b} .

CHIRAL CHERENKOV RADIATION AT HIGH ENERGY

KT (2018)



• Photon spectrum:
$$\frac{dW^{e \rightarrow \gamma + e}}{dx} = \frac{\alpha \lambda b \cos \beta}{4} \left(\frac{(1-x)^2 + 1}{x} - \frac{2m^2}{\lambda b \cos \beta E} \right)$$

Kinematic constraints:
$$x \leq \frac{1}{1 + \frac{m^2}{b \lambda \cos \beta E}}$$

$$\lambda \cos \beta > 0$$

⇒ • The right/left-polarized photons emitted in the forward/backward direction.

CHIRAL CHERENKOV RADIATION IN A WEYL SEMIMETAL

Magnetic Weyl semimetals as a source of circularly polarized THz radiation

Jeremy Hansen, Kazuki Ikeda, Dmitri E. Kharzeev, Qiang Li, Kirill Tuchin (May 17, 2024)

e-Print: [2405.11076](https://arxiv.org/abs/2405.11076) [cond-mat.mtrl-sci]

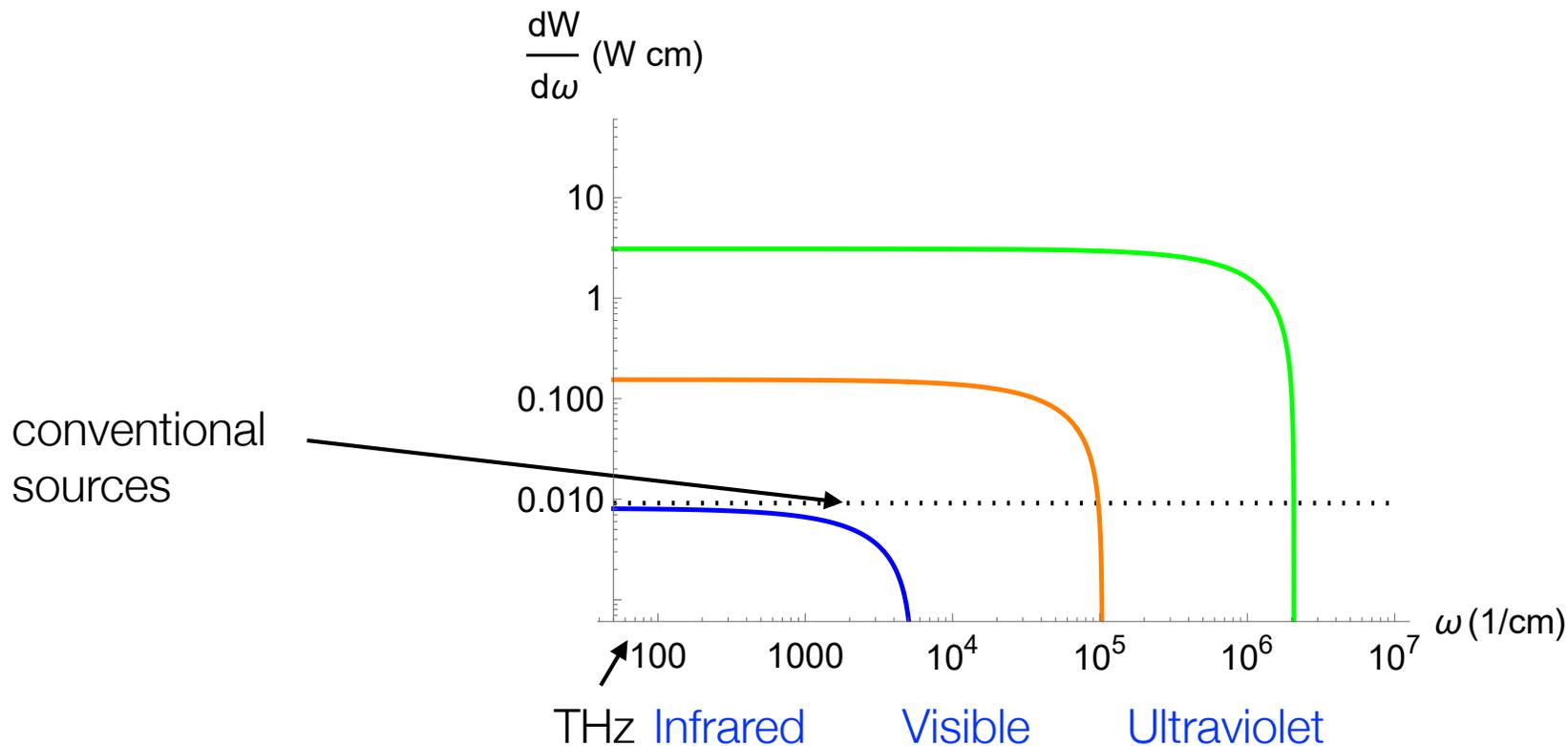


Fig. 3. Intensity spectrum of Chiral Cherenkov radiation by a pulse of $N = 10^{11}$ electrons of energy $E = 3$ MeV, $\beta = 0$ in a Weyl semimetal with $b = 3.8$ eV (green), $b = 0.19$ eV (orange), and $b = 0.01$ eV (blue). The spectra are given by Eq. (14) with $\epsilon_b = 10$, dotted line: bremsstrahlung spectrum in $\text{Co}_3\text{Sn}_2\text{S}_2$ computed according to [9].

AXION ELECTRODYNAMICS

Sikivie (84), Wilczek (87), Carroll et al (90)

$$\mathcal{L}_{\text{MCS}} = \mathcal{L}_{\text{QED}} + c_A \theta(x) \vec{E} \cdot \vec{B}$$

$$\nabla \cdot \mathbf{B} = 0,$$

$$\nabla \cdot \mathbf{E} = \rho - c \nabla \theta \cdot \mathbf{B},$$

$$\nabla \times \mathbf{E} = -\partial_t \mathbf{B},$$

$$\nabla \times \mathbf{B} = \partial_t \mathbf{E} + \mathbf{j} + c(\partial_t \theta \mathbf{B} + \nabla \theta \times \mathbf{E}),$$

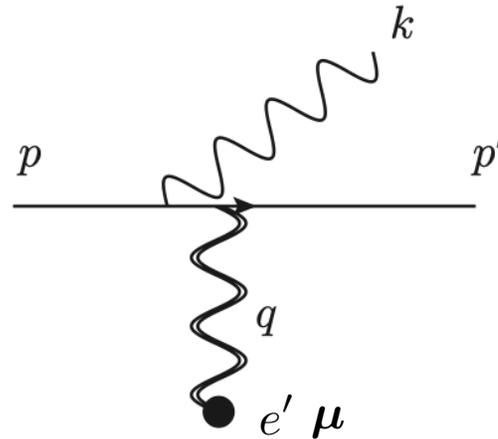
Anomalous Hall Effect

Chiral magnetic effect: $\mathbf{j} = b_0 \mathbf{B}$

- If $b_0 \approx \text{const}$, then a fast proton moving in the classical axion field will emit highly polarized chiral Cherenkov radiation.
- A new way to search for axions.

BREMSSTRAHLUNG IN CHIRAL MEDIA

Hansen, KT (2022)



- Photon/gluon propagator (static limit):

$$D_{00}(\mathbf{q}) = \frac{i}{\mathbf{q}^2},$$

$$D_{0i}(\mathbf{q}) = D_{i0}(\mathbf{q}) = 0,$$

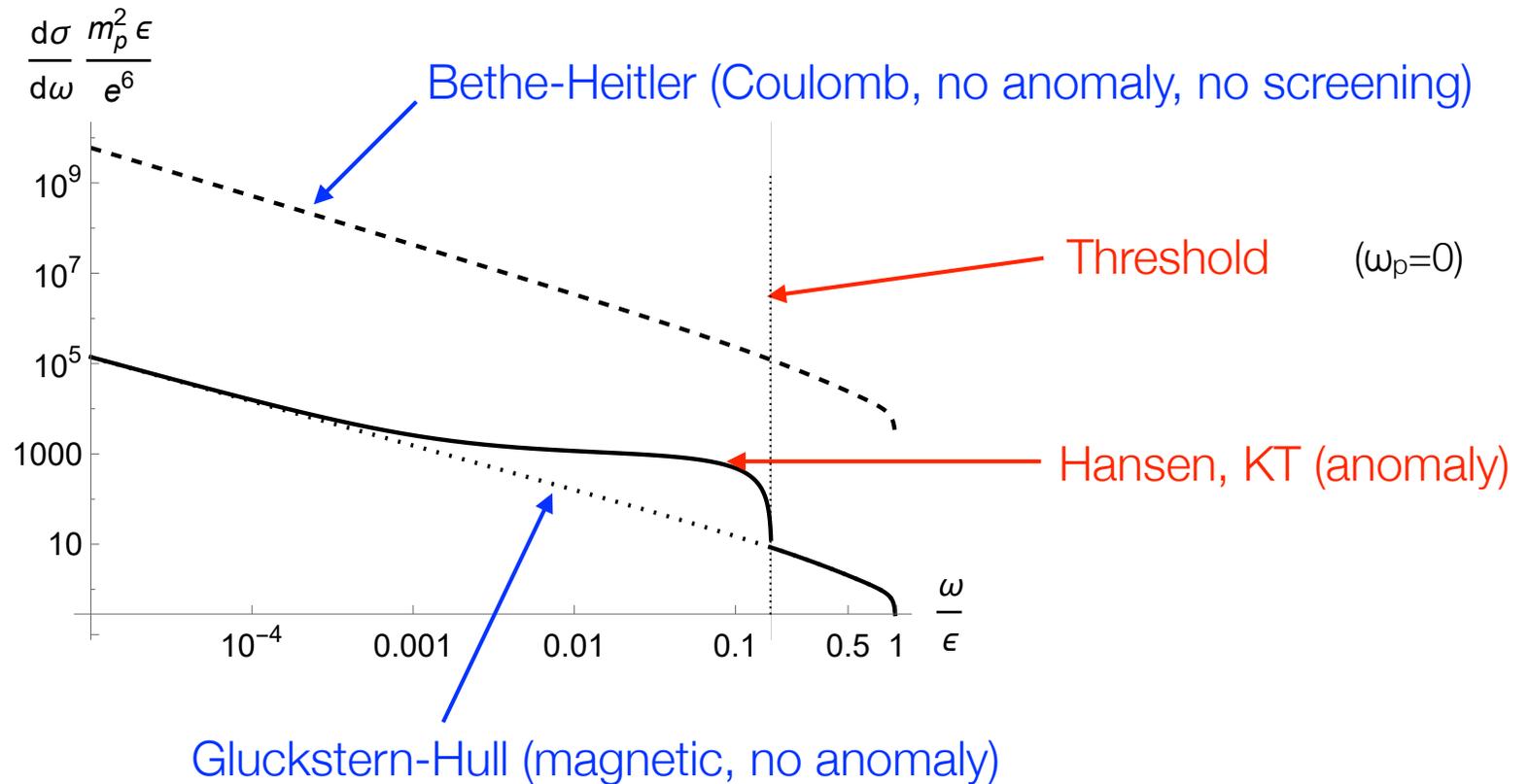
$$D_{ij}(\mathbf{q}) = -\frac{i\delta_{ij}}{\mathbf{q}^2 - b_0^2} - \frac{\epsilon_{ijk}q^k}{b_0(\mathbf{q}^2 - b_0^2)} + \frac{\epsilon_{ijk}q^k}{b_0\mathbf{q}^2}$$

resonance!

- D_{ij} couples only to the magnetic moment of the target $\mathbf{J}(\mathbf{x}) = \nabla \times (\boldsymbol{\mu}\delta(\mathbf{x}))$

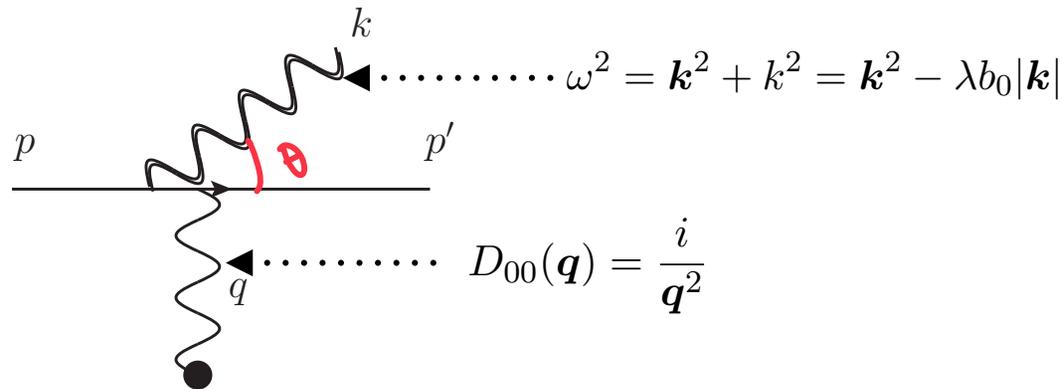
MAGNETIC MOMENT CONTRIBUTION TO BREMSSTRAHLUNG

Hansen, KT (2022)



ELECTRIC MONOPOLE CONTRIBUTION TO BREMSSTRAHLUNG

Hansen, KT (2023)



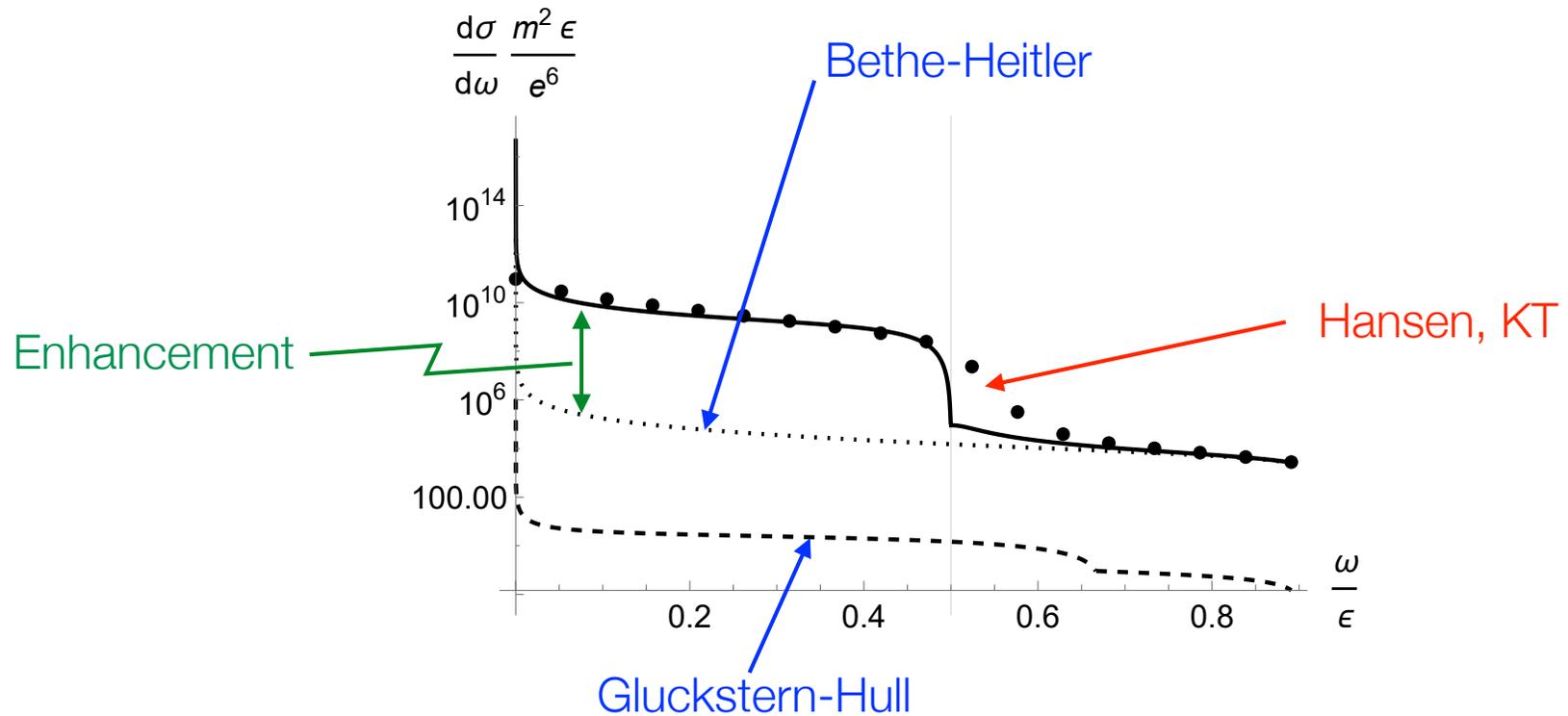
- Fermion propagator:
 $b_0 \lambda > 0$

$$\frac{1}{2p \cdot k - k^2 + iE/\tau} = \frac{1}{\omega E \left(\frac{m^2}{E\omega} \frac{\omega - \omega^*}{E - \omega^*} + \theta^2 + \frac{i}{\omega\tau} \right)}$$
 with $\omega^* = \frac{\lambda b_0 E^2}{\lambda b_0 E + m^2}$
- The resonance emerges when $\omega < \omega^*$ due to the anomaly in the photon/gluon dispersion relation.
- The photon/gluon propagator shows similar behavior: $\mathbf{q}_{\min}^2 = \frac{1}{4} \frac{\omega^2 E^2}{E'^2} \left[\frac{m^2(\omega - \omega^*)}{\omega E(E - \omega^*)} + \theta^2 \right]^2$

RADIATION ENHANCEMENT

Weak screening limit $\mu \ll m$

μ -Debye mass



- Only one photon polarization ($b_0 \lambda > 0$) is enhanced!

SUMMARY

- In the presence of the chiral magnetic effect, the jets exhibit polarization and a distinctive transverse momentum dependence.
- Relativistic heavy-ion collisions can serve as a testing ground not only for the Early Universe but also for the Dark Matter.

Additional material

QUANTIZATION OF EM FIELD

KT (2018)

$$\mathbf{A}(\mathbf{x}, t) = \underbrace{\sum_{\mathbf{k}\lambda} (a_{\mathbf{k}\lambda} \mathbf{A}_{\mathbf{k}\lambda} + a_{\mathbf{k}\lambda}^\dagger \mathbf{A}_{\mathbf{k}\lambda}^*)}_{\text{Transverse EM waves}} + \underbrace{\sum_{\mathbf{k}\nu} (a_{\mathbf{k}\nu} \mathbf{A}_{\mathbf{k}\nu} + a_{\mathbf{k}\nu}^\dagger \mathbf{A}_{\mathbf{k}\nu}^*)}_{\text{Longitudinal EM waves (exist only if there is spatial dispersion)}}$$

Transverse EM waves

Longitudinal EM waves (exist only if there is spatial dispersion)

$$\mathbf{A}_{\mathbf{k}\lambda} = \mathbf{e}_{\mathbf{k}\lambda} \left(\frac{k v_{\mathbf{k}\lambda}}{2\omega_{\mathbf{k}\lambda}^2 \epsilon_{ij} e_{\mathbf{k}\lambda i}^* e_{\mathbf{k}\lambda j}} V \right)^{1/2} e^{i\mathbf{k}\cdot\mathbf{x} - i\omega_{\mathbf{k}\lambda} t}$$

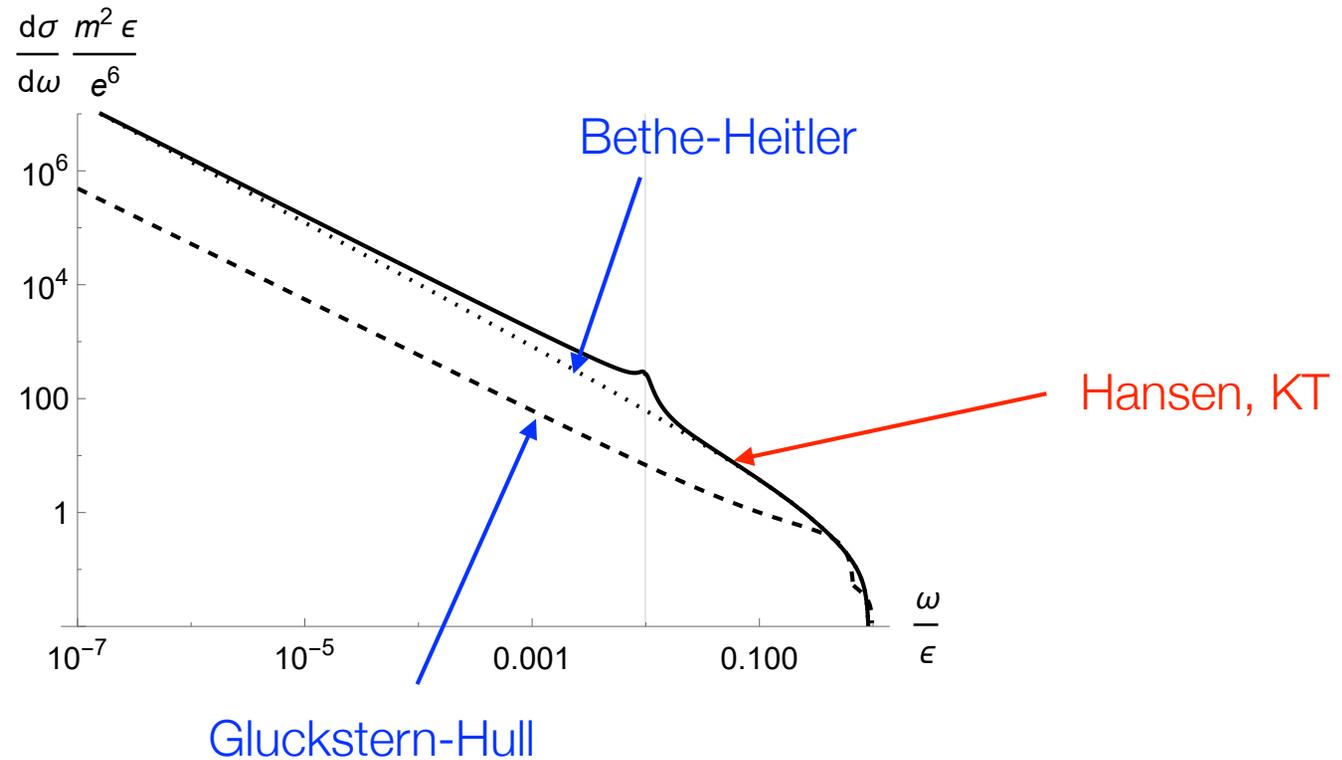
$$\mathbf{A}_{\mathbf{k}\nu} = \hat{\mathbf{k}} \left(\frac{k^2}{\omega_{\mathbf{k}\nu}^2 k_i k_j \partial \epsilon_{ij} / \partial \omega_{\mathbf{k}\nu}} V \right)^{1/2} e^{i\mathbf{k}\cdot\mathbf{x} - i\omega_{\mathbf{k}\nu} t}$$

Chiral Cherenkov radiation: $\frac{dW}{d\Omega d\omega} = \frac{\alpha Q^2}{16\pi} \sum_{\lambda} \delta(\omega + \epsilon' - \epsilon) \frac{k^3}{\epsilon \epsilon' \omega^2 \epsilon_{ij} e_{\mathbf{k}\lambda i}^* e_{\mathbf{k}\lambda j}} \sum_{ss'} |\mathcal{M}_0|^2$

At high energies equations simplify: $e_{\mathbf{k}\pm i}^* e_{\mathbf{k}\pm j} \rightarrow \frac{1}{2} \left(\delta^{ij} - \frac{k_i k_j}{k^2} \right)$

STRONG SCREENING $\mu \gg m$

Hansen, KT (2023)



- The effect of anomaly is reduced by screening.