## General Analysis of Anti-Deuteron Dark Matter Search

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### **Outline**

- **Introduction**
- 2 Anti-D Cosmic Ray Flux
- Experimental Reach for Various Final States
- General Bounds/Features of DM Related to its Detections
- $ar{D}$  Detection Prospect for Specific Models
- Conclusions

### **Search Paths for Dark Matter**

Existence of DM ✓ – Macroscopic effects: galaxy rotation curve, gravitational lensing...

What is DM? Microscopic feature?—Little is known... Familiar search Paths:

- Direct Detection: DM scatters off target nucleus, better control/estimation of background (CDMS, XENON...)

  But rate may be highly suppressed: current bound SI elastic  $\sigma_{\chi p} \lesssim 10^{-7} \mathrm{pb}$  for  $10-100 \mathrm{GeV}$  DM, could get more stringent in coming years (XENON100/1T, Super-CDMS)
- Indirect Detection: Cosmic Ray SM particles produced from DM annihilation, s-wave annihilation
   (σ<sub>ann</sub>ν)<sub>thermal</sub> = 1pb√ (Ω<sub>DM</sub>)
   But most IdDt channels (e<sup>+</sup>, γ, p̄): large astrophysical bkg, uncertainties, hard to 'confirm' as DM origin (e.g.controversies after PAMELA, FERMI excess)

## Low Background Channel for IdDt?⇒Low energy D! (Bottino, Donato, Fornengo and Salati, 1998)

- Conventional DM: color multiplicity→significant BR(ann) to hadrons ('Conservative' about PAMELA excess).
   Advantages compared with p̄:
- Higher threshold energy for secondary astrophysical production: (pH),  $(pH_e)$  collision,  $E_{th}(\bar{p}) = 7m_p$ ,  $E_{th}(\bar{D}) = 17m_p$ , suppression from cosmic ray p number distribution  $N_p \sim E_p^{-2.7}$ .  $K_{\bar{D}} \sim 2 \text{GeV}$
- Suppressed tertiary production of low E D̄: 'slow-down' during inelastic scattering off galactic nucleus: p̄√,
   Not for D̄! 'Fragility': E<sub>binding</sub>(D̄) = 2.2MeV⇒ Breaking apart instead of losing energy

High sensitivity experiments coming soon!
-AMS-02 (2010), GAPS (LDB2011, ULDB2014, SAT)

#### **Our Goal**

Most existing anti-D related DM study: signal for particular DM models, e.g. SUSY  $\tilde{\chi}_0$  (Donato, Fornengo, Salati, 1999; Baer and Profumo 2005, etc.)

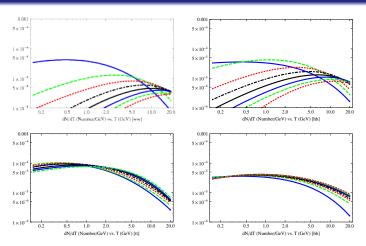
**Our goal:** Take a broader view— +general analysis for general DM candidates

- Anti-D flux from various SM final states, mass reach at AMS-02, GAPS
- Generic scalar, fermion, vector DM models: correlation between thermal relic density, DiDt and IdDt, operator analysis

### **Injection Spectrum**

- $\bar{D}$  injection spectrum:  $m_{DM}$ , final states composition ( $\bar{t}t$ ,  $\bar{b}b$ ,  $h^0h^0$ , gg,  $W^+W^-$ )

  -hadronization simulated by PYTHIA6.4
- Formation of  $\bar{D}$  from  $\bar{p} \bar{n}$  (coalescence model): in  $\bar{n}$  rest frame,  $K_{\bar{p}} < B$ , or  $|\vec{k}_{\bar{n}} \vec{k}_{\bar{p}}| < (2m_pB)^{\frac{1}{2}} \sim p_0 \sim 70 \text{ MeV} \Rightarrow \bar{D}!$  more accurately,  $p_0$  by fitting ALEPH Z decay data:  $p_0 = 160 \text{MeV}$
- Different Spectral features for different final states—colored  $(\bar{b}b, \bar{t}t)$ : hadronize in rest frame, peak at low K even at large  $m_{DM}$ —favored by  $\bar{D}$  search; color-neutral  $(h^0h^0, W^+W^-)$ : hadronize in boosted frame, peak at higher K esp. at high  $m_{DM}$



**Figure:** The anti-D injection spectrum as a function of Kinetic Energy, T, for  $W^+W^-$ , hh(115 GeV),  $\bar{t}t$ ,  $b\bar{b}$  final states.  $m_{DM}=100 \text{ GeV(blue/solid)}$ , 200 GeV(green/dashed), 300 GeV(red/dottd), 400 GeV(blue/solid), 600 GeV(blue/solid), 700 GeV(green/dashed), 800 GeV(hlue/solid)

## Anti-D Flux: Propagation from galactic halo to us

• 2D diffusion model. The diffusion equation for charged cosmic rays (Uncertainty in model parameters: MIN, MED, MAX):

$$\frac{d}{dt}\psi(r,z,E) = Q(r,z,E) - 2h\delta(z)\Gamma_{ann}(E)(n_H + 4^{\frac{2}{3}}n_{He})\psi(r,z,E) 
+ K(E)\left(\frac{\partial^2}{\partial z^2} + \frac{1}{r}\frac{\partial}{\partial r}r\frac{\partial}{\partial r}\right)\psi(r,z,E) - V_C\frac{\partial}{\partial z}\psi(r,z,E)$$

primary source Q obtained from DM  $\bar{D}$  injection spectrum  $(\frac{dN}{dT})$ 

$$Q(r, z, T) = \frac{1}{2} \langle \sigma v \rangle \left( \frac{\rho(r, z)}{m_{DM}} \right)^{2} \frac{dN}{dT}.$$

$$\rho_{Ein}(r) = \rho_{\odot} \exp \left[ -2 \left[ \left( \frac{r}{r_{c}} \right)^{\alpha} - \left( \frac{r_{\odot}}{r_{c}} \right)^{\alpha} \right] / \alpha \right]$$

Solar Modulation:

$$\Phi_{\bigoplus}(T_{\bigoplus}) = \frac{2mT_{\bigoplus} + T_{\bigoplus}^2}{2mT_{\bot} + T_{\bigoplus}^2} \Phi(T), \quad T = T_{\bigoplus} + e\phi_F.$$

# Experimental Reach for Certain Final States ( $BR = 1, \langle \sigma v \rangle = 1 \text{pb}$ )

Mass reach: the largest DM mass (GeV) for which the anti-D flux yields  $N_{crit}$ -number for  $2\sigma$  or  $5\sigma$  signal at certain experiment.

Experiment	σ̄q	tt	$h^0 h^0$	W <sup>+</sup> W <sup>-</sup>	N <sub>crit</sub>
AMS-02 high $(2\sigma)$	110	< <i>m</i> <sub>t</sub>	< <i>m</i> <sub>h</sub>	$< m_W$	1
AMS-02 low $(2\sigma)$	150	220	150	140	1
GAPS (LDB) $(2\sigma)$	150	220	150	120	1
GAPS (ULDB) $(2\sigma)$	360	560	300	200	1
GAPS (SAT) $(2\sigma)$	700	1000	550	270	4
AMS-02 high $(5\sigma)$	50	< <i>m</i> <sub>t</sub>	< <i>m</i> <sub>h</sub>	$< m_W$	6
AMS-02 low $(5\sigma)$	70	$< m_t$	$< m_h$	$< m_W$	4
GAPS (LDB) $(5\sigma)$	75	$< m_t$	$< m_h$	$< m_W$	3
GAPS (ULDB) $(5\sigma)$	150	220	150	120	5
GAPS (SAT) (5 $\sigma$ )	360	550	300	200	14

## General Bounds/Features of DM related to its detections

• Features of general DM: spin (0, 1/2, 1), interaction with SM (operator), mass $\Rightarrow$ 

$$\begin{array}{l} \Omega_{DM} \rightarrow \langle \sigma | v | \rangle_{therm} = 1 \mathrm{pb}, \, \langle \sigma | v | \rangle_{ann} \, (IdDt), \\ \sigma_{SI} \lesssim 10^{-7} \mathrm{pb} \, (XENON, \, CDMS \, bound), \, \, \sigma_{SD} \, (DiDt) \\ \Rightarrow \frac{\langle \sigma | v | \rangle_{therm}}{\sigma_{SI}} \geq 10^{7} \end{array}$$

Correlation between ⟨σ|v|⟩<sub>therm</sub> and σ<sub>SI</sub> via crossing symmetry of Feynman diagram⇒Tension
 E.g. DM χ interacts with quarks, leptons, W/Z with 'unbiased' universal couplings, mediator couplings to DM and SM state g<sub>1</sub>, g<sub>2</sub>. To relate to both ⟨σ|v|⟩<sub>therm</sub> and DiDt, focus on e.g. u quark. Effective Fermi coupling for the related operator χ<sup>†</sup>χq̄q

$$G = \frac{g_1 g_2}{[(4m_Y^2 - M^2)^2 + \Gamma_M^2 M^2]^{1/2}}$$

BR(u) for annihilation  $\sim 10\% \Rightarrow$ 

$$\langle \sigma | v | \rangle_{therm}^{u} = \frac{3(g_1 g_2)^2}{4\pi [(4m_{\chi}^2 - M^2)^2 + \Gamma_M^2 M^2]} = 10^{-37} \text{cm}^2.$$

Crossing the Feynman diagram  $\Rightarrow$  associated process/rate for DiDt(SI)

$$\sigma_{\chi p} = \frac{1}{4\pi} \frac{m_p^2}{(m_\chi + m_p)^2} \frac{(g_1 g_2)^2}{M^4} \left( \sum_{q=u,d,s} \frac{m_p}{m_q} f_{Tq}^p + \sum_{q=c,b,t} \frac{m_p}{m_q} \frac{2}{27} f_{TG}^p \right)^2$$

$$\approx \frac{1}{\pi} \frac{m_p^2}{m_\chi^2} \frac{(g_1 g_2)^2}{M^4} \sim 10^{-41} \text{cm}^2$$

 $f_{TG}^{p}, f_{Tq}^{p} \propto$ gluon and quark matrix element in the nucleon However, current DiDt bound $\Rightarrow \sigma_{\chi p} \lesssim 10^{-43} \mathrm{cm}^{2}$  for EW mass DM  $\Rightarrow$  naive estimation  $\sim O(100)$  real  $\frac{\langle \sigma | v | \rangle_{therm}}{\sigma_{Sl}}$  (more severe if null result in near future XENON100/1T...)

## Realistic Models: Mechanisms Affecting $\frac{\langle \sigma|v| \rangle_{therm}}{\sigma_{Sl}}$ -1

- Enhance  $\langle \sigma | v | \rangle_{therm}$ :
  - S-Channel Resonance
  - Coannihilation with mass degenerate partner, particularly useful when self-annihilation p-wave suppressed
- Suppress SI coupling
  - Suppression from Flavor Dependent Couplings: Suppressed coupling to light quark, while other efficient channels (t, lepton, W/Z) maintains  $\langle \sigma | v | \rangle_{therm}$ . 'Classic' example--Yukawa coupling via h-like mediator: Go back to SI  $\sigma_{yp}$ , replace the universal  $g_2$  by  $y_a$ :

$$\sigma_{\chi\rho} = \frac{1}{4\pi} \frac{m_{\rho}^2}{(m_{\chi} + m_{\rho})^2} \frac{(g_1)^2}{M^4} \left( \sum_{q=u,d,s} \frac{m_{\rho}}{m_q} y_q f_{Tq}^{\rho} + \sum_{q=c,b,t} \frac{m_{\rho}}{m_q} y_q \frac{2}{27} f_{TG}^{\rho} \right)^2$$

$$\approx \frac{1}{\pi} \frac{m_{\rho}^2}{m_{\phi}^2} \frac{(g_1)^2}{M^4} (\frac{m_{\rho}}{v})^2 \cdot 0.2 \approx 10^{-45} \text{cm}^2$$

around the reach of XENON100/XENON1T, Super-CDMS!

## Realistic Models: Mechanisms Affecting

$$\frac{\langle \sigma | v | \rangle_{therm}}{\sigma_{SI}}$$
-2

Operator dependent kinematic suppression:

small transferred 
$$p \sim \text{keV} \Rightarrow \epsilon_{\text{v}} = \left(\frac{\text{v}_{DM}}{c}\right)^2 \sim 10^{-6}$$
; low  $p_q$  in nucleon: $\epsilon_{QCD} = \left(\frac{\Lambda_{QCD}}{m_{DM}}\right)^2 \sim 10^{-6}$ 

- Inelastic splitting: DM has heavier 'excited' partner, inelastic scattering dominant;  $\Delta m \Rightarrow$  kinematic barrier, suppressed by  $n_{DM}$  at high v. In general  $\Delta m \gtrsim 1 \text{MeV}$  evade all DiDt bounds. Recently well known for explaining DAMA with  $\Delta m \sim 100 \text{keV}$ .
- Annihilation to Dark Sector States: DM dominantly couples to dark sector, only via small mixing to SM.
   GeV-dark sector recently well explored in light of PAMELA, FERMI anomaly.
- Non-Thermal DM: axions, gravitino LSP. Mostly 'super-weakly' interacting at both DiDt and IdDt

### Operator Properties Relevant for Dark Matter Detection

- Motivation: operator dependence of  $\epsilon_V$ ,  $\epsilon_{QCD}$ ,  $\epsilon_Y$  for DiDt and p-wave/helicity suppression for IdDt
- Study general scalar, fermion (Majorana, Dirac), vector DM. All 4-point SM-DM interaction operator can be written in form of ODMOSM, where O is bilinear operator
- All interesting information (potential suppressions) easily extracted from bilinear properties and CP, J conservation. (Tables listed next page)
- Useful tool for model building, as well as systematic understanding of existing models (later...)

#### Fermion:

	r crinion.									
ſ		ΨΨ	$\bar{\Psi}\gamma^5\Psi$	$\bar{\Psi}\gamma^{\mu}\Psi$	$\bar{\Psi}\gamma^{\mu}\gamma^{5}\Psi$	$\bar{\Psi}\sigma^{\mu\nu}\Psi$	$\bar{\Psi}\sigma^{\mu\nu}\gamma_5\Psi$	$(\bar{\Psi}\gamma^{\mu}\partial^{\nu}\Psi)_{\pm}$	$(\bar{\Psi}\gamma^{\mu}\gamma^{5}\partial^{\nu}\Psi)_{\pm}$	
ſ	SI	$\epsilon \gamma$	0	<b>√</b>	$\epsilon_{ m v}$	$\epsilon_{ m v}$	$\epsilon_{ m v}$	€QCD	$\epsilon_{ m v}$	
	SD	0	$\epsilon_{\rm v}\epsilon_{\rm Y}$	$\epsilon_{ m v}$	✓	✓	✓	$\epsilon_{ m v}$	$\epsilon_{QCD}$	
ſ	С	+	+	_	+	_	_	Ŧ	±	
Π	Р	+	_	$(-)^{\mu}$	$-(-)^{\mu}$	$(-)^{\mu,\nu}$	$-(-)^{\mu,\nu}$	$(-)^{\mu,\nu}$	$-(-)^{\mu, u}$	
	s-wave	0	✓	<b>√</b>	✓	✓	✓	+:√,-:0	+:0, -: ✓	

Scalar:

		$\phi^{\dagger}\phi$	$(\phi^{\dagger}\partial_{\mu}\phi)_{\pm}$	$(\phi^{\dagger} \partial_{\mu} \partial_{\nu} \phi)_{\pm}$
٠ ا	С	+	±	±
	Р	+	$(-)^{\mu}$	$(-)^{\mu,\nu}$
	s-wave	<b>√</b>	+: ✓, −: 0	$+: \checkmark, -: 0$

#### Vector boson:

ſ		VV	$(VV)^{\mu\nu}_{\pm}$	$(\epsilon VV)^{\mu\nu}_{\pm}$	$(V\partial V)^{\mu}_{\pm}$	$(\epsilon V \partial V)^{\mu}_{\pm}$	$(V\partial\partial V)^{\mu\nu}_{\pm}$	$(\epsilon V \partial \partial V)^{\mu\nu}_{\pm}$	$(V\partial^2 V)_{\pm}$	
ſ	С	+	±	±	±	±	±	±	±	
ſ	Р	+	$(-)^{\mu,\nu}$	$-(-)^{\mu,\nu}$	$(-)^{\mu}$	$-(-)^{\mu}$	$(-)^{\mu,\nu}$	$-(-)^{\mu,\nu}$	+	
Π	s-wave	<b>√</b>	<b>√</b>	<b>√</b>	$+: \checkmark, -: 0$					

### Anti-D detection prospect for specific models

Predicted number of anti-deuterons detected in various experiments for a set of dark matter models. Promising at GAPS -ULDB, SAT

Model	m <sub>DM</sub> (GeV)	σ v	ξw	ξq	ξ <sub>t</sub>	ξh	$N_{2\sigma} = 1$ $N_{5\sigma} = 4$ (ULDB)	$N_{2\sigma} = 5$ $N_{5\sigma} = 14$ (SAT)	σ <sub>SI</sub>	$\sigma_{SD}$
SUSY F.P (1)	190	0.67	0.2	0.02	0.73	0	4	47	10-8	10-4
SUSY F.P (2)	772	0.33	0.55	0	0.38	0	0	1	10 <sup>-8</sup>	10 <sup>-5</sup>
SUSY coann	148	0.17	0	1	0	0	1	11	10 <sup>-8</sup>	10 <sup>-6</sup>
SUSY A-funnel	163	0.6	0	0.92	0	0	2	30	10 <sup>-8</sup>	10 <sup>-6</sup>
UED <i>B</i> <sup>(1)</sup>	900	0.6	0	0.19	0.16	0.02	0	0	10 <sup>-8</sup>	10-6
UED B <sup>(1)</sup> coann.	600	0.6	0	0.19	0.16	0.02	0	1	10 <sup>-8</sup>	10 <sup>-6</sup>
LHTP	200	0.8	1	0	0	0	0	9	10 <sup>-12</sup>	10-10
LZP $\nu_R^0$	300	1	0.06	0	0.94	0	3	38	10 <sup>-9</sup>	10-7
Singlet (scalar)	200	1	0	0	0	1	2	33	10 <sup>-8</sup>	0
Doublet/Sing	et 75	0.1	1	0	0	0	3	46	0	$10^{-4}$

### **Conclusions**

- Anti-D is a unique low background ldDt channel for DM
- With current day  $\langle \sigma | v | \rangle_{ann} = 1$  pb, near future experiments (AMS-02, 3-phase of GAPS) have good reach for various annihilation final states
- General tension between  $\langle \sigma | v | \rangle_{therm}$  and bound on  $\sigma_{SI}$  is studied, basic mechanisms listed as solution. Operator analysis for various DM/interaction: for a variety of models significant  $\bar{D}$  signal even when DiDt rate highly suppressed
- Detection prospects for various well-motivated models is studied: promising at GAPS-ULDB, SAT