Quantum Closures and Riemann Solvers for Neutrino Moment Transport

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How do we undertake supernova and neutron star merger simulations which account for neutrino flavor transformation?

- The issue is oscillations introduce a new length scale
 - The neutrino oscillation wavelength in dense matter is ~10 microns.
 - The best simulations have spatial grid zones ~100 m 1 km.
 - To make it even worse, the time steps are set by the light crossing time of a grid zone.
- Throwing silicon at the problem is not the solution. To make simulations with transformations feasible we will have to make approximation and/or get creative:
 - e.g. Nagakura & Zaizen rescaled the neutrino Hamiltonian down by a factor of 10⁻⁴ then extrapolated their results back to the proper strength.

Nagakura & Zaizen PRL **129** 261101 (2022)

see also Xiong et al PRD 107 083016 (2023)

Oscillations with moments

- Many classical simulation codes, e.g. FLASH, evolve the neutrino field using moments.
 - the number of moments evolved is usually small, e.g. M1 transport evolves 1+3=4 moments in 3D.
- It is possible to compute neutrino transformation with moments.

 Strack and Burrows, PRD 71 093004 (2005)
 Zhang and Burrows, PRD 88 105009 (2013)

 Myers et al, PRD 105 123036 (2022)
 Grohs et al, PLB, 846, 138210 (2023)

 Froustey et al, PRD 109 043046 (2024)
 Grohs et al., ApJ 963 11 (2024)

 Froustey et al., 111 063022 (2025)
 Kneller et al., PRD 111 063046 (2025)

 Grohs et al, PRD 111 083018 (2025)
 Froustey arXiv:2505.16961 (2025)



 We define a quantum moment of the single-particle distribution matrix *f* as

$$\mathcal{M}_n(q) = \int q \cos^n \theta \quad f \, d \, \Omega_q$$

- Where n is an integer, q is the energy of the neutrino, θ the angle relative to a given direction (e.g. the radial)
- The first few moments have well-known names
 - n = 0 is the energy density E
 - n = 1 is the energy flux F
 - n = 2 is the pressure P

Assuming spherical symmetry, the moments evolve according to

$$\frac{\partial E}{\partial t} + \frac{\partial F}{\partial r} + \frac{2F}{r} = -i[H_V + H_M + H_E, E] + i[H_F, F]$$

$$\frac{\partial F}{\partial t} + \frac{\partial P}{\partial r} + \frac{3P - E}{r} = -i[H_v + H_M + H_E, F] + i[H_F, P]$$

- the H's are contributions to the neutrino Hamiltonian,
- the absorption / emission / collisions have been omitted
- The infinite tower of equations can be truncated at what ever level one desires.
 - Usually one considers two schemes: a one-moment (M0) and a twomoment (M1)
- We need a relationship between the moments to close the equations.
- This relationship is called 'The Closure'

Quantum Closures

- Since we are using quantum moments, we need a Quantum Closure.
- Start from the ansatz that two quantum moments, e.g. E and P, are related by

P = L E R

- L and R are arbitrary matrices
- Since E and P are Hermitian they have eigenvalue matrices

$$E = U_E \Lambda_E U_E^{\dagger} \qquad P = U_P \Lambda_P U_P^{\dagger}$$

 the U's are unitary 'mixing' matrices, the Λ's are the matrices of the eigenvalues. • We can factorize U_{F} and U_{P} as

$$U_E = Y_E \Phi_E \qquad U_P = Y_P \Phi_P$$

- where Y_E and Y_P are unique unitary matrices which can be written in terms of the elements of E and P, and Φ_E and Φ_P are arbitrary diagonal unitary matrices.
- Define the quantum Eddington factor matrix to be

$$X = \Lambda_P \Lambda_E^{-1}$$

• It can be shown the quantum closure has to be of the form

$P = L E L^{\dagger}$

• where L is

$$L = \mathbf{Y}_P^{\dagger} X^{1/2} \mathbf{Y}_E$$

• all the arbitrary phases in the Φ 's can be redefined away.

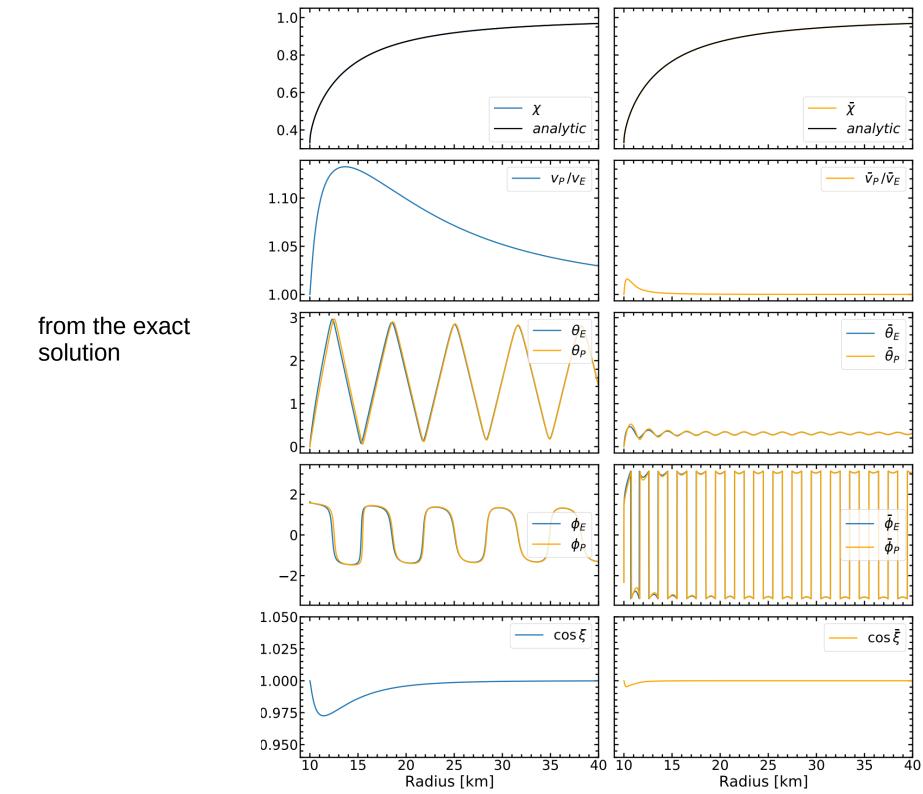
• A quantum closure can be understood as a three-step relation:

$$P = \mathbf{Y}_{P}^{\dagger} \Big(X^{1/2} \Big(\mathbf{Y}_{E} E \mathbf{Y}_{E}^{\dagger} \Big) X^{1/2} \Big) \mathbf{Y}_{P}$$

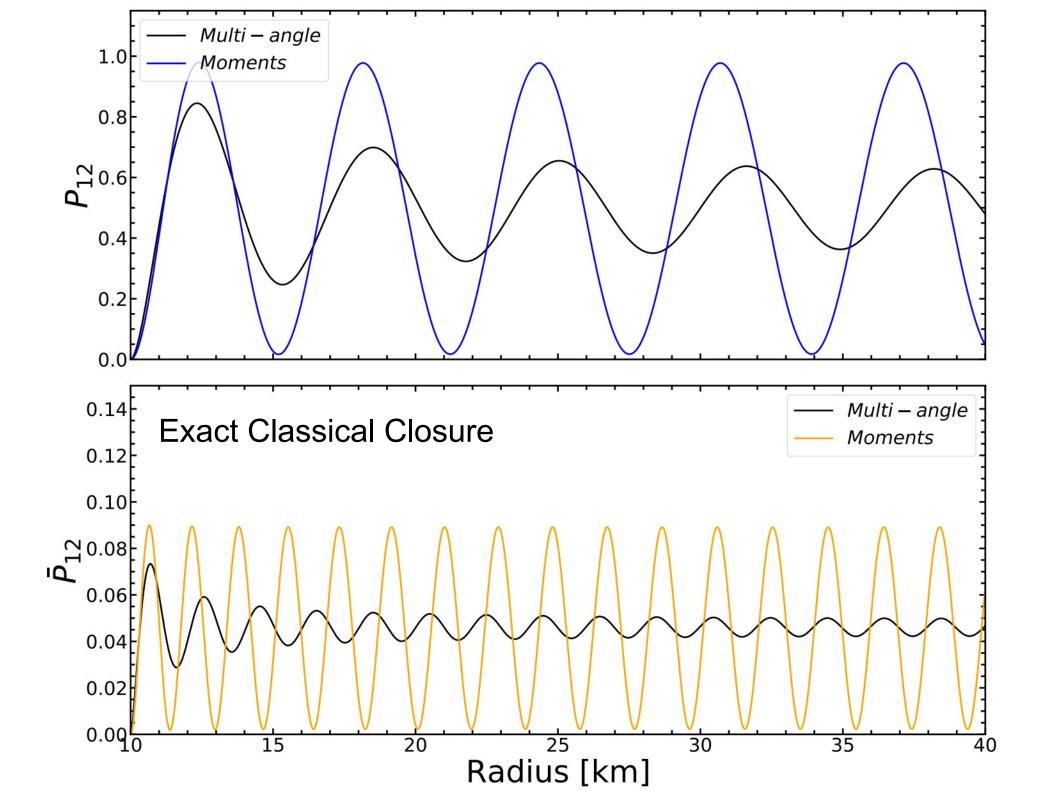
- Y_E 'rotates' E to its eigenvalue matrix
- X^{1/2} rescales the eigenvalues
- Y_{p} 'rotates' away from a diagonal matrix

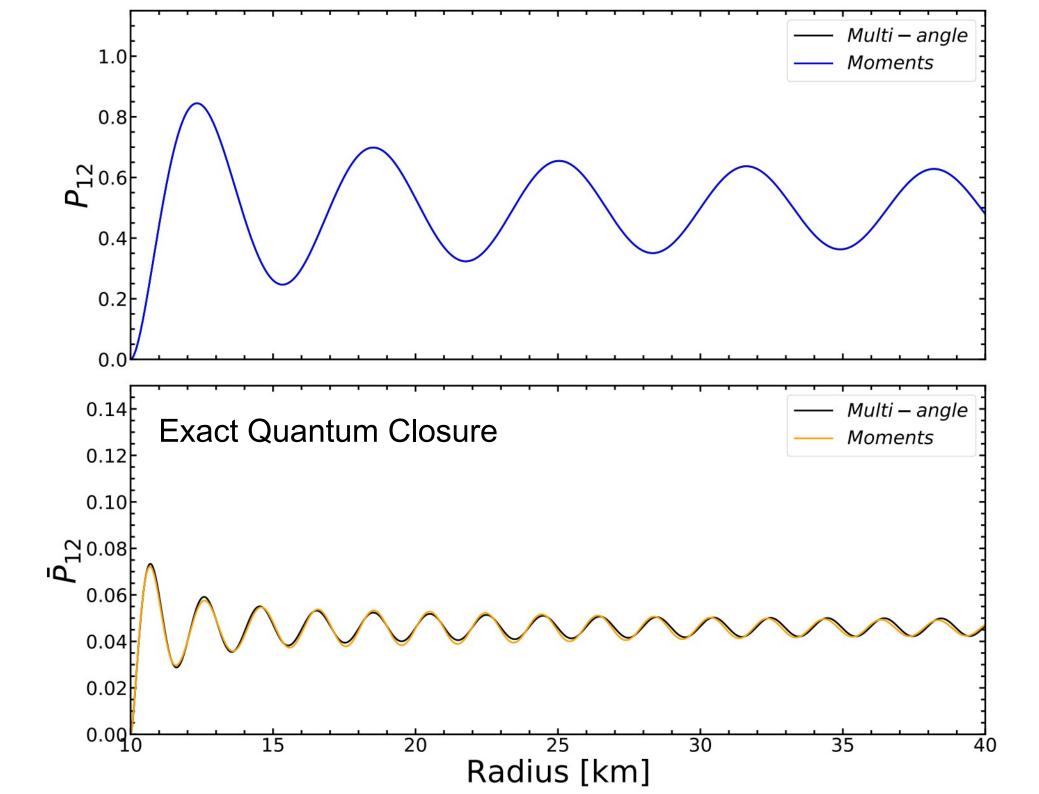
Is the Quantum Closure Complete?

- If one uses the 'correct' closure in a moment calculation, the correct result must be recovered.
- Consider the MSW problem for neutrinos emitted from a spherical source (a lightbulb).
 - There is an exact solution for a single neutrino.
- We set the matter density and choose mixing parameters so that the neutrinos are on the MSW resonance.



 Now that we know the correct closure from the exact solution, we can do a moment calculation using this closure.





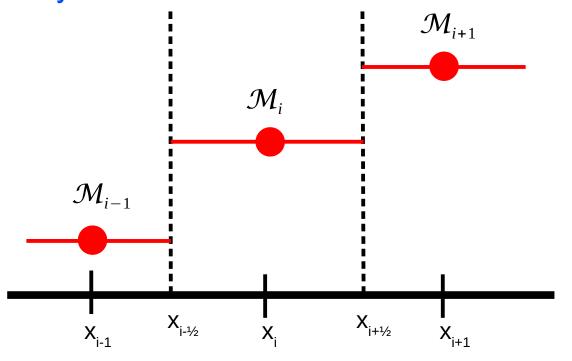
Quantum Advection

Consider again the moment transport equations.

$$\frac{\partial E_q}{\partial t} + \frac{\partial F_q}{\partial r} + \frac{2 F_q}{r} = -i[H_V + H_M + H_E, E_q] + i[H_F, F_q]$$
$$\frac{\partial F_q}{\partial t} + \frac{\partial P_q}{\partial r} + \frac{3 P_q - E_q}{r} = -i[H_V + H_M + H_E, F_q] + i[H_F, P_q]$$

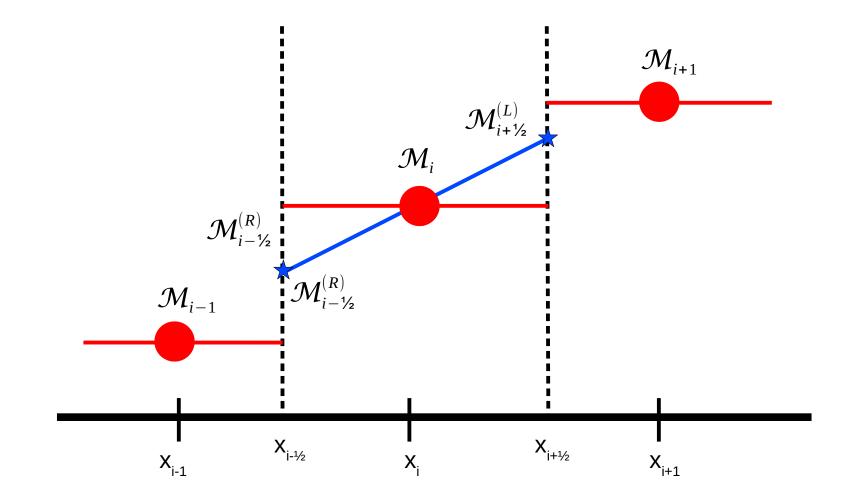
- To numerically solve these equations we move everything except the time derivatives to the right-hand side.
- We discretize the space into finite volumes and evolve the zone-averages of E and F within each grid zone.
- The zone averages 'live' at the cell centers.

 We end up with a set of 'Riemann Problems' for E and F: the initial value problem of the flow of some quantity across a discontinuity.

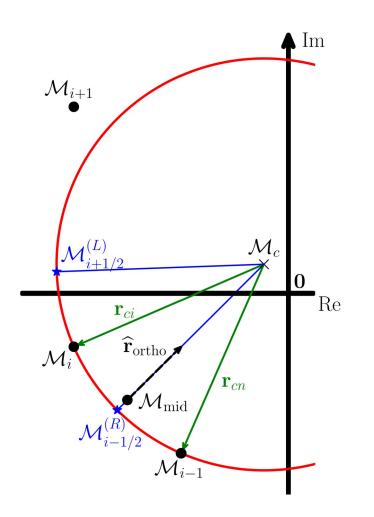


- There exist many algorithms collectively known as 'Riemann Solvers' – for the classical version of this problem.
- For quantum moments, the off-diagonal elements the coherences – are complex numbers.
- How do you build a Riemann Solver for quantum moments?

- Riemann Solvers have two parts: the 'Reconstruction' and the 'Riemann Solution'.
- The Reconstruction step generates the values of the moments either side of the discontinuity from the values at the zone centers.

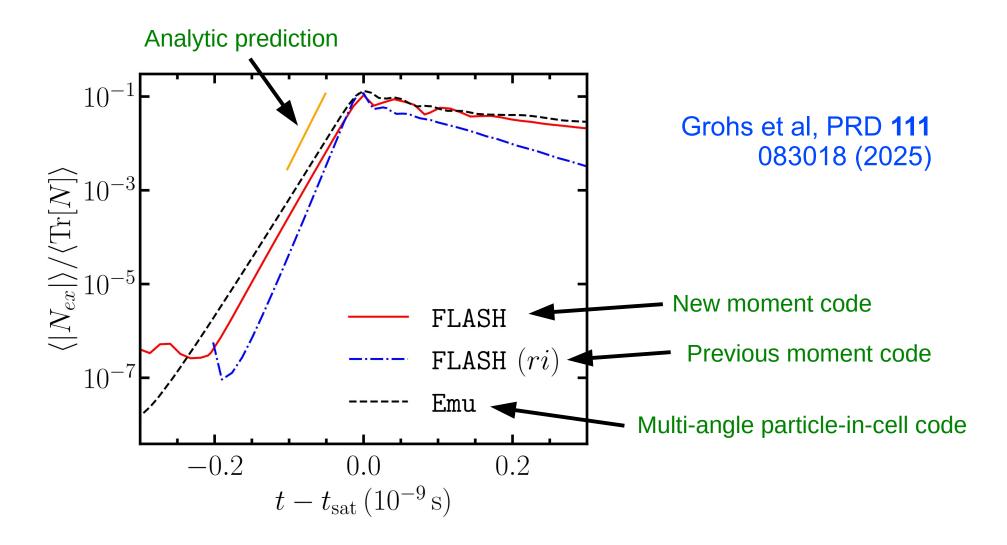


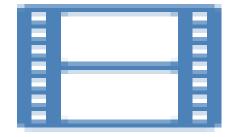
 We tried many ideas. The one that works most robustly is based on the ideas behind 'minmod' and uses interpolation using circular arcs in the complex (Argand) plane.



Are moment-based approaches any good?

 We have undertaken extensive testing of our version of FLASH with quantum moments vs EMU, a 'multi-angle' code.





- This animation is computed using our version of FLASH with quantum moments.
- It is the flavor transformation in a small volume using data taken from a binary neutron star merger by Foucart et al.



- Moments are a more efficient way of doing neutrino flavor transformation calculations.
- We know how to write a complete quantum closure, and if the correct closure is used, the results from moment calculations are exact.
- We have adapted classical Riemann solvers for quantum moments and they match results from 'multi-angle' codes.