

# Complex Singlet Extended SM and CoGeNT (and Xenon100)

Mathew McCaskey  
University of Wisconsin-Madison

Collaborators: Vernon Barger and Gabe  
Shaughnessy (arXiv:1005.xxxx)

# Outline

- Motivation
- Intro to the complex singlet extended SM (csxSM)
- Constraints
- Results
  - Spin independent scattering cross section
  - Higgs signatures
- The debate: CoGeNT vs. Xenon100
- Conclusions

# Motivation

- Recent results from the CoGeNT collaboration favor a light DM scenario: arXiv:1002.4703
  - $M_{DM} \sim 7 - 11 \text{ GeV}$
  - $\sigma_{SI} \sim 10^{-41} - 10^{-40} \text{ cm}^2$
- In the context of the complex singlet extended SM (csxSM): arXiv:0811.0393
  - Satisfy CoGeNT, WMAP, LEP constraints.
  - Study Higgs phenomenology

# The csxSM

$$V_{\text{csxSM}} = \frac{m^2}{2} H^\dagger H + \frac{\lambda}{4} (H^\dagger H)^2 + \frac{\delta_2}{2} H^\dagger H |\mathcal{S}|^2 + \frac{b_2}{2} |\mathcal{S}|^2 + \frac{d_2}{4} |\mathcal{S}|^4 \\ + \left( \frac{|b_1|}{4} e^{i\phi_{b_1}} \mathcal{S}^2 + |a_1| e^{i\phi_{a_1}} \mathcal{S} + c.c. \right)$$

$$\mathcal{S} = S + iA$$

- Real component,  $S$ , obtains a vacuum expectation value  $v_S$
- Complex component,  $A$ , is stable (DM particle)
- For viable phenomenology we set  $\phi_{b_1} = \pi$  and  $\phi_{a_1} = 0$

# Particle Content

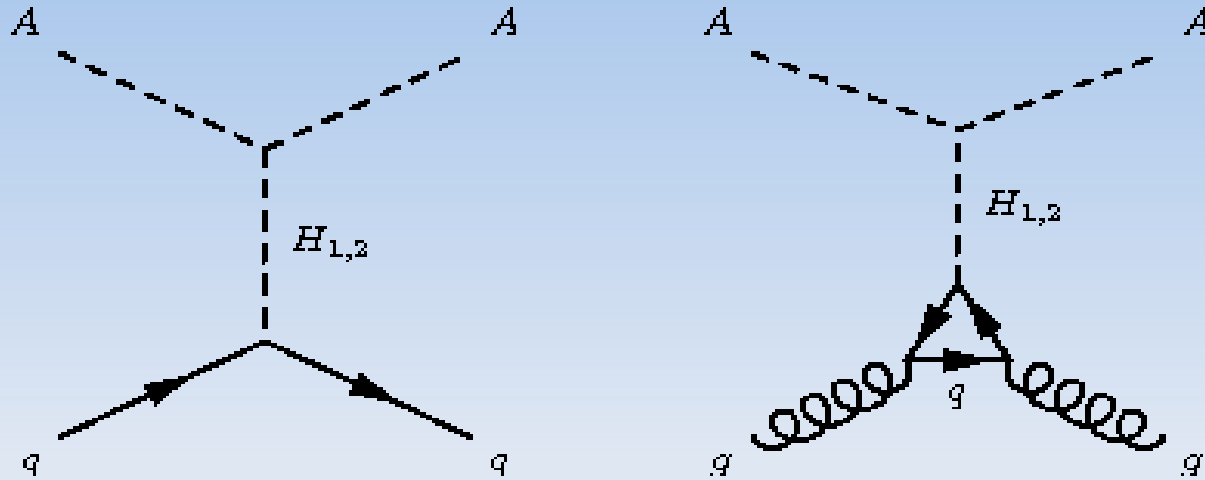
$$M_{H_{1,2}}^2 = \frac{1}{4} \left( \lambda v^2 + d_2 v_S^2 - \frac{2\sqrt{2}a_1}{v_S} \mp \sqrt{\left( \lambda v^2 - d_2 v_S^2 + \frac{2\sqrt{2}a_1}{v_S} \right)^2 + 4\delta_2^2 v^2 v_S^2} \right)$$

$$M_A^2 = b_1 - \frac{\sqrt{2}a_1}{v_S}$$

$$\tan 2\phi = \frac{2\delta_2 v v_S}{\lambda v^2 - d_2 v_S^2 + \frac{2\sqrt{2}a_1}{v_S}}$$

- $H_1$  and  $H_2$  couplings to SM particles are reduced by  $\cos(\phi)$  and  $-\sin(\phi)$  respectively.

# Spin-independent scattering cross section



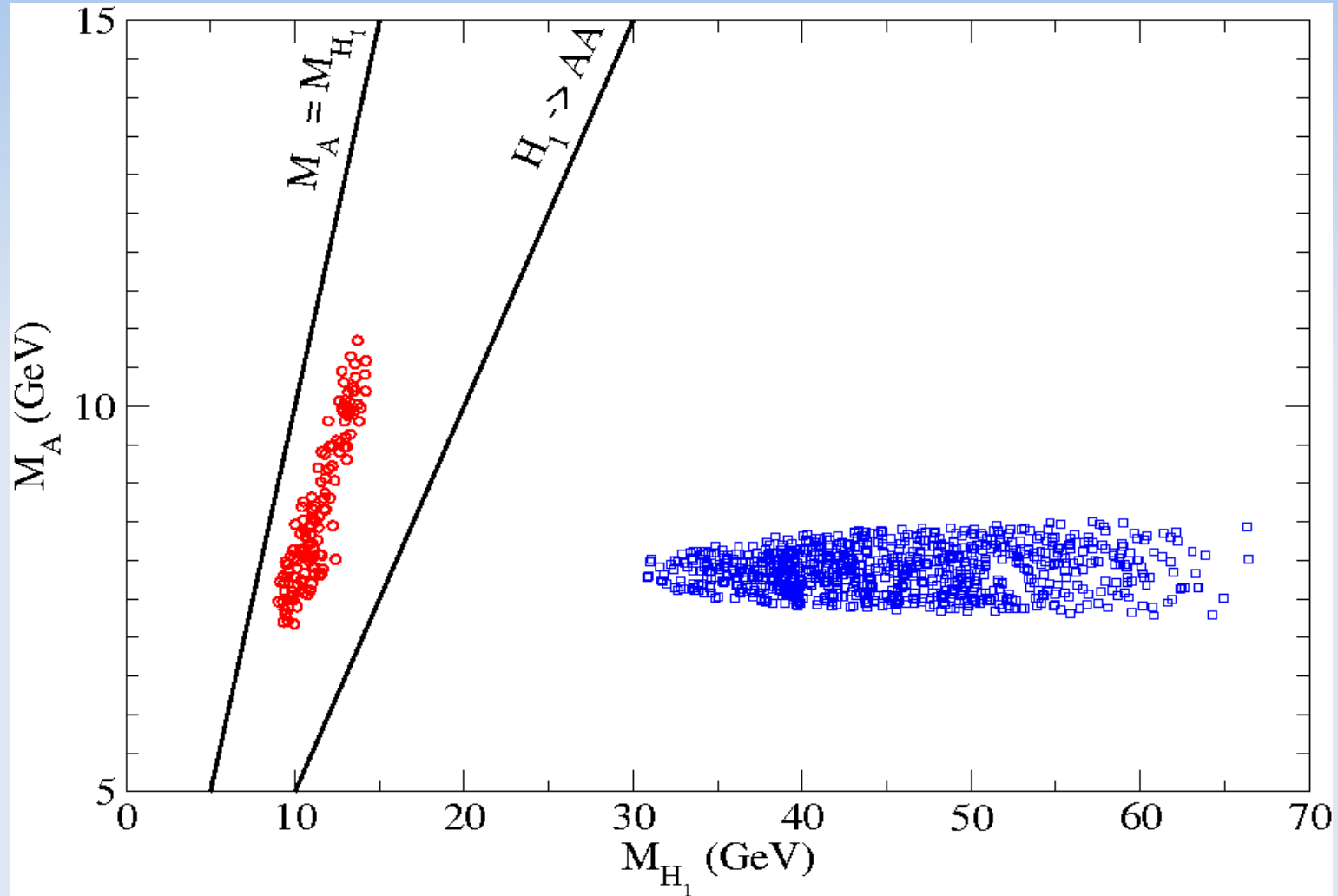
$$\sigma_{SI} = \frac{m_p^4}{2\pi v^2 (m_p + M_A)^2} \left( \frac{g_{AAH_1} \cos \phi}{M_{H_1}^2} - \frac{g_{AAH_2} \sin \phi}{M_{H_2}^2} \right)^2 \left( f_{pu} + f_{pd} + f_{ps} + \frac{2}{27} (3f_G) \right)^2$$

- Barger, Keung, Shaughnessy. arXiv:0806.1962
- Spin-independent cross sections agree with those calculated by micrOmegas.

# Constraints

- WMAP constraints on the relic density
  - $0.099 < \Omega h^2 < 0.122$
  - Calculated using micrOmegas
- LEP constraints (OPAL and DELPHI)
  - Higgs decay to SM and invisible modes
  - ZZh coupling
- Cascade decays
  - $H_2 \rightarrow 4 b, 2b 2\tau, 4 \tau$
- EWPO
  - $H_2 < 180 \text{ GeV}$

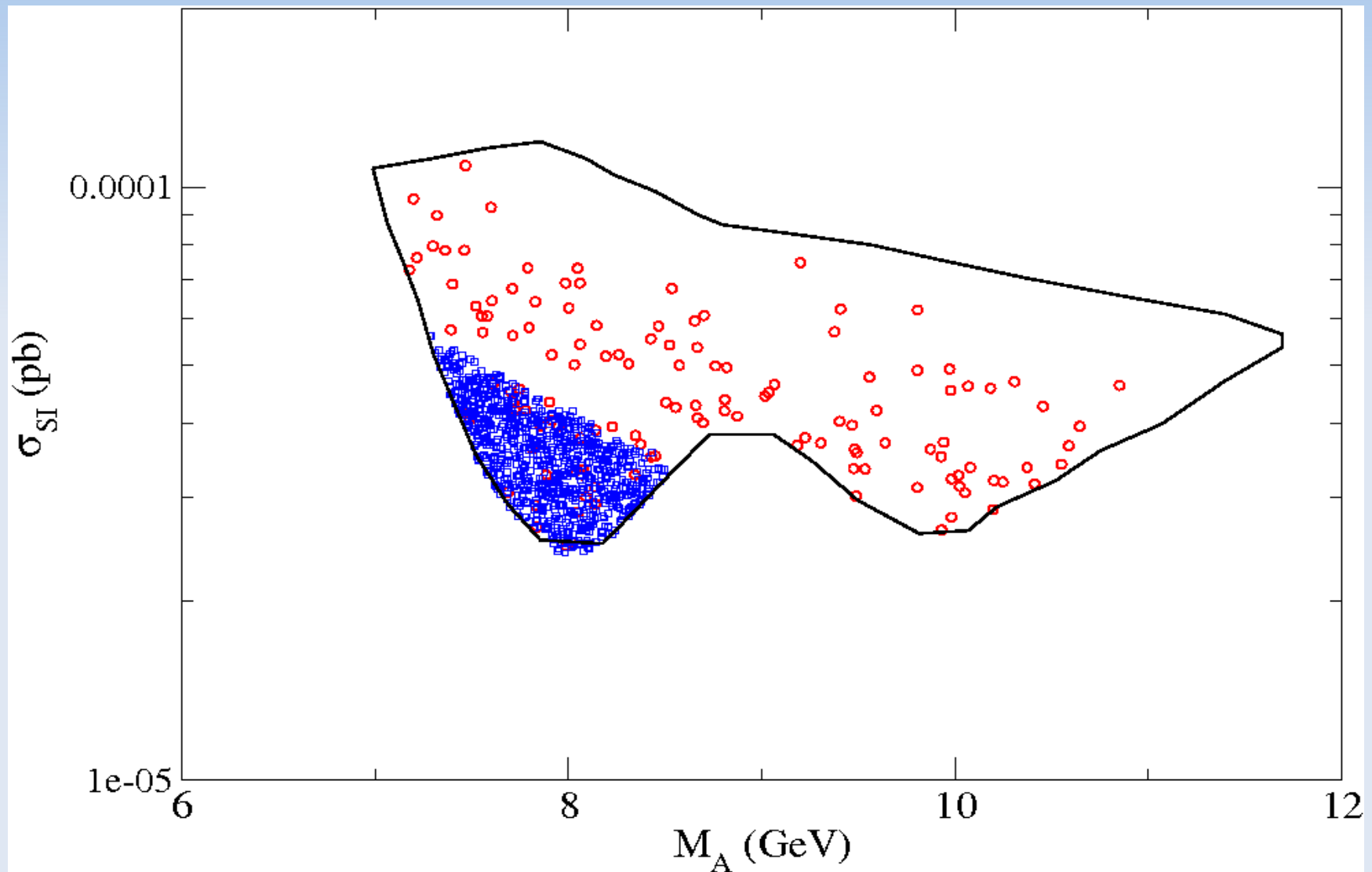
# Results: Mass spectrum



- $110 \text{ GeV} < M_{H_2} < 180 \text{ GeV}$

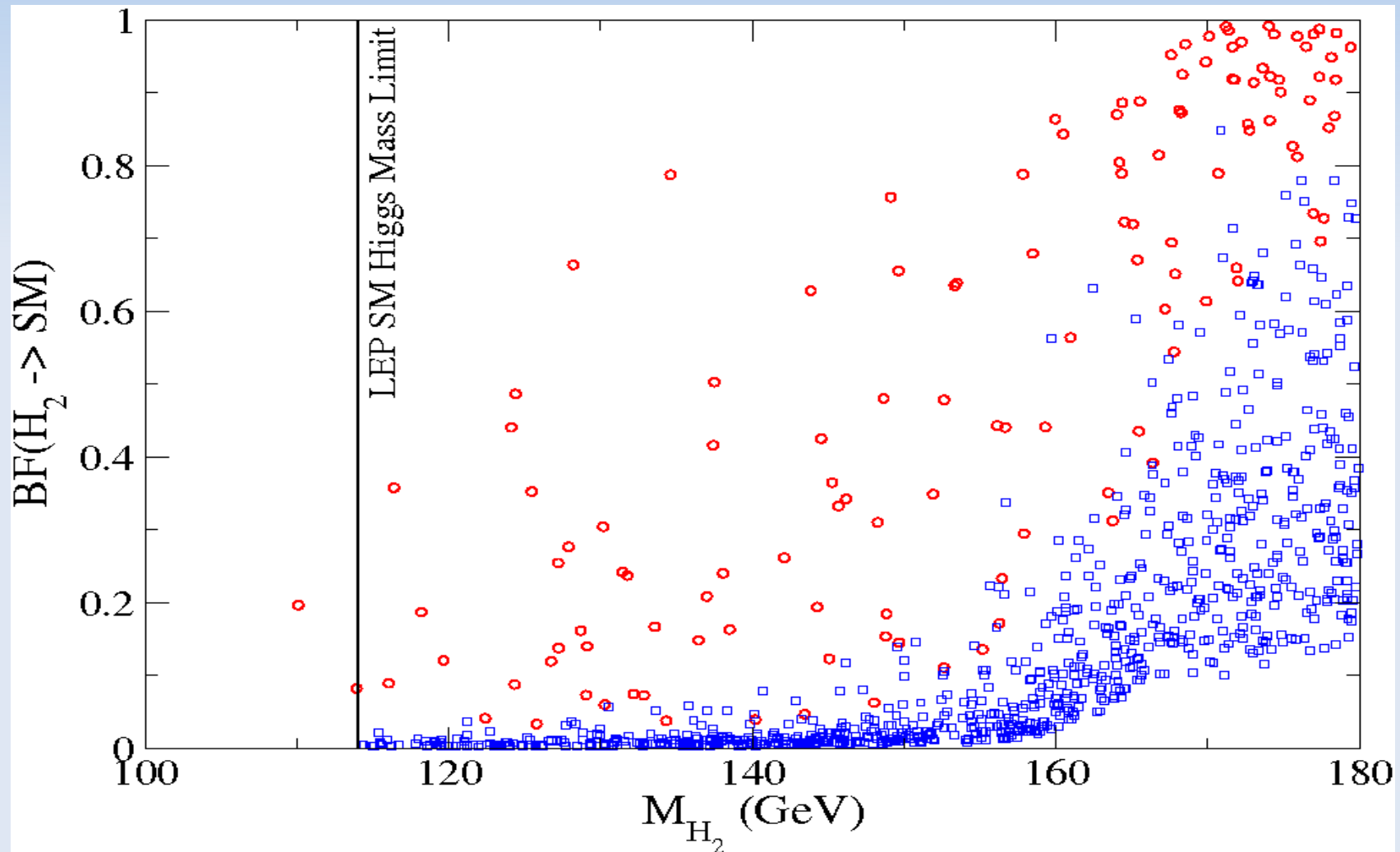


# Results: $\sigma_{SI}$

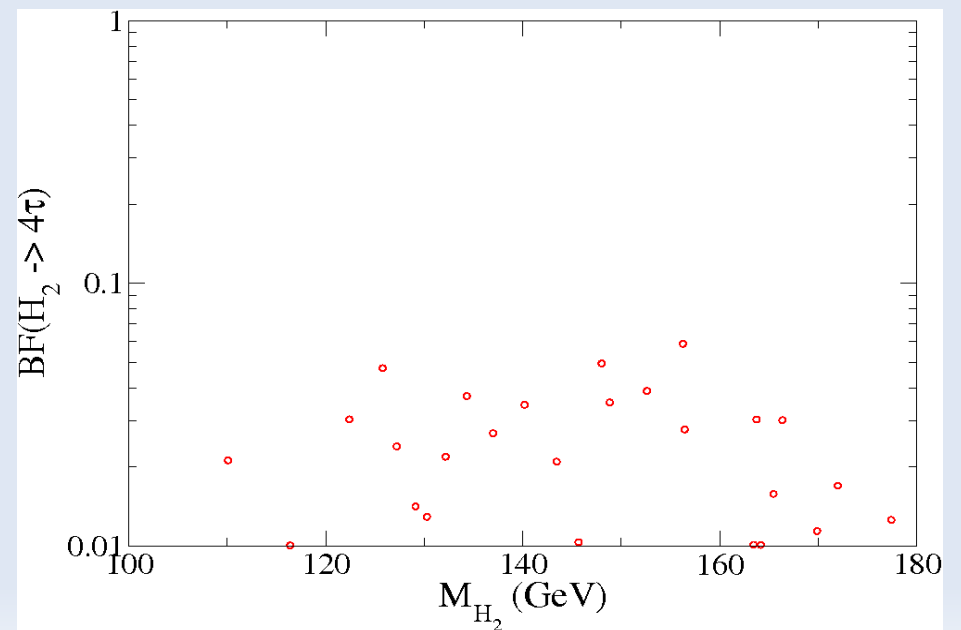
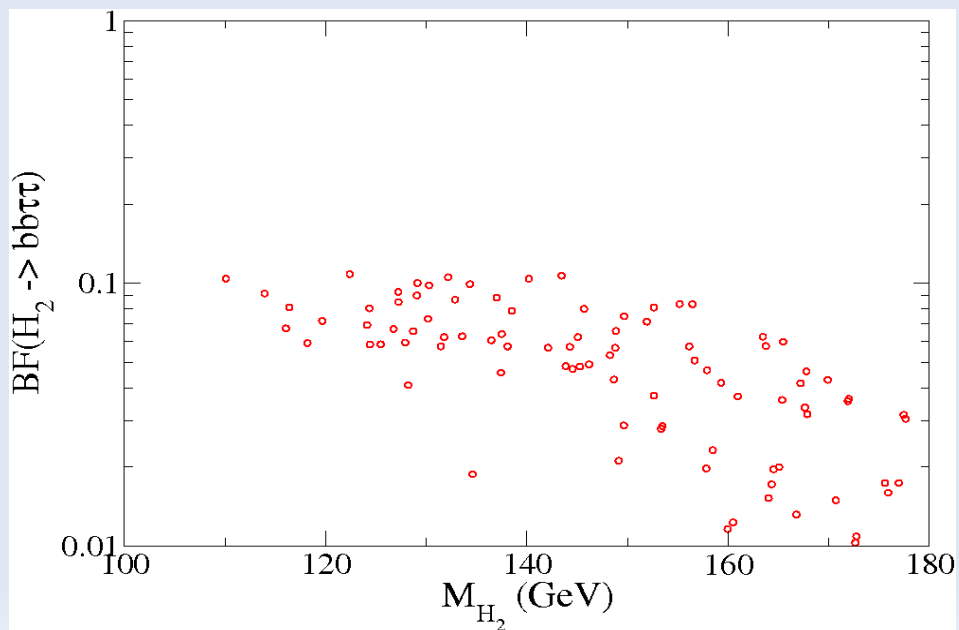
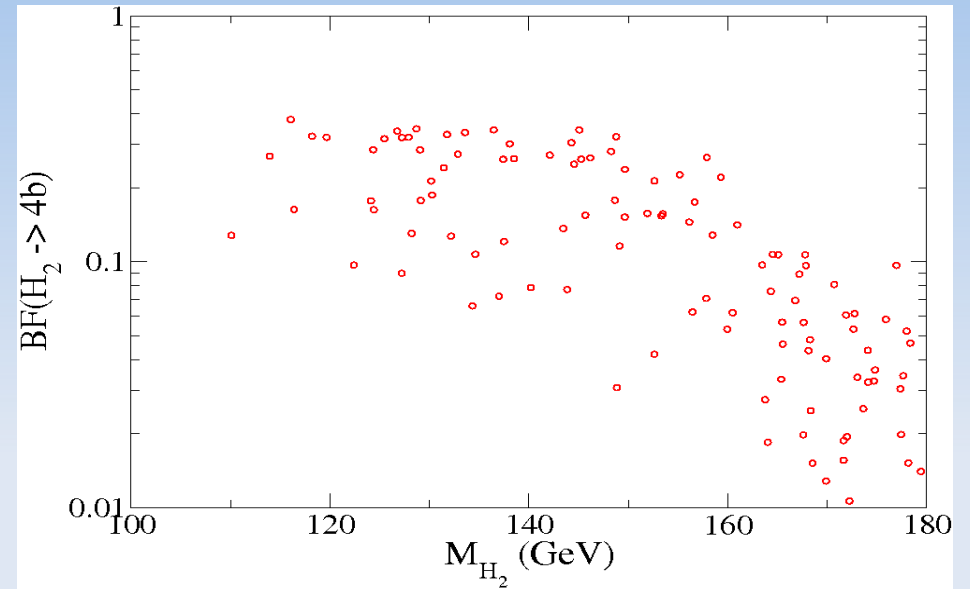
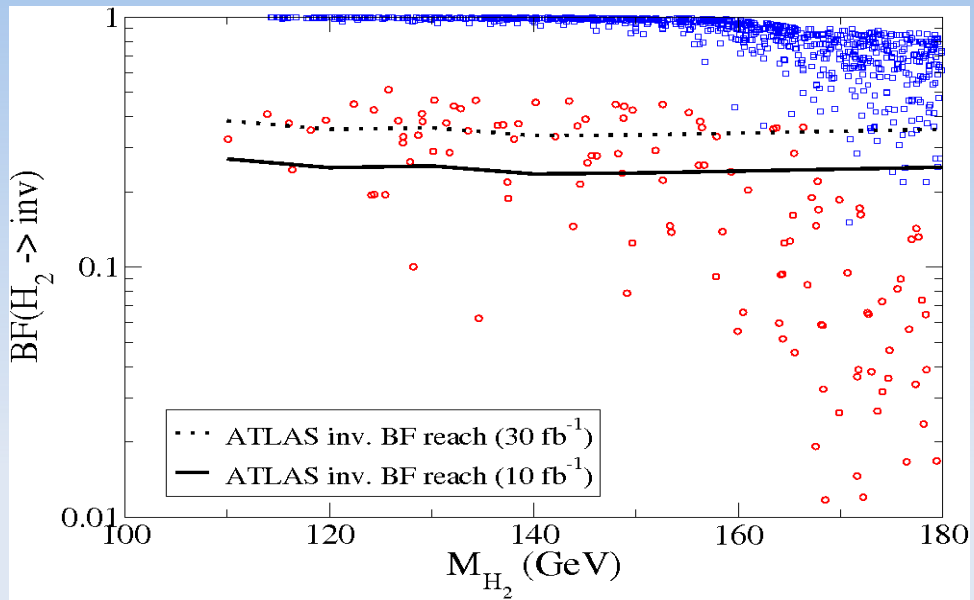


# Higgs signatures

- $H_1$  decays either to DM (blue) or SM (red)



# Cascade Decays



# The Debate: Part 1

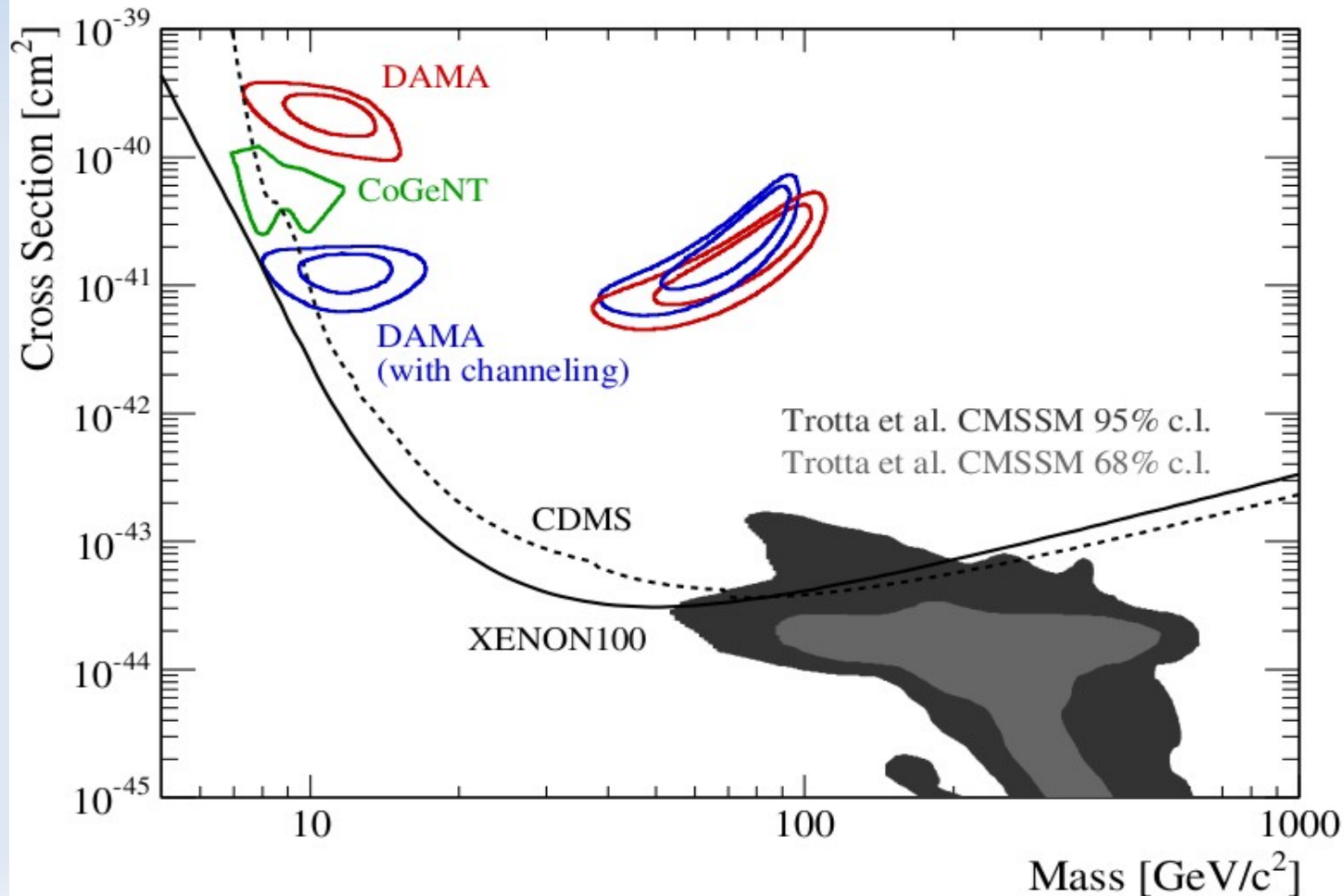
arXiv:1005.0380 [pdf, other]

## First Dark Matter Results from the XENON100 Experiment

E. Aprile, K. Arisaka, F. Arneodo, A. Askin, L. Baudis, A. Behrens, E. Brown, J. M. R. Cardoso, B. Choi, D. B. Cline, S. Fattori, A. D. Ferella, K.-L. Giboni, K. Hugenberg, A. Kish, C. W. Lam, J. Lamblin, R. F. Lang, K. E. Lim, J. A. M. Lopes, T. Marrodán Undagoitia, Y. Mei, A. J. Melgarejo Fernandez, K. Ni, U. Oberlack, S. E. A. Orrigo, E. Pantic, G. Plante, A. C. C. Ribeiro, R. Santorelli, J. M. F. dos Santos, M. Schumann, P. Shagin, A. Teymourian, D. Thers, E. Tziaferi, H. Wang, C. Weinheimer (XENON100 Collaboration)

Comments: 4 pages, 5 figures.

Subjects: **Cosmology and Extragalactic Astrophysics (astro-ph.CO)**



# The Debate: Part 2

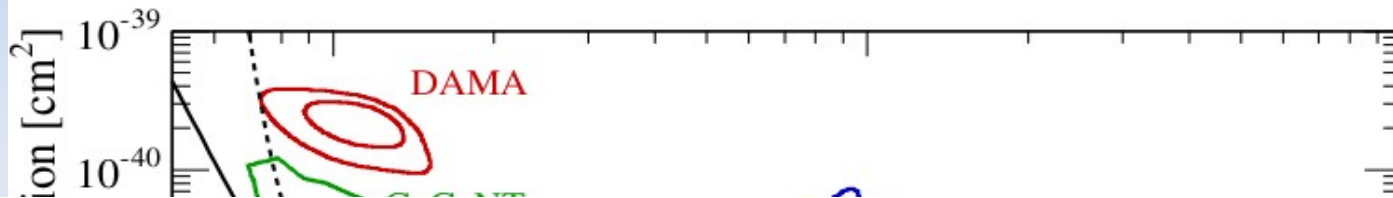
[arXiv:1005.0380](#) [pdf, other]

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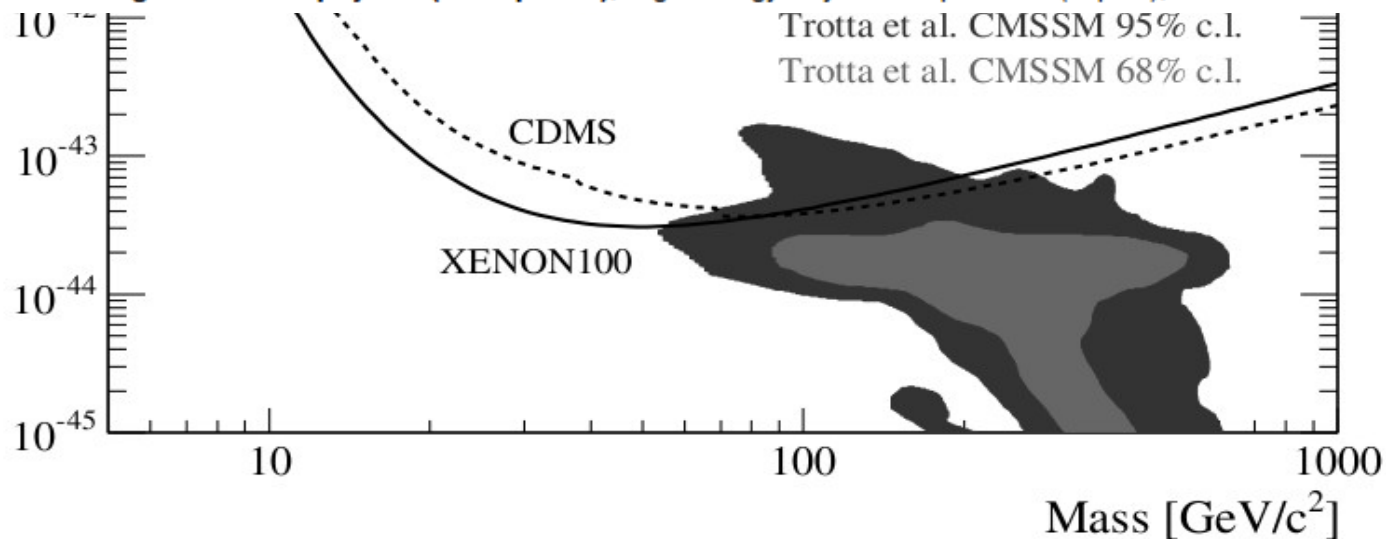
[arXiv:1005.0838](#) [pdf, ps, other]

## Comments on "First Dark Matter Results from the XENON100 Experiment"

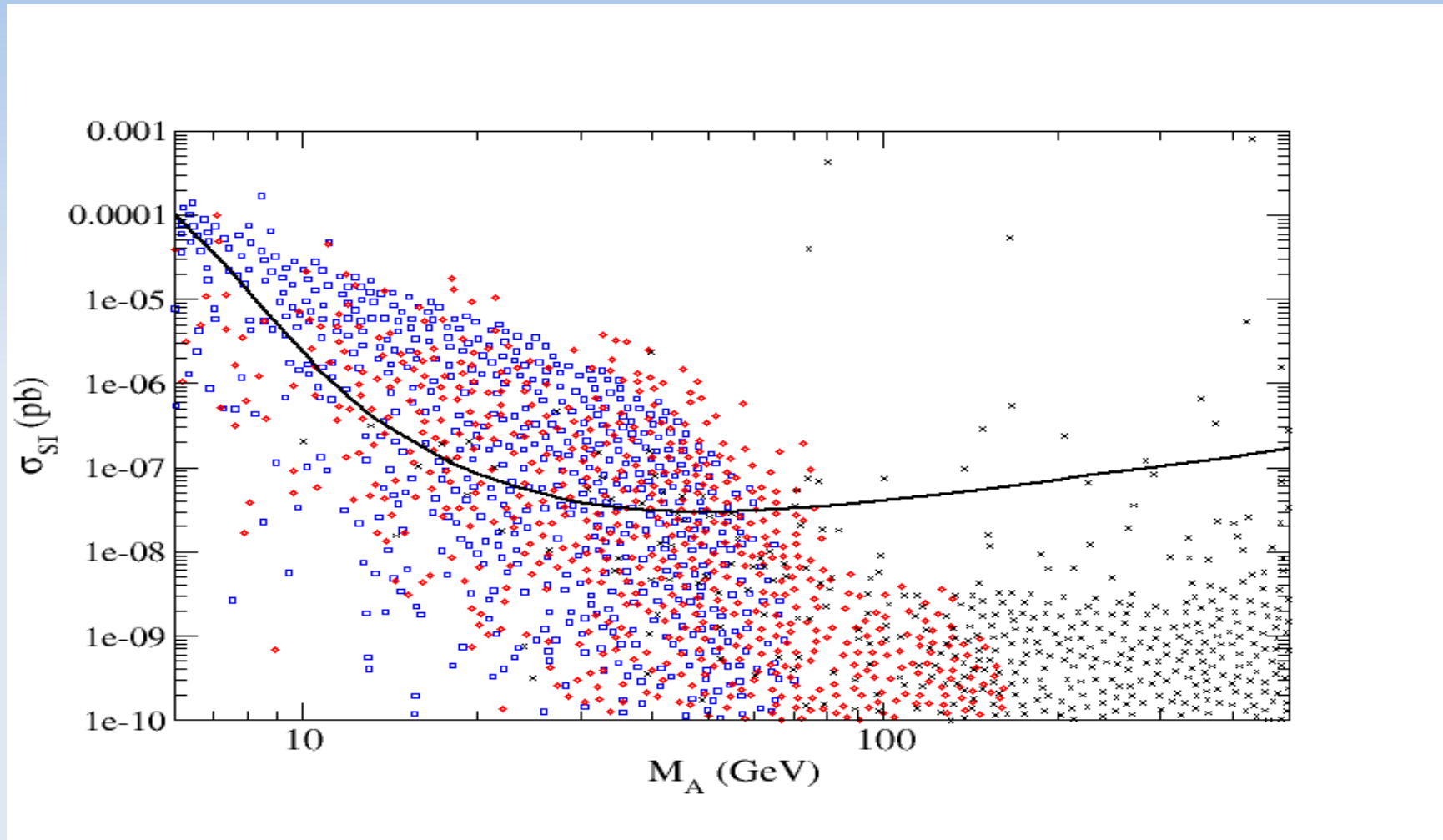
J.I. Collar, D.N. McKinsey

Comments: Important information added to a footnote. Figures enlarged, references fixed.

Subjects: **Cosmology and Extragalactic Astrophysics (astro-ph.CO)**; High Energy Physics - Experiment (hep-ex); Instrumentation and Detectors (physics.ins-det)



# Preliminary Results with Xenon100



- Need to analyze the Higgs phenomenology (work in progress)

# Conclusions

- CoGeNT
  - Spin-independent scattering cross section consistent with CoGeNT
  - Light singlet-like Higgs mostly decays to DM
  - Heavy SM-like Higgs can be probed at the LHC (both SM decays and non-SM decays)
- Xenon100
  - The csxSM can also be consistent with Xenon100
- Either way we still have some very interesting Higgs phenomenology that can potentially be seen at the LHC!

# The End

Thank you!



# References

- CoGeNT (C. E. Aalseth et al., (2010), arXiv:1002.4703)
- Hadronic matrix elements (J. R. Ellis, A. Ferstl, and K. A. Olive, Phys. Lett. B481, 304 (2000))
- micrOmegas (G. Belanger et al., (2010), arXiv:1004.1092.)
- WMAP (D. Larson et al., (2010), arXiv:1001.4635)
- LEP (A. Sopczak, (2005), arXiv:hep-ph/0502002)
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- DELPHI, J. Abdallah et al., Eur. Phys. J. C32, 475 (2004), arXiv:hep-ex/0401022
- Cascade decays: R. Dermisek and J. F. Gunion, Phys. Rev. D73, 111701 (2006)
- V. Barger, P. Langacker, and G. Shaughnessy, Phys. Rev. D75, 055013 (2007)
- M. Carena, T. Han, G.-Y. Huang, and C. E. M. Wagner, JHEP 04, 092 (2008)
- EWPO (C. Amsler et al., Physics Letters B 667, 1 (2008), Review of Particle Physics)
- ATLAS inv. BF: O. J. P. Eboli and D. Zeppenfeld, Phys. Lett. B495, 147 (2000)

Backup slides

# Parameter scan

$$7 \text{ GeV} < \sqrt{b_1} < 100 \text{ GeV},$$

$$7 \text{ GeV} < M_A < 12 \text{ GeV},$$

$$0 < \lambda < 2,$$

$$-2 < \delta_2 < 2,$$

$$0 < d_2 < 4,$$

$$10 \text{ GeV} < v_S < 1000 \text{ GeV},$$

$$\left( 14 \text{ GeV} < \sqrt{b_1} < 100 \text{ GeV} \right)$$

$$\left( 7 \text{ GeV} < M_A < 12 \text{ GeV} \right)$$

$$\left( 0.4 < \lambda < 1.1 \right)$$

$$\left( 0.01 < |\delta_2| < 1.1 \right)$$

$$\left( 0 < d_2 < 4 \right)$$

$$\left( 10 \text{ GeV} < v_S < 400 \text{ GeV} \right)$$

# Couplings

$$g_{AAH_1} = (\delta_2 v \cos \phi + d_2 v_S \sin \phi) / 2$$

$$g_{AAH_2} = (d_2 v_S \cos \phi - \delta_2 v \sin \phi) / 2$$

- Hadronic matrix elements (J. R. Ellis, A. Ferstl, and K. A. Olive, Phys. Lett. B481, 304 (2000))

$$f_{pu} = 0.02, \quad f_{pd} = 0.026, \quad f_{ps} = 0.118, \quad f_G = 0.836$$