

MODEL INDEPENDENT BOUNDS ON KINETIC MIXING

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Work in Progress with Eder Izaguirre and Jay G. Wacker
arxiv:1005:xxxx

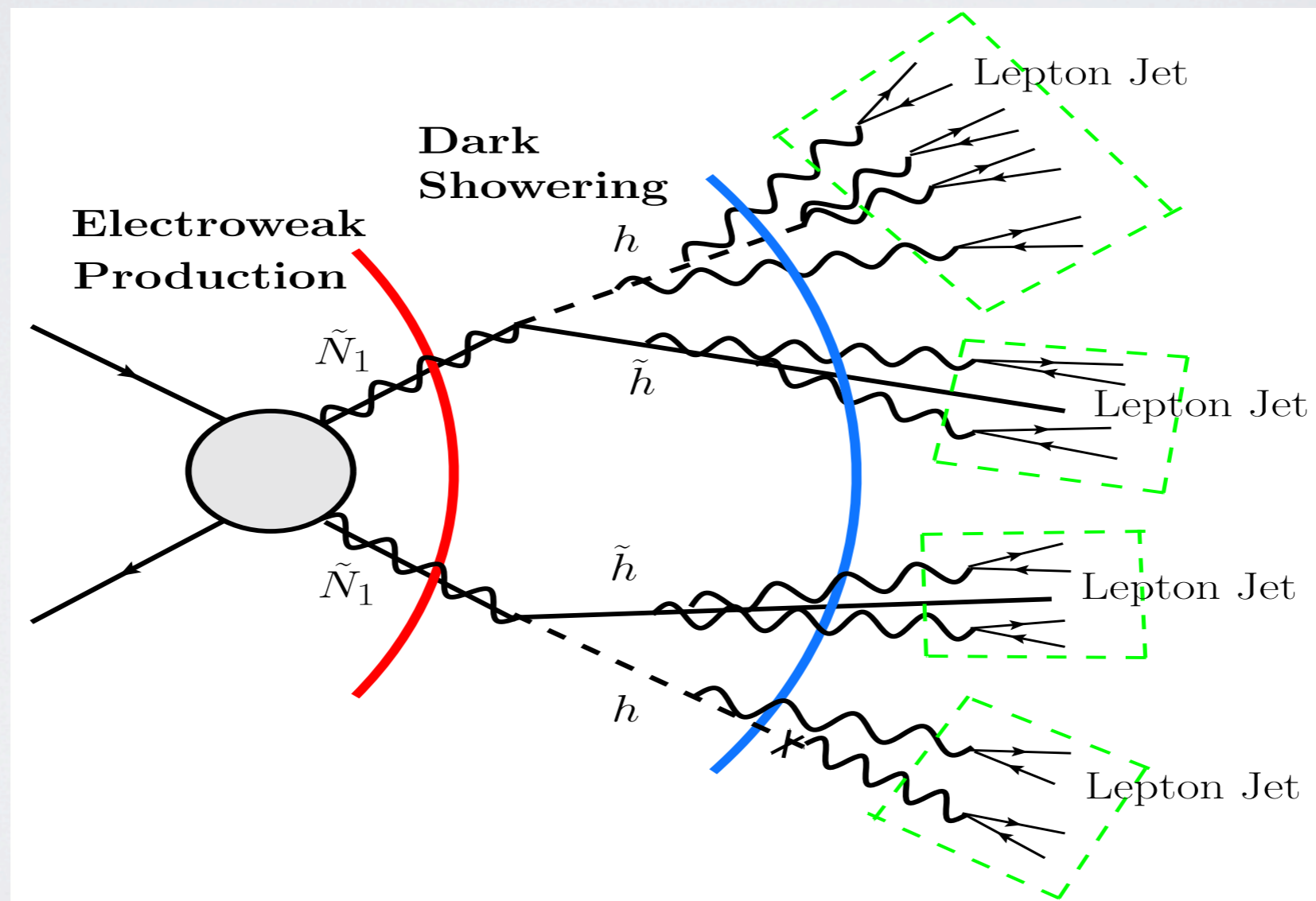
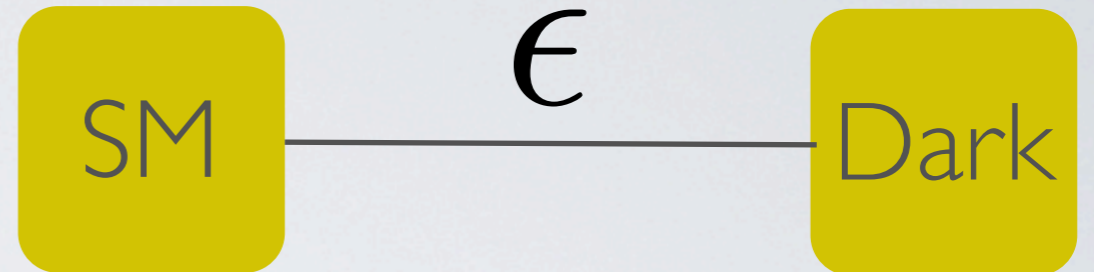
Why Kinetic Mixing?

$$\mathcal{L} \supset \frac{1}{2} \epsilon F'_{\mu\nu} F_Y^{\mu\nu}$$

- Renormalizable interaction
 - does not decouple
 - allowable by symmetries
- Numerous astrophysical and direct detection anomalies - Dark Matter?
 - PAMELA/DAMA
 - MeV \sim GeV scale dark sector

Why Kinetic Mixing?

- Gateway to SM
 - Pamela - Leptons

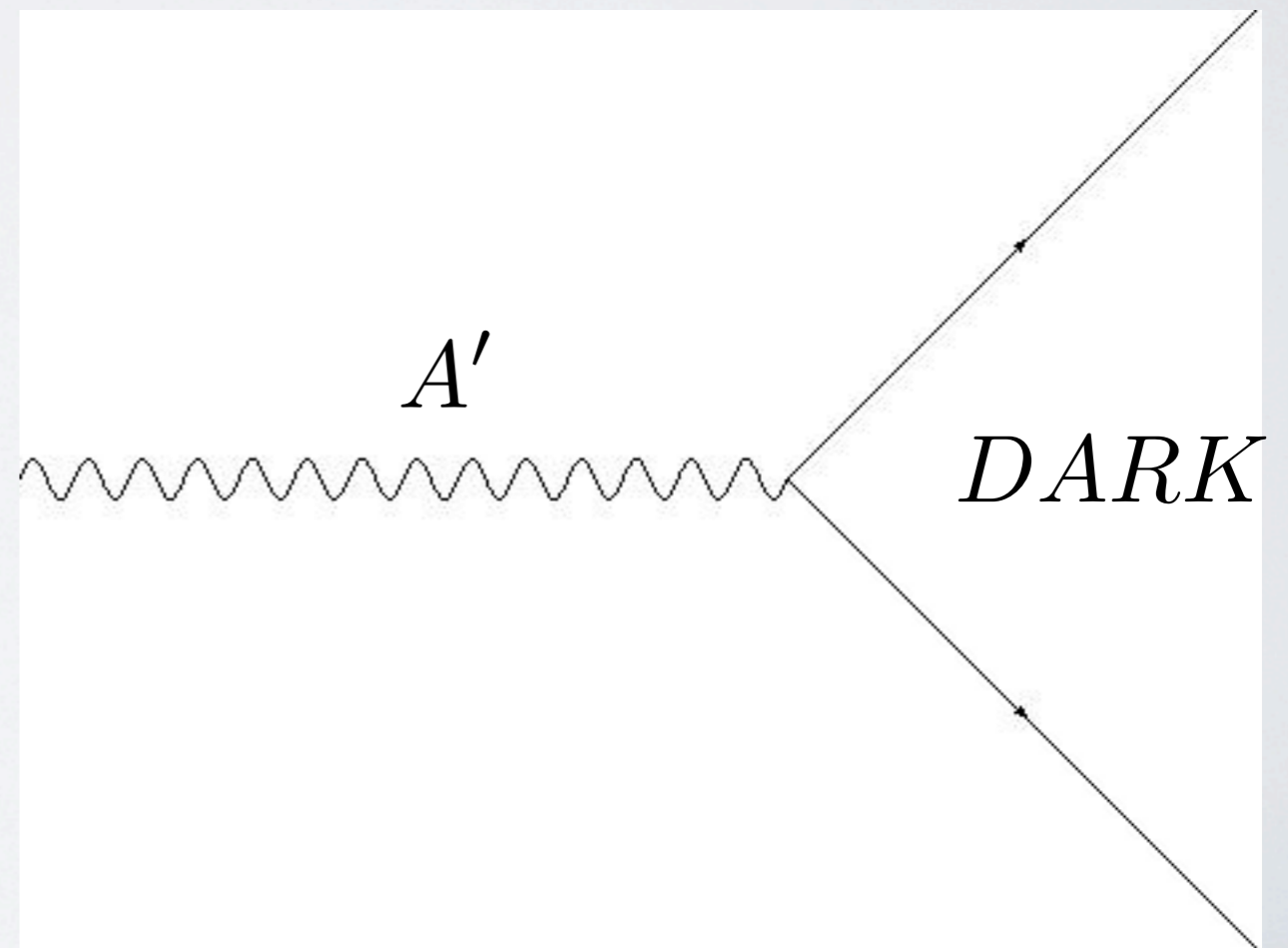


Taken from arxiv:0909.0290

Model Independent Kinetic Mixing

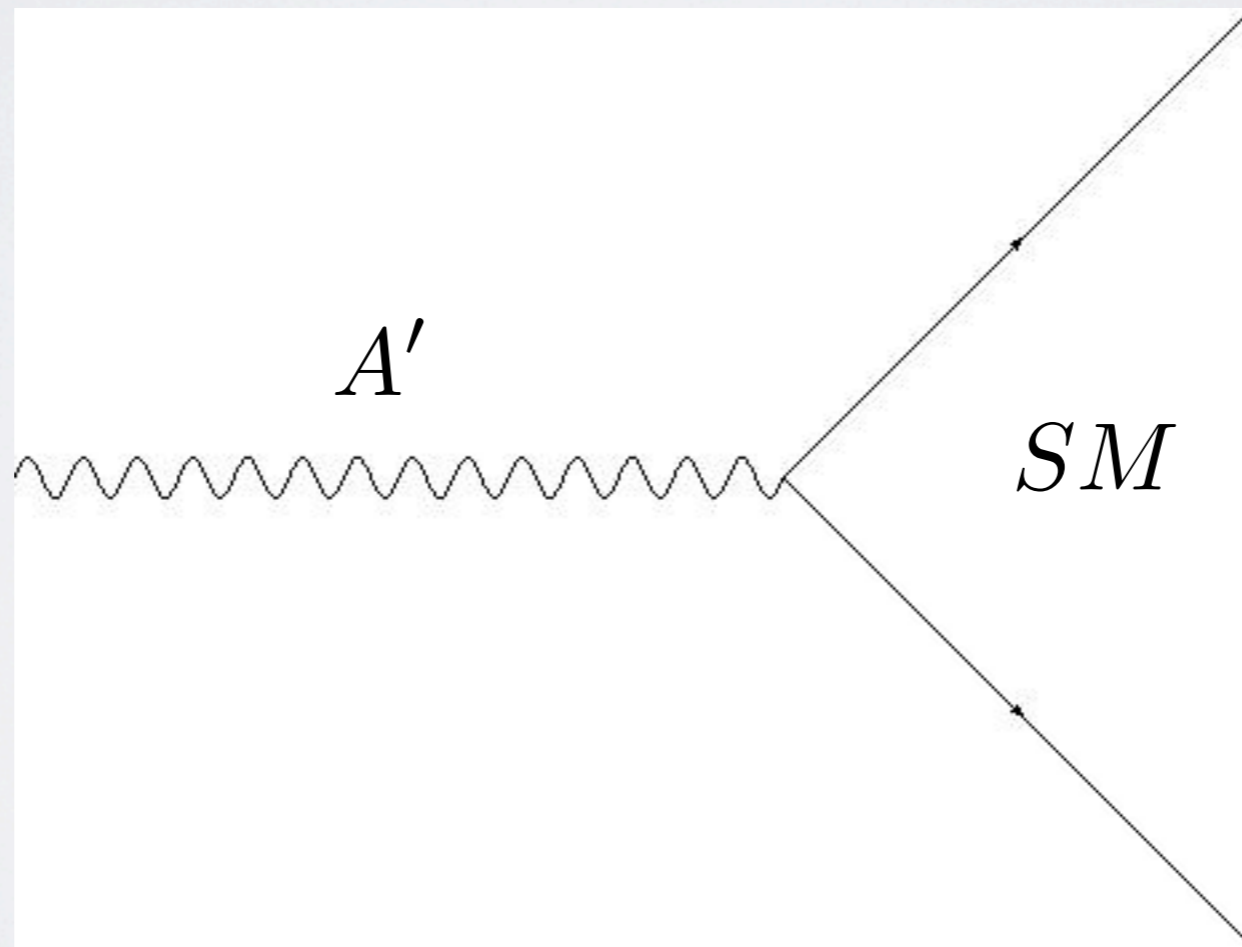
$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{2} \epsilon F'_{\mu\nu} F_Y^{\mu\nu} + \frac{1}{2} m_{A'}^2 A'_\mu{}^2$$

- No assumptions about decays
- Γ_{DARK} completely unknown
- Any couplings allowed with dark sector

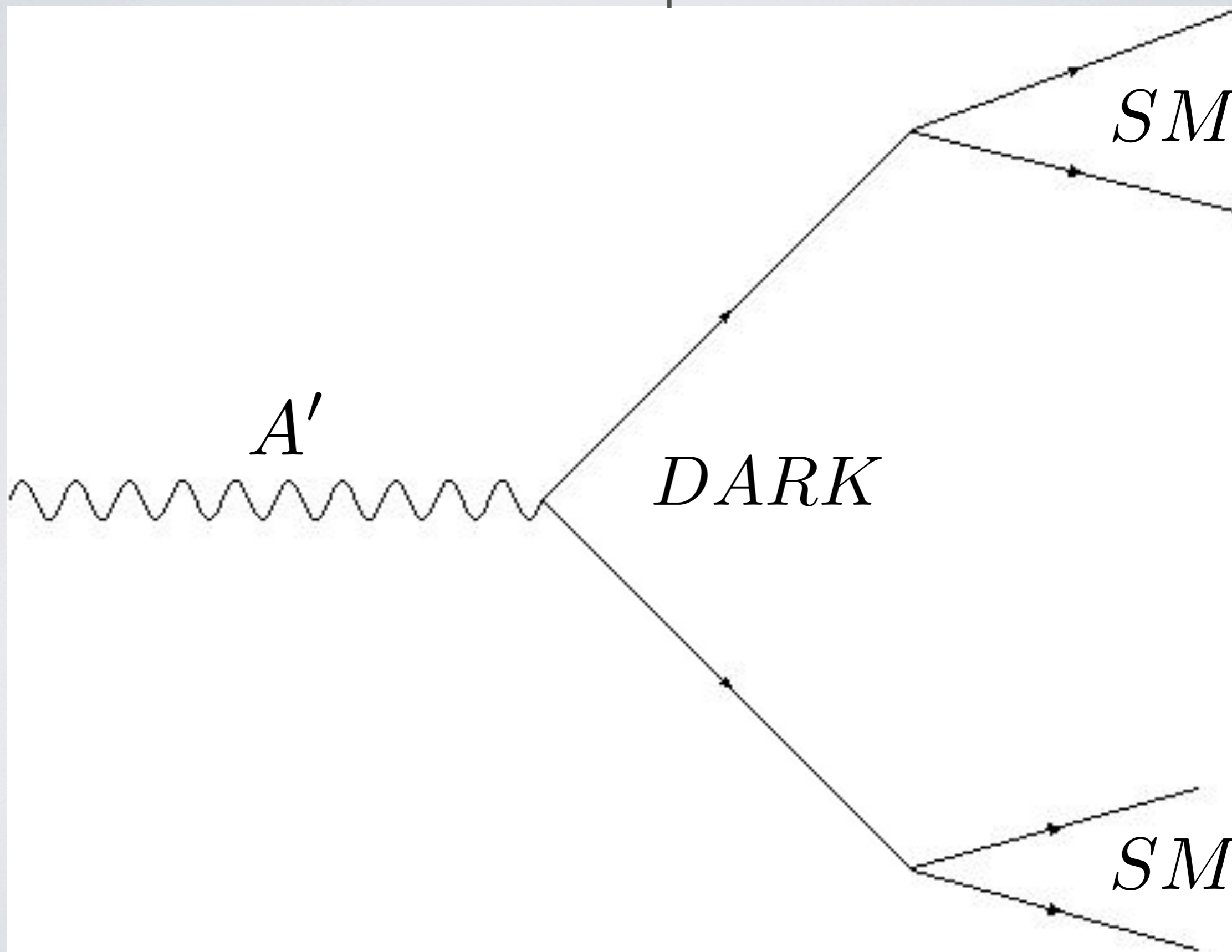


Couplings to the Standard Model

- Determined by diagonalization
 - SM epsilon charged $\Gamma_{SM} = \mathcal{O}(\epsilon^2)$

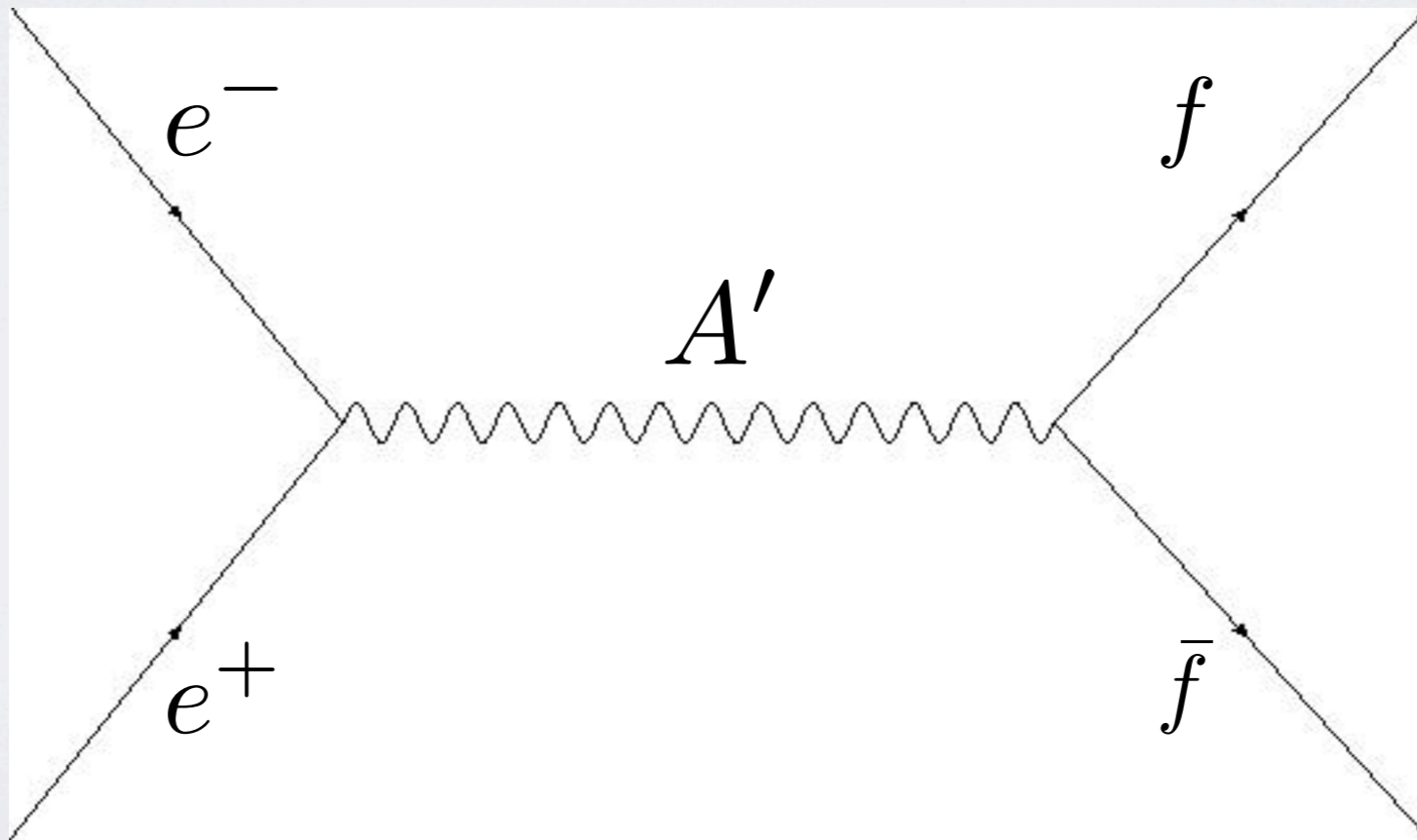


Model Dependent



On Shell Production

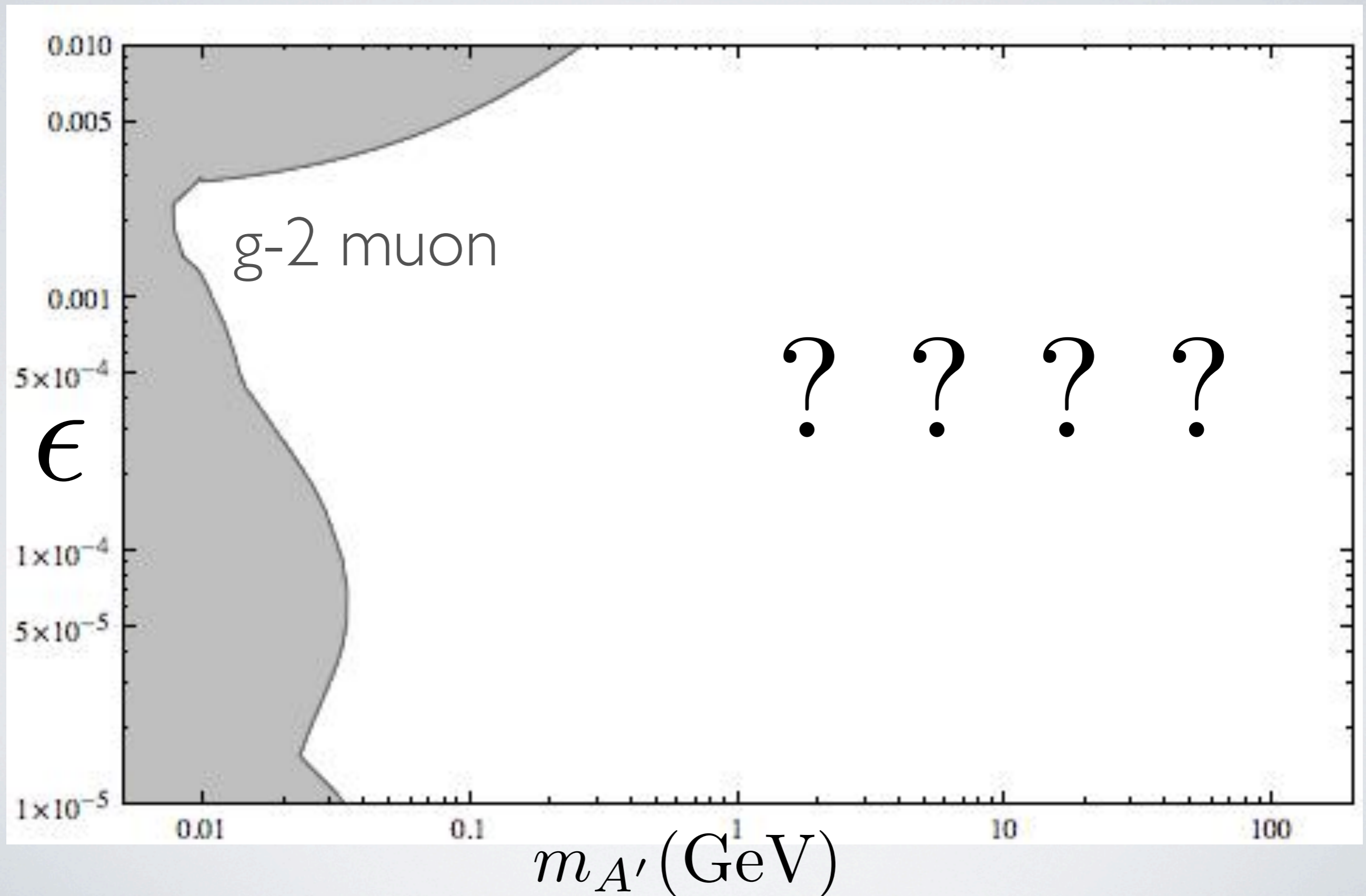
- At experimental energies: resonant s channel
- Necessary model dependence
 - Γ_{SM} small
 - $\Gamma_{DARK} = 0$ or $\Gamma_{DARK} = 1\%M_{A'}$ as reference



Current Bounds

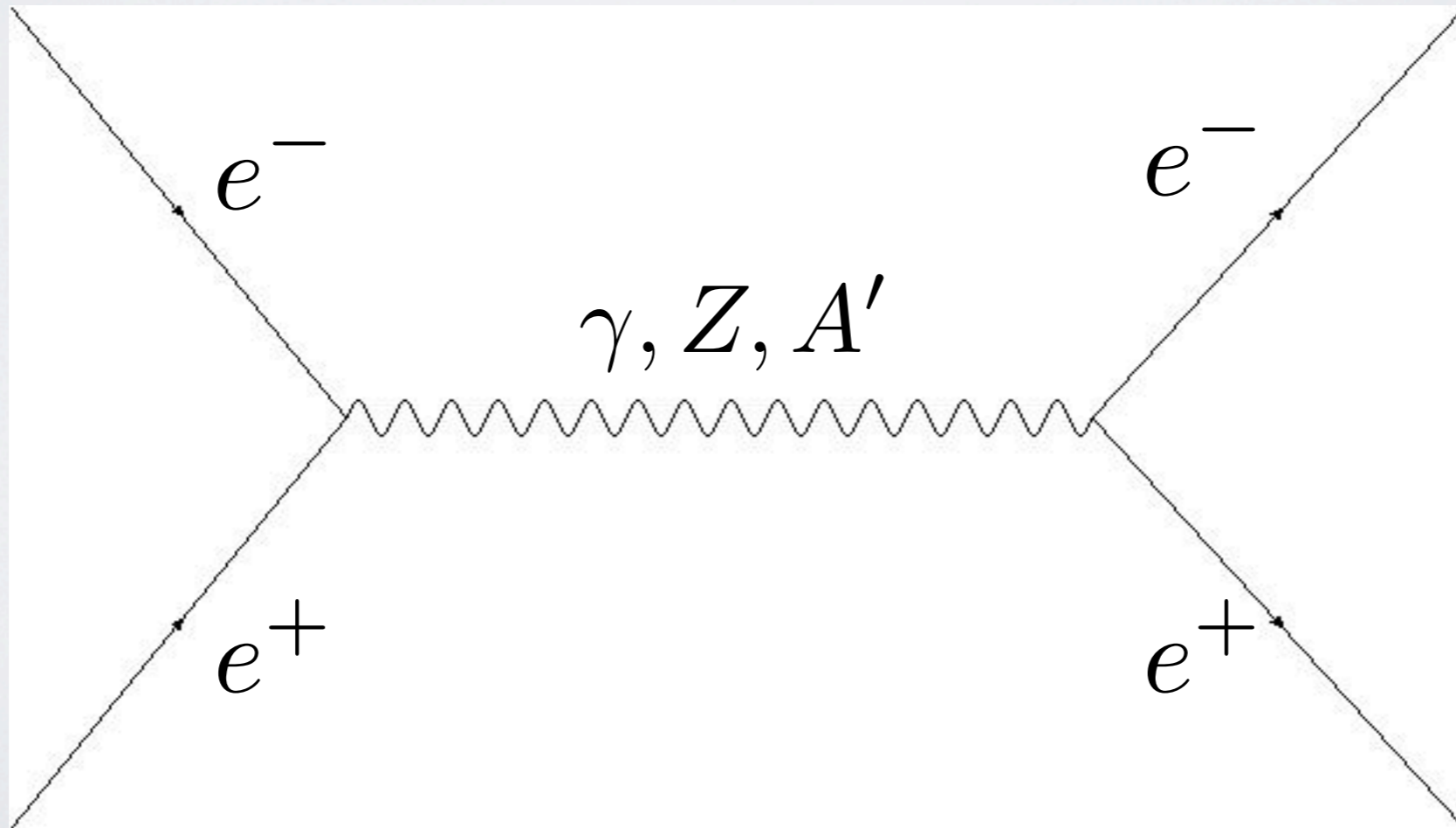
- Low Mass: $g-2$ of muon
 - good until ~ 200 MeV
- High Mass: S & T parameters
- In between: experiments determine SM parameters

Current Bounds



Model Independent Kinetic Mixing

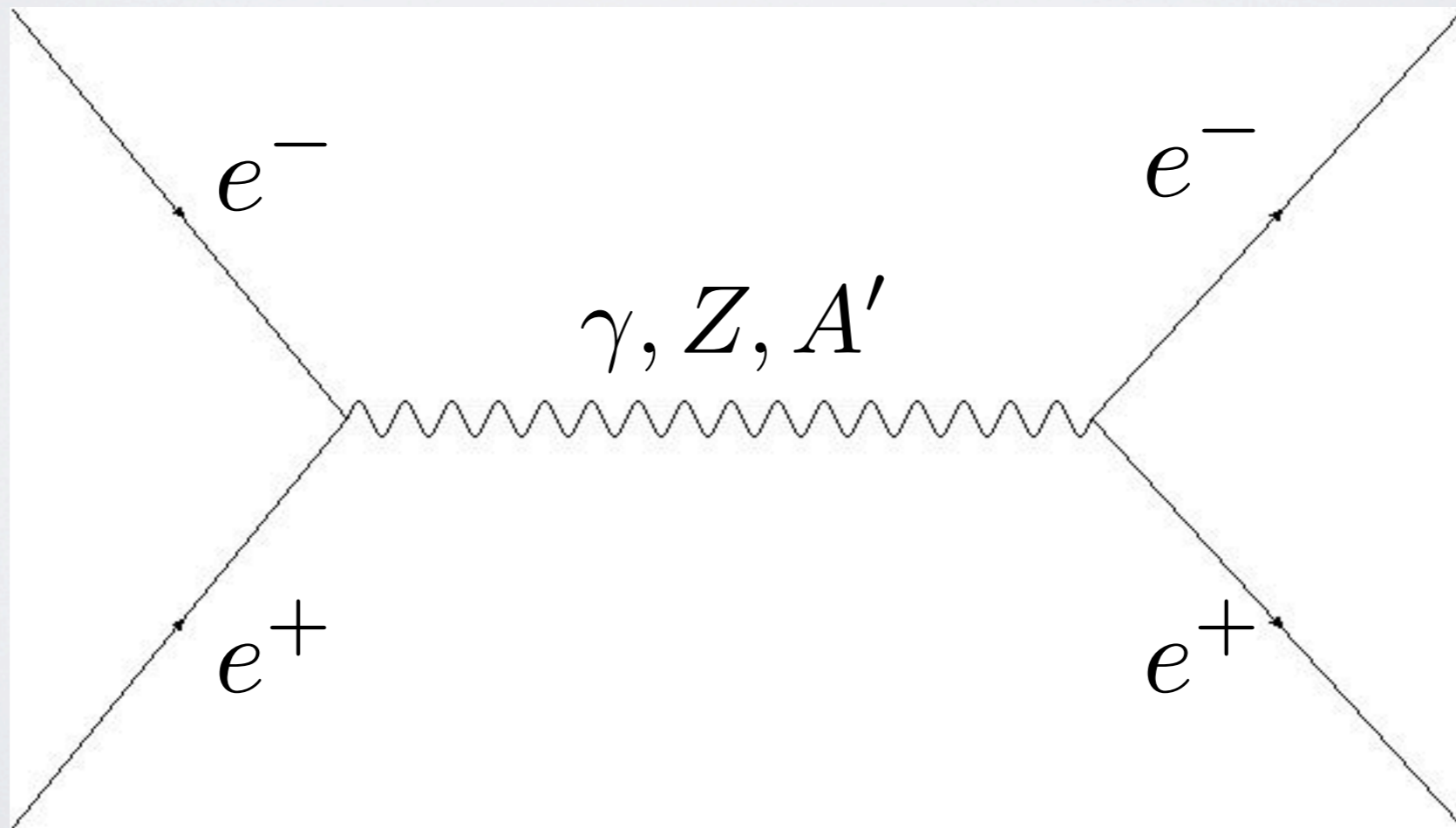
- Electroweak constraints
 - Bhabha scattering



Model Independent Kinetic Mixing

- Electroweak constraints
 - asymmetry measurements

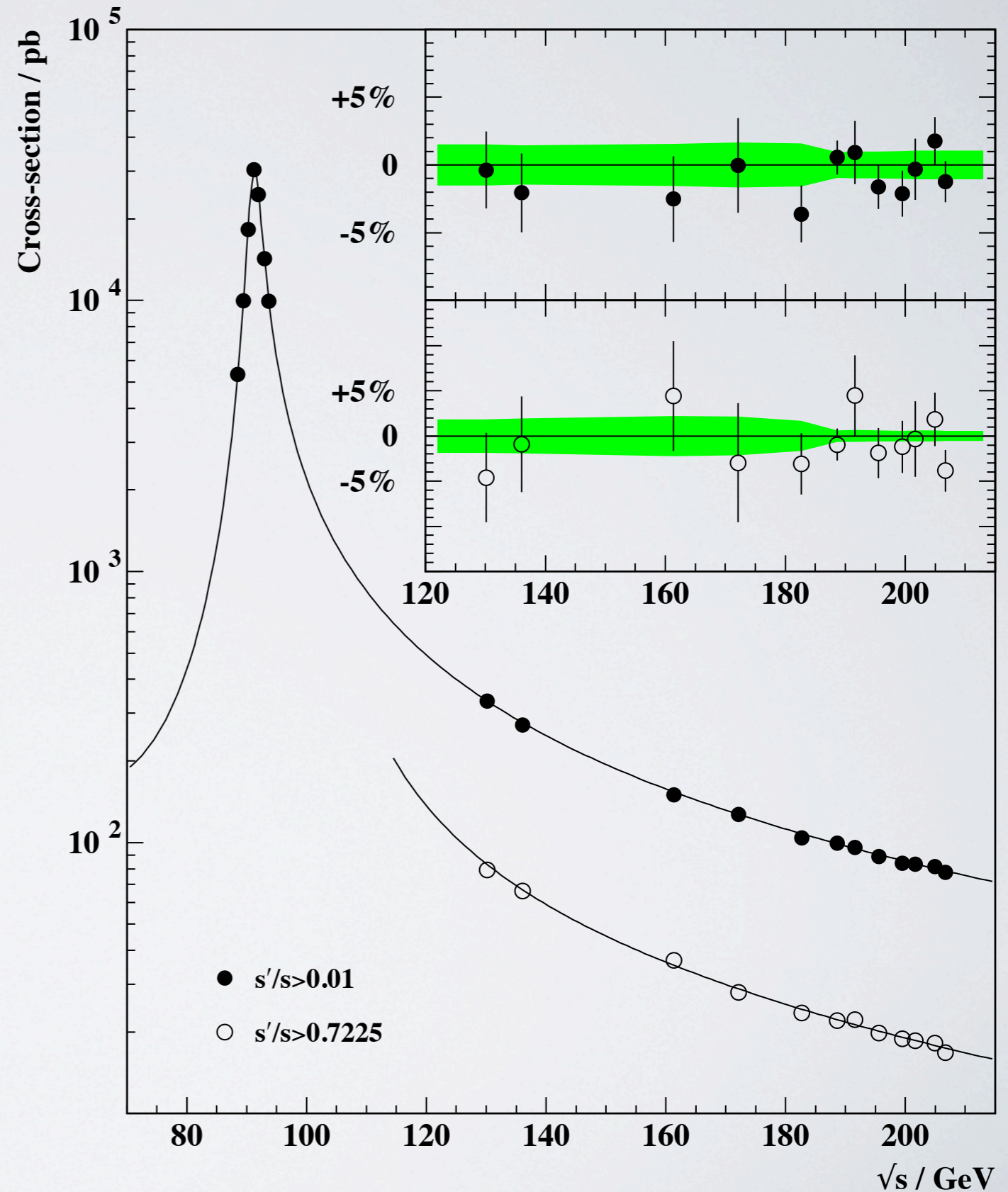
$$\frac{d\sigma}{d\Omega} \propto c_1 \cos \theta + c_2 \cos^2 \theta$$



Experiments

OPAL $e^+e^- \rightarrow \text{hadrons}$

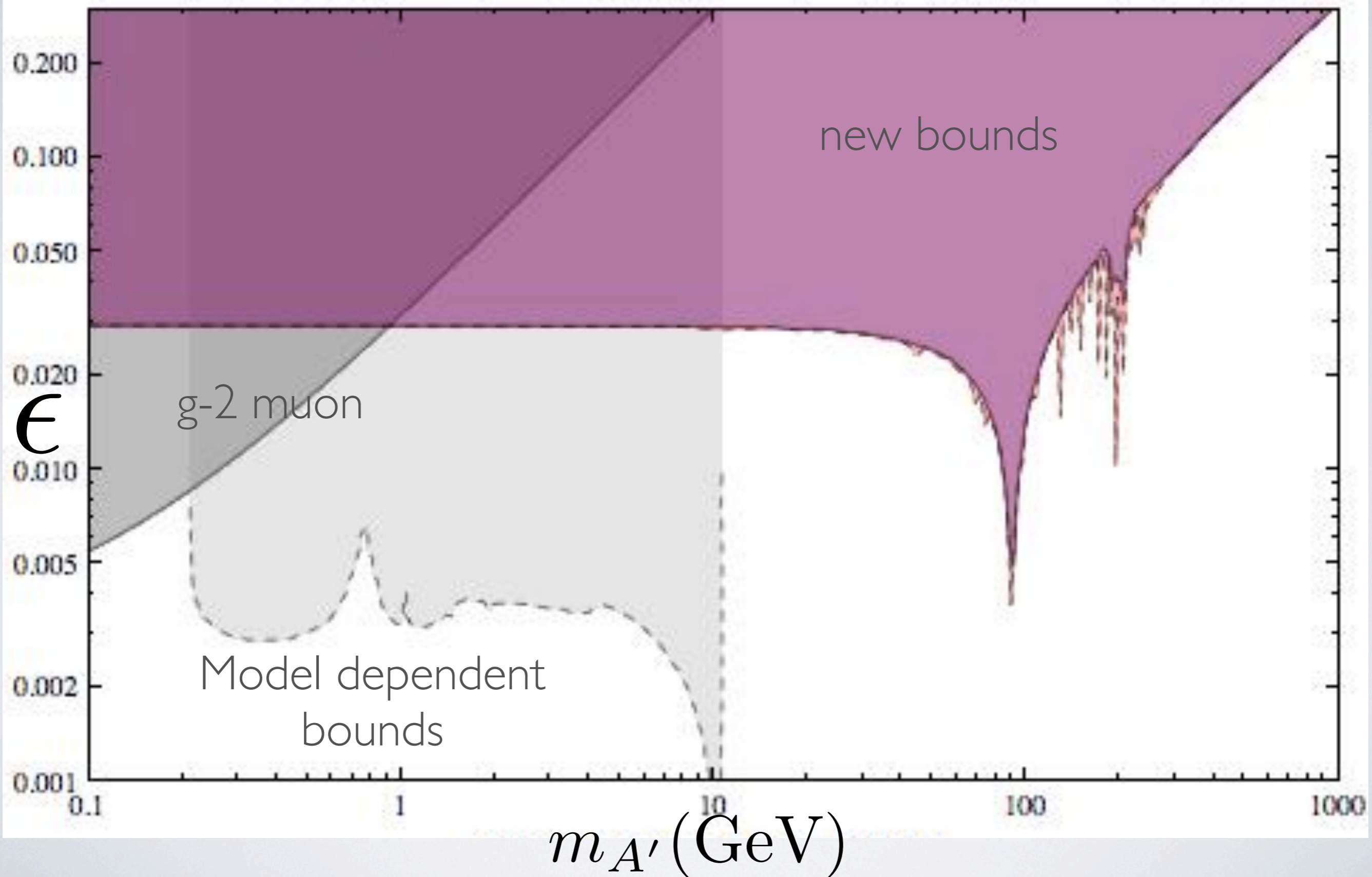
- LEP II (OPAL/DELPHI)
 - Differential bhabha scattering
 - Hadronic cross sections



Experiments

- LEP I / SLD
 - Differential bhabha scattering
 - Hadronic cross section
 - charm, bottom, muon and tau asymmetry measurements
 - Mass of the Z
- TASSO / CELLO / TRISTAN
 - Differential bhabha scattering with energies 14 - 58 GeV

Results



Conclusion

- Kinetic Mixing: well motivated SM extension
 - plethora of unexplained anomalies
- Model Independent bounds are important and are $\epsilon \sim 10^{-2}$