

Neutrino oscillation in the LHC era

Phenomenology 2010 Symposium
University of Wisconsin, Madison
10-12 May, 2010

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Outline

Introduction

Neutrino Oscillation in the Standard Model

Neutrino oscillation beyond the SM

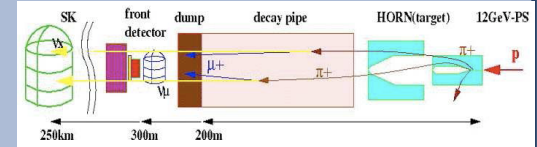
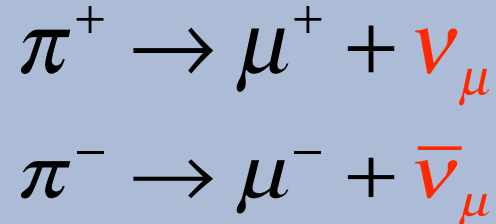
Conclusions

1) Introduction

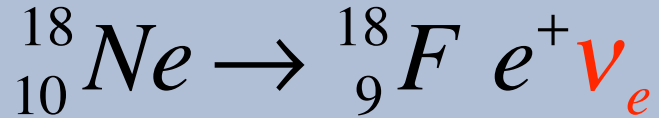
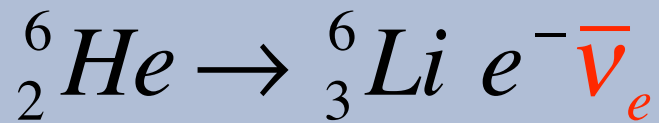
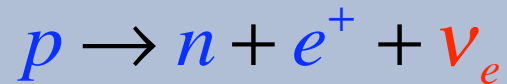
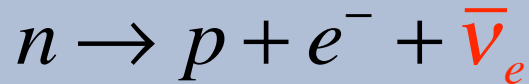
- ◆ In frame of the Standard Model (SM)
 - ⊗ only left handed neutrinos (right-handed antineutrinos) are produce and detected,
 - ⊗ the neutrino flavour states are well defined, they are orthogonal, phenomenon of neutrino oscillation is well described,
- ◆ At higher energies (e.g. at LHC), beyond the SM can appears and can be tested (e.g. supersymmetry), in such models, standard neutrino oscillation theory does not work,
- ◆ Generally there are many models with Non Standard neutrino Interaction (NSI) where the existing theory of oscillation can not be applied,
- ◆ Here we propose the general approach for neutrino oscillation, valid for any NSI where, neutrinos in positive and negative helicity states can be produced, and neutrino flavour states are only approximately defined,
- ◆ We found the necessary and sufficient conditions for the NSI, under which the standard neutrino oscillation theory is correct.

Different processes for neutrino production

✧ Neutrino superbeams, Off axis neutrino beams

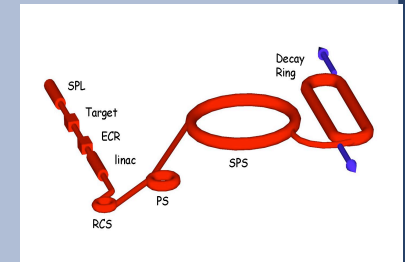


✧ Beta beams

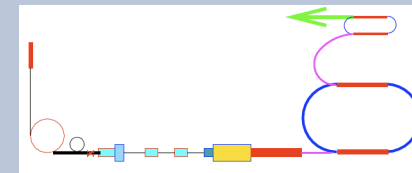
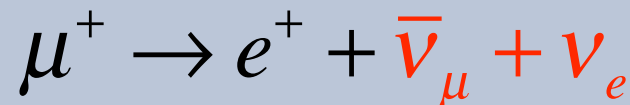
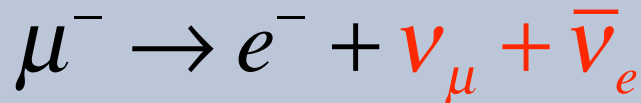


Average $E_{cms} = 1.937$ MeV

Average $E_{cms} = 1.86$ MeV

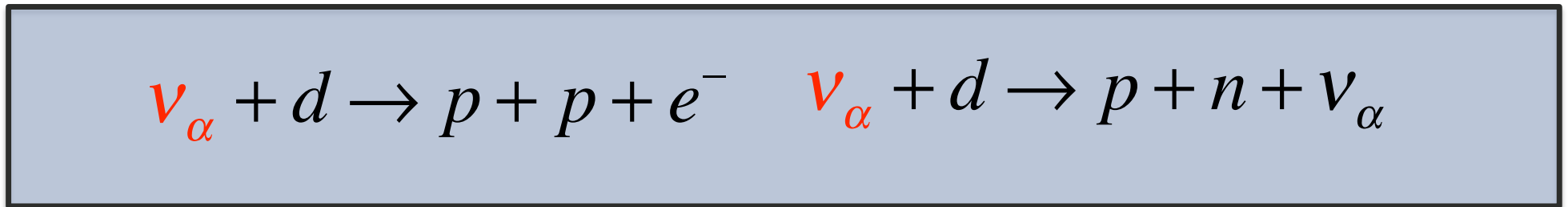
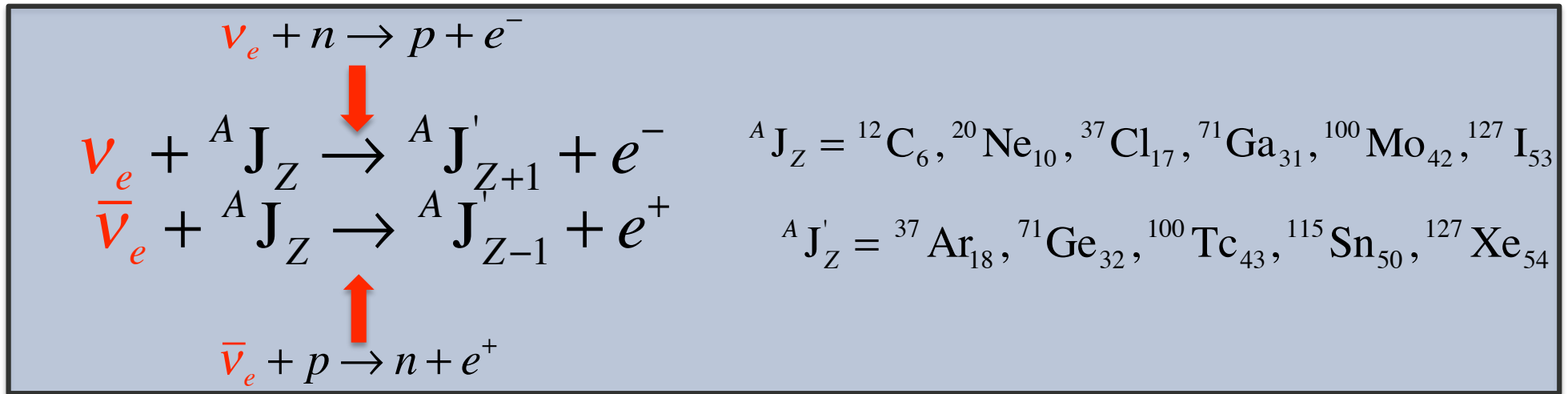


✧ Neutrino factories



Processes for neutrino detection

- 1) Water Cerenkov Detector
- 2) Liquid Argon Detector
- 3) Iron Calorimeter
- 4) Emulsion Detector



Neutrino oscillation in the SM

For production and detection processes - charge current

$$L_{CC} = \frac{e}{2\sqrt{2} \sin \theta_W} \sum_{\alpha, i} l_{\alpha} \gamma^{\mu} (1 - \gamma_5) U_{\alpha i} \nu_i W_{\mu}^{-} + h.c$$

Relativistic (anti)neutrinos
are produced in pure
Quantum Mechanical
flavour state

$$|\nu_{\alpha} \downarrow\rangle = \sum_i U_{\alpha i}^* |\nu_i \downarrow\rangle$$

Neutrinos always with
negative helicity

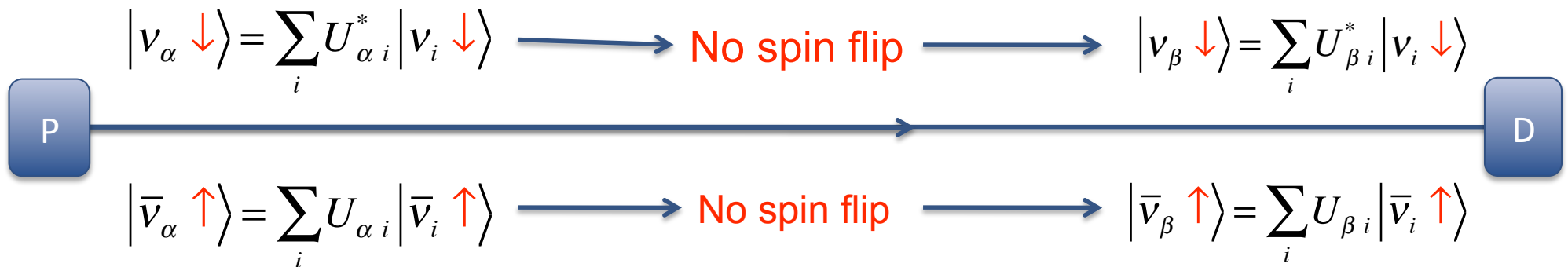
Z. Maki, M. Nakagawa, S. Sakata,
Prog.Theor.Phys. 28(1962)870

$$|\bar{\nu}_{\beta} \uparrow\rangle = \sum_i U_{\beta i} |\bar{\nu}_i \uparrow\rangle$$

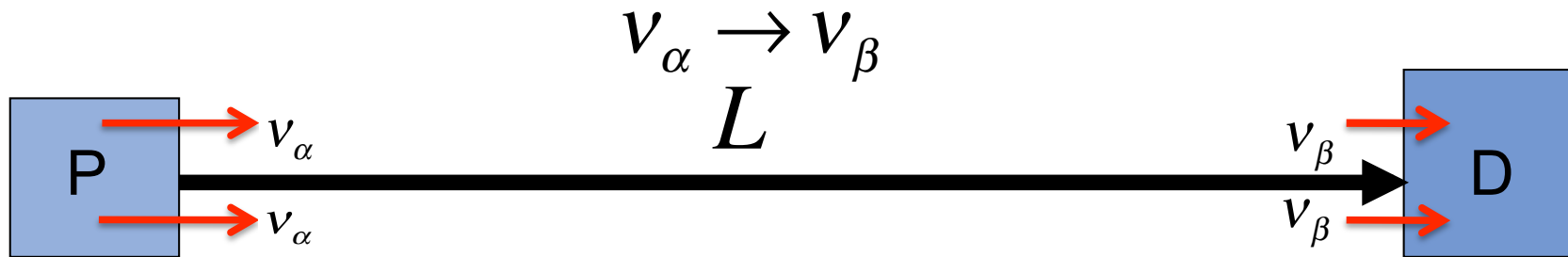
Antineutrinos always
with positive helicity

Neutrino propagation in the vacuum or in a matter – - neutral current

$$L_{NC} = \frac{e}{4 \sin \theta_W \cos \theta_W} \sum_{i=1,2,3} \bar{\nu}_i \gamma^\mu (1 - \gamma_5) \nu_i Z_\mu$$



Neutrinos oscillation rate in a detector is described by the factorized formula:



$$\Delta N_D(L, E) = j_\alpha(E) P_{\alpha \rightarrow \beta}(L, E) \sigma_\beta(E) N_D$$

Number of the β neutrinos with energy E , which reach detector in a unit time

Flux of the initial neutrinos of the α type

Probability of the α to β neutrino conversion

Detection cross section for β neutrinos

Number of active scattering centres in a detector

Oscillation rate is the same for Dirac and Majorana neutrinos

There are models which predict
NonStandard neutrino Interaction (NSI)
in the weak scale range, which can modify

- ❖ neutrino production process,
- ❖ oscillation inside matter, and
- ❖ detection process

- ❖ Models which try to resolve problem of neutrino mass
e.g. see-saw
- ❖ Charged Higgs,
- ❖ Right handed currents,
- ❖ Supersymmetric models.

LHC era

**What we should change to describe future
experiment with NSI of neutrinos??**

Oscillation beyond the SM

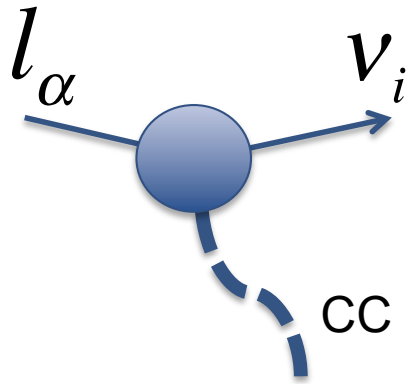
For the neutrino production (detection - for low energy)

dimensional six operators -
four-fermions effective Hamiltonian
e.g for muon decay (in neutrino factory)

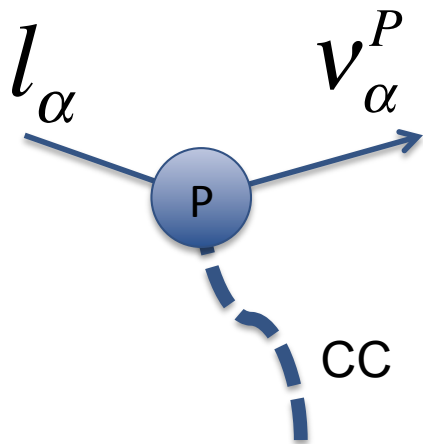
$$H = \frac{4G_F}{\sqrt{2}} \sum_{\substack{\delta=S,V,T \\ \varepsilon,\eta=L,R}} \sum_{i,k=1}^3 (g_{\varepsilon,\eta}^{\delta})_{i,k} (\bar{l}_{\alpha} \Gamma^{\delta} \nu_{i,\varepsilon}) (\bar{\nu}_{k,\eta} \Gamma_{\delta} l_{\beta}) + h.c.$$

First suggestion that a result of neutrino oscillation depends on three ingredients, on the production process, on the propagation inside matter and on their detection, appeared in 1995

Y. Grossman,
Phys.Lett.B359,141(1995)

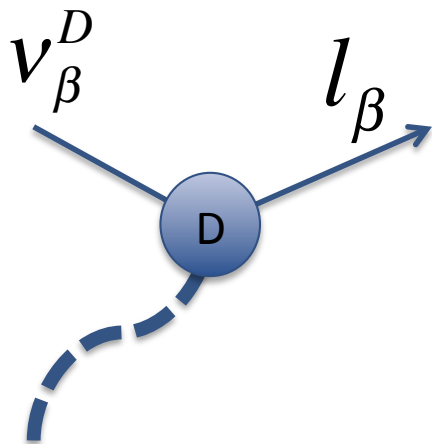


Neutrino masses are unknown- but are very small, experiments cannot observe neutrinos as mass eigenstates. But the mass basis $|\nu_i\rangle$ is well defined. Such states are process independent.



Neutrinos are produced by charged current interaction. This process defines neutrino flavour. Such states are process dependent:

$$|\nu_\alpha^P\rangle = \sum_i U_{\alpha,i}^P |\nu_i\rangle$$



Detection process measures different state – the detection flavour neutrino states:

$$|\nu_\beta^D\rangle = \sum_i U_{\beta,i}^D |\nu_i\rangle$$

Production and detection
flavour mixing matrices are
 constructed from the
 production and detection
 interaction Hamiltonians.

$$|U_{\alpha, i}^P|^2 \propto \left| \langle \nu_i; f_P | H^P | l_\alpha; i_P \rangle \right|^2$$

$$|U_{\beta, i}^D|^2 \propto \left| \langle l_\beta; f_D | H^D | \nu_i; i_D \rangle \right|^2$$

The probability of finding neutrinos in a $|\nu_\beta^D\rangle$ states in
 the original $|\nu_\alpha^D\rangle$ beam at the time t is given by

$$P_{\alpha \rightarrow \beta}(t) = \left| \langle \nu_\beta^P | e^{-iHt} | \nu_\alpha^D \rangle \right|^2$$

Two types of such approaches can be found in the literature:

NSI and internal or external wave packets: [Kayser\(81\)](#), [Giunti, Kim, Lee\(91\)](#), [Rich\(93\)](#).

Field theory and of mass shell neutrino
 propagation: [Grimus and Stckinger\(96\)](#), [MZ\(98\)](#),
[Cardall\(00\)](#), [Giunti\(02\)](#), [Beuthe\(03\)](#).

In these approaches, as in the Standard Model:

- 1) Production and detection states are pure Quantum Mechanical states

$$\left| \nu_{\alpha}^P \right\rangle \quad \left| \nu_{\beta}^D \right\rangle$$

- 2) It is possible to define flavour change probability

$$P_{\alpha \rightarrow \beta}(t) = \left| \left\langle \nu_{\beta}^P \left| e^{-iHt} \right| \nu_{\alpha}^D \right\rangle \right|^2$$

which factorize:

$$\Delta N_D(L \approx t, E) = j_{\alpha}(E) P_{\alpha \rightarrow \beta}(t) \sigma_{\beta}(E) N_D$$

In the proper approach - neutrino states are calculated in the standard way

State of the neutrinos produced in the process

$$A \rightarrow B + \bar{l}_\alpha + \nu_i$$

is described by the density matrix (if initial particle (**A**) is not polarized and polarizations of the final particles (**B, I**) are not measured):

$$\rho_{\lambda, i; \mu, k}^\alpha = \frac{1}{N_\alpha} \sum \int f_i^\alpha(\lambda_A; \lambda_B, \lambda_l, \lambda) f_k^{\alpha*}(\lambda_A; \lambda_B, \lambda_l, \mu)$$

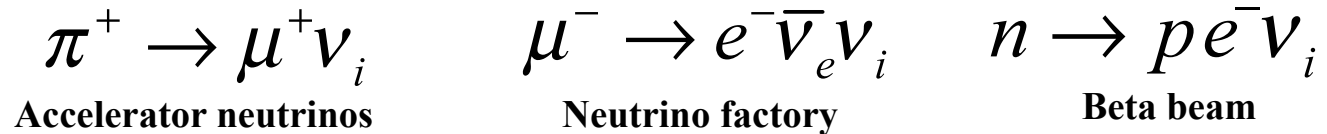
$$\text{Tr}(\rho^\alpha) = 1$$

where $f_i^\alpha(\lambda_A; \lambda_B, \lambda_l, \lambda)$ is the amplitude for the production process $A \rightarrow B + \bar{l}_\alpha + \nu_i$.

**We need the density matrix in the laboratory (detector) frame
= Lorentz boost + time evolution**

$$\rho_{CM}^{\alpha}$$

Calculated in the CM of decaying particle:



Lorentz boost,
Wigner spin
rotation

$$\rho_{LAB}^{\alpha}(L=0) \cong \rho_{CM}^{\alpha}$$

M.Ochman,R.Szafron and MZ,
arXiv:0707.4089

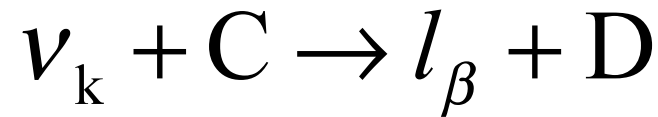
Neutrino propagation in the
vacuum or in a matter

$$\rho^{\alpha}(T=L) = e^{-iHT} \rho^{\alpha}(L=T=0) e^{iHT}$$

$$\rho_{LAB}^{\alpha}(L=T \neq 0)$$

H – vacuum or matter Hamiltonian

Any detection process:



A^β is the amplitude for the detection process

$$A^\beta(\Omega) \equiv A_k^\beta(\lambda_k, \lambda_C; \lambda_\beta, \lambda_D; \Omega)$$

$$\sigma_{\alpha \rightarrow \beta}(L, E) = \frac{1}{64\pi^2 s} \frac{p_f}{p_i} \frac{1}{2s_C + 1} \sum_{spins} \int d\Omega A^\beta(\Omega) \rho_{LAB}^\alpha(L=T) A^{\beta*}(\Omega)$$

Generally the $\sigma_{\alpha \rightarrow \beta}(E, L)$ cross section does not factorize

$$\sigma_{\alpha \rightarrow \beta}(E, L) \neq P_{\alpha \rightarrow \beta}(E, L) \sigma_\beta(E)$$

There is no factorization for the detection rate

$$N(E, L) = \rho_\alpha(E) \sigma_{\alpha \rightarrow \beta}(E, L) N_T$$

For Dirac neutrinos

For muon decay

In the SM

distinguishable

$$\bar{\nu}_e = \nu(\lambda = +1)$$

$$\nu_\mu = \nu(\lambda = -1)$$

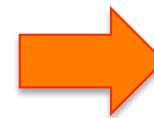
$$|\bar{\nu}_e\rangle = \sum_{i=1}^3 U_{ei} |\bar{\nu}_i\rangle$$

$$|\nu_\mu\rangle = \sum_{i=1}^3 U_{\mu i}^* |\nu_i\rangle$$

Pure
QM
STATES

Beyond the SM

$$\left. \begin{array}{l} \bar{\nu}_e \rightarrow \nu(\lambda = +1) \\ \nu_\mu \rightarrow \nu(\lambda = -1) \end{array} \right\}$$



Mixed QM STATES



**Density matrix
required**

For Majorana neutrinos

In the SM
or
beyond

$$\left. \begin{array}{l} \nu_e \rightarrow \nu(\lambda = +1) \\ \nu_\mu \rightarrow \nu(\lambda = -1) \end{array} \right\}$$



QM mixed STATE



$$\rho_{-1, i; -1, k}^\alpha, \rho_{+1, i; +1, k}^\alpha, \rho_{-1, i; +1, k}^\alpha, \rho_{+1, i; -1, k}^\alpha$$

Non-standard description

If only positive(negative) helicity neutrinos(antineutrinos) are produced -- Theorem:

The necessary and sufficient condition for pure initial state of produced neutrinos with negative helicities is the factorization for spin and mass production amplitudes

$$f_i^\alpha(\lambda_A; \lambda_B, \lambda_l, \lambda = -1) = g^\alpha(\lambda_A; \lambda_B, \lambda_l, \lambda = -1) * h_i^\alpha \equiv g_\mu^\alpha h_i^\alpha$$

If we introduce the shortcut notation

$$\mu = (\lambda_A; \lambda_B, \lambda_l)$$

Then:

$$f_i^\alpha(\mu) = g^\alpha(\mu, \lambda = -1) * h_i^\alpha \equiv g_\mu^\alpha h_i^\alpha$$

Then the density matrix is given by:

$$\rho^\alpha(i, \lambda = -1; k, \mu = -1) \equiv \rho_{i,k}^\alpha = \frac{h_i^\alpha h_k^{\beta*}}{\sum_i |h_i^\alpha|^2} = \chi_i^\alpha \chi_k^{\alpha*}$$

where


$$\chi_i^\alpha = \frac{h_i^\alpha}{\sqrt{\sum_i |h_i^\alpha|^2}}$$

which is equivalent to the pure QM state:

$$|v_\alpha\rangle = \sum_i \chi_i^\alpha |v_i\rangle$$

$$|v_\alpha\rangle = \sum_i U_{\alpha i}^* |v_i\rangle$$

$$\langle v_\alpha | v_\beta \rangle = \sum_i \chi_i^\alpha \chi_i^{\beta*} \neq \delta_{\alpha\beta} (= 1 \text{ for } \alpha = \beta)$$

 **vSM**

$$\chi_i^\alpha = U_{\alpha i}^*$$

which are normalized but not necessarily orthogonal

Factorization for the final oscillation rate

$$\sigma_{\alpha \rightarrow \beta}(L, E) =$$

$$\frac{1}{64\pi^2 s} \frac{p_f}{p_i} \frac{1}{2s_f + 1} \sum_{i,k,\lambda,\lambda_C,\lambda_l,\lambda_D} \int d\Omega A_i^\beta(\lambda, \lambda_C, \lambda_l, \lambda_D; \Omega) \rho^\alpha(i, \lambda; k, \lambda; L, E) A_k^{\beta*}(\lambda, \lambda_C, \lambda_l, \lambda_D; \Omega)$$

The density matrix after oscillation

$$\rho^\alpha(i, \lambda = -1; k, \mu = -1; L, E) = \rho_{i,k}^\alpha(L, E) e^{-\frac{\delta m_{i,k}^2 L}{2E}}$$

If the detection amplitudes factorize

$$A_i^\beta(\lambda = -1, \lambda_C, \lambda_l, \lambda_D; \Omega) = e_\eta^\beta(\theta, \varphi) * k_i^\beta$$

Then the final cross section factorize

$$\sigma_{\alpha \rightarrow \beta}(L, E) =$$

$$\begin{aligned} \frac{1}{32\pi s} \frac{p_f}{p_i} \frac{1}{2s_f + 1} \sum_{i,k,\eta} \int d\Omega (e_\eta^\beta * k_i^\beta) (\rho_{i,k}^\alpha e^{-\frac{\delta m_{i,k}^2 L}{2E}}) (e_\eta^{\beta*} * k_k^{\beta*}) = \\ = P_{\alpha \rightarrow \beta}(L, E) \sigma_\beta(E) \end{aligned}$$

If there is factorization for the initial and final states?

$$\sigma_{\alpha \rightarrow \beta}(L, E) = \frac{1}{64\pi^2 s} \frac{p_f}{p_i} \frac{1}{2s_C + 1} \sum_{spins} \int d\Omega A^\beta(\Omega) \rho_{LAB}^\alpha(L=T) A^{\beta*}(\Omega)$$

The density matrix after oscillation

$$\rho^\alpha(i, \lambda = -1; k, \mu = -1; L, E) = \chi_i^\alpha \chi_k^{\alpha*} e^{-\frac{\delta m_{i,k}^2}{2E} L}$$

If the detection amplitudes factorize

$$A_i^\beta(\lambda = -1, \lambda_C, \lambda_l, \lambda_D; \Omega) = e_\eta^\beta(\theta, \varphi) * k_i^\beta$$

Then the final cross section factorize

$$\begin{aligned} \sigma_{\alpha \rightarrow \beta}(L, E) &= \\ \frac{1}{64\pi^2 s} \frac{p_f}{p_i} \frac{1}{2s_f + 1} \sum_{i,k,\eta} \int d\Omega (e_\eta^\beta * k_i^\beta) (\chi_i^\alpha \chi_k^{\alpha*} e^{-\frac{\delta m_{i,k}^2}{2E} L}) (e_\eta^{\beta*} * k_k^{\beta*}) &= \\ &= P_{\alpha \rightarrow \beta}(L, E) \sigma_\beta(E) \end{aligned}$$

The oscillation probability is given by:

$$P_{\alpha \rightarrow \beta}(L, E) = \sum_{i, k} k_i^\beta \chi_i^\alpha \chi_k^{\alpha*} k_k^{\beta*} e^{-\frac{\delta m_{i, k}^2}{2E} L}$$

And the final detection cross section:

$$\sigma_\beta(E) = \frac{1}{64\pi^2 s} \frac{p_f}{p_i} \frac{1}{2s_f + 1} \sum_\eta \int d\Omega \left| e_\eta^\beta(\theta, \varphi) \right|^2$$

The sum over all final flavours

$$\sum_{\beta=e, \mu, \tau} P_{\alpha \rightarrow \beta}(L, E) = \sum_{i, k} \sum_{\beta=e, \mu, \tau} (k_i^\beta k_k^{\beta*}) \rho_{i, k}^\alpha e^{-\frac{\delta m_{i, k}^2}{2E} L} \stackrel{\text{if}}{=} \sum_i \rho_{i, i}^\alpha = 1$$

The probability is conserved if the final states are orthogonal

$$\sum_{\beta=e, \mu, \tau} (k_i^\beta k_k^{\beta*}) = \delta_{i, k}$$

☞ If dominant and subdominant neutrino helicity states are produced and detected then the description of neutrino oscillation is not standard.

☞ If only dominant neutrino helicity states are produced and detected the standard description of the neutrino oscillation is recovered then and only then if the production and detection amplitudes factorizes for spin and mass parts

Pure or mixed initial neutrino state \longleftrightarrow $\begin{cases} \text{is } Tr(\rho^2) < 1 \\ \text{or } Tr(\rho^2) = 1 \end{cases}$

©If the left-handed V_L and right-handed V_R chiral neutrino operators are present in NI – both type of neutrino helicities are produced -- **mixed states.**

©If only one left handed operators V_L describes the neutrino interaction then neutrino state depends on the number and structure of the helicity production amplitudes:

- If only one helicity amplitude describes production process, then, independently of the NI, neutrino **state is pure**, family lepton number can not be conserved,
- If there are more then one helicity amplitude, but all have the same structure -- **state is pure**,
- If there are more then one helicity amplitudes, and at least two of them have different structure -- **state is mixed.**

Dirac and Majorana neutrinos in oscillation experiments

- Production process (charge currents are responsible)
 - ❖ If only one neutrino appears in a production process (**as in pion or in beta decays**) – production process does not distinguish both types of neutrinos.
 - ❖ If two neutrinos (neutrinos + antineutrinos, as in the muon decay) are produced, the interference terms in the spin amplitudes, which are present for the Majorana neutrinos, and do not occur for Dirac neutrinos, can distinguish between two types of neutrinos in a production process.
- Propagation in matter distinguishes both types of neutrinos (neutral currents are crucial).

4) Conclusions

- 1) neutrino production states represented by density matrix can be
pure or mixed
depending on the production mechanism,
- 2) final neutrino detection rates generally do not factorize,
- 4) Dirac and Majorana neutrinos oscillate in different way,
- 5) coherent and incoherent oscillation can be defined.

Present bounds on NI parameters give possibility that future precise experiments can see some effects