

Υ Decays Into Light Scalar Dark Matter

Gagik Yeghiyan

Wayne State University, Detroit, MI, USA

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Based on

G. K. Yeghiyan, Phys. Rev. D 80, 115019 (2009)

Υ meson: $b\bar{b}$, $J=1$

We consider

$$\Upsilon(1S) \rightarrow \Phi\Phi^*$$

$$\Upsilon(3S) \rightarrow \Phi\Phi^* \gamma$$

within the class of light (mass \sim few GeV or less) scalar DM models where these decays are due to exchange of heavy virtual states (with a mass $\gg M_Y$)

Other possibility that these decays are mediated by a light (preferably resonant) degree of freedom have been considered earlier:

Gunion, Hooper, McElrath, PRD 73, 015001 (2006):

P Fayet et al (in a number of papers).

GeV or lighter DM models: have in general tension with satisfying $\Omega_{\text{DM}} h^2 \sim 0.11$

may be avoided if DM annihilates due to exchange of light resonances (beyond the scope of the paper)

or, even if no light mediators, still may have no tension with $\Omega_{\text{DM}} h^2 \sim 0.11$ for spin-0 DM, e.g

- **WIMPless miracle (J. Feng, J. Kumar, et al.) :**
within the MSSM with gauge mediated SUSY breaking, scalar DM particle in the hidden sector, **or**
- **Type-II 2HDM with light scalar DM,**
if $v_2/v_1 = \tan \beta \gg 1$, **(Bird, Kowalewski, Pospelov)**

Try to test this class of models, in particular using Υ decays with missing energy

Most recent experimental data - BaBar:

$$B_{\text{exp}}(\Upsilon(1S) \rightarrow \text{invisible}) < 3 \cdot 10^{-4}$$

and

$$B_{\text{exp}}(\Upsilon(3S) \rightarrow \gamma + \text{invisible}) < 3 \cdot 10^{-6} \text{ for } s^{1/2} \leq 7 \text{ GeV}$$

- **May constrain light spin-0 DM models even if no light propagators are exchanged in $\Upsilon(1S) \rightarrow \Phi\Phi^*$ and $\Upsilon(3S) \rightarrow \Phi\Phi^* \gamma$.**
- **Bounds are derived on the parameters which cannot be tested otherwise**

Problem: bound on $\Upsilon(3S) \rightarrow \gamma + \text{invisible}$ is derived assuming that the photon energy is monochromatic- may be used to make only preliminary estimates of possible constraints on light DM models with non-resonant DM production.

The approach:

integrate out heavy degrees of freedom, use low-energy effective theory of four-particle interactions:

$$H_{eff} = \frac{2}{\Lambda_H^2} \sum_i C_i O_i$$

$$O_1 = m_b (\bar{b} b) (\Phi^* \Phi), \quad O_2 = im_b (\bar{b} \gamma_5 b) (\Phi^* \Phi),$$

$$O_3 = (\bar{b} \gamma^\mu b) (\Phi^* i \vec{\partial}_\mu \Phi), \quad O_4 = (\bar{b} \gamma^\mu \gamma_5 b) (\Phi^* i \vec{\partial}_\mu \Phi)$$

$$\vec{\partial} = 1/2(\vec{\partial} - \overleftarrow{\partial})$$

- **The most general effective Hamiltonian to the LO in $1/m_b$ and $1/\Lambda_H$ expansions.**
- **One can perform model-independent analysis, then applying the derived results to particular models**

$$B(\Upsilon(1S) \rightarrow \Phi\Phi^*) = \frac{\Gamma(\Upsilon(1S) \rightarrow \Phi\Phi^*)}{\Gamma_{\Upsilon(1S)}} = \frac{C_3^2}{\Lambda_H^4} \frac{f_{\Upsilon(1S)}^2}{48\pi\Gamma_{\Upsilon(1S)}} \left[M_{\Upsilon(1S)}^2 - 4m_{\Phi}^2 \right]^{3/2}$$

$B(\Upsilon(1S) \rightarrow \Phi\Phi) = 0$ **if $\Phi = \Phi^*$ - may be inferred without any derivation**

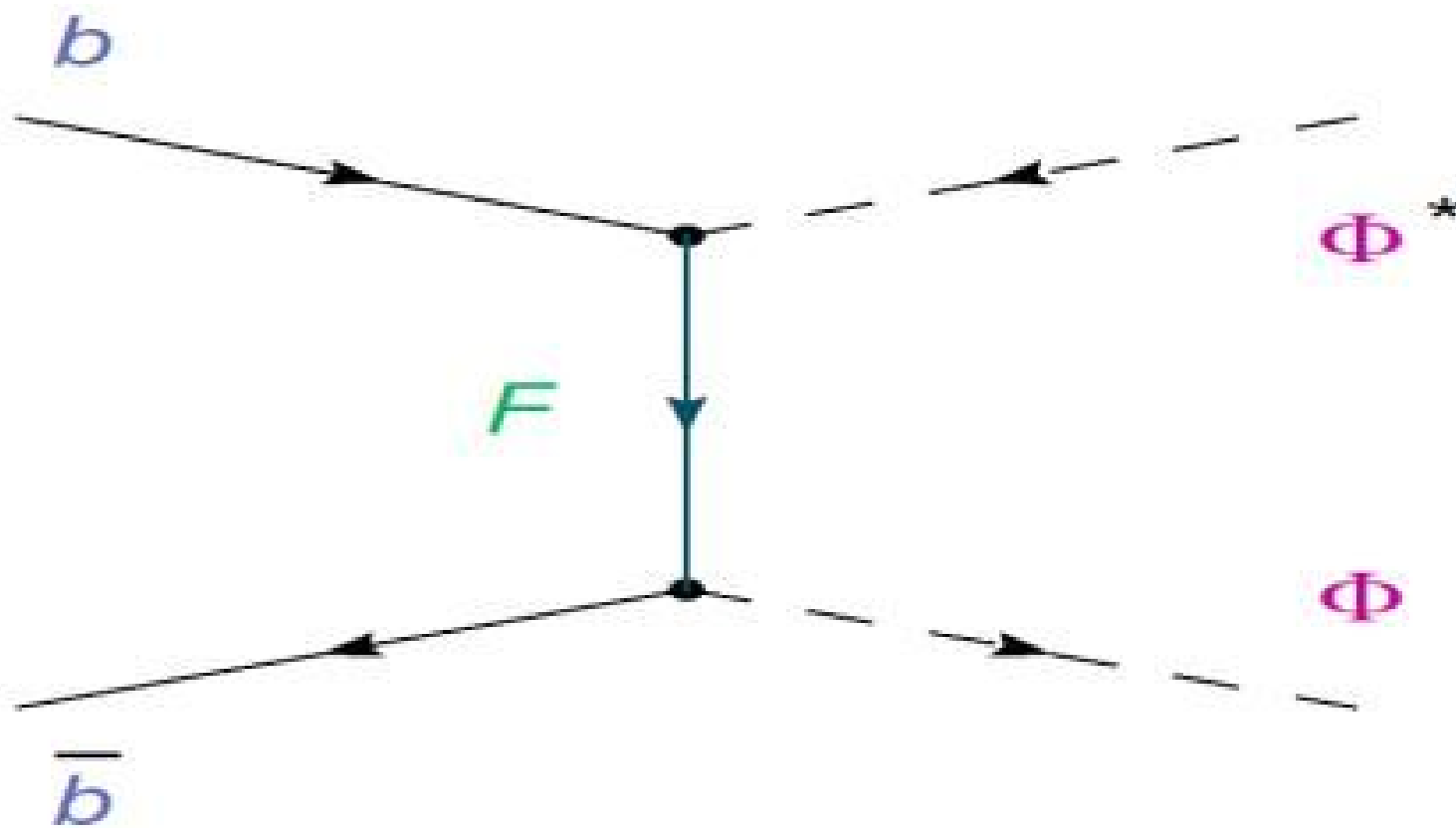
By angular momentum conservation:

$\Phi\Phi$ must be a **P-wave** state - impossible because of the **Bose-Einstein symmetry** of identical particles.

- **Signal for $\Upsilon \rightarrow$ invisible would mean that DM particle, if being a light scalar, has a complex field nature.**
- **No signal implies some constraints on the models with light complex spin-0 DM.**

$B_{\text{exp}}(\Upsilon(1S) \rightarrow \text{invisible}) < 3 \cdot 10^{-4}$ leads to

$$|C_3| < 0.75 \left(\frac{\Lambda_H}{100\text{GeV}} \right)^2 \left(1 - \frac{4m_{\Phi}^2}{M_{\Upsilon(1S)}^2} \right)^{-3/4}$$



Mirror Fermion Models

$$- \mathcal{L} = \Phi (\lambda_{b_L} \bar{F}_{b_R} b_L + \lambda_{b_R} \bar{F}_{b_L} b_R) + h.c. + \dots$$

e.g. MSSM with gauge mediated SUSY breaking with DM in the hidden sector and F -s as connectors

Two scenarios are considered

- Chiral scenario: e.g. $\lambda_{b_R} = 0$
- Non – chiral scenario: $\lambda_{b_R} = \lambda_{b_L} = \lambda_b$

$$- \mathcal{L} = \Phi (\lambda_{b_L} \bar{F}_{b_R} b_L + \lambda_{b_R} \bar{F}_{b_L} b_R) + h.c. + \dots$$

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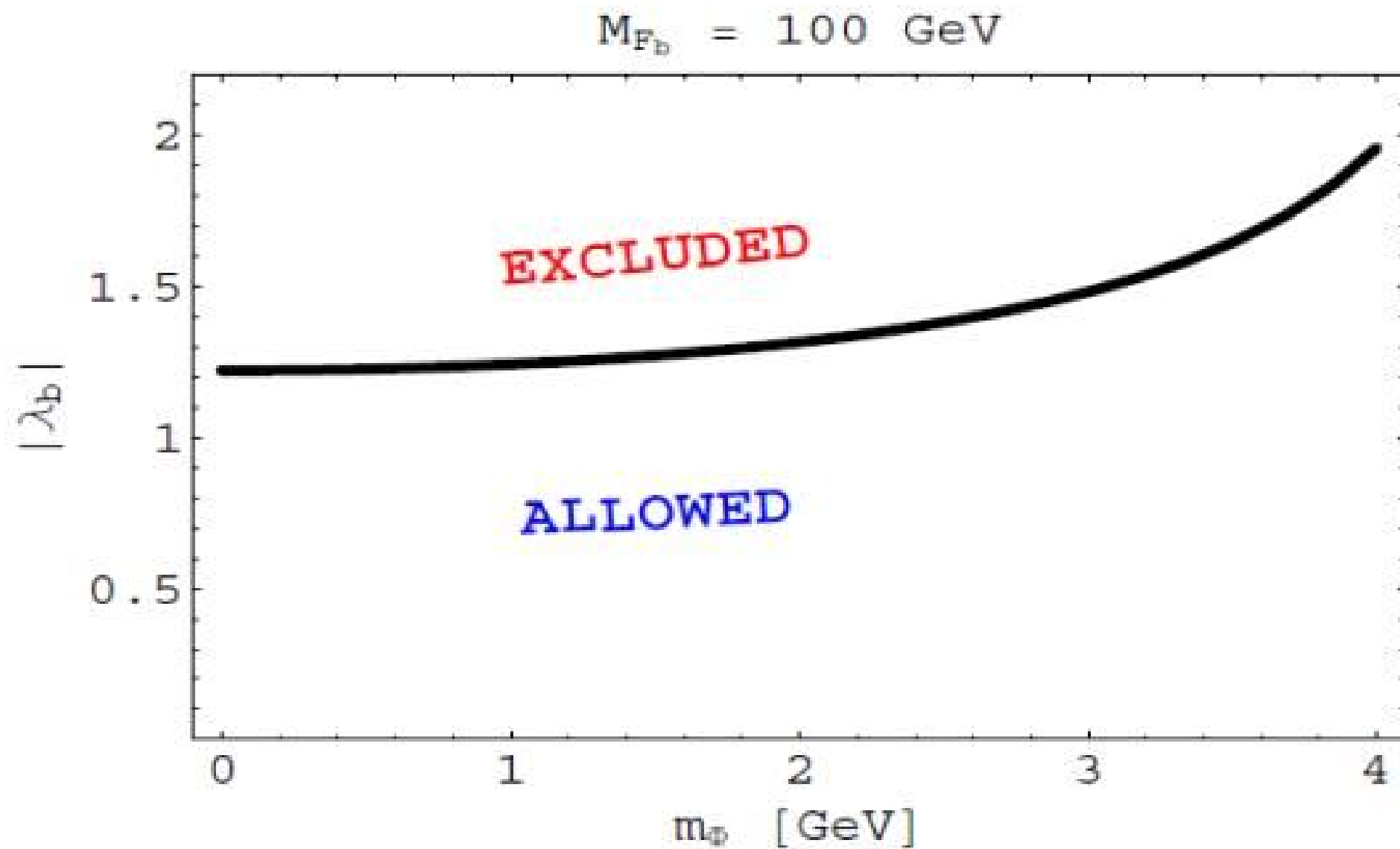
$$|\lambda_{b_L}| < 1.73 \left(\frac{M_{F_b}}{100 \text{GeV}} \right) \left(1 - \frac{4m_\Phi^2}{M_{\Upsilon(1S)}^2} \right)^{-3/8}$$

Non – chiral scenario: $\lambda_{b_R} = \lambda_{b_L} = \lambda_b$

$$|\lambda_b| < 1.22 \left(\frac{M_{F_b}}{100 \text{GeV}} \right) \left(1 - \frac{4m_\Phi^2}{M_{\Upsilon(1S)}^2} \right)^{-3/8}$$

$$\lambda_{bR} = 0: \quad |\lambda_{bL}| < 1.73 \left(\frac{M_{F_b}}{100 \text{ GeV}} \right) \left(1 - \frac{4m_\Phi^2}{M_{\Upsilon(1S)}^2} \right)^{-3/8}$$

$$\lambda_{bR} = \lambda_{bL} = \lambda_b \quad |\lambda_b| < 1.22 \left(\frac{M_{F_b}}{100 \text{ GeV}} \right) \left(1 - \frac{4m_\Phi^2}{M_{\Upsilon(1S)}^2} \right)^{-3/8}$$



$$\lambda_{bR} = 0: \quad |\lambda_{bL}| < 1.73 \left(\frac{M_{F_b}}{100 \text{ GeV}} \right) \left(1 - \frac{4m_{\Phi}^2}{M_{Y(1S)}^2} \right)^{-3/8}$$

$$\lambda_{bR} = \lambda_{bL} = \lambda_b \quad |\lambda_b| < 1.22 \left(\frac{M_{F_b}}{100 \text{ GeV}} \right) \left(1 - \frac{4m_{\Phi}^2}{M_{Y(1S)}^2} \right)^{-3/8}$$

- **Bounds on λ_b – couplings are derived for the first time**
- **DM scattering off nuclei and/or annihilation lead to constraints on λ_u and λ_d but not λ_b**
- **$B \rightarrow K + \text{invisible}$ and $B_s \rightarrow \text{invisible}$ depend on λ_t or a combination of λ_b and λ_s but not on λ_b alone**
- **In other words, study of $Y(1S) \rightarrow \Phi \Phi^*$ leads to bound on parameters of the model that could not be constrained otherwise**

Non – chiral scenario: $\lambda_{bR} = \lambda_{bL} = \lambda_b$

$B_{\text{exp}}(\Upsilon(1S) \rightarrow \Phi \Phi^*) < 3 \cdot 10^{-4}$ leads to

$$|\lambda_b| < 1.22 \left(\frac{M_{F_b}}{100 \text{ GeV}} \right) \left(1 - \frac{4m_{\Phi}^2}{M_{\Upsilon(1S)}^2} \right)^{-3/8}$$

May be improved as $B(\Upsilon \rightarrow \Phi \Phi^* \gamma) \propto (M_{F_b}/m_b)^2$

Preliminary estimate from

$B_{\text{exp}}(\Upsilon(3S) \rightarrow \gamma + \text{invisible}) < 3 \cdot 10^{-6}$

$|\lambda_b| < 0.5, |\lambda_b| < 0.65, |\lambda_b| < 0.9$ respectively for

$M_{F_b} = 100 \text{ GeV}, M_{F_b} = 200 \text{ GeV}, M_{F_b} = 400 \text{ GeV}$

One estimates also to have rigorous bounds from

$B(\Upsilon \rightarrow \Phi \Phi \gamma) \propto \tan^2 \beta$ on the parameters of

type-II 2HDM with light self-conjugate scalar DM.

Experimental studies of $\Upsilon \rightarrow \gamma + \text{invisible}$ for the photons with non-monochromatic energies are encouraged

Conclusions and Summary

- $\Upsilon(1S) \rightarrow \Phi \Phi^*$ and $\Upsilon(3S) \rightarrow \Phi \Phi^* \gamma$ have been studied within the models where DM production is due to exchange of heavy degrees of freedom.
- $B(\Upsilon(1S) \rightarrow \text{invisible}) < 3 \cdot 10^{-4}$ leads to constraints on the model parameters that cannot be tested by the other DM search experiments
- Bounds may also be derived from study of $\Upsilon \rightarrow \Phi \Phi^* \gamma$. We encourage experimental groups to analyze data for $Y \rightarrow \gamma + \text{invisible}$ for non-monochromatic photon energy as well.