

Nuclear Matter Equation of State from In-Medium Similarity Renormalization Group



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Outline

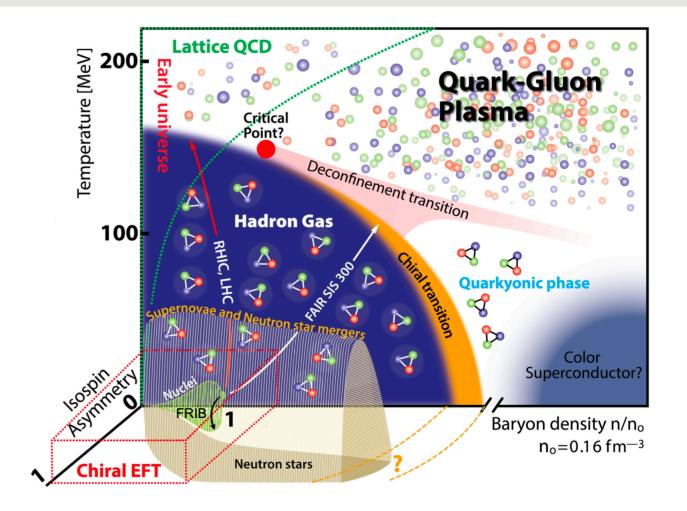
- Introduction
- IMSRG-EOS Framework
- EOS Results
- PMM Emulator
- Summary and Outlook



Introduction - Nuclear Matter and Equation of State

 Nuclear Matter: an idealized system of interacting nucleons in the thermodynamical limit

- Why is it interesting to us:
- Testing ground for many body methods
- Strongly related to dense astronomical objects
 like neutron stars, offers a link between nuclear
 physics and astrophysical observables

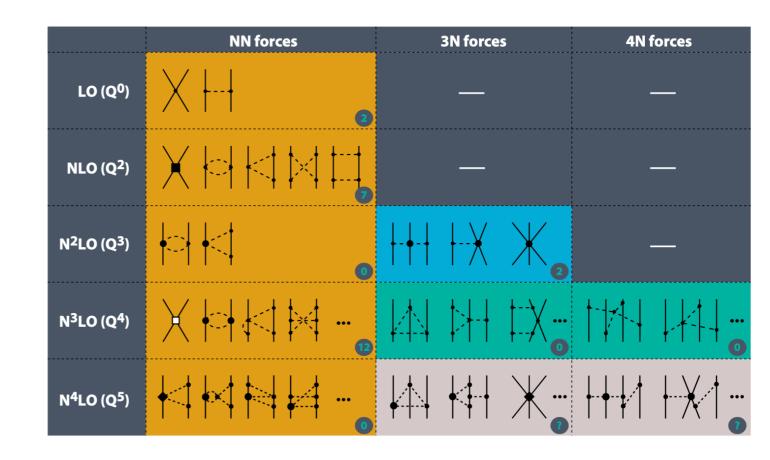




C. Drischler et al, *Chiral Effective Field Theory and the High-Density Nuclear Equation of State*

Introduction - Chiral EFT & Many Body Methods

- Chiral effective field theory:
 - Consistent NN, NNN, ... interactions
 - Systematic low-momentum expansion
 - Link with underlying QCD
- Many Body Methods: QMC, CC, MBPT, IMSRG...





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C. Drischler et al, *Chiral Effective Field Theory and the High-Density Nuclear Equation of State*

Basic Idea

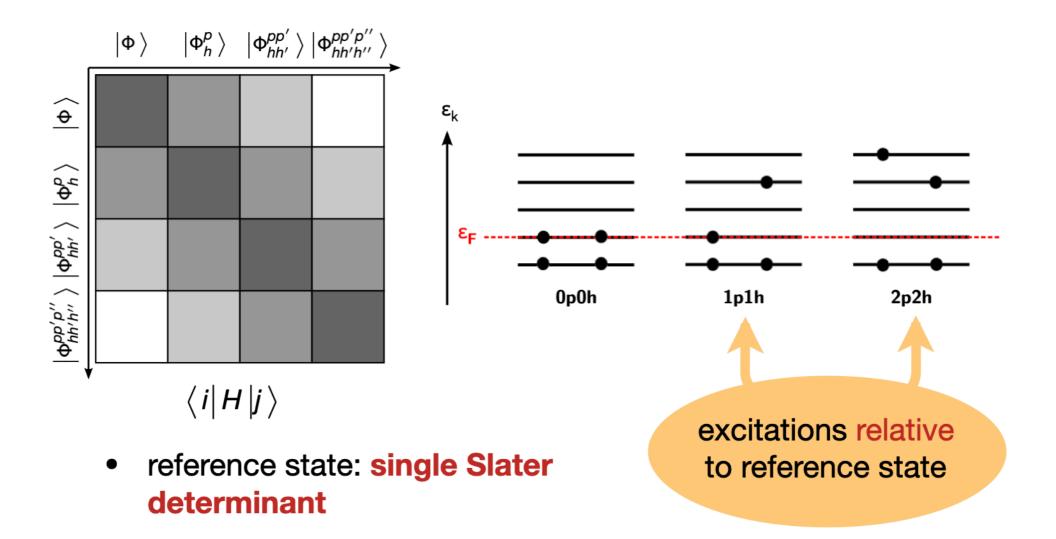
continuous unitary transformation of the Hamiltonian to banddiagonal form w.r.t. a given "uncorrelated" many-body basis



Normal-Ordered Hamiltonian

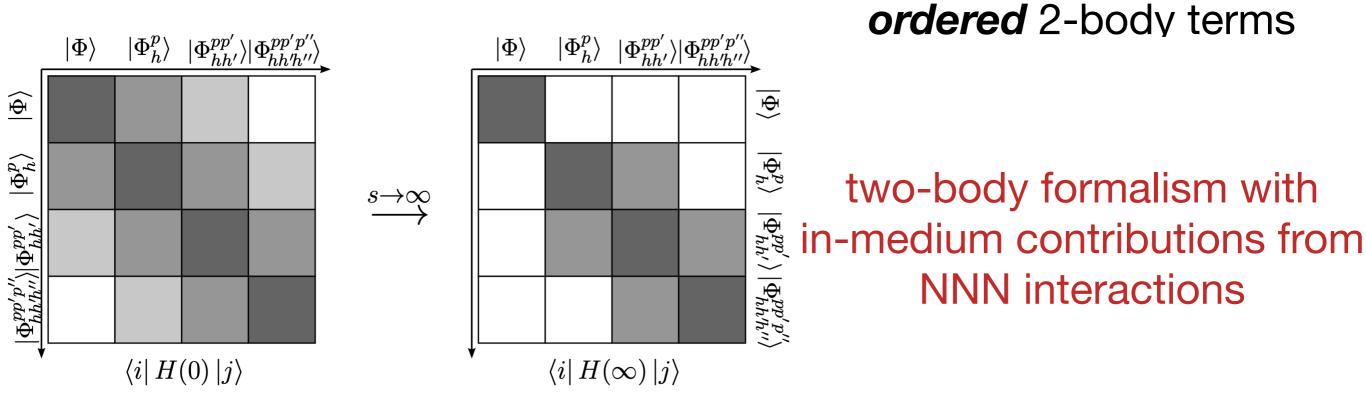
$$H = E_0 + \sum_{kl} f_l^k : A_l^k : + \frac{1}{4} \sum_{klmn} \Gamma_{mn}^{kl} : A_{mn}^{kl} : + \frac{1}{36} \sum_{ijklmn} W_{lmn}^{ijk} : A_{lmn}^{ijk} :$$
$$A_{j_1\dots j_N}^{i_1\dots i_N} \equiv a_{i_1}^{\dagger} \dots a_{i_N}^{\dagger} a_{j_N} \dots a_{j_1}$$
$$\langle \Phi | : A : |\Phi \rangle = 0$$







$$\frac{d}{ds}H(s) = [\eta(s), H(s)], \quad \text{e.g.,} \quad \eta(s) \equiv [H_d(s), H_{od}(s)]$$





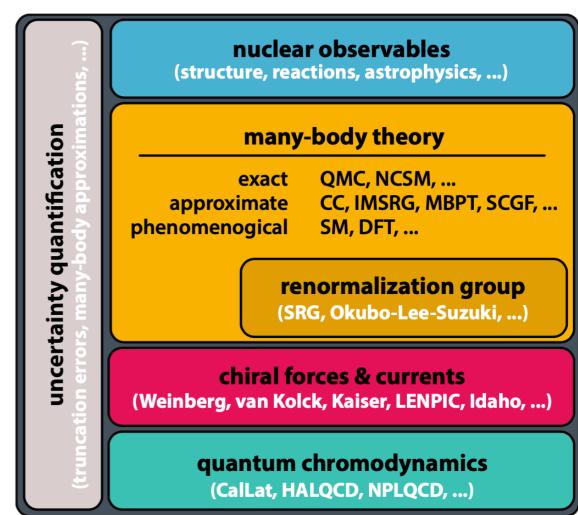
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IM-SRG(2): Truncate

H(s), η(s) to *normal*

IMSRG-EOS Framework - Overview

- Physical System: Nucleons in a finite box
- Framework:
- Single particle basis (plane waves w/Periodic boundary condition) -> Many particle basis
- 2. Input from chiral EFT -> Hamiltonian Matrix Elements
- 3. IMSRG Evolution of the Hamiltonian, NO2B level





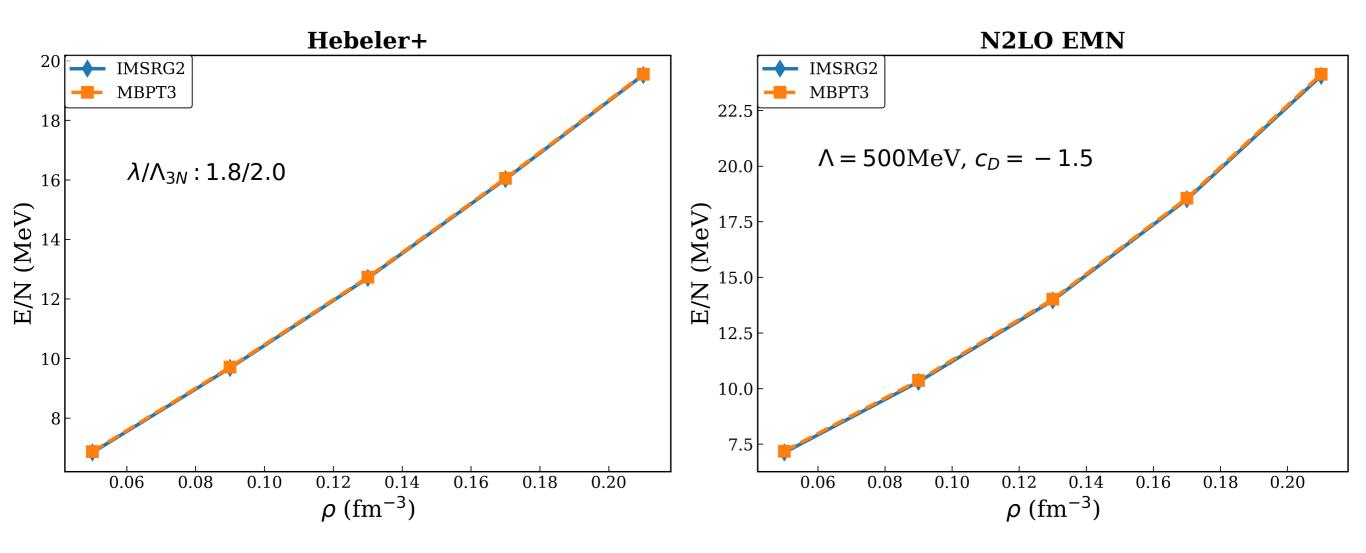
C. Drischler et al, A Brief Account of Steven Weinberg's Legacy in ab initio Many Body Theory

Preliminary Results - Interactions

- Hebeler Interactions:
 - Based on chiral EFT, but not fully consistent NN and 3N interactions
 - Starts from N3LO EM 500 MeV NN potential
 - NN interaction is softened by SRG evolution
 - NNLO 3N interaction adjusted to fit the triton binding energy and He charge radius
 - Denoted by λ/Λ_{3N} , where λ is the SRG flow parameter, Λ_{3N} is the 3N cutoff
- N2LO EMN Interactions:
 - Not SRG-evolved
 - consistent NN and 3N interaction
 - c_D and c_E are fitted to the ³H and empirical saturation properties

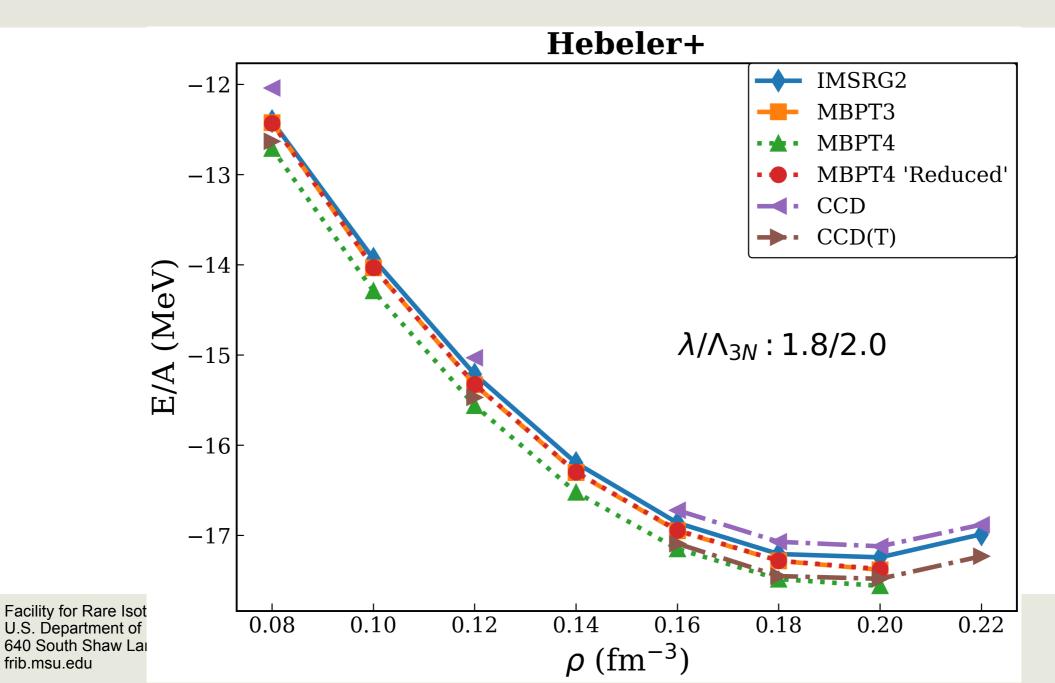


Preliminary Results - PNM EOS





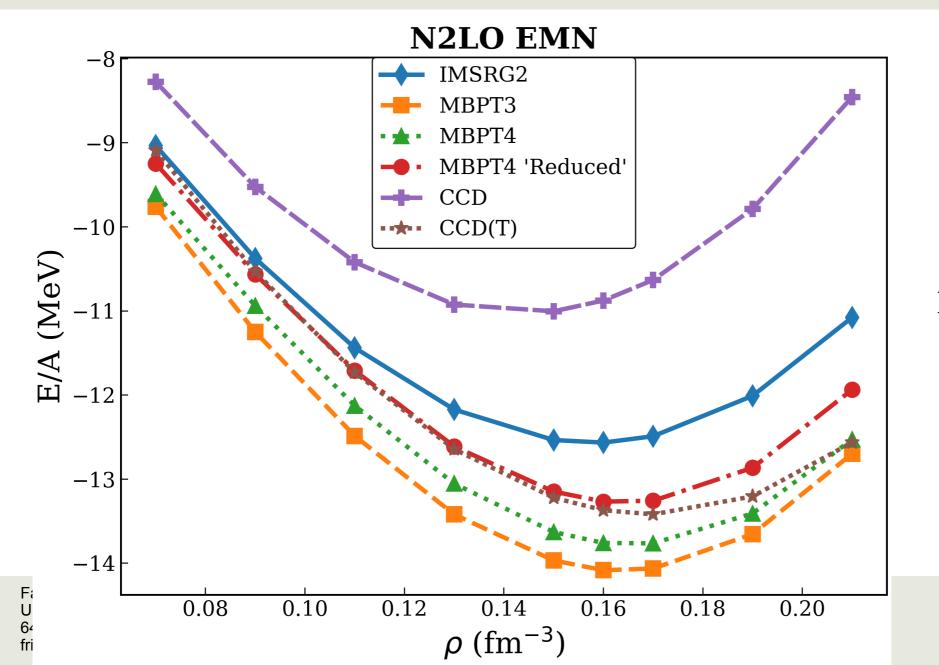
Preliminary Results - SNM EOS with Hebeler Interaction



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Preliminary Results - SNM EOS with N2LO EMN Interaction

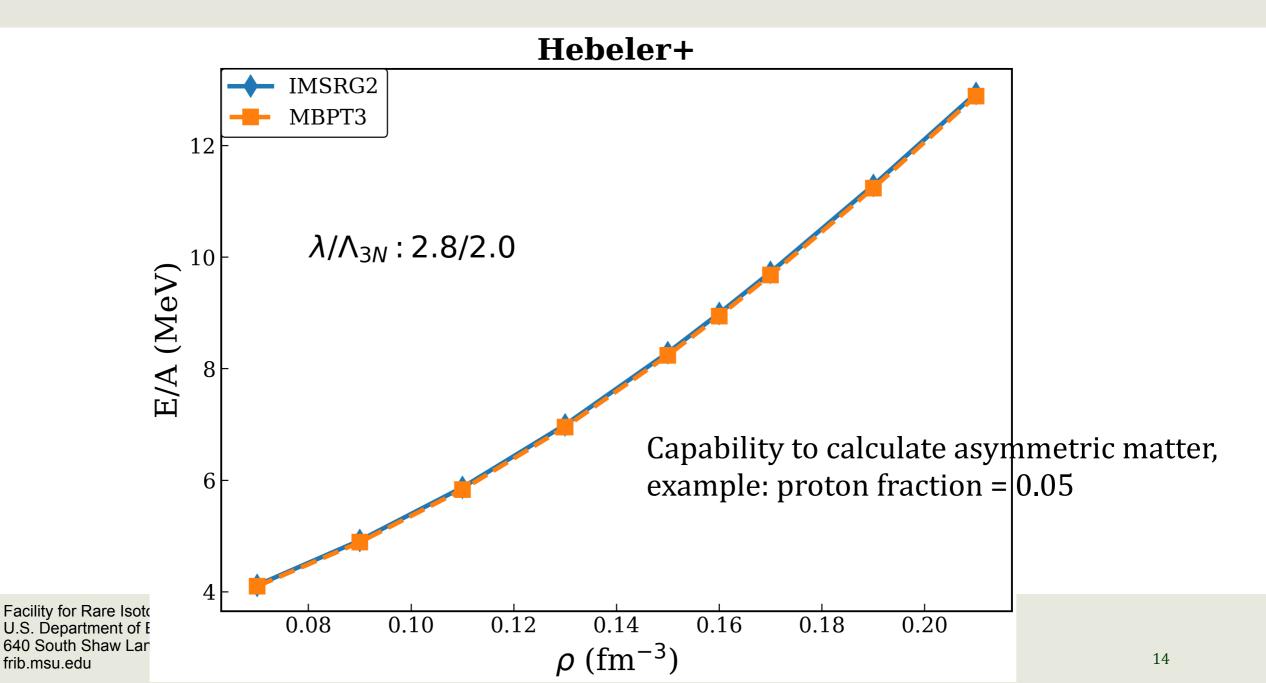


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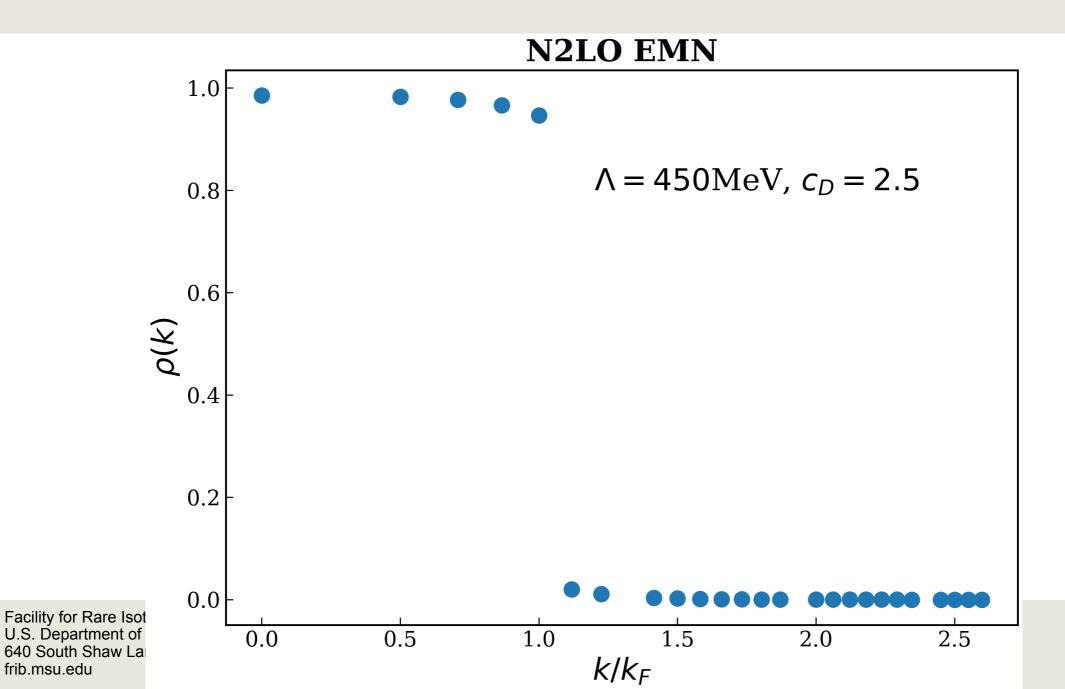
IMSRG(2) fall between CCD and CCD(T)

Preliminary Results - ANM EOS with Hebeler Interactions

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Preliminary Results - Single Nucleus Momentum Distribution with N2LO EMN Interaction

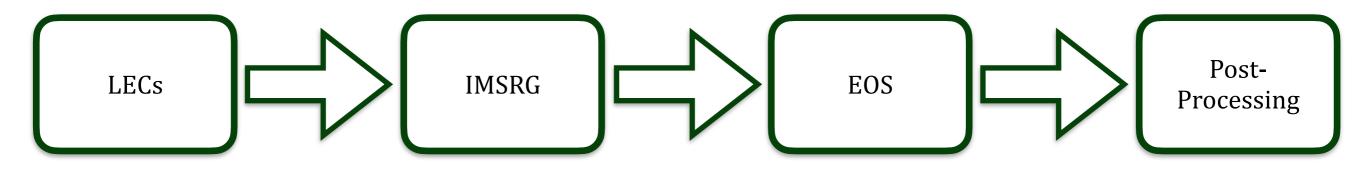


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Emulators - Why Do We Need Emulators?

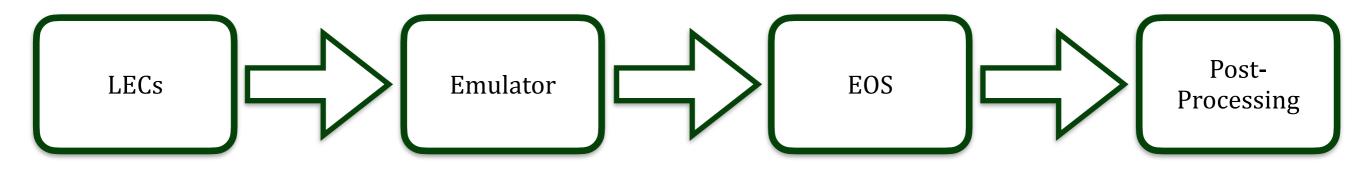
- Full IMSRG calculations are accurate but expensive
- Uncertainty quantification, sensitivity studies and bayesian analysis need millions of samples
- Replace full IMSRG solver with accurate enough but much faster emulator that predicts EOS from LECs in order to propagate uncertainties in LEC to EOS





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Emulators - Why Parametric Matrix Model (PMM)?

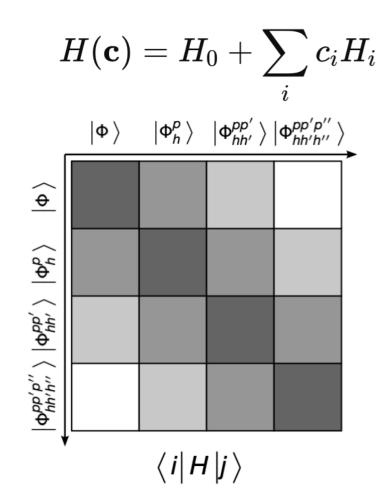
What is PMM?

Physics (Equation) - Based Machine Learning Algorithm



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Emulators - Why Parametric Matrix Model (PMM)?



$$M(\{c_l\}) = \underline{M_0} + \sum_{l=1}^p c_l \underline{M_l}$$

$\lceil r \rceil$	n_{11}	m_{12}	•••	m_{1n}
r	n_{21}	m_{22}	•••	m_{2n}
	• •	•	۰.	:
$\lfloor r \rfloor$	n_{n1}	m_{n2}	•••	m_{nn}

Infinite Dimensions in Many-Body Hilbert Space

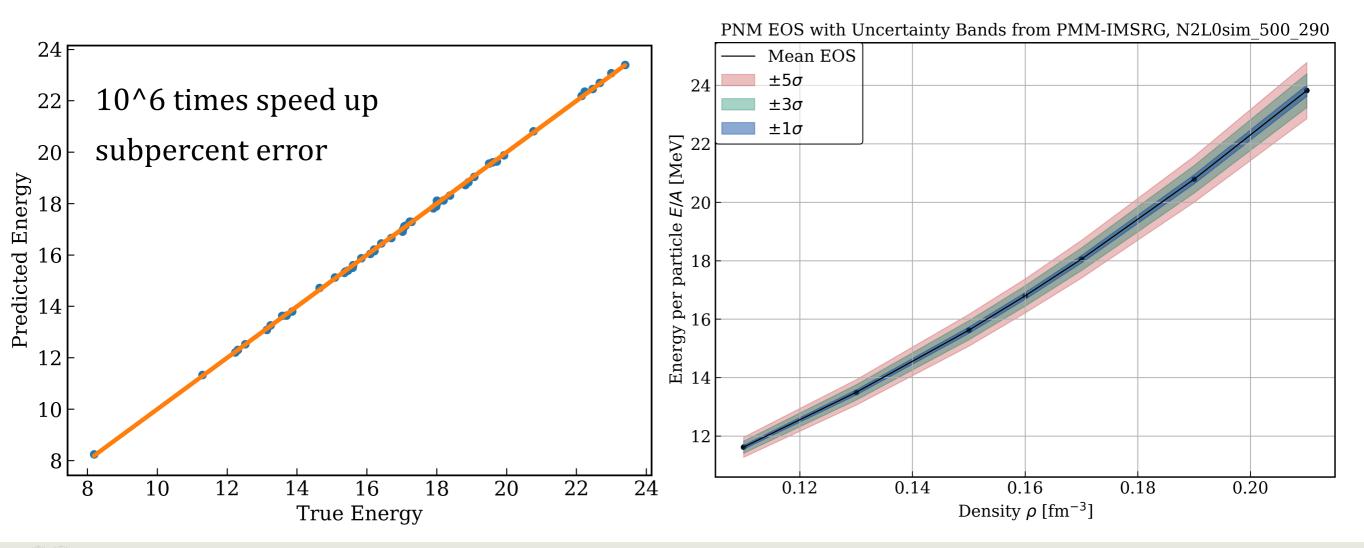


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Finite Dimensions in Learned Subspace

Patrick. Cook et al, Parametric Matrix Models

Emulators - Preliminary Results



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Summary

We've built a nuclear many-body modeling infrastructure based on IMSRG for EOS calculations.

- Converged calculations for a range of proton fractions
- Good agreement with other many body methods MBPT and CC for perturbative system (PNM)
- For more correlated system (SNM), the agreement is still good for softer interactions (Hebeler)
- Noticeable discrepancies starts to occur for relatively harder interactions (N2LO EMN)

We've constructed a fast and accurate emulator for the IMSRG framework based on a machine learning technique called parametric matrix model (PMM), which enables systematic propagation of nuclear interaction uncertainties to our EOS.



Outlook

- More detailed uncertainty quantification at T = 0 for different chiral EFT interactions with the PMM emulator
 - Bayesian analysis of EFT truncation errors (BUQEYE)
 - Comparison of different many body methods
- Provide astrophysical available ab initio EOS with tunable LECs using PMM emulator and reliable interpolators in both density and proton fraction
- Possible extensions
 - Finite T (see Smith et al. <u>https://arxiv.org/abs/2407.00576</u>)
 - Approximate IMSRG(3) (see Stroberg et al., <u>https://arxiv.org/abs/2406.13010</u>)
 - Response via EOM (and other) techniques



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C. Drischler et al, *Quantifying uncertainties and correlations in the nuclear-matter* equation of state

S. R. Stroberg et al, *IMSRG with flowing 3 body operators, and approximations thereof* Isaac G. Smith et al, *The IMSRG at Finite Temperature*

Thank you for your attention!



• flow equation for Hamiltonian $H(s) = U(s)HU^{\dagger}(s)$:

$$\frac{d}{ds}H(s) = \left[\eta(s), H(s)\right], \quad \eta(s) = \frac{dU(s)}{ds}U^{\dagger}(s) = -\eta^{\dagger}(s)$$

• choose $\eta(s)$ to achieve desired behavior, e.g.,

$$\eta(\mathbf{s}) = \begin{bmatrix} \mathbf{H}_{d}(\mathbf{s}), \mathbf{H}_{od}(\mathbf{s}) \end{bmatrix}$$

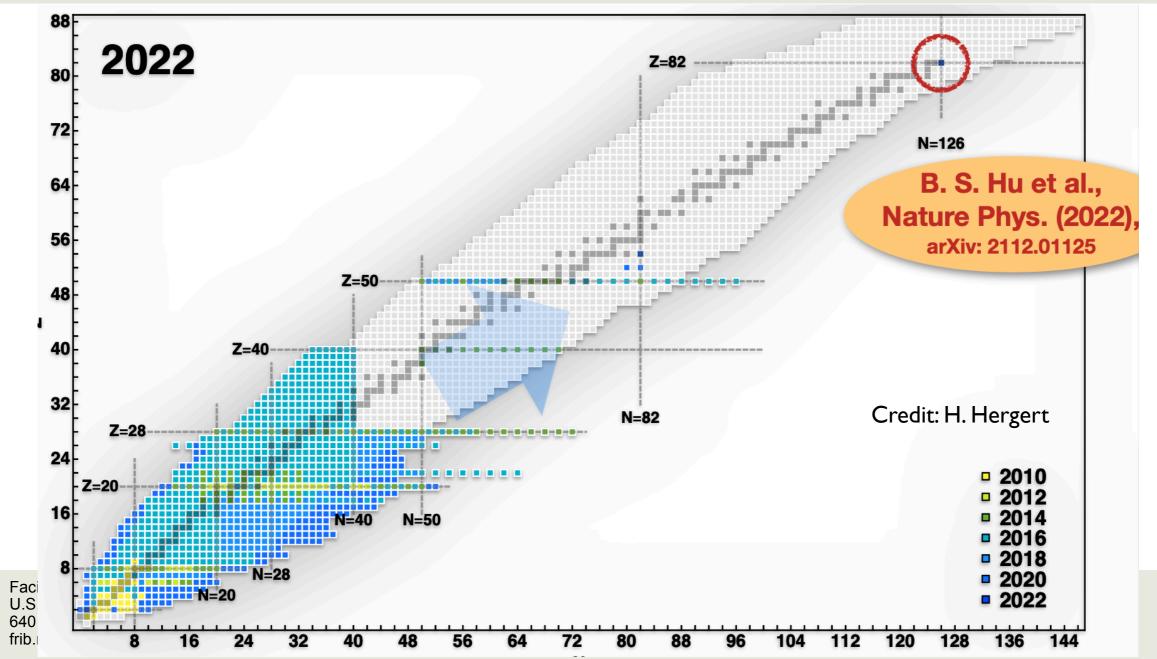
to suppress (suitably defined) off-diagonal Hamiltonian

$$\lim_{s \to \infty} H_{od}(s) \longrightarrow 0$$



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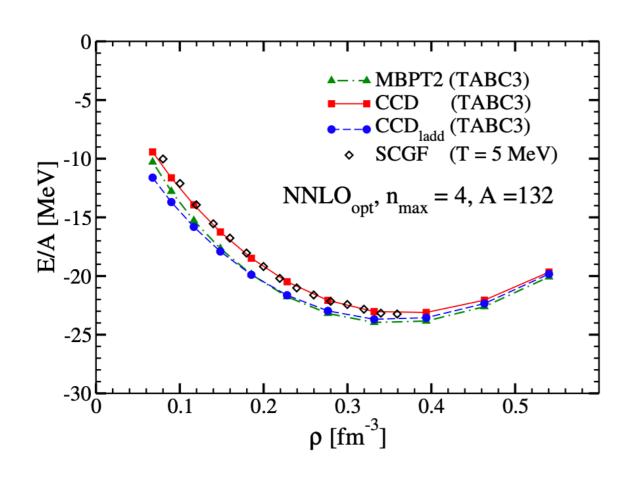
Introduction - Application of Ab Initio Methods in Nuclear Structure

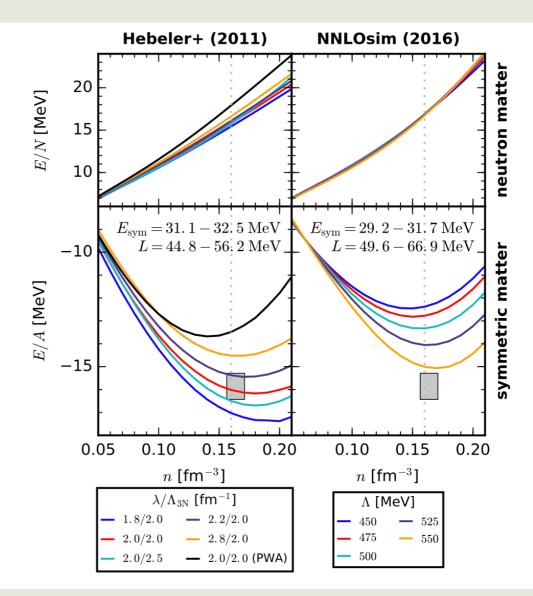


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Introduction - Application of Other Ab Initio Methods in NM EOS

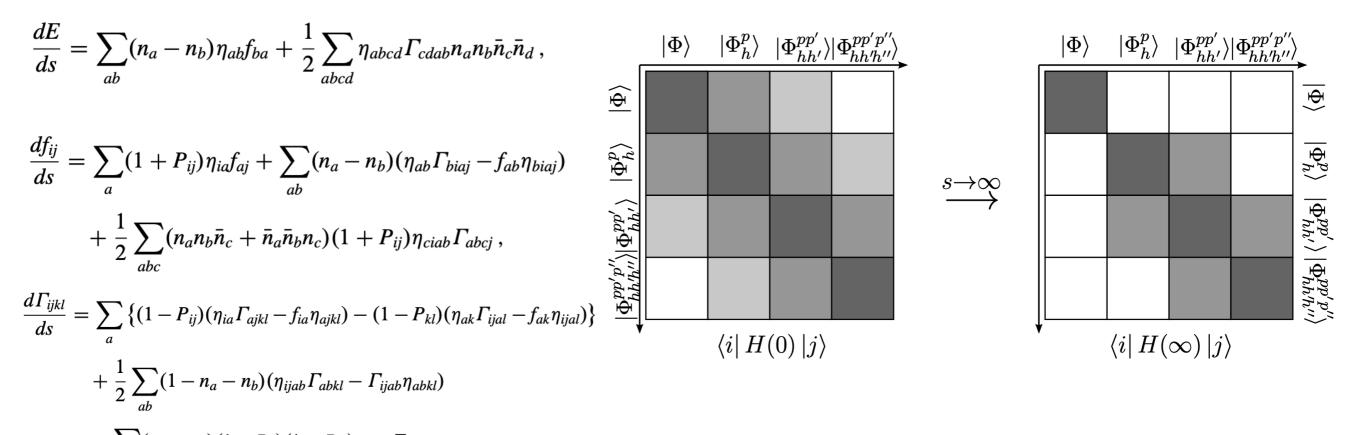




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G. Hagen et al, Coupled-cluster calculations of nucleonic matter C. Drischler et al, *Chiral Interactions up to Next-to-Next-to-Leading* Order and Nuclear Saturation 26

IMSRG(2) flow equations:



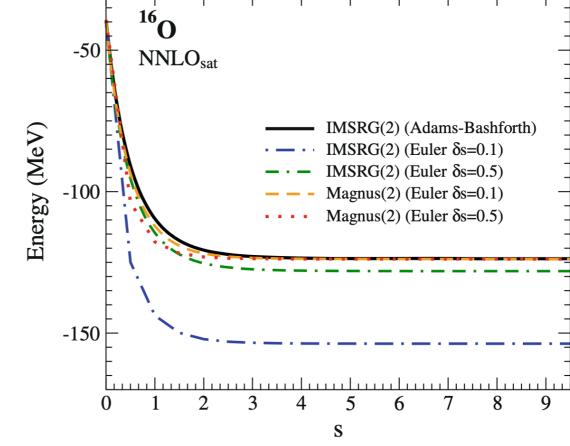
$$+\sum_{ab}(n_a-n_b)(1-P_{ij})(1-P_{kl})\eta_{aibk}\Gamma_{bjal}.$$



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Bridging the Scales from Quarks to Neutron Stars
H. Hergert et al, Nuclear Structure from the In-Medium Similarity
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Introduction - Magnus Formulation

¹⁶**0** $\hat{U}(s) \equiv e^{\hat{\Omega}(s)}$ -50 **NNLO**_{sat} $\frac{d\hat{\Omega}}{ds} = \sum_{k=0}^{\infty} \frac{B_k}{k!} \operatorname{ad}_{\hat{\Omega}}^k(\hat{\eta})$ Energy (MeV) -100 $\operatorname{ad}_{\hat{O}}^{0}(\hat{\eta}) = \hat{\eta}$ $\operatorname{ad}_{\hat{O}}^{k}(\hat{\eta}) = [\hat{\Omega}, \operatorname{ad}_{\hat{O}}^{k-1}(\hat{\eta})]$ -150 $\hat{H}(s) \equiv e^{\hat{\Omega}(s)}\hat{H}(0)e^{-\hat{\Omega}(s)} = \sum_{k=0}^{\infty} \frac{1}{k!} \operatorname{ad}_{\hat{\Omega}(s)}^{k} \left(\hat{H}(0)\right)$ 2 3 0 5 4 6 S



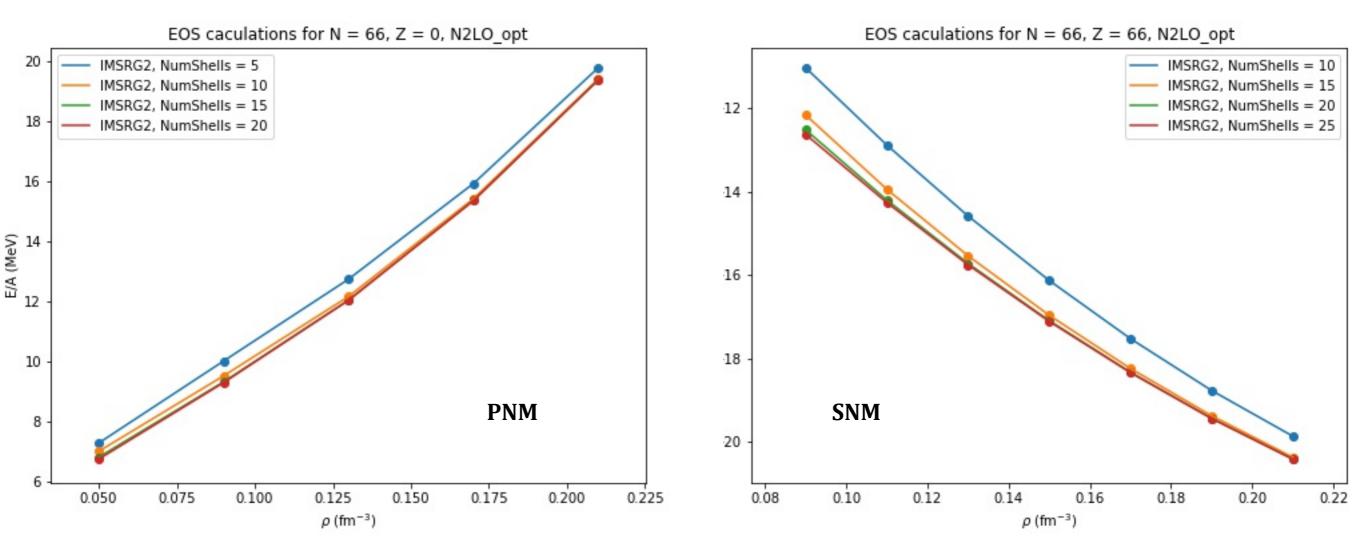
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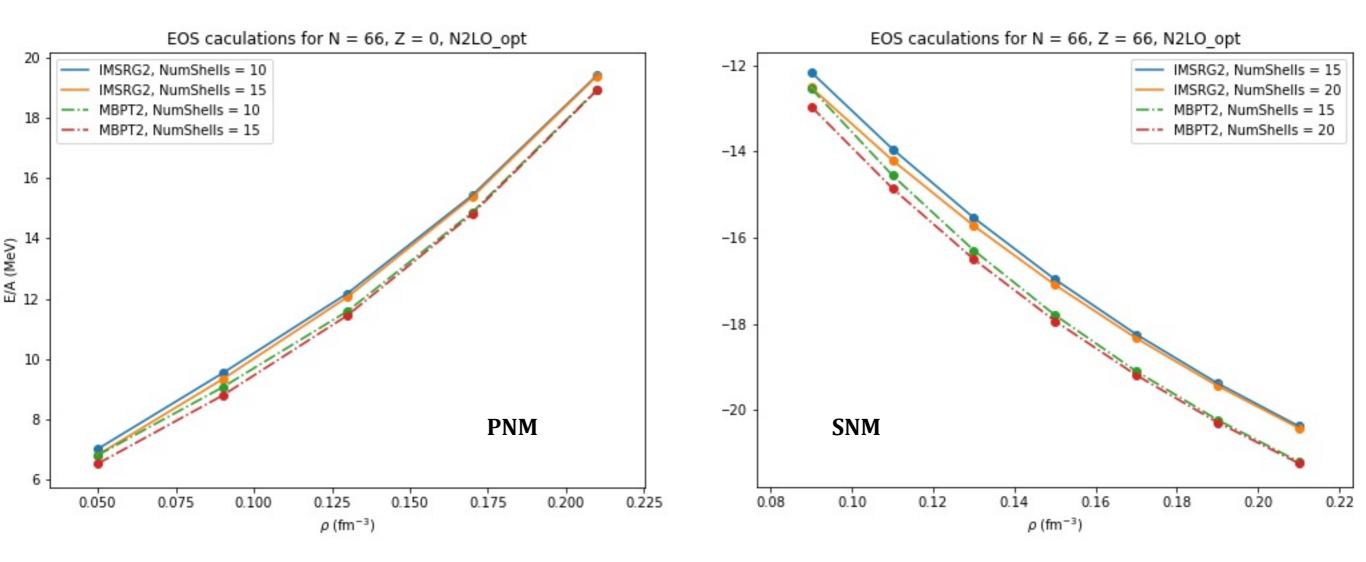
M Hjorth-Jensen et al, An Advanced Course in Computational Nuclear Physics: Bridging the Scales from Quarks to Neutron Stars H. Hergert et al, Nuclear Structure from the In-Medium Similarity 28

Results - Basis Convergence



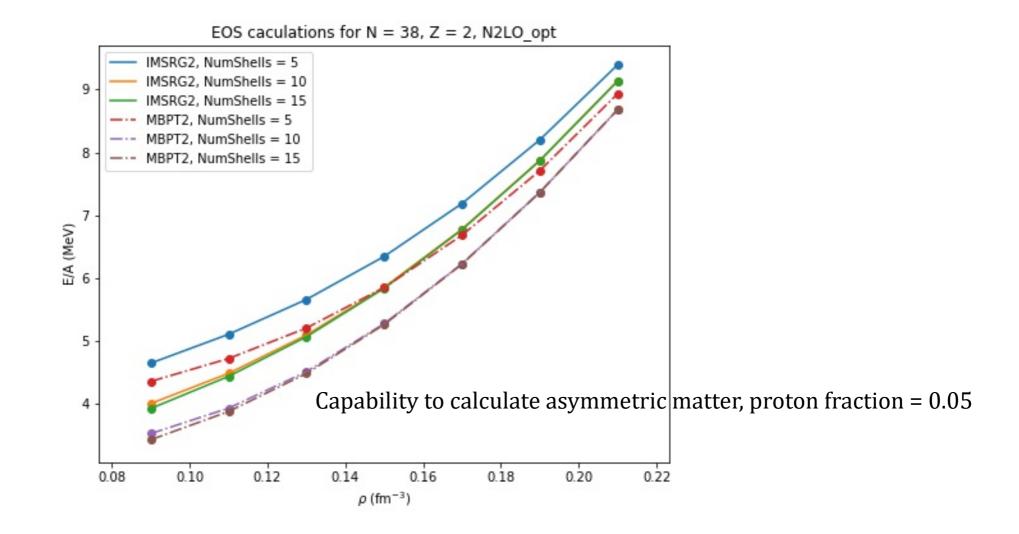


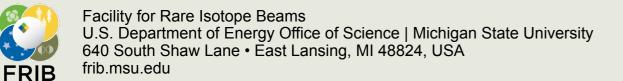
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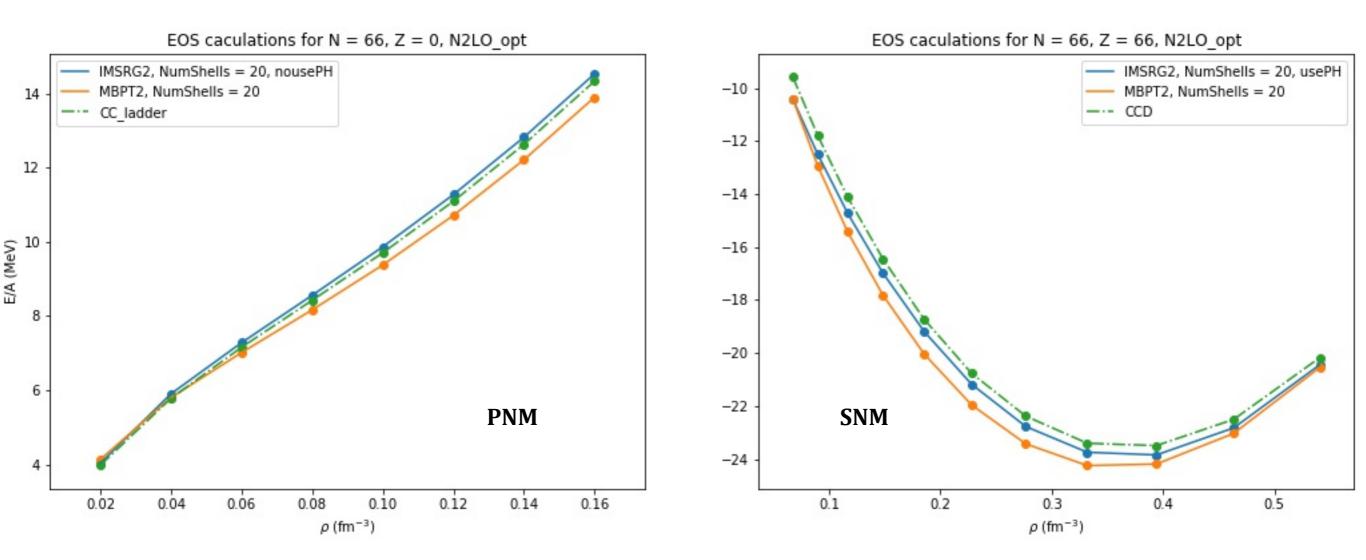


Results - Basis Convergence





Results - Benchmark

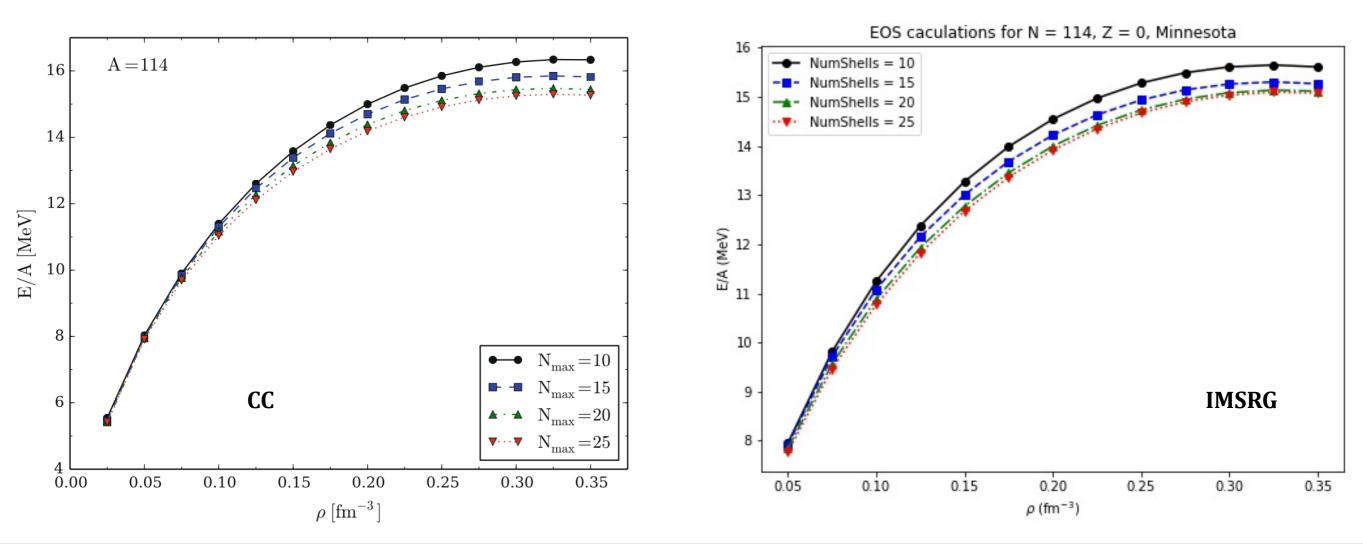




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G. Hagen et al, Phys. Rev. C 89, 014319

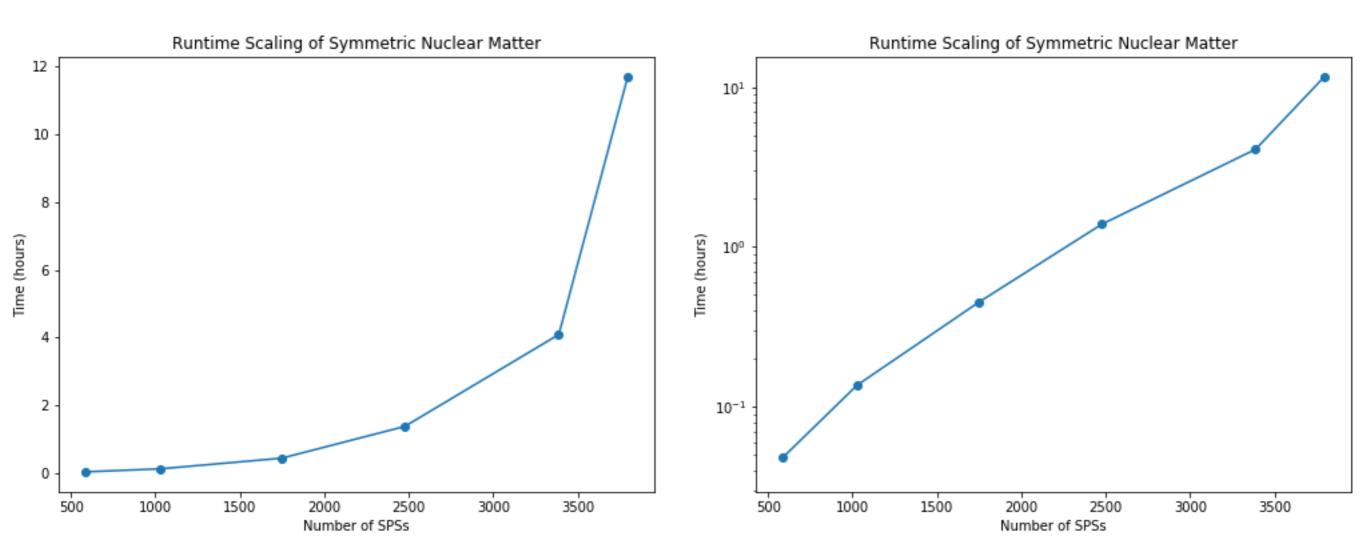
Results - Benchmark





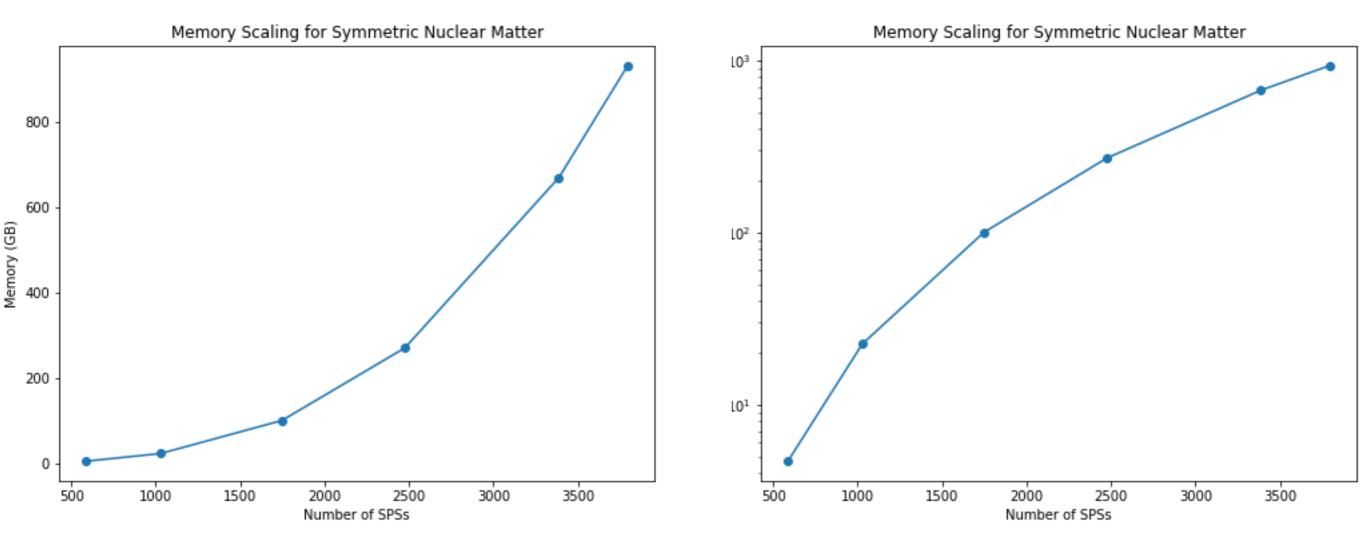
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Results - Scaling





Results - Scaling





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Results - Finite Size Effect

