

Warped Penguins

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PHENO 10

The Penguins

- Penguin Diagrams are loop-induced processes.
- We focus on $\mu \rightarrow e \gamma$,



References

FCNC in RS

R.Kitano (00);S.J. Huber (03); K.Agashe, G.Perez, A.Soni (04)(05);
G.Moreau, K.Agashe, A.E.Belchman, F. Petriello (06); J.I.Silva-Marcos (06);
A.L. Fitzpatrick, L.Randall, G.Perez (08); K. Agashe, A. Azatov, and L. Zhu
(09)...

Feynman Rules in 5D

L.Randall, M. D. Schwartz (01); M. Puchwein, Z. Kunszt (04); R. Catino, A.
Pomarol (04)...

Reduceing the Flavor Constraints

J. Santiago (08); M. Blanke, A. J. Buras, B. Duling, S. Gori, A. Weiler (09); A.
J. Buras, B. Duling, and S. Gori (09); M. E. Albrecht, M. Blanke, A. J. Buras,
B. Duling, K. Gemmler (09); K. Agashe (09); C. Csaki, C. Delaunay, C.
Grojean, Y. Grossman (09)...

Problem: $\mu \rightarrow e \gamma$ in RS is divergent!?



Previous Analyses

$\mu \rightarrow e \gamma$ with brane higgs loop is **UV sensitive!**

- Tree-level process: $\mathcal{M}_{\text{tree}} + \mathcal{M}_{\text{loop}}(\Lambda_c)$.
- Loop-induced process: $\mathcal{M}_{\text{loop}}(\Lambda_c) + \dots$
- Cannot get a physical result.

Problem: $\mu \rightarrow e \gamma$ in RS is divergent!?



We Claim

One loop $\mu \rightarrow e \gamma$ in RS is **finite!**

- Can get a physical result for the amplitude.
- Gives interesting bounds on Yukawa.

The divergence in the KK picture

Naive NDA

		KK	SM
Loop Integral	$(\sum) d^4k$	5	4
Fermi-Propg	$(1/k)^2$	-2	-2
Higgs-Propg	$(1/k)^2$	-2	-2
Gauge Symm	$(k/k^2) \rightarrow (p/k^2)$	-1	-1
k -Power		0	-1



Hard to see the finiteness in KK!

- Infinitely large mass matrix including the KK-towers.
- Complications from summing the KK-modes.

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Show the finiteness with a full 5D calculation!

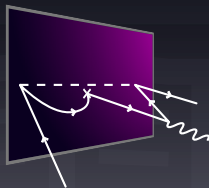
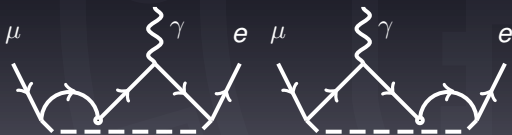
- No need to do the KK-sum.
- The only cutoff is the 4D momentum.

Results of the 5D Calculation

Finiteness & Branching ratio:

- The **finiteness** comes from the Lorentz invariance & chiral BC.
- The **branching ratio** given by the explicit 5D-loop calculation is

$$\text{Br}(\mu \rightarrow e \gamma) \simeq 2 \cdot 10^{-7} \left(\frac{3\text{TeV}}{M_{\text{KK}}} \right)^4 \left(\frac{Y_*}{2} \right)^4$$



The Yukawa bounds from FCNC

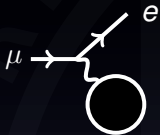
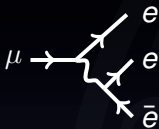
How to use this result?

$$\text{Br}(\mu \rightarrow e \gamma) \simeq 2 \cdot 10^{-7} \left(\frac{3\text{TeV}}{M_{\text{KK}}} \right)^4 \left(\frac{Y_*}{2} \right)^4$$

We can use the flavor bounds on the **Yukawa** coupling.

Use the branching ratio to constraint the Yukawa

There are two different FCNC processes:



gives a lower bound on Y .



gives an upper bound on Y .

They constraint the Yukawa from both sides!

The Yukawa bounds

Combine the two results

There is a tension between the $\mu \rightarrow e$ & $\mu \rightarrow e \gamma$ bounds!

For $M_{kk} = 3 \text{ TeV}$,

$$Y_*(\mu \rightarrow e) > 3.7 \quad Y_*(\mu \rightarrow e \gamma) < 0.2$$

These flavor bounds are important for model building.

The Yukawa bounds

Combine the two results

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These flavor bounds are important for model building.

To relax the tension: Custodial Symmetry, A_4 Symmetry, ...

Match the **5D** & **KK** results.

Finiteness & Amplitude

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Finiteness & **Amplitude**

Matching the KK & 5D results

Calculating the amplitude in **KK** picture, if we do

$$\mathcal{M} = \sum_{n=1}^N \int_0^\infty d^4k \hat{\mathcal{M}}^{(n)}(k)$$

the result is different from the **5D** result.

Matching the KK & 5D results

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5D \neq KK !?

Matching the KK & 5D results

The reason is,

Wrong UV limit!

To match the 5D & KK, the momentum cutoff \lesssim KK cutoff.

One way to do this is

$$\lim_{N \rightarrow \infty} \sum_{n=1}^N \int_0^{M_{KK}} d^4 k \hat{\mathcal{M}}^{(n)}(k)$$

This gives the same result as in 5D. **5D = KK**

Conclusion

- One loop $\mu \rightarrow e\gamma$ in RS is **finite**.
- A tension between the $\mu \rightarrow e$ & $\mu \rightarrow e\gamma$ Yukawa bounds.
- The **nontrivial UV limit** in KK calculation is important for matching the KK & 5D results.

$\mu \rightarrow e \gamma$ in 5D

The diagrams of $\mu \rightarrow e \gamma$:

- Get the leading order result by doing **mass insertions**.
- Rotate all the **flavor mixing** to the **Yukawa**.
- It is an **L-R** operator: $a \bar{L} \sigma^{\mu\nu} F_{\mu\nu} E$
- From the **gauge symmetry**:

$$\epsilon_\mu \mathcal{M}^\mu \sim \epsilon_\mu \bar{u} (p^\mu + p_e^\mu - (m_\mu + m_e) \gamma^\mu) u$$

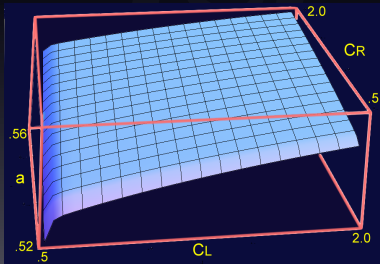
- To get **a**, we only need to calculate the coefficient of p^μ .

The Result

The result of the amplitude can be written as follows:

$$a_{kl} \times R'^2 \frac{e}{16\pi^2} \frac{v}{\sqrt{2}} \left(f_{Li} Y_{ik} Y_{kl}^\dagger Y_{lj} f_{-E_j} \right) \bar{L}_i^{(0)} \sigma^{\mu\nu} E_j^{(0)} F_{\mu\nu}^{(0)}$$

The loop integral gives $a_{kl} \simeq 0.5$.



$$\text{Br}(\mu \rightarrow e \gamma)_{\text{thy}} > 8.2 \cdot 10^{-7} a^2 \left(\frac{3\text{TeV}}{M_{KK}} \right)^4 \left(\frac{Y_*}{2} \right)^4$$

$$\text{Br}(\mu \rightarrow e \gamma)_{\text{exp}} < 1.2 \cdot 10^{-11}$$

for $M_{KK} = 3\text{TeV}$, we have

$$Y_* < 0.2$$

The fermion propagator in 5D

$$(-\not{p} + i\gamma^5 \partial_5 + m)\Delta(p, z, z') = i\delta(z - z')$$

Trick:

$$(p^2 - \partial_5^2 + m^2)F(p, z, z') = i\delta(z - z')$$

$$\Delta(p, z, z') = (\not{p} - i\gamma^5 \partial_5 + m)F(p, z, z')$$

In flat XD case:

$$F(p, z, z') = A(p, z') \cos(pz) + B(p, z') \sin(pz)$$

The chiral BC

The chiral BC constraints the form of the amplitude!



$$\Psi_L = \begin{pmatrix} \chi_L \\ \bar{\psi}_L \end{pmatrix}$$

$$\Psi_R = \begin{pmatrix} \chi_R \\ \bar{\psi}_R \end{pmatrix}$$

$$\Delta = \begin{pmatrix} \Delta_{\psi\chi} & \Delta_{\psi\psi} \\ \Delta_{\chi\chi} & \Delta_{\chi\psi} \end{pmatrix} = \begin{pmatrix} D_+ F_- & \sigma^\mu p_\mu F_+ \\ \bar{\sigma}^\mu p_\mu F_- & D_- F_+ \end{pmatrix}$$

How to calculate the 5D loop?

Feynman's trick with Bessel functions?? We don't know that...

However, we can

- **Taylor expand** the propagator into powers of the **external momentum** (p^μ, q^μ). $\frac{1}{(k^2 + 2k \cdot q)} = \frac{1}{k^2} \left(1 - \frac{2k \cdot q}{k^2} + \dots \right)$.

- Isolate the p^μ terms. Get the coefficient. $\not{p}\gamma^\mu = 2p^\mu - \gamma^\mu \not{p}$.

- Solve the numerical integral, get the coefficient of $\mu \rightarrow e \gamma$.

$$\mathbf{a} \times R'^2 \frac{e}{16\pi^2} \frac{v}{\sqrt{2}} \left(f_L Y Y^\dagger Y f_E \right) \bar{u} \not{a} u; \quad \mathbf{a} = \int dx \int dy \text{ (scalar function)}.$$

Why is $\mu \rightarrow e \gamma$ finite?

Lorentz invariance + chiral BC gives the finiteness:

The propagators are composed of two parts

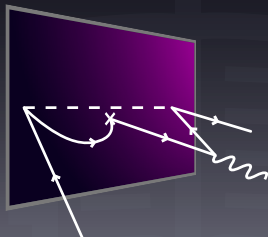
- The Dirac operator part (gives the γ^μ structure).
- The Bessel function part (contains the 5D profile).

In the UV limit, the photon vertex is pulled back to the IR brane.

- Can just look at the Dirac operator part.

Each loop contains two sectors:

- photon emission
- mass insertion



Why is $\mu \rightarrow e \gamma$ is finite?

- γ emission: Lorentz inv + chiral BC gives $(\not{k}\gamma^\mu\not{k} - k^2\gamma^\mu)$
- mass insertion: The chiral BC gives \not{k} .
- Combining the two m-insertion amp, the leading order becomes

$$\mathcal{M}_{(a)} + \mathcal{M}_{(b)} \sim \not{k} (\not{k}\gamma^\mu\not{k} - k^2\gamma^\mu) + (\not{k}\gamma^\mu\not{k} - k^2\gamma^\mu) \not{k} = 0$$

- From NDA, only the leading order term can be divergent.

$\mu \rightarrow e \gamma$ is finite!