

# Neutrino Mass and Flavor Unification

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# Neutrino Mass- What we do and don't know ?

- **Masses:**  $\Delta m_{sol}^2 \cong 7.67 \times 10^{-5} eV^2$  ;  $\Delta m_{Atm}^2 \cong 2.39 \times 10^{-3} eV^2$
- **Mixings:**  $\sin^2 \theta_{12} \cong .312$ ;  $\sin^2 \theta_{23} \cong .466$  ;  $\sin^2 \theta_{13} \leq .04$
- **Overall mass scale:**  $< .1- 1 eV$  (roughly)
- **To be determined** (expts in progress or planning)

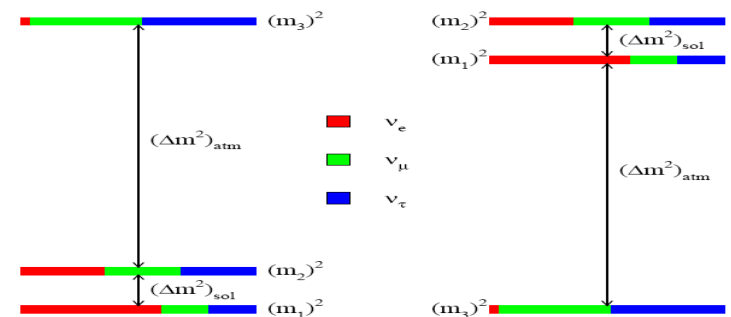
(i) Majorana or Dirac ?

(ii) Mass ordering: NH or IH

(iii) Value of  $\theta_{13}$

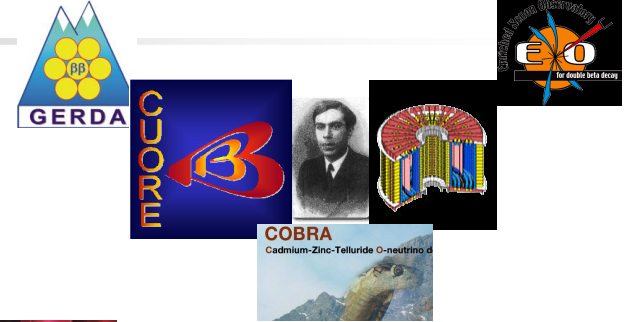
(iv) Any possible CP violation ?

(v) Leptonic unitarity



# Many experiments on the way

(i) Majorana or Dirac  $\beta\beta_{0\nu}$



(ii) Absolute mass scale:



(iii) Mass ordering:



(iv) Value of  $\theta_{13}$



(v) CP phase



# Challenge for Theory

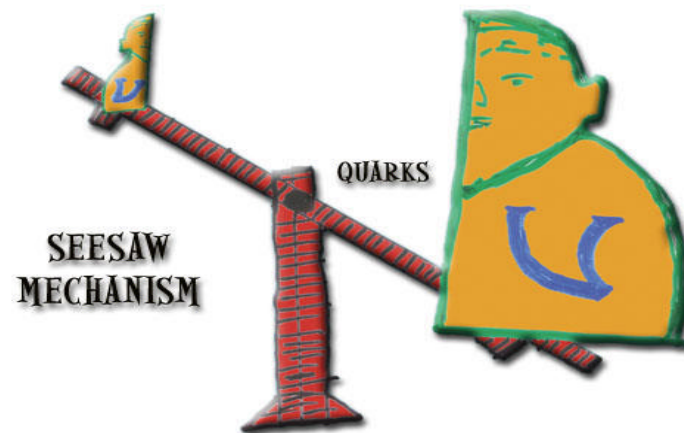
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- **Neutrino mass** → New physics beyond SM:
  - Two important issues:
  - **New mass scale**  
Is it accessible at the LHC ?
- **Understanding flavor** : Can we have a unified understanding quark- lepton mixings ?

# 1. Mass scale

WHY  $m_\nu \ll m_{q,l}$  ?

- **Seesaw Paradigm:** Add heavy right handed neutrinos  $N$  to SM and Seesaw:



**Seesaw scale is the new physics scale !!**

- Two different seesaws depending on whether  $N$  Majorana or Dirac

# Type I Seesaw

- Majorana  $N$

$$L_Y = h_\nu \bar{L} H N_R + M_R N N$$

- $M_R$  Breaks **B-L** : New scale

and new symmetry:

$$m_\nu \cong -\frac{h_\nu^2 v_{wk}^2}{M_R}$$

- After EWSB

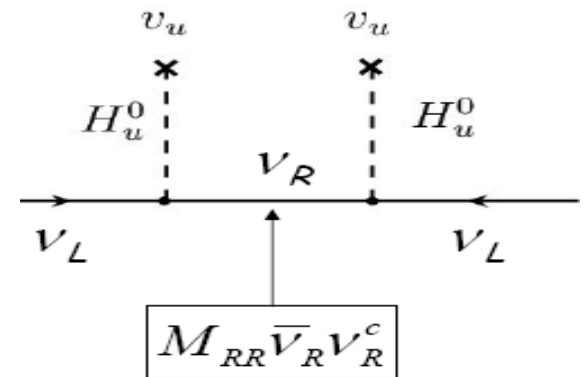
-Neutrino majorana-  $\beta\beta_{0\nu}$  a key test:

$$\frac{m_D}{M_R} \equiv \frac{h_\nu v_{wk}}{M_R}$$

a parameter for collider test

$$\sim 10^{-7} - 10^{-10}$$

type I



# Inverse Seesaw

- Mostly Dirac  $N$  i.e. add another singlet  $S$

$$L_Y = h_\nu \bar{L} H N_R + M S N_R + \mu S S \quad \mu \ll M$$

$$\begin{pmatrix} \nu_L, \nu_R, S \end{pmatrix} \begin{pmatrix} 0 & h\nu_{wk} & 0 \\ h\nu_{wk} & 0 & M \\ 0 & M & \mu \end{pmatrix}$$

$$m_\nu \cong -m_D^T M^{-1} \mu M^{-1} m_D$$

(RNM'86; RNM, Valle'86)

- Seesaw testing parameter  $\frac{m_D}{M} \sim 10^{-3}$  or larger;

# Possible Seesaw scale

- **Neutrino masses do not determine seesaw scale**

- $m_D \approx m_t \rightarrow M_R \approx 10^{14} \text{ GeV} ; M_U \sim 10^{16} \text{ GeV}$

**Both**  $m_D \approx m_t$  **and high seesaw scale** indication for SUSYGUTs; **No collider signals ! (Common Lore)**

- $m_D \approx m_e \rightarrow$  B-L scale at **TeV** ★

- Inverse seesaw  $\rightarrow$  B-L at TeV even with  $m_D \approx m_t$  ★

**Both collider accessible**



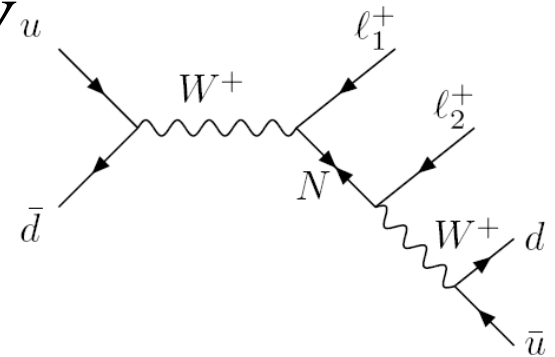
# Searching for TeV seesaw

- Manifestations via  $\nu - N$  mixing

$$\frac{m_D}{M}$$

**(a) collider production** for  $M_N < TeV^u$

del Aguila, Saavedra, Pittau'06; Han, Atre, Pascoli, Zhang'09



**(b) Leptonic non-unitarity:**

(Antusch, Gavela, Biggio, Fernandez,...)

- **Both negligible for type I seesaw but observable for inverse seesaw  $M_N \sim TeV$**
- **Situation different with gauge forces !!**

# How plausible are new gauge forces ?

- **Type I** : why is seesaw scale so far below Planck scale: **Local B-L** symmetry - a plausible answer !
- **Inverse seesaw case**:

■ **Why**

$$\begin{pmatrix} 0 & hv_{wk} & 0 \\ hv_{wk} & 0 & M \\ 0 & M & \mu \end{pmatrix}$$

**why not**

$$\begin{pmatrix} 0 & hv_{wk} & h'v_{wk} \\ hv_{wk} & M' & M \\ h'v_{wk} & M & \mu \end{pmatrix}$$

- **New Gauge symmetry can explain this !!**

# A concrete realization: Left-Right Sym model

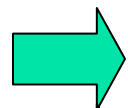
■  $N_R \rightarrow$  Gauge group:  $SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$

■ New  $W'$  and  $Z'$

■ Fermion assignment

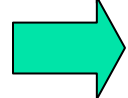
$$\begin{array}{ccc}
 W_L^\pm & & W_R^\pm \quad Z, Z', \gamma \\
 \begin{pmatrix} u_L \\ d_L \\ \nu_L \\ e_L \end{pmatrix} & \begin{matrix} P \\ \Leftrightarrow \\ P \\ \Leftrightarrow \end{matrix} & \begin{pmatrix} u_R \\ d_R \\ \nu_R \\ e_R \end{pmatrix}
 \end{array}$$

■ Two Avatars of Low scale LR:



type I

$$\phi(2,2,0) ; \Delta_R(1,3,+2) \oplus \Delta_L(3,1,+2)$$



Inverse seesaw

$$\phi(2,2,0) + \chi_L(2,1,-1) + \chi_R(1,2,-1)$$

# BOUND ON LR SCALE

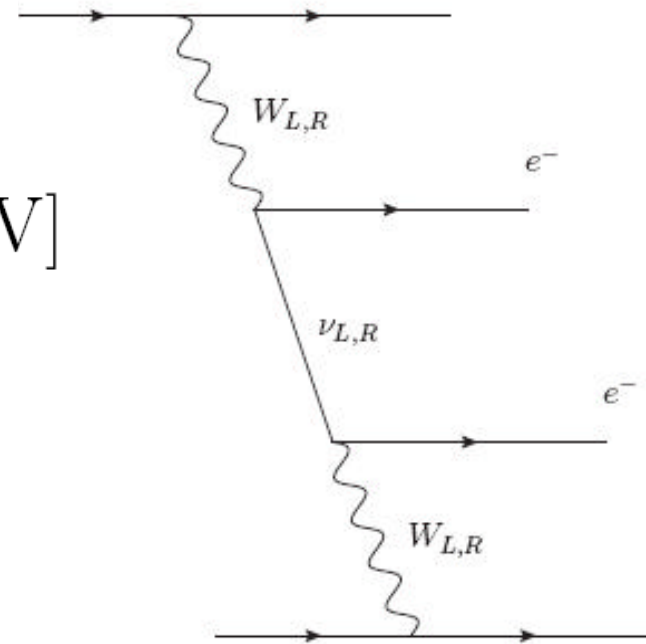
- Most stringent bounds come from CP viol. Observables e.g.  $\varepsilon, \varepsilon', d_n^e$  ; clearly depends on how CP is introduced:  
**Two minimal scenarios**
- **Parity defined as usual: ( $\psi_L \leftrightarrow \psi_R$ ) minimal model:**  
 $\theta_L^{CKM} = \theta_R$  ; 2 CP phases  $M_{W_R} \geq 4TeV$  (An, Ji, Zhang, RNM '07)
- **Parity as C (as in SUSY i.e.  $\psi \leftrightarrow \psi^c$  )** (Maezza, Nesti Nemevsek, Senjanovic'10)  
 $M_{W_R} \geq 2.5TeV$
- **With SUSY: bounds weaker: > 1-2 TeV** (An, Ji, Zhang'08)
- **Collider (CDF, D0) 640-750 GeV.**

# Bounds from Nu-less double beta decay

- New contributions from WR-N exchange (**only for Case I**) (RNM, 86; Hirsch, Klapdor, Panella 96)
- **Diagram:**

$$\rightarrow m_{W_R} \geq 1.1 \left( \frac{\langle m_N^{(V)} \rangle}{1\text{TeV}} \right)^{(-1/4)} [\text{TeV}]$$

**From Ge76:**

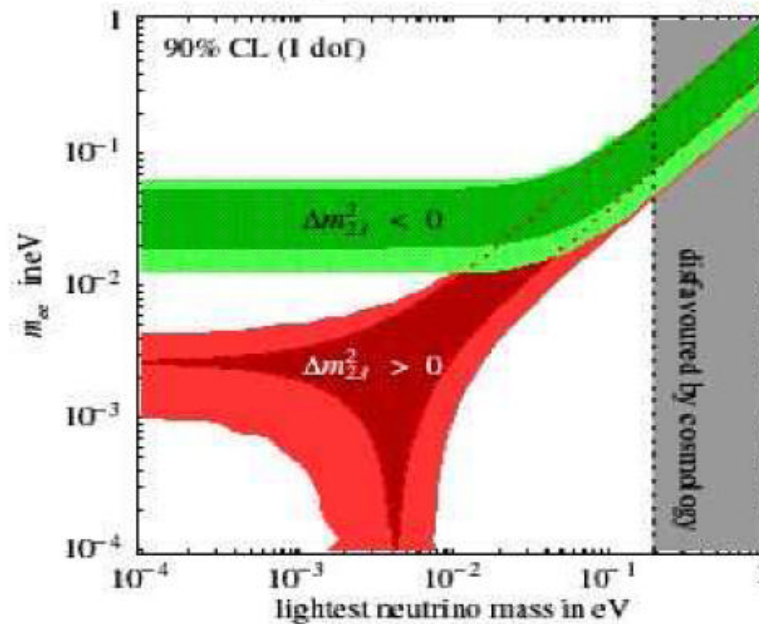


# TeV Seesaw signal from $\beta\beta_{0\nu}$

- **Nu contribution:**

Inverse hierarchy

Normal hierarchy



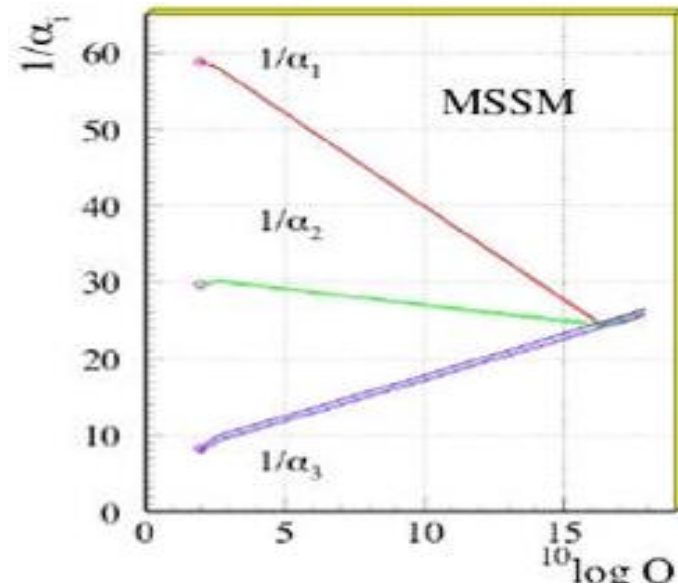
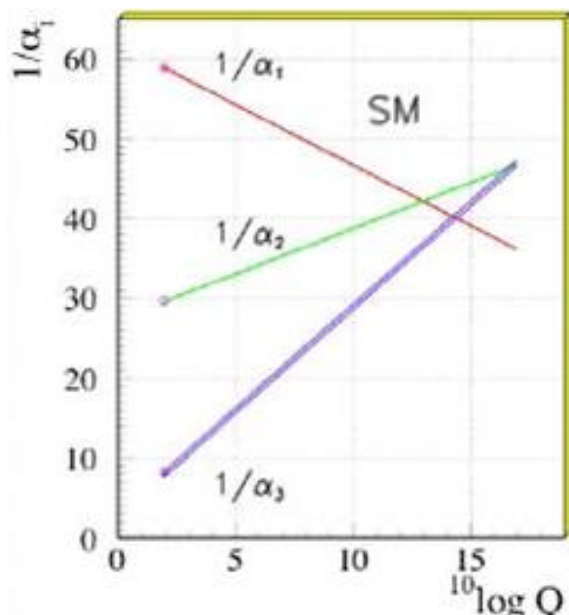
## Punch line:

- Suppose long baseline  $\rightarrow \Delta m_{31}^2 > 0$
- and nonzero signal for  $\beta\beta_{0\nu}$  (+ RP if susy)

**$\rightarrow$  could be a signal of TeV WR and type I**

# Grand unification prospects of TeV seesaw

- Major selling point of MSSM: **gauge couplings unify with TeV scale new physics**



- How does seesaw physics affect unification ?**
- Can seesaw physics coexist w/MSSM at TeV ?**

# SO(10)- Natural GUT for LR Seesaw

- {16 }- spinor for all matter  
-includes RH neutrino

$$\begin{pmatrix} u & u & u & \nu \\ d & d & d & e \end{pmatrix}_{L,R}$$

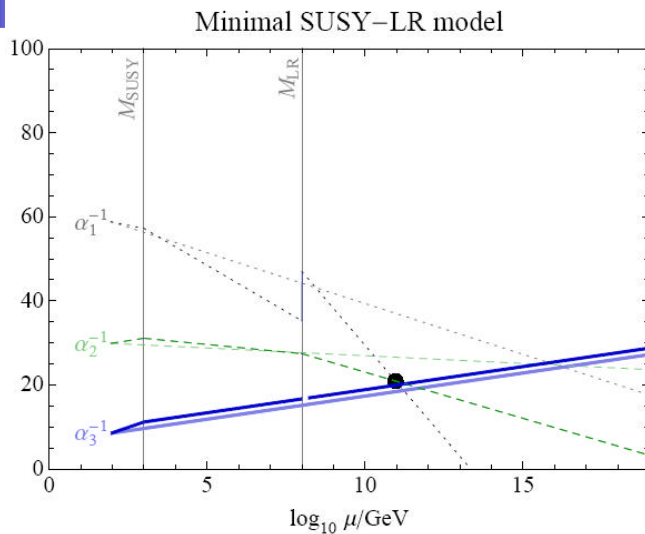
- SO(10) contains left-right group

$$SO(10) \supset SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L} \otimes SU(3)_c$$

- A long standing question: does TeV scale LR unify to SO(10) ?
- If yes LHC can be a machine that probes grand unification !



# TeV type I does not grand unify:



$M_R$ (GeV)	$\mu_0$ (GeV)
$10^3$	$7.76 \times 10^{13}$
$10^5$	$4.56 \times 10^{14}$
$10^7$	$2.56 \times 10^{15}$
$10^8$	$6.16 \times 10^{15}$
$10^9$	$1.44 \times 10^{16}$
$10^{10}$	$3.46 \times 10^{16}$
$10^{11}$	$8.31 \times 10^{16}$

Landau Pole

(Kopp, Lindner, Niro, Underwood'09)

(Parida, Sarkar, Majee, Raichaudhuri'09)

## ■ Type I: only unification route

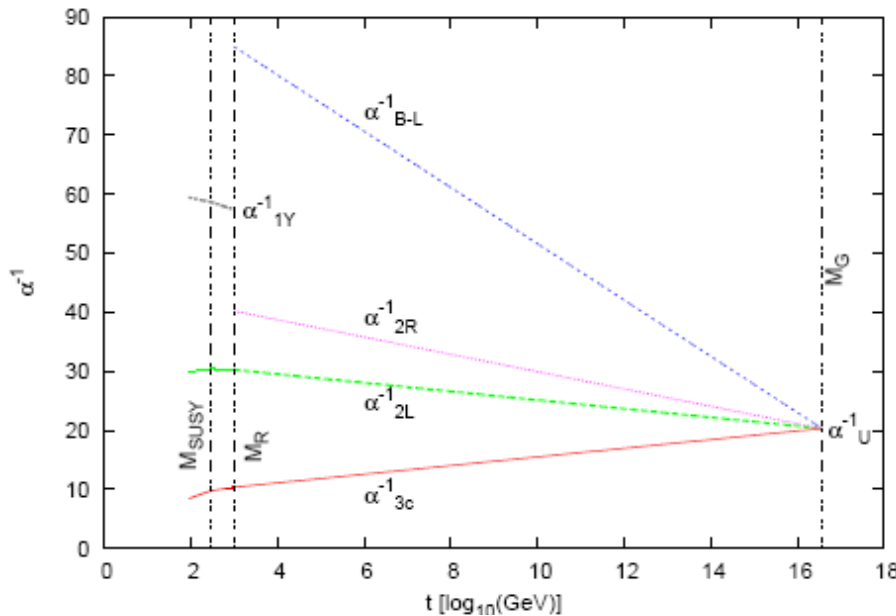
$$\bullet SO(10) \xrightarrow{M_G} 3_c 2_L 1_Y (\text{MSSM}) \xrightarrow{M_{\text{SUSY}}} 3_c 2_L 1_Y (\text{SM}) \xrightarrow{M_Z} 3_c 1_Q$$

$$M_{U,BL} \cong 2 \times 10^{16} \text{ GeV}$$

no new TeV physics besides MSSM

# TeV Inverse Seesaw (LR) does unify

- **New result! Inverse seesaw does unify –TeV WR and Z'**



- does give realistic fermion mass fits;
- OK with proton decay

$$m_{\tilde{q}_{1,2}} \geq TeV$$

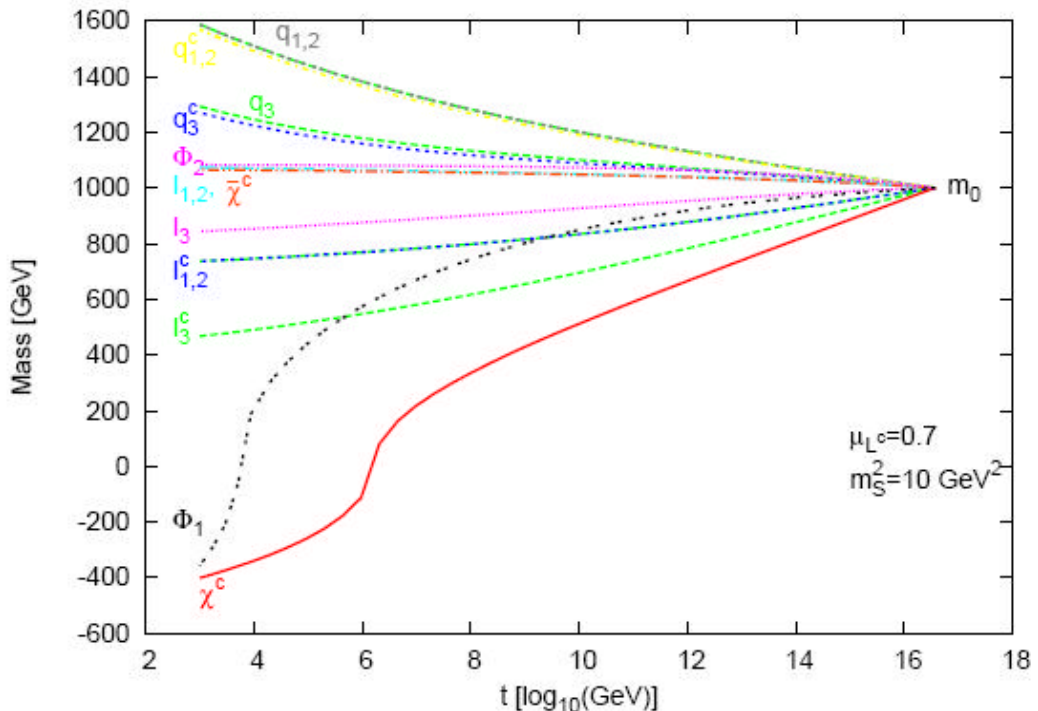
- $SO(10) \xrightarrow{M_G} 3_c 2_L 2_R 1_{B-L} \xrightarrow{M_R} 3_c 2_L 1_Y (\text{MSSM}) \xrightarrow{M_{\text{SUSY}}} 3_c 2_L 1_Y (\text{SM}) \xrightarrow{M_Z} 3_c 1_Q$

$$M_U \cong 10^{16} GeV; M_{BL,R} \cong TeV$$

(Dev, RNM, 09; PRD; arXiv: 1003:6102);

# Radiative Breaking of B-L, SM in TeV seesaw SO(10)

- Positive spartner mass square at GUT scale-
- RGE turns them negative much like SM with large t-mass
- (Dev, RNM'10)



# Other advantages of a Low seesaw scale

- **Leptogenesis**- : High scale leptogenesis + hierarchical NR (as in type I or II SO(10) →

$$M_N \geq 10^9 \text{ GeV}$$

vs

$$T_R < 10^6 - 10^9 \text{ GeV}$$

adequate leptogenesis

gravitino reheat constraint

→ Low scale resonant leptogen as alternative:

- **Suppression of proton decay**: High scale seesaw SO(10) with 16-Higgs has proton decay problem from operator;

→ suppressed for TeV  
B-L breaking.

$$\frac{(16_m)^3 16_H}{M_{Pl}}$$

# Type I vs Inverse seesaw at LHC

- WR and Z':  $u\bar{d} \rightarrow W_R \rightarrow l^+ N$  ;  $u\bar{u} \rightarrow Z' \rightarrow NN$

(Keung, Senjanovic; Han, Perez, Huang, Li, Wang; Del Aguila, Aguilar-Saavedra; de Blas, Azuelos,

- N-decay: (a)  $\nu N$  mixing and/or (b)  $W_R$  exchange

- **type I**  $\theta_{\nu N} \ll 10^{-3}, M_{W_R} < 4TeV$  (a) negligible;  $N \rightarrow l^\pm jj$

- **Signal: like sign dileptons+jets; no missing E**

- **Inv. Seesaw:**  $\theta_{\nu N} > 10^{-3} \rightarrow N \rightarrow l^- jj, l^- l^+ \nu$

- **Trileptons; no same sign dileptons; dominantly**  $\tau^\mp l^\pm l^\pm$

- Background from  $t\bar{t}nj, WZnj, WWnj, \dots$

# Other signals

- **TeV type I Seesaw requires B-L=2 Higgs:**

$$\Delta = \begin{pmatrix} \frac{1}{\sqrt{2}} \Delta^+ & \Delta^{++} \\ \Delta^0 & -\frac{1}{\sqrt{2}} \Delta^+ \end{pmatrix} \quad \Delta^{++} \rightarrow \mu^+ \mu^+, ee, \tau\tau$$

- Doubly charged Higgs  $\rightarrow$  **Four lepton signals at LHC**
- **Leptonic non-unitarity for inverse seesaw:**
- Can be probed in Short baseline neutrino expts  $\nu_\mu - \nu_\tau$   
also  $\mu \rightarrow e + \gamma$        $\mathbf{B} \sim 10^{-15}$
- (Ilakovac, Pilaftsis; Antusch, Gavela, Biggio, Fernandez, Malinsky, Ohlsson, Zhang,)
- $\delta U_{PMNS} \sim \left( \frac{m_D}{M_N} \right)^2 > 10^{-6}$        $|\eta| < \begin{pmatrix} 2.0 \times 10^{-3} & 3.5 \times 10^{-5} & 8.0 \times 10^{-3} \\ 3.5 \times 10^{-5} & 8.0 \times 10^{-4} & 5.1 \times 10^{-3} \\ 8.0 \times 10^{-3} & 5.1 \times 10^{-3} & 2.7 \times 10^{-3} \end{pmatrix}$       **Current bds:**



## 2. Understanding Flavor

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(i) Mass hierarchies

(ii) Strange mixing patterns:

Leptons:

Quarks:

$$\mathbf{U}_{\text{PMNS}} \approx \begin{pmatrix} \frac{\sqrt{6}}{3} & \frac{\sqrt{3}}{3} & 0 \\ -\frac{\sqrt{6}}{6} & \frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \\ \frac{\sqrt{6}}{6} & -\frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \end{pmatrix} \quad \mathbf{U}_{\text{CKM}} = \begin{pmatrix} 1 & \lambda & \lambda^3 \\ -\lambda & 1 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix}$$

(Harrison, Perkins, Scott; He, Zee, Xing,..)

# Mass Texture

- Up-quark and charged lepton diagonal basis:

$$M_d = m_b \begin{pmatrix} \lambda^5 & \lambda^3 & \lambda^3 \\ \lambda^3 & \lambda^2 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix} \quad M_\nu \cong \begin{pmatrix} \varepsilon_1 & \varepsilon_3 & \varepsilon_3 \\ \varepsilon_3 & 1 + \varepsilon_1 & -1 + \varepsilon_3 \\ \varepsilon_3 & -1 + \varepsilon_3 & 1 + \varepsilon_1 \end{pmatrix}$$

- $\varepsilon_i \sim \lambda$  = Cabibbo angle






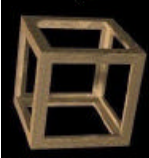
# Strategy for texture

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- Key idea: SM has a large sym for zero fermion masses :  **$[SU(3)]^5$** ;
- **Choose subgroup**: Discrete subgroup with 3-d. rep.
- **Replace Yukawa's by scalar fields (flavons)**;
- **Minima of the flavon theory determines Yukawas**:

# Application to Neutrinos

- Successful Family symmetries for TBM:

$$S_{2(\mu-\tau)} \subset S_3, S_4, A_4, \Delta(3n^2), \dots$$


## Flavon fields are triplets:

(Ma, Rajasekaran; Babu, Ma, Valle, King, Ross; Altarelli, Feruglio, Chen, Mahanthappa; Everett, Ramond; Luhn, Nasri, Yu, RNM, Hagedorn, Morissi, Bazzocchi, Merlo, Toorop, Tanimoto, Shimizu, Ishimori,...)

**How can we unify with quarks ?**

**Grand unified theories:**

# A new high scale Ansatz for unifying quark-lepton flavor

**A**

$$M_u = M_0 + \delta_u$$

$$M_d = rM_0 + \delta_d$$

$$M_l = rM_0 + \delta_l$$

(Dutta, Mimura, RNM'PRD-09)

$$m_\nu = f\nu_L$$

$$\delta_{u,d,l} \ll M_0$$

■  $f$  diagonal.

■ Anarchic  $M_0$ , quark mixings small while lepton mixings large.

**B**

Rank 1  $M_0$

→ explains mass hierarchies +

$$m_b \cong m_\tau$$



# A SUSY SO(10) realization

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- Fermions in {16}:

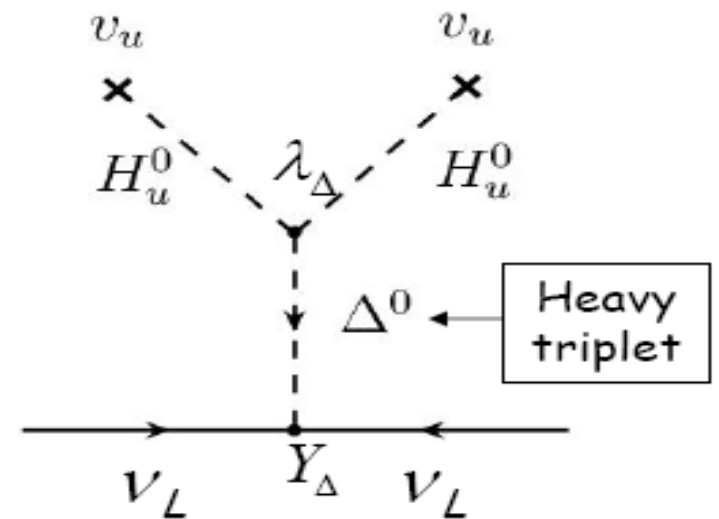
- $16_m \times 16_m = \{10\}_H + \{120\}_H + \{126\}_H$

- Fermion masses from Yukawas as in SM:

$$L_Y = h 16 \cdot 16 \cdot 10_H + f 16 \cdot 16 \cdot \overline{126}_H + h' 16 \cdot 16 \cdot [12010]_H$$

# Neutrino mass formula in GUT scale B-L in SO(10):

$$m_\nu = f v_\Delta - M_D \frac{1}{f v_{BL}} M_D^T$$



- **Type II seesaw:**

# SO(10) with GUT scale B-L → unified approach to flavor

- → fermion mass formulae:

$$Y_u = h + r_2 f + r_3 h'$$

$$Y_d = r_1(h + f + h')$$

$$Y_e = r_1(h - 3f + c_e h')$$

$$m_\nu \cong f \nu_\Delta$$

(Babu, Mohapatra'92)

Bajc, Senjanovic, Vissani'03

- For  $f, h' \ll h$ , → yields ansatz part A at  $M_U$ ;
- Rank from flavor symmetry: e.g.  $S_4, \Delta(27), \dots$

# An $S_4$ xSO(10)- example

- Solar mass

$$\frac{m_{solar}}{m_{atm}} \cong \lambda \cong \theta_c$$

Dutta, Mimura, RNM arXiv:0911.2242

- Bottom-tau:  $m_b \approx m_\tau$  and

$$m_\mu = -3m_s$$

- Leading order PMNS:  $\mathbf{U} =$

$$\begin{pmatrix} \frac{\sqrt{6}}{3} & \frac{\sqrt{3}}{3} & 0 \\ -\frac{\sqrt{6}}{6} & \frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \\ \frac{\sqrt{6}}{6} & -\frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \end{pmatrix}$$

- Corrections: Testable

Bjorken, King, Pakvasa Ferrandis; Chen, Mahanthappa

$$\theta_{13} = \frac{\theta_c}{3\sqrt{2}} \cong 0.05$$

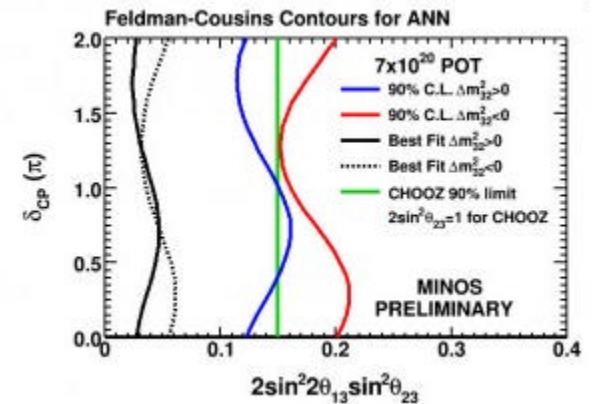
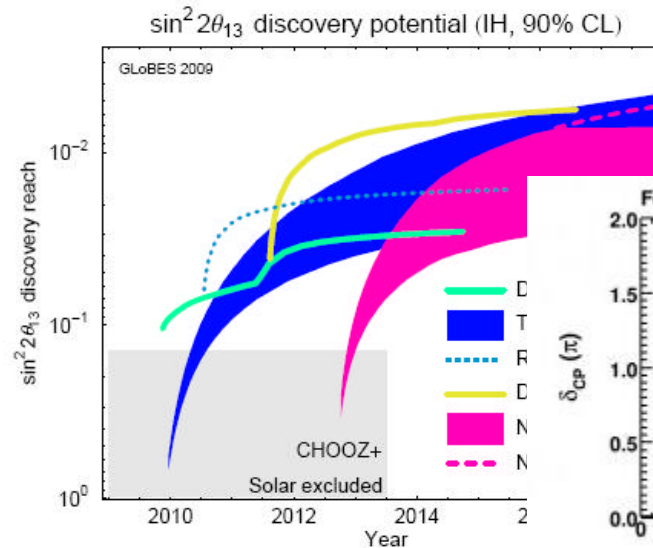
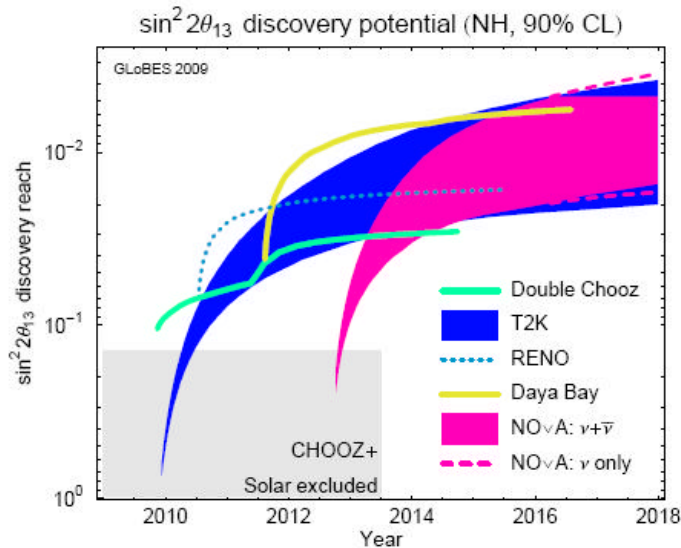
- Double beta mass 3 meV.

# Prospects for measuring

$$\theta_{13}$$

- Reactor, Long base line e.g. MINOS, T2K, NoVA:

(Lindner, Huber, Schwetz, Winter'09)



**Our prediction**

$$\sin^2 2\theta_{13} > 0.01$$

**MINOS**





## Conclusion:

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- **TeV scale WR compatible with  $SO(10)$  GUT; Can be tested at LHC; New dark matter candidate:**
- **New unified approach to flavor based on typeII+  $SO(10)$ - testable via  $\theta_{13}$ . Can also be extended to inverse seesaw case.**

# New Dark matter in TeV scale Inverse seesaw:

- If super-partner of RH neutrino is the lightest, it will be stable due to R-parity- become DM.

$$W = W_{MSSM} + h_\nu LHN + f\nu_R NS + \mu SS$$

- Soft breaking:

$$-L = -L_{soft}^{MSSM} + M^2 \tilde{N}\tilde{N} + M_S^2 \tilde{S}\tilde{S} + A\tilde{L}H\tilde{N} + B\tilde{N}\tilde{S}$$

- **Lightest linear combination is dark matter:**

(Fornengo, Arina, Bazzochi, Romao, Valle'08) (Matchev, Lee, Nasri'08)

■ Minimal Type I case:

Usual Bino-Higgsino

# Dark matter in TeV scale

## Inverse seesaw:

- Inverse seesaw case: New DM :**  $\tilde{\nu}^c$  **:Two**  
**contributions to relic density:**  
**Z' exchange No or small Z' effect**

