# Ov2β sensitivity of the XLZD rare-event observatory Based on 2410.19016

Chami Amarasinghe (UCSB) On behalf of the XLZD collaboration CIPANP – June 2025

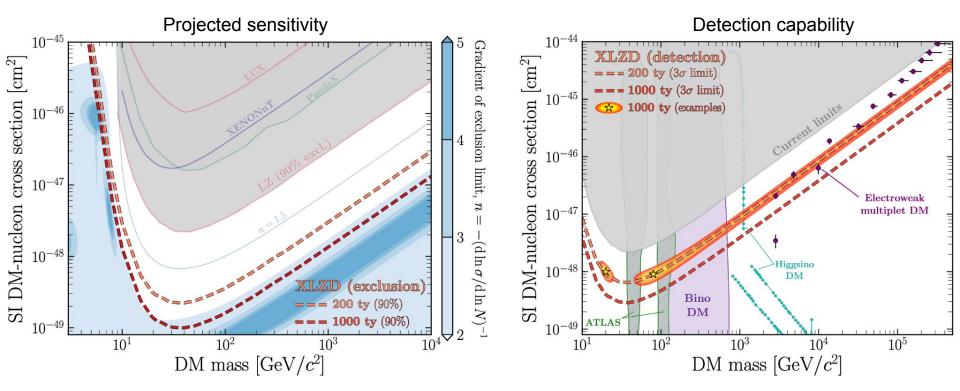




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### Liquid xenon detectors as definitive tools for WIMP searches

Based on the XLZD design book: 2410.17137



A multi-tonne liquid xenon time projection chamber will capable of reaching the neutrino fog and constraining DM parameters

### Large scale enables a broad rare-event program

#### Liquid target is scalable

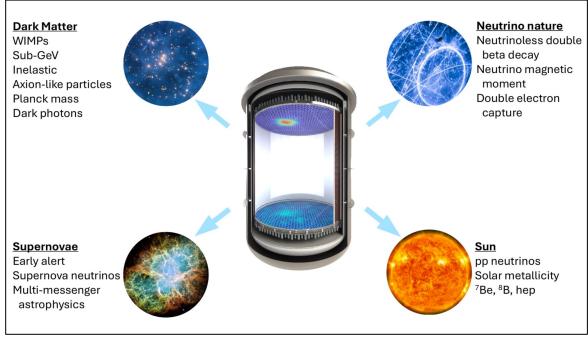
- LZ, XENONnT, PANDAX demonstrating dual-phase TPC at 10t scale

#### **Background controls**

- Continuous purification
- Self-shielded fiducial volume
- Material selection

#### LXe TPC response

- Good energy resolution
- ER/NR discrimination
- Low energy threshold
- Multiple scatter rejection



Physics reach of liquid xenon rare event observatory

### The XLZD Collaboration

76 institutions

**17** countries

440+ members

Next generation liquid xenon rare event observatory

- select best options from LZ, XENONnT, DARWIN RnD
- combine expertise in radiopurity, HV, cryogenics
- risk management using LZ, XENONnT, and test setups

XLZD sensitivity to  $0v2\beta - 2410.19016$  studied scenarios:

- Active volume 60t baseline (extendable to 80t pending Xe market)
- Nominal and optimistic cases for
  - Energy resolution, background reduction, site, ...

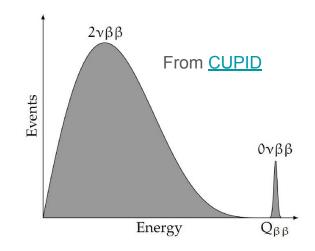


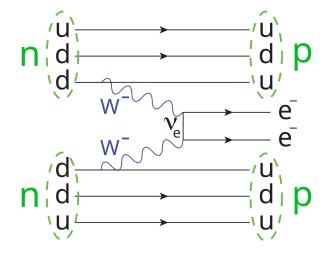


# <sup>136</sup>Xe 0v2b decay and signature

- <sup>136</sup>Xe 2v2β decay occurs and is background to WIMP searches
- Hypothesized 0v2β process
  - Violates lepton number and B-L
  - Implies neutrinos are Majorana
- $0v2\beta$  at the tail of  $2v2\beta$

For <sup>136</sup>Xe,  $\mathbf{Q}_{\beta\beta}$  = 2457.83 ± 0.37 keV





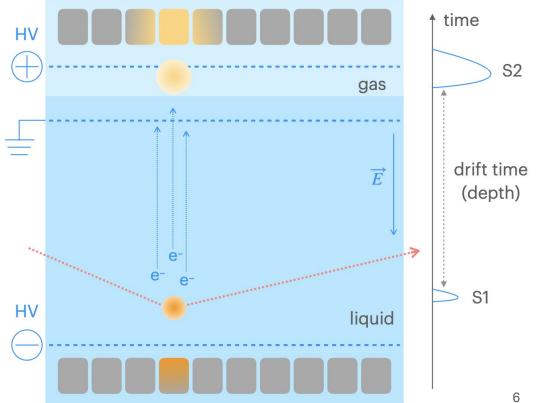
# Signal reconstruction in LXe TPCs

Best energy resolution from S1 + S2. Near  $Q_{RR}$ 

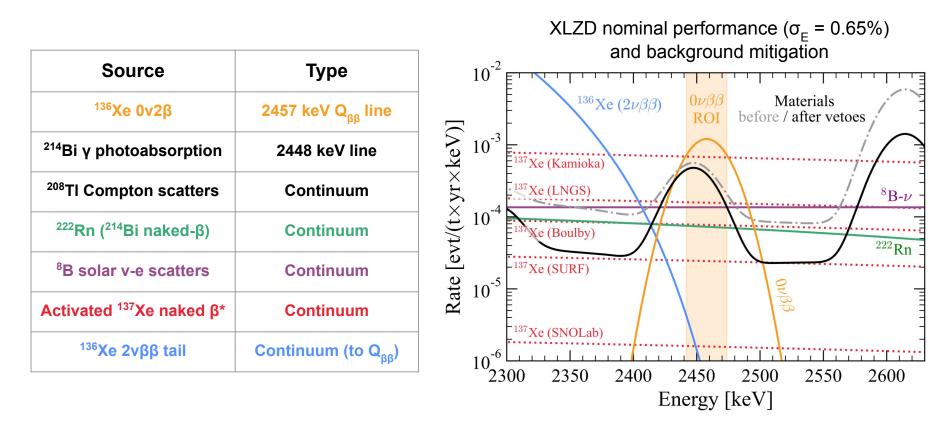
- LZ: σ/μ = 0.67% <u>JINST 18.04 C04007</u>
- XENON1T: σ/μ = 0.80% <u>EPJ C 80 1-9</u>
- XLZD targets:
  - Nominal 0.65% 0
  - Optimistic 0.60% Ο

Good position resolution  $\rightarrow$  multiple scatter rejection

- Z-position from time between S1 and S2
  - 3mm resolution nominal 0
  - 2mm optimistic <u>Saltao thesis</u> 0
- (X, Y) from S2 light pattern on top PMT array



### Components of search space



### **Background mitigation summary**

Common mitigation for all background sources is good energy resolution!

Source	Туре	Further mitigation	
<sup>136</sup> Xe 0v2β	2457 keV Q <sub>ββ</sub> line	_	
<sup>214</sup> Bi γ photoabsorption	2448 keV line Radiopure materials, self-shielding		
<sup>208</sup> TI Compton scatters	Continuum	Radiopure materials, self-shielding, veto	
<sup>222</sup> Rn ( <sup>214</sup> Bi naked-β)	Continuum Radiopure materials, distillation, Bi–Po tag		
<sup>8</sup> B solar v-e scatters	Continuum	Irreducible	
Activated <sup>137</sup> Xe naked $\beta^*$	Continuum	Deep site, shielding	
<sup>136</sup> Xe 2vββ tail	Continuum (to Q <sub>ββ</sub> )	Irreducible	

# Material radiopurity

Minimize  $\gamma$  &  $\beta$  backgrounds (<sup>214</sup>Bi and <sup>208</sup>TI)

#### Techniques:

- Strict material choices, e.g.
  - a. low-radioactivity PMTs
  - b. custom field-shaping resistors
  - c. cleanest batches of materials from  $\underline{LZ}$ , <u>XENON</u>, and other assays
- Clean assembly protocols

#### Performance assumptions:

- Realistic: 25% of LZ's external background rate
  - based on current assays
- Optimistic: 10% of LZ level

	m LZ (967 kg $ imes$ 1000 d)		$\begin{array}{c} \text{XLZD} \\ (8.2\text{t}\times10\text{yr}) \end{array}$	
Component	Nominal	Reduced	Projected	
TPC PMTs	2.95	0.98	0.61	
PMT structures	2.75	0.54	0.33	
Field-cage resistors	2.46	0	0	
Internal sensors	1.81	0.22	0.14	
PMT bases	1.52	0.39	0.24	
Cryostat	1.26	0.82	0.51	
PMT cables	1.01	0.16	0.10	
Field-cage rings	0.97	0.40	0.25	
OD tank supports	0.73	0	0	
OD foam	0.71	0	0	
Skin PMTs	0.69	0.06	0.04	
Other skin parts	0.68	0.05	0.03	
Other components	3.56	1.42	0.88	

<sup>214</sup>Bi 2448 keV gamma SS counts

LZ projection from *Phys. Rev. C. 102:014602* 

5.05

21.10

Total

3.15

### Self-shielding and veto systems

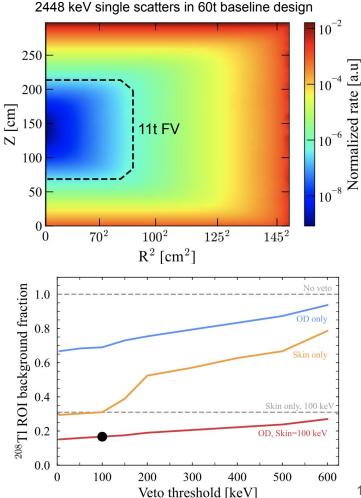
Reject external γ-ray backgrounds (<sup>214</sup>Bi and <sup>208</sup>TI)

Fiducialize to escape the <sup>214</sup>Bi 2448 keV peak

- LXe high density 10 cm Compton interaction length
- 60t design: 8.2t (nominal) and 11t (optimistic)
- 80t design: 13.6t (nominal) and 17.2t (optimistic)

Veto system to reject <sup>208</sup>TI Compton background

- Xenon skin and outer-detector (100 keV thresholds)
- Tag multiple-scattering Compton or associated gamma
- 83.3% rejection of <sup>208</sup>Tl in ROI



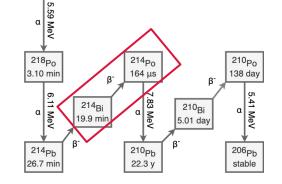
# Further reducing β backgrounds

<sup>214</sup>Bi β-decay

Cosmogenic <sup>137</sup>Xe production



• Further reduction from Bi-Po tagging



<sup>214</sup>Po is  $\alpha$  emitter – tag  $\beta$ + $\alpha$  sequence and veto event

#### **Performance assumptions:**

222<sub>Rn</sub> 3.82 d

- Nominal: 99.95% Bi–Po tagging efficiency
- Optimistic: 99.99%

Siting matters: deeper  $\rightarrow$  less muons  $\rightarrow$  less activation

Assume laboratory neutrons can be fully mitigated

#### **Performance assumptions:**

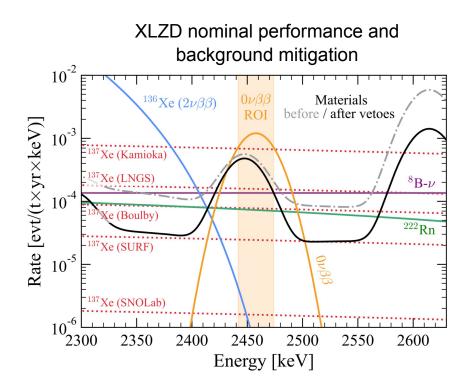
- Nominal: LNGS (3600 m.w.e.)
- Optimistic: SURF (4300 m.w.e)

	D	epth	$\mu$ flux	$^{137}$ Xe rate	SS ROI rate
Site	[m]	[m w.e.]	$[/(m^2 \cdot d)]$	$[/(t \cdot yr)]$	$[\mathrm{evt}/(\mathrm{t}{\cdot}\mathrm{yr}{\cdot}\mathrm{keV})]$
SNOLAB	2070	5890	< 0.3	0.007	$1.29 \times 10^{-6}$
$\mathbf{SURF}$	1490	4300	4.6	0.142	$2.72{ imes}10^{-5}$
Boulby	1300	3330	14.6	0.404	$7.73 { imes} 10^{-5}$
LNGS	1400	3800	29.7	0.822	$1.57 \times 10^{-4}$
Kamioka	1000	2700	128	3.54	$6.78 \times 10^{-4}$

# Summary of projection scenarios

- Nominal
  - existing detector performances,
  - LNGS installation,
  - 25% of LZ gamma background rates
- Optimistic
  - improved detector performance,
  - SURF installation,
  - ambitious gamma reduction (10%)

	Scenario		
Parameter	Nominal	Optimistic	
$^{222}$ Rn concentration [ $\mu$ Bq/kg]	0.1		
BiPo tagging efficiency [%]	99.95	99.99	
External $\gamma$ -ray [% LZ]	25	10	
Installation site	LNGS	SURF	
Energy resolution [%]	0.65	0.60	
SS/MS vert. separation [mm]	3	2	



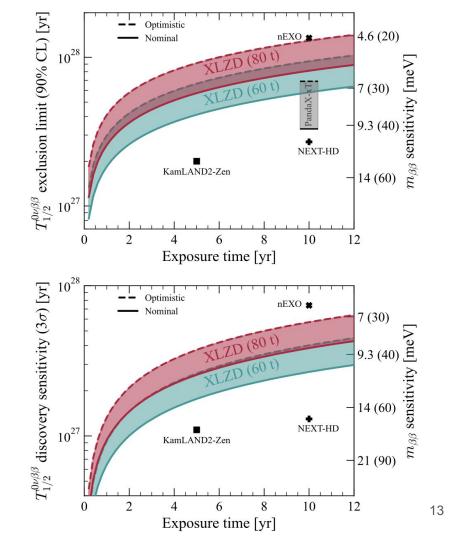
# Sensitivity projections

Counting experiment in an optimized fiducial volume

Sensitivity is primarily driven by target mass

Impact of design drivers on 3 $\sigma$  discovery sensitivity (Loss from reverting each optimistic input to nominal, all else fixed)

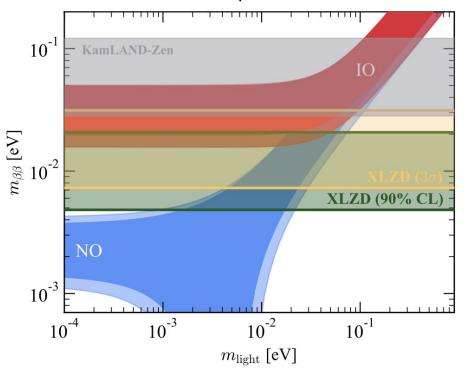
- Gamma background: 15.5%
- Experiment depth: 11%
- BiPo tagging efficiency: 6%
- Single-/multiple-scatter discrimination: 4.3%
- Energy resolution: 2.5%



### Sensitivity to effective Majorana mass

XLZD probes neutrino mass ordering, assuming only Majorana neutrino exchange

Band widths set by uncertainties of nuclear matrix elements



XLZD 80t optimistic scenario

## Conclusions

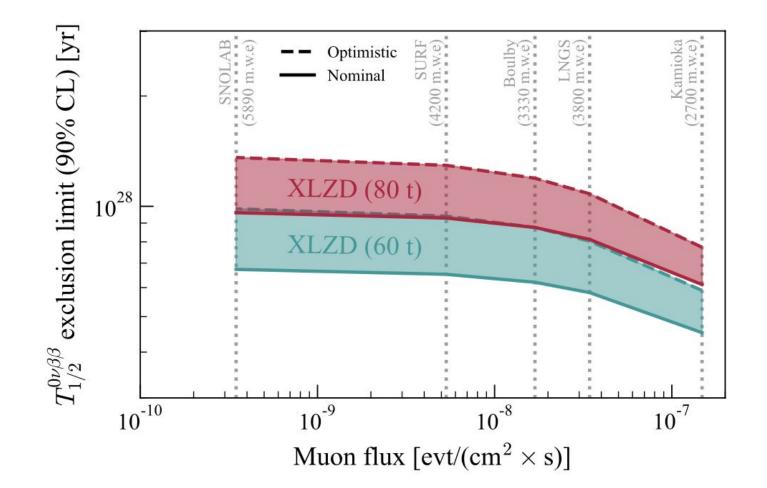
XLZD is a scalable rare-event observatory — with world-leading discovery potential for both dark matter and  $0\nu\beta\beta$ 

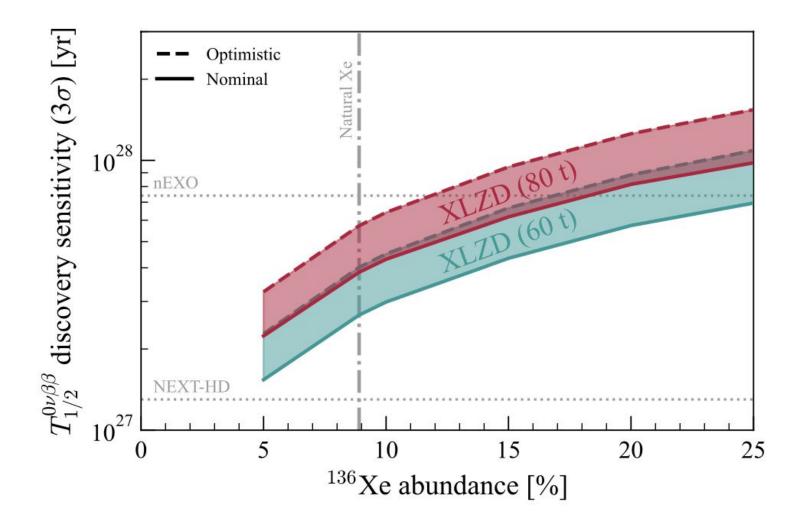
#### XLZD builds on mature dual-phase xenon technology

- 60–80t of active mass
- good background control and resolution

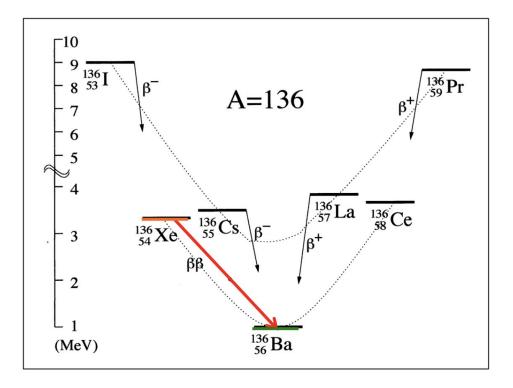
#### With natural xenon 80t optimistic scenario for XLZD can

- Discover  $0\nu\beta\beta$  at  $3\sigma$  for half-lives up to  $5.7 \times 10^{27}$  yr
- Exclude half-lives up to 1.3×10<sup>28</sup> yr at 90% CL
- Exclude the inverted neutrino mass ordering and explore the normal ordering

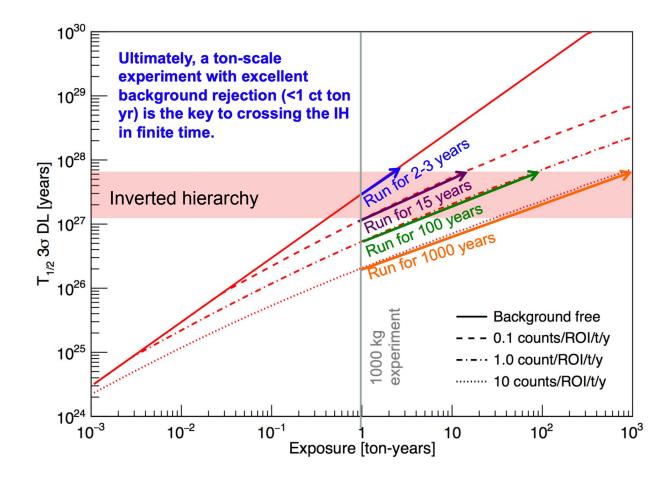




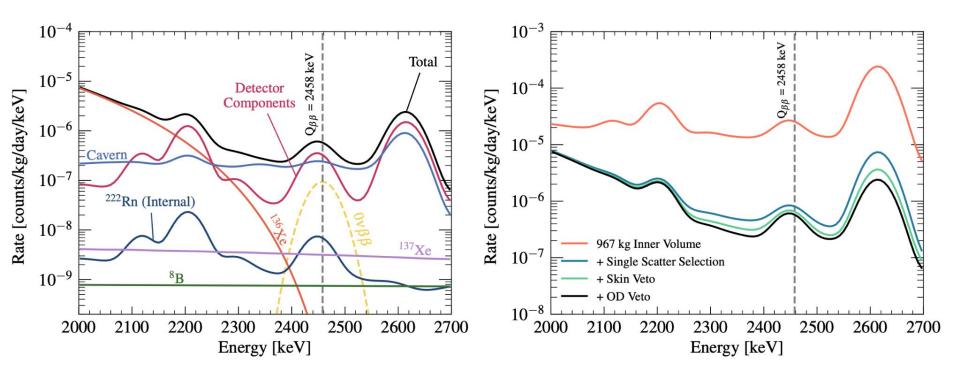
### A = 136 isobar



#### Physics of 0v2b TASI: https://arxiv.org/pdf/2108.09364



### LZ 0V2B simulation



### Material screening

- Voltage grading resistors EXO200 <u>https://arxiv.org/pdf/1202.2192</u>
- PMT base capacitors XENON1T https://arxiv.org/pdf/1705.01828
- Low radioactivity PMTs -

https://indico-tdli.sjtu.edu.cn/event/1861/contributions/11609/attachments/451 7/7175/XeSAT2024\_PMT12699.pdf

- Cleanest PTFE, stainless steel, copper, Kovar, and Kapton
  - LZ assay: <u>https://arxiv.org/pdf/2006.02506</u>
  - XENON1T assay: <u>https://arxiv.org/pdf/1202.2192</u>