Which long-baseline neutrino experiment?

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Options for a U.S. long-baseline program

NOVA - 1:

- off-axis beam from FNAL, L=810 km to 25 kt
 TASD
- · beam energy is tuned to 1st osc. max.
- neutrino mode only

Sensitive only to $\theta_{13} \Rightarrow$ upgrade required

Possible upgrades:

NOVA - II:

upgrade FNAL proton infrastructure (HINS)
run in neutrino and antineutrino modes
2nd detector (liquid Ar TPC at original site)
2nd detector (50 Kt water CherenKov)
at same L but different OA angle (2nd osc. max.)
at same L/E but shorter L (diff. matter effect)

Wide band beam:

upgrade in FNAL proton infrastructure (HINS)
run in neutrino and antineutrino modes
300 kt water Cherenkov detector at DUSEL
energy spectrum information

Off-axis studied extensively. Consider WBB

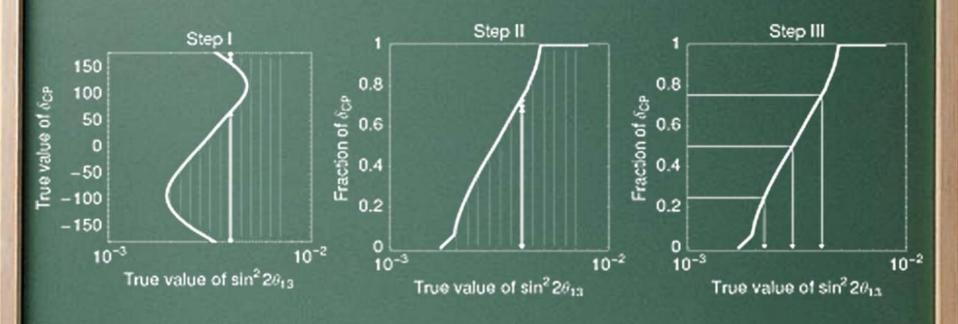
Performance indicators:

1. θ_{13} discovery potential - exclusion of $\theta_{13} = 0$

2. discovery of mass hierarchy - suppose Δm_{31}^2) o how well can Δm_{31}^2 (o be excluded?

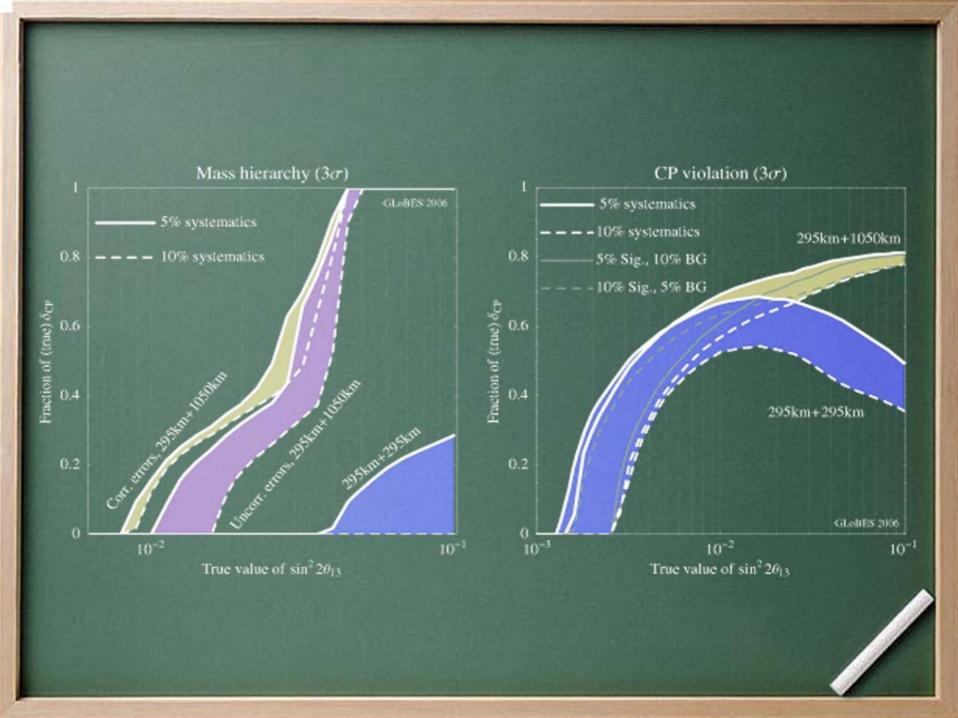
3. CP violation - exclusion of CP conserving values, $\delta = 0, \pi$

Definition of CP fraction



T2KK

• 4MAW protons from Tokai same decay tunnel as for T2K • two 270 Kt WC detectors • two baselines, 295 km and 1050 km same off-axis angle • 4 years neutrinos + 4 years antineutrinos • π° rejection as in T2K



NOVA*

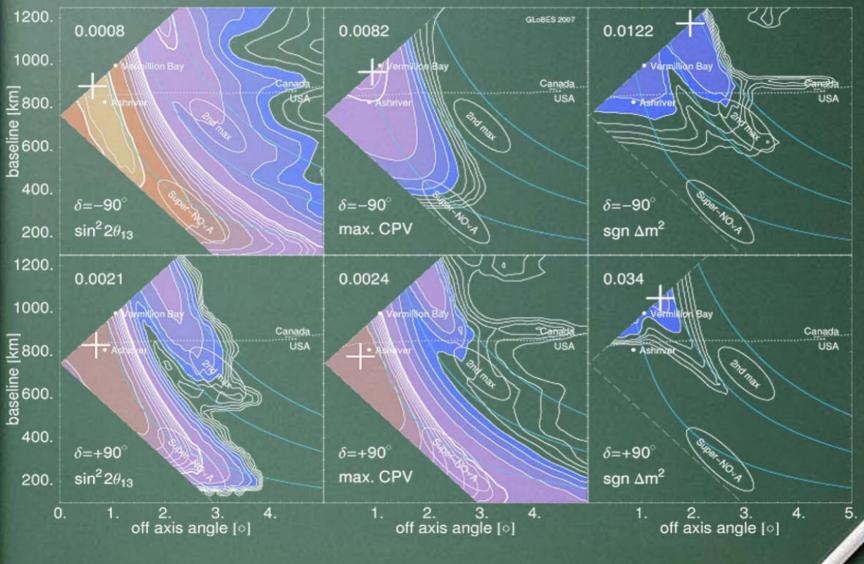
• 1.13 MAW from MAI at FNAL

- same decay tunnel as for MINOS and NOVA
- · 100 Kt LArTPC
- 3 years neutrinos + 3 years antineutrinos of
 25kt NOvA at Ash River
- +3 years neutrinos + 3 years antineutrinos of
 both

LArTPC

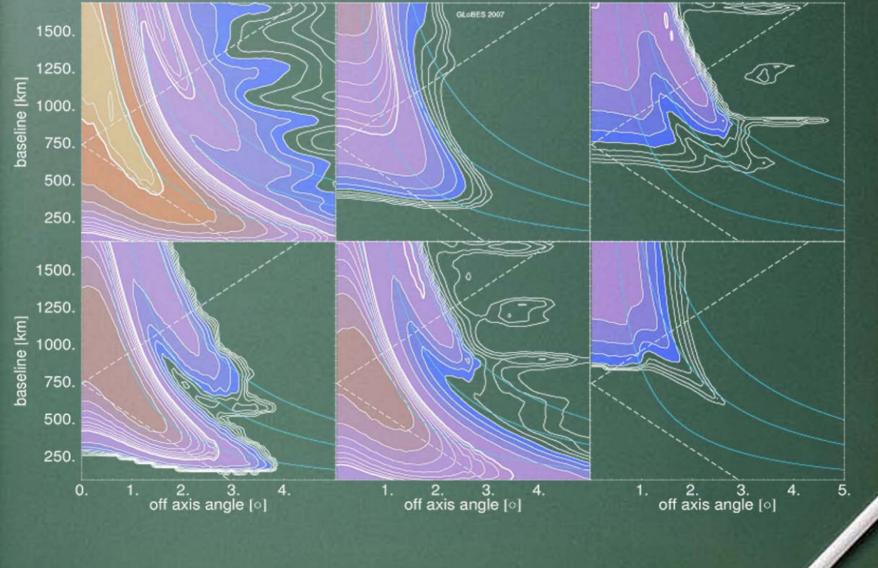
80% efficiency
no NC background
5% energy resolution for QE events
20% energy resolution for non-QE events

Where to put NOVA*?



and the second

On-axis or off-axis?

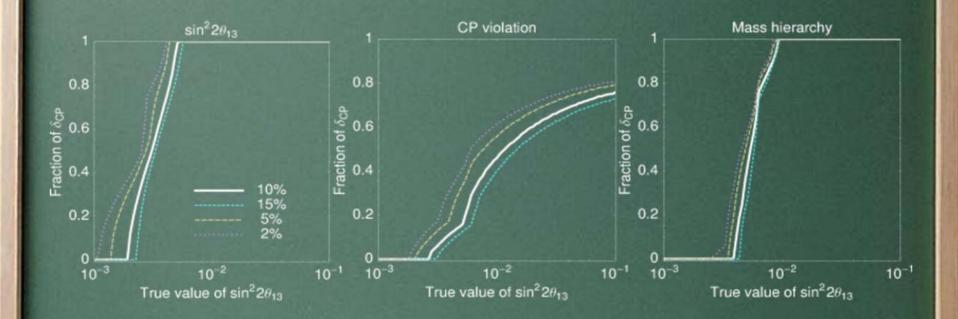




• $I(v) + 2(\overline{v}) \wedge W \text{ at } 28 \text{ GeV}$ · 300 kt WC detector on axis • L = 1300 km• 5 years neutrinos + 5 years antineutrinos · performance based on full detector MC • improved π^{o} rejection

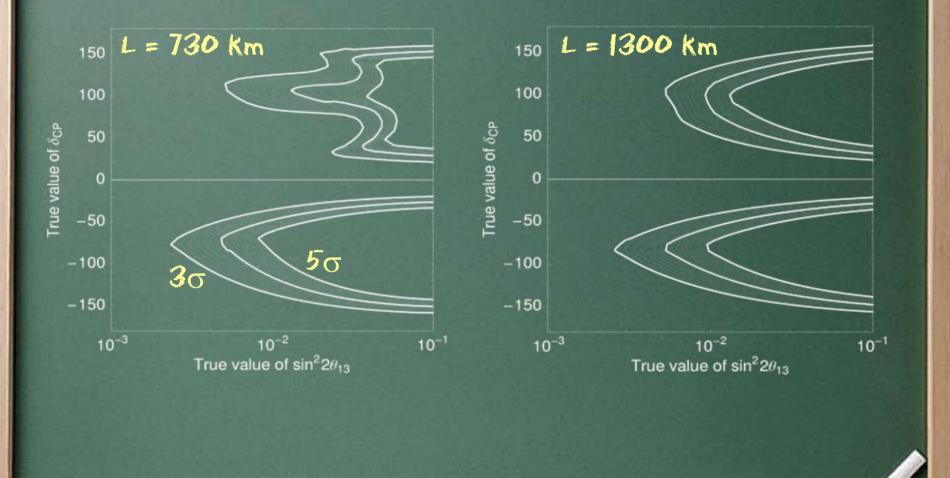


30 sensitivities

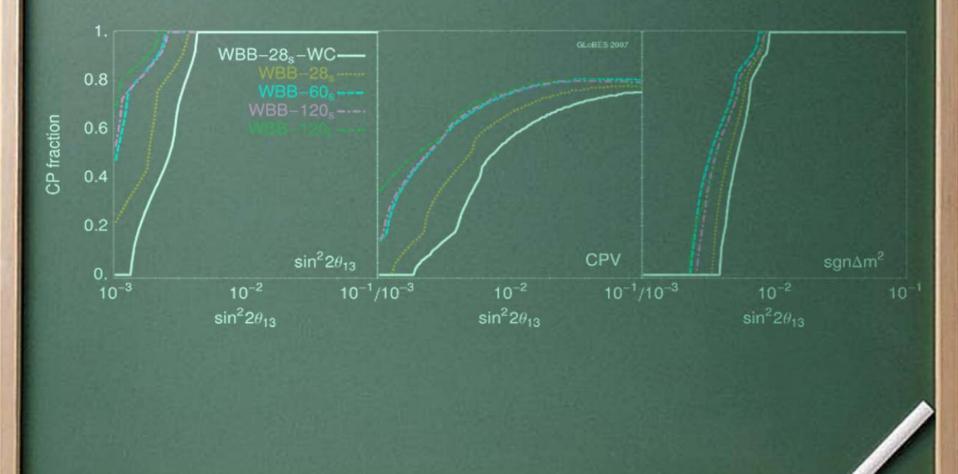


Barger, DiercKxsens, Diwan, Huber, Lewis, Marfatia, Viren (hep-ph/0607177)

Discovery potential for CP violation



Proton energies and decay tunnel length



Different proposals assume different

- uptime/year
- number of years
- detector size
- beam power

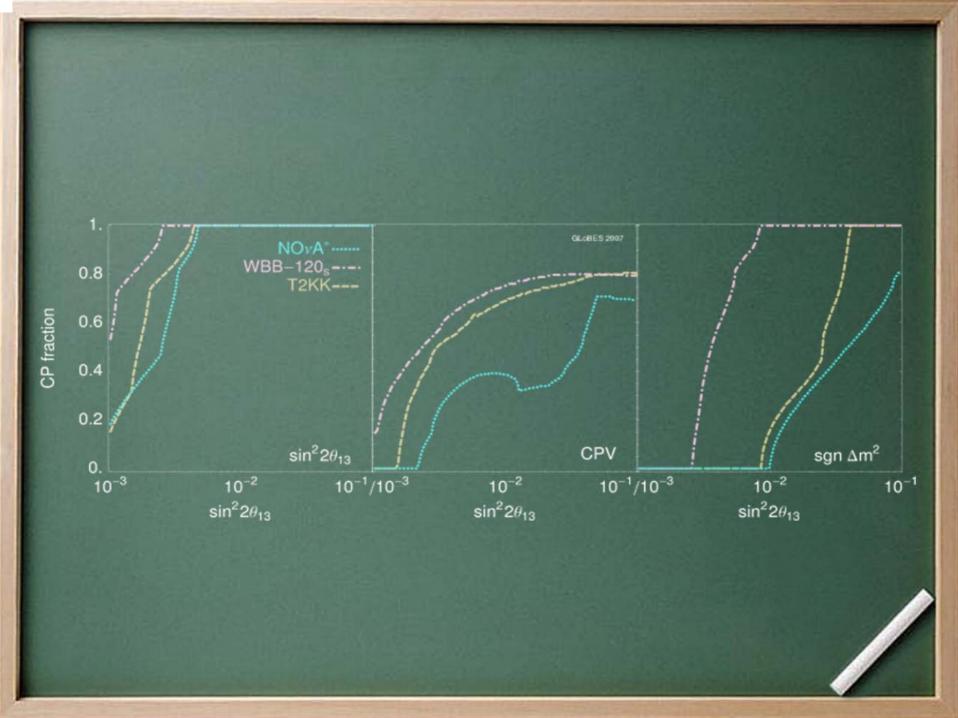
Exposure = detector mass (MAt) x target power (MAW)

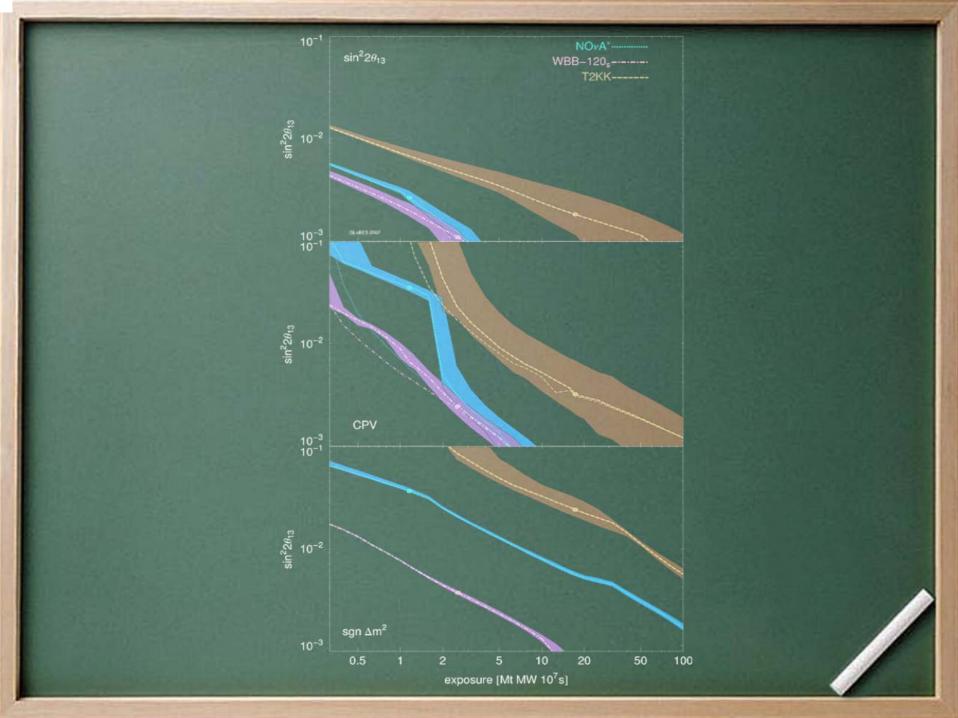
x running time $(10^7 s)$

is a measure of integrated luminosity

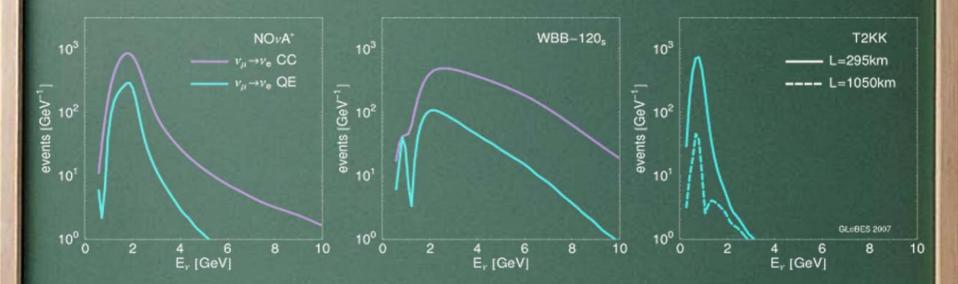
Very approximately, $cost \propto exposure$

Setup	t_{ν} [yr]	$t_{\bar{ u}}$ [yr]	P_{Target} [MW]	$L [\mathrm{km}]$	Detector technology	$m_{\rm Det} \; [{\rm kt}]$	L
$NO\nu A^*$	3	3	$1.13~(u/ar{ u})$	810	LArTPC	100	1.15
WBB-120 _S	5	5	$1 (\nu) + 2 (\bar{\nu})$	1290	LArTPC	100	2.55
T2KK	4	4	$4 (\nu/\bar{\nu})$	295 + 1050	Water Cherenkov	270 + 270	17.28
β -beam	4	4	n/a	730	Water Cherenkov	500	n/a
NuFact	4	4	4	3000 + 7500	Magn. iron calor.	50 + 50	n/a

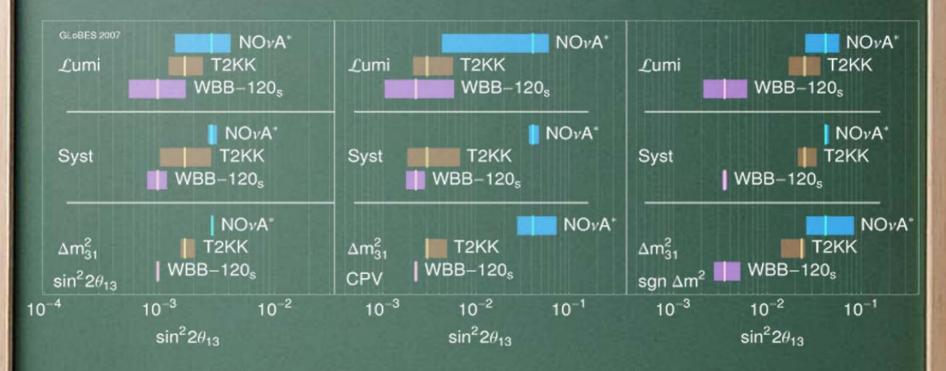




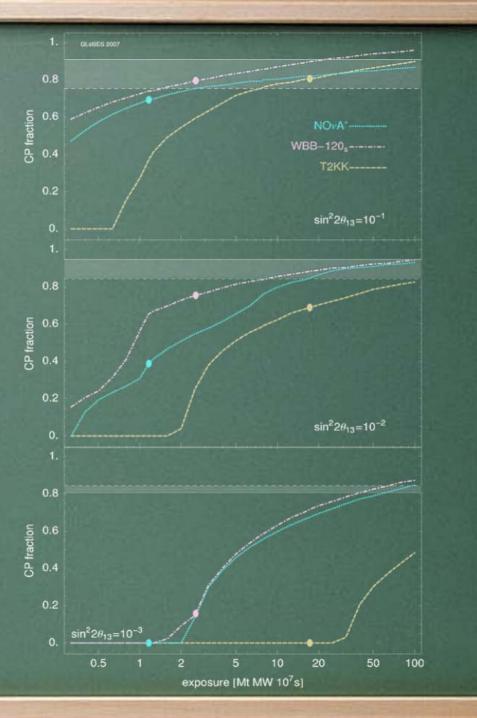
Event rates for 1 MAT MAW 107s and $\sin^2 2\theta_{13} = 0.04$



Robustness



• Exposure from 2 to 0.5 nominal value • Systematics from 2% to 10% • Δm_{31}^2 from 2-3x10⁻³ eV² CPV



· Optimal NOVA* location in the U.S. is Ash River . To be competitive for CPV, it is crucial that NOVA* gets enough exposure () 2 rat raw 10^7 s) • WBB experiments can make all 3 measurements and have the most robust performance • WBB-LAr better than WBB-WC if cost/kt of LAr is smaller than cost/4 kt of water · Every strategy requires MAW beams, O.I MAt detectors and 10 years of running (0.5 billion \$) • For $\sin^2 2\theta_{13}$ > 0.01 no need for a NuFact or beta-beam