



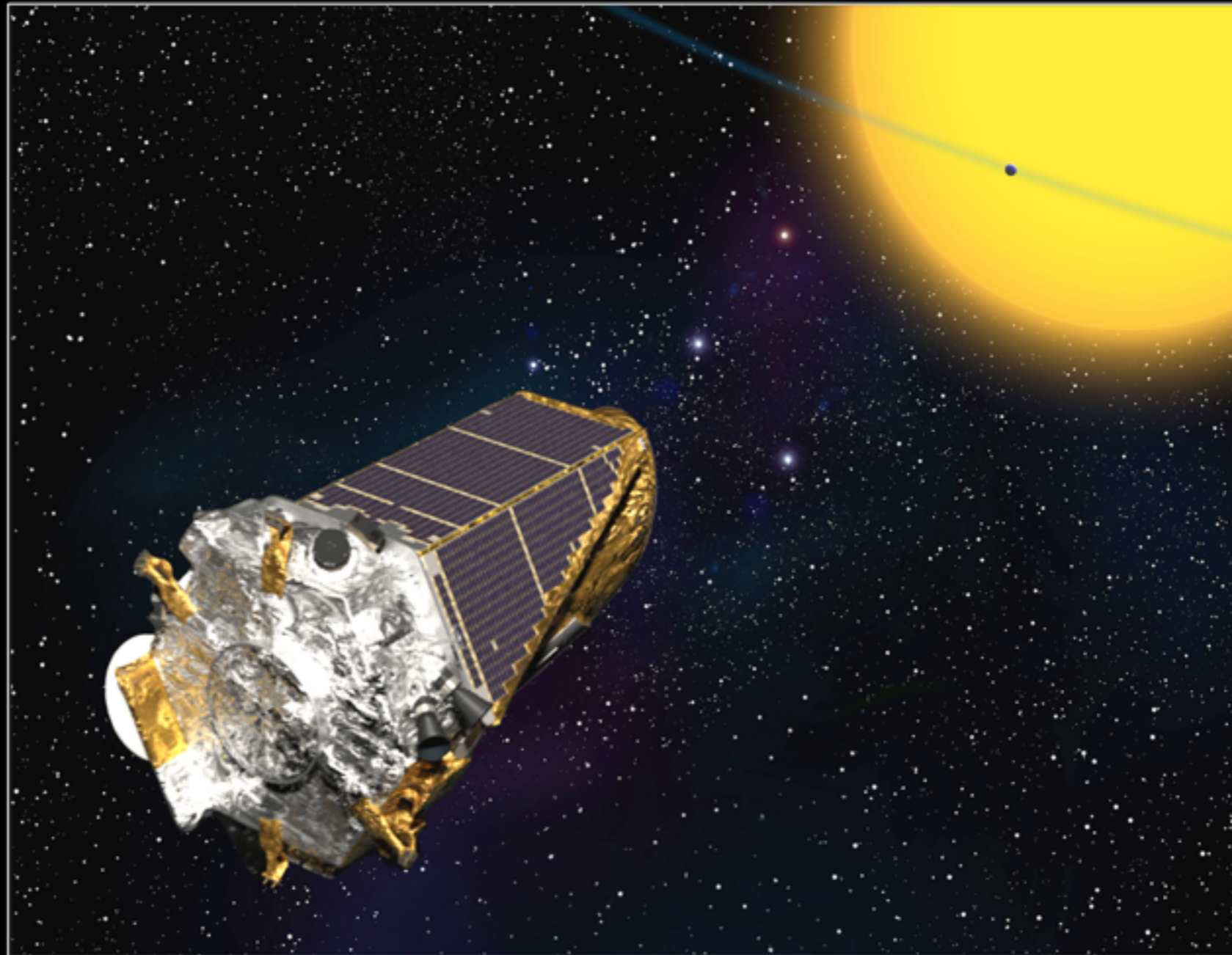
**Catching Shadows:
Kepler's Year-Two Transit Census**

Kepler



**Natalie Batalha
San Jose State University**

Are potentially habitable, earth-size planets abundant in our galaxy?



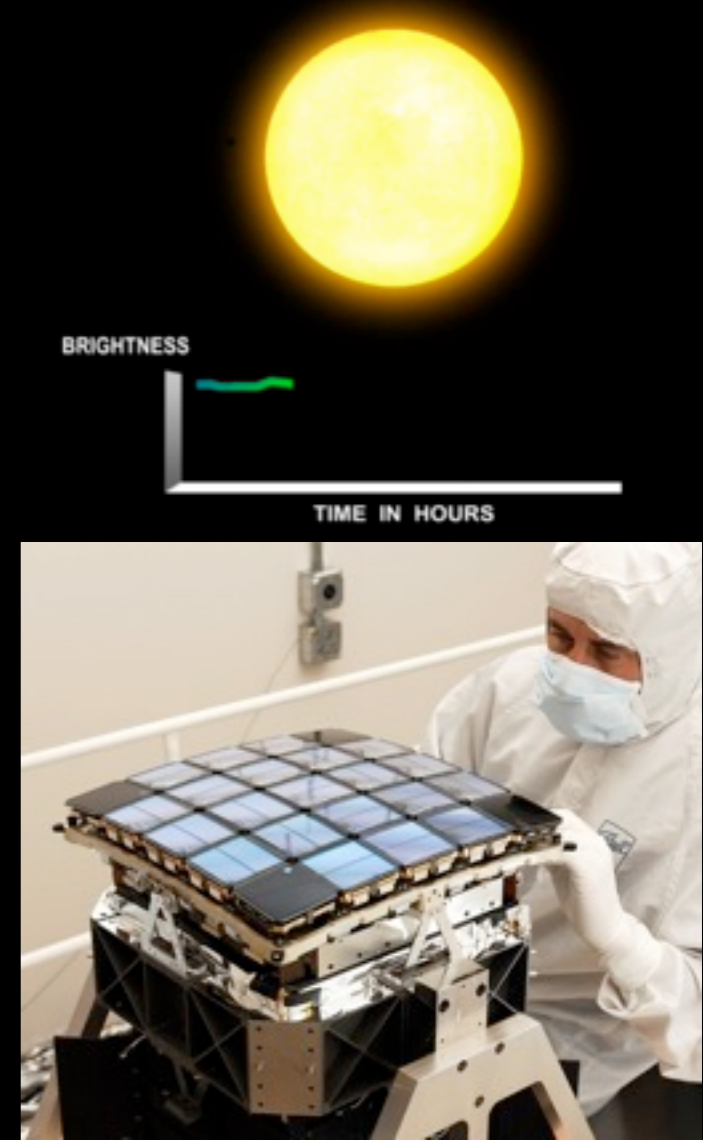
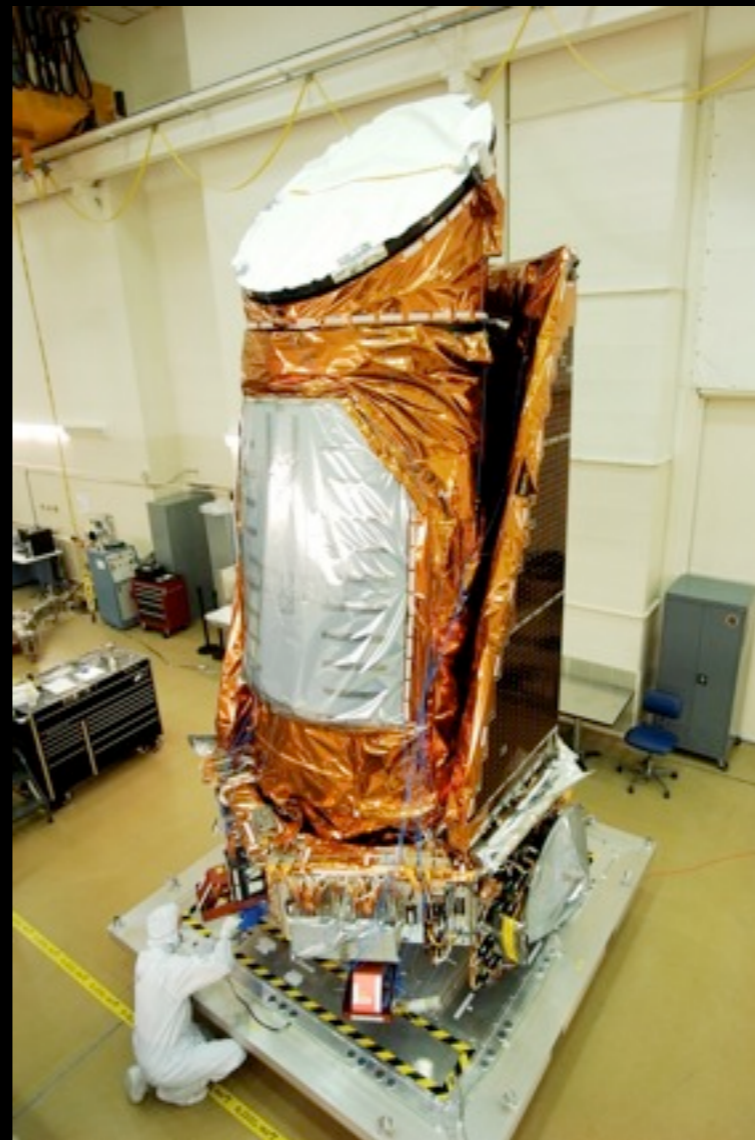
Design



KEPLER: A Wide Field-of-View Photometer that monitors $\geq 100,000$ Stars for 3.5 yrs with precision to find Earth-size planets in the Habitable Zone

Transit Detection using:

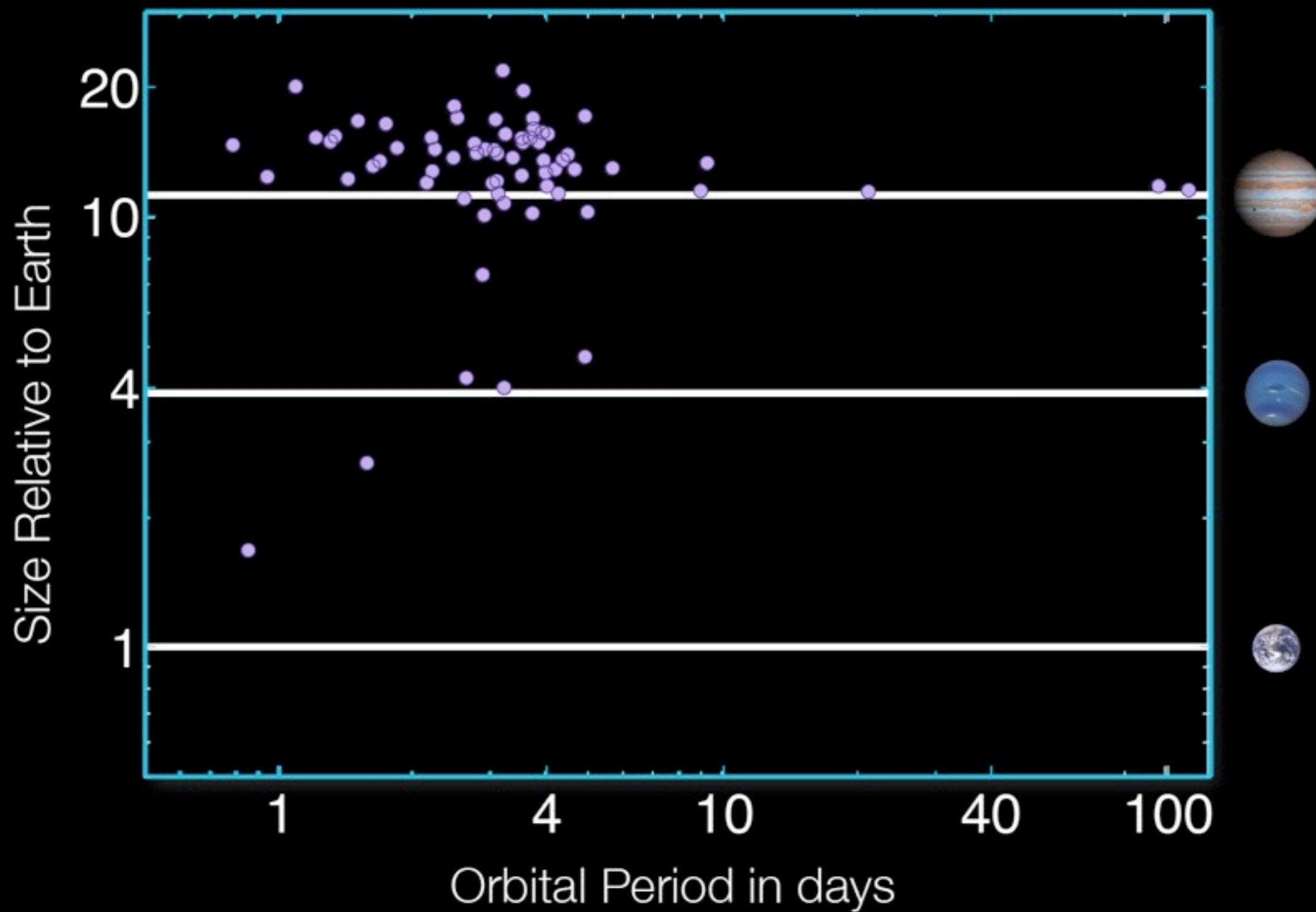
- 0.95 meter aperture
- Wide FOV: 100 sq deg
- 42 CCDs
- ≥ 3.5 years
- Fixed pointing
- Heliocentric orbit
- 170k targets: 30 min
- 512 targets: 1 min



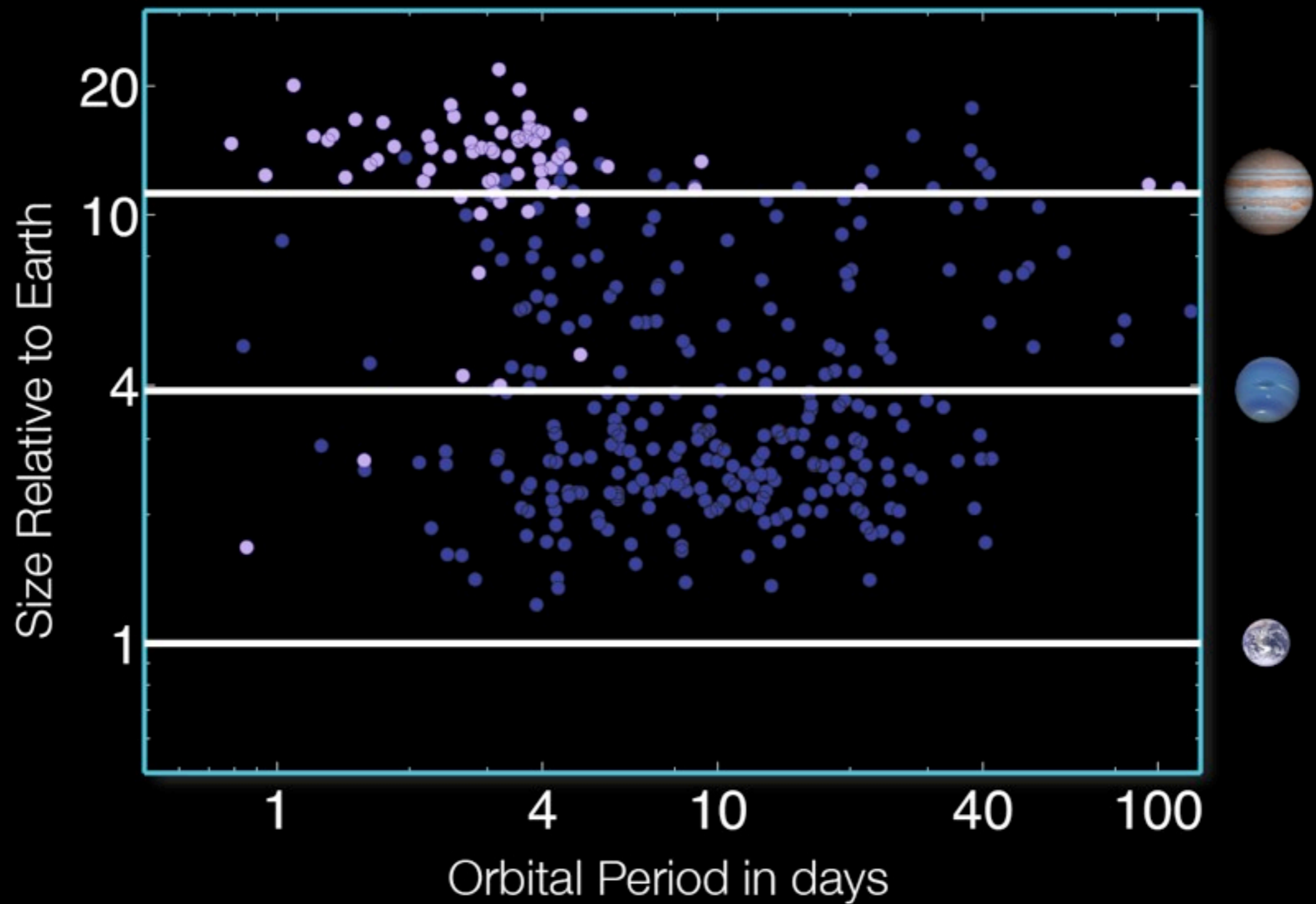
March 6, 2009



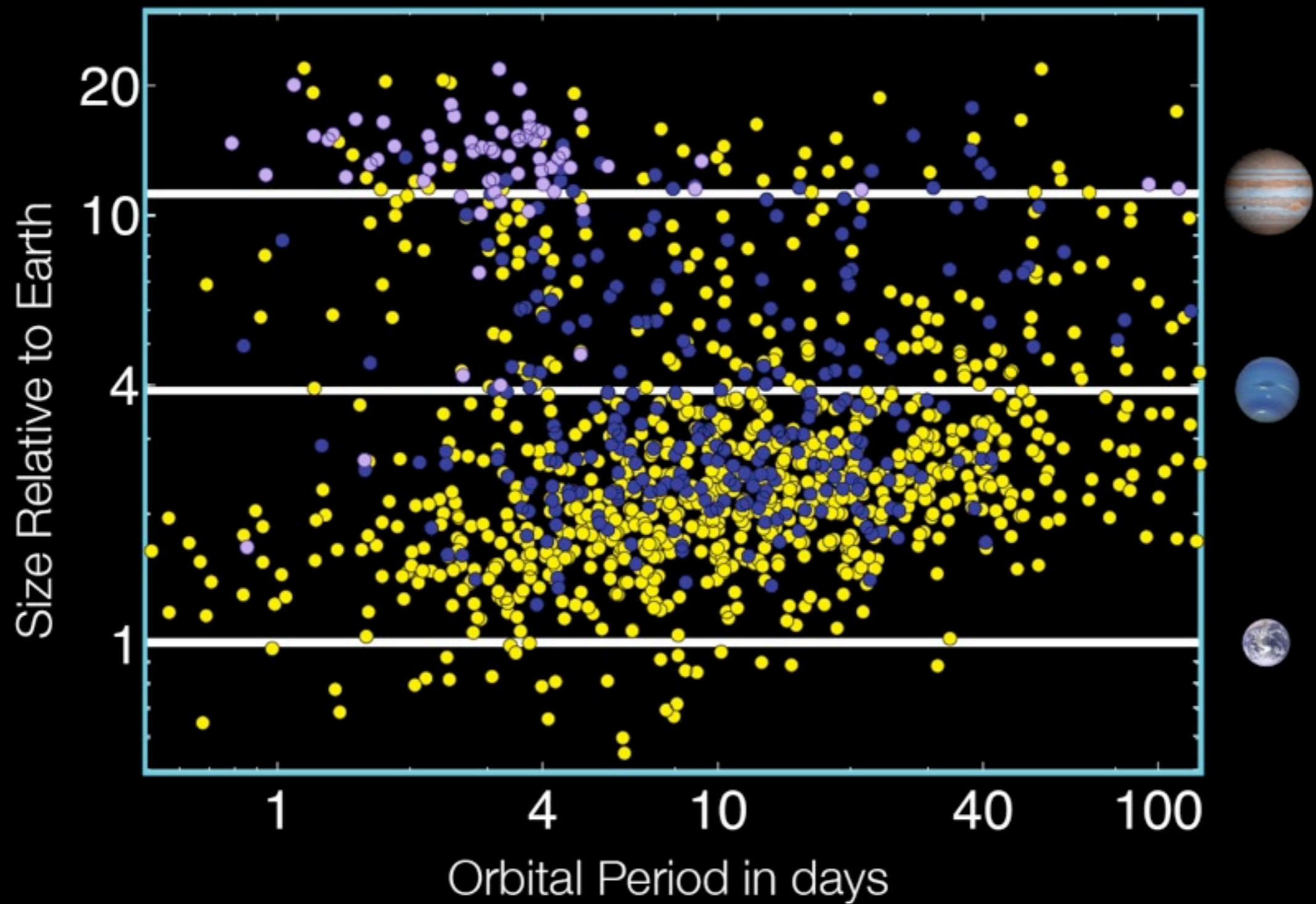
Pre-Kepler Transiting Planets - 2009



Kepler Candidates as of June 2010



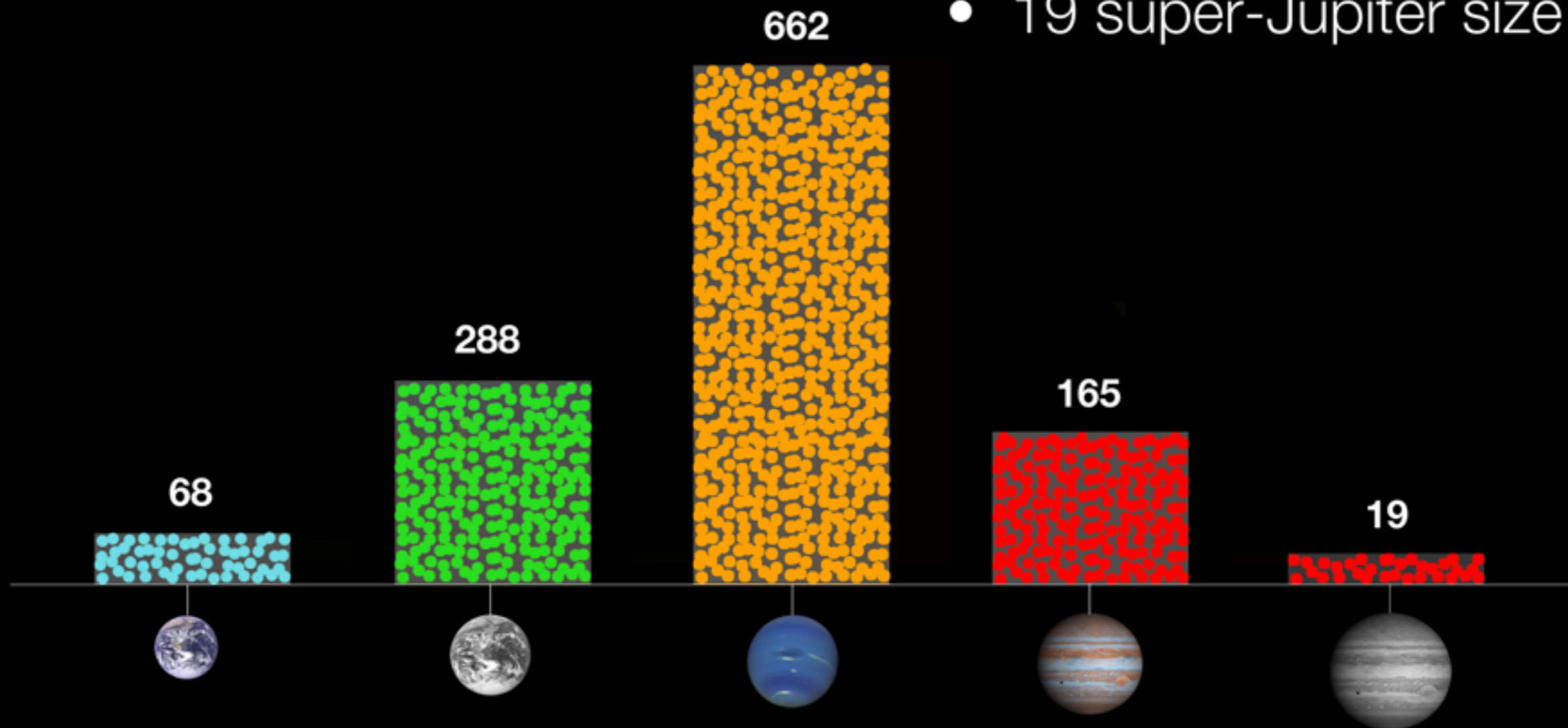
Kepler Candidates as of February 1, 2011



Borucki et al. 2011, arXiv: 1102.0541

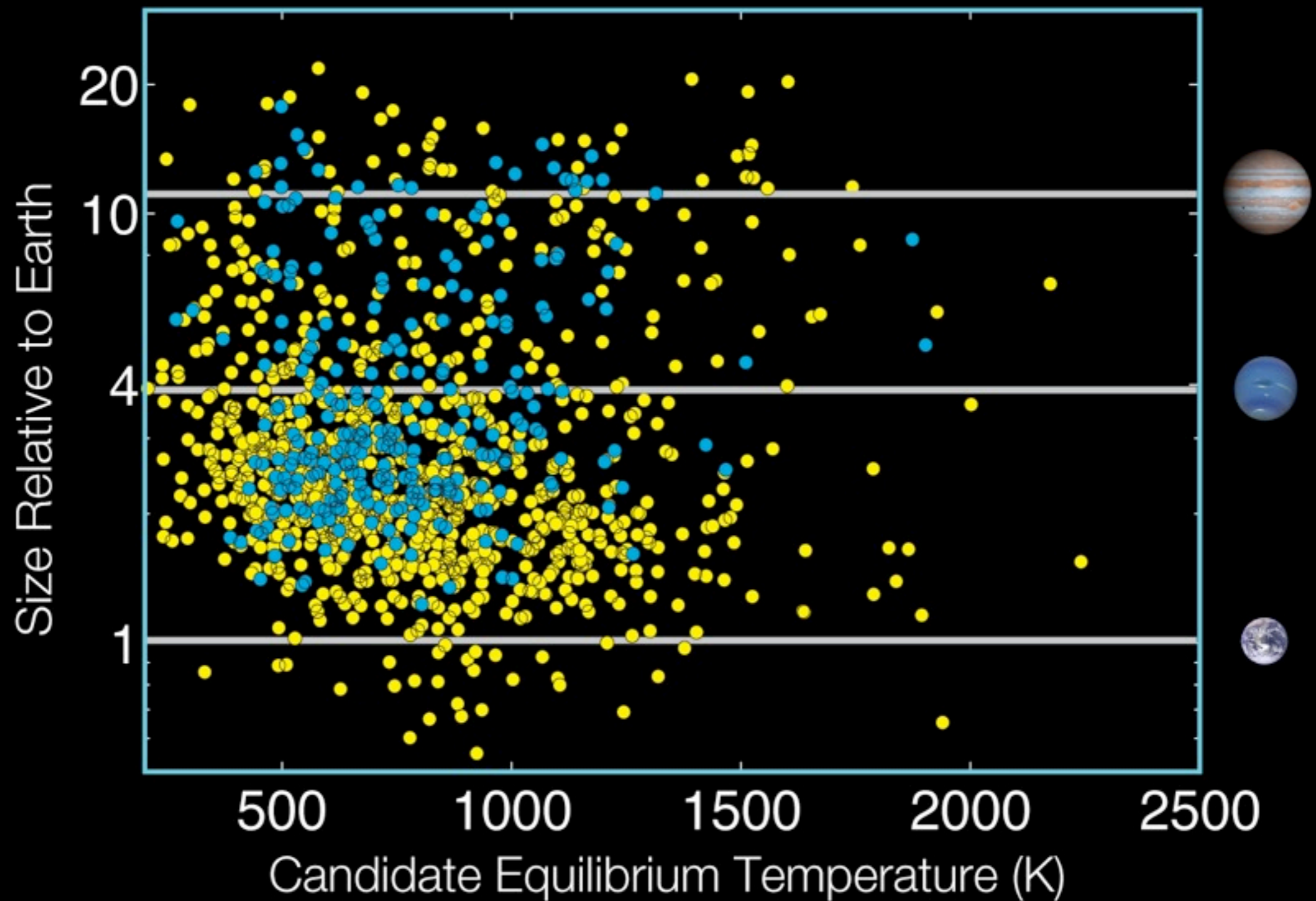
Numbers of Planet Candidates

- 68 Earth-size
- 288 super-Earth size
- 662 Neptune size
- 165 Jupiter size
- 19 super-Jupiter size



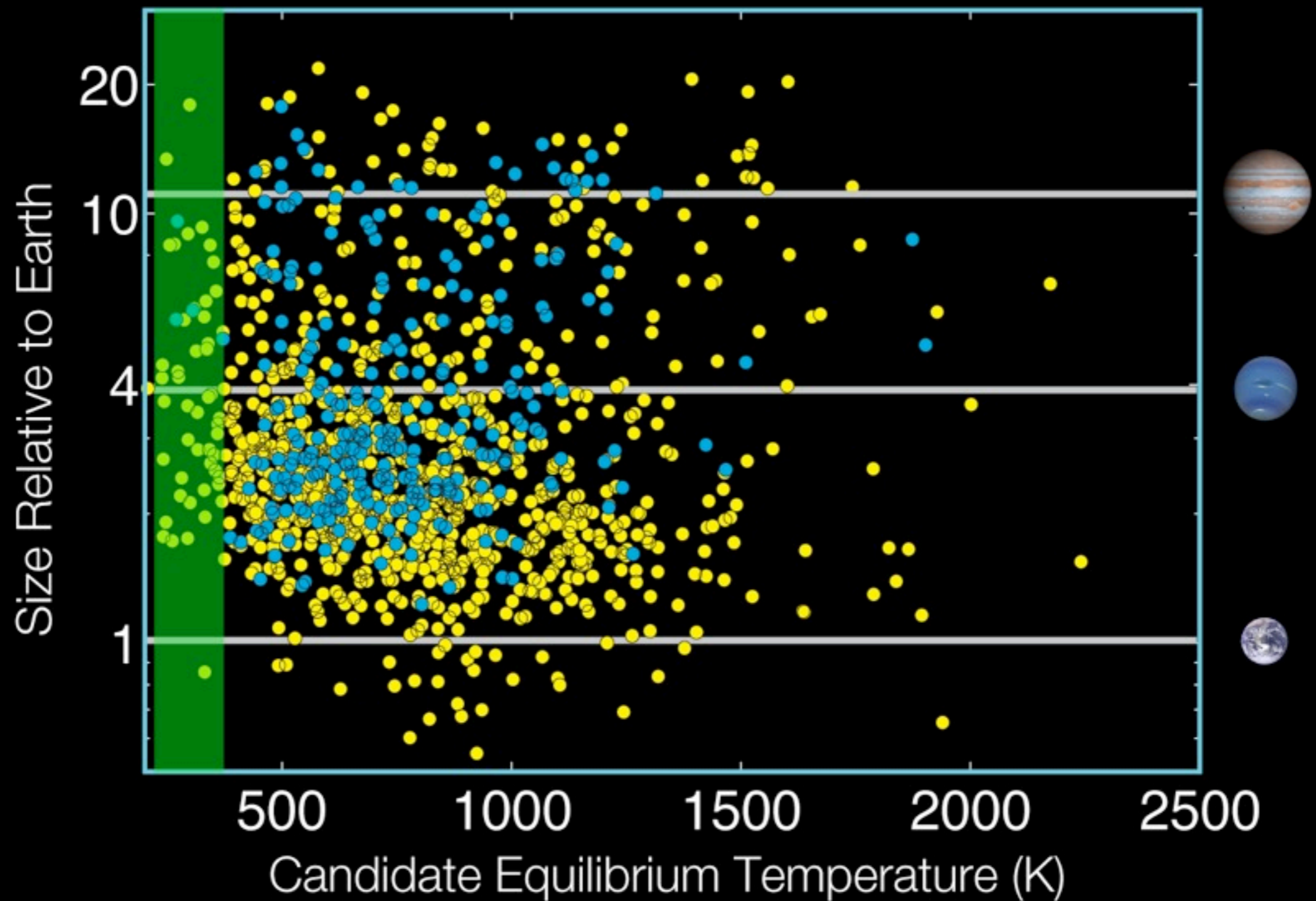
Borucki et al. 2011, arXiv: 1102.0541

Kepler Candidates as of February 1, 2011



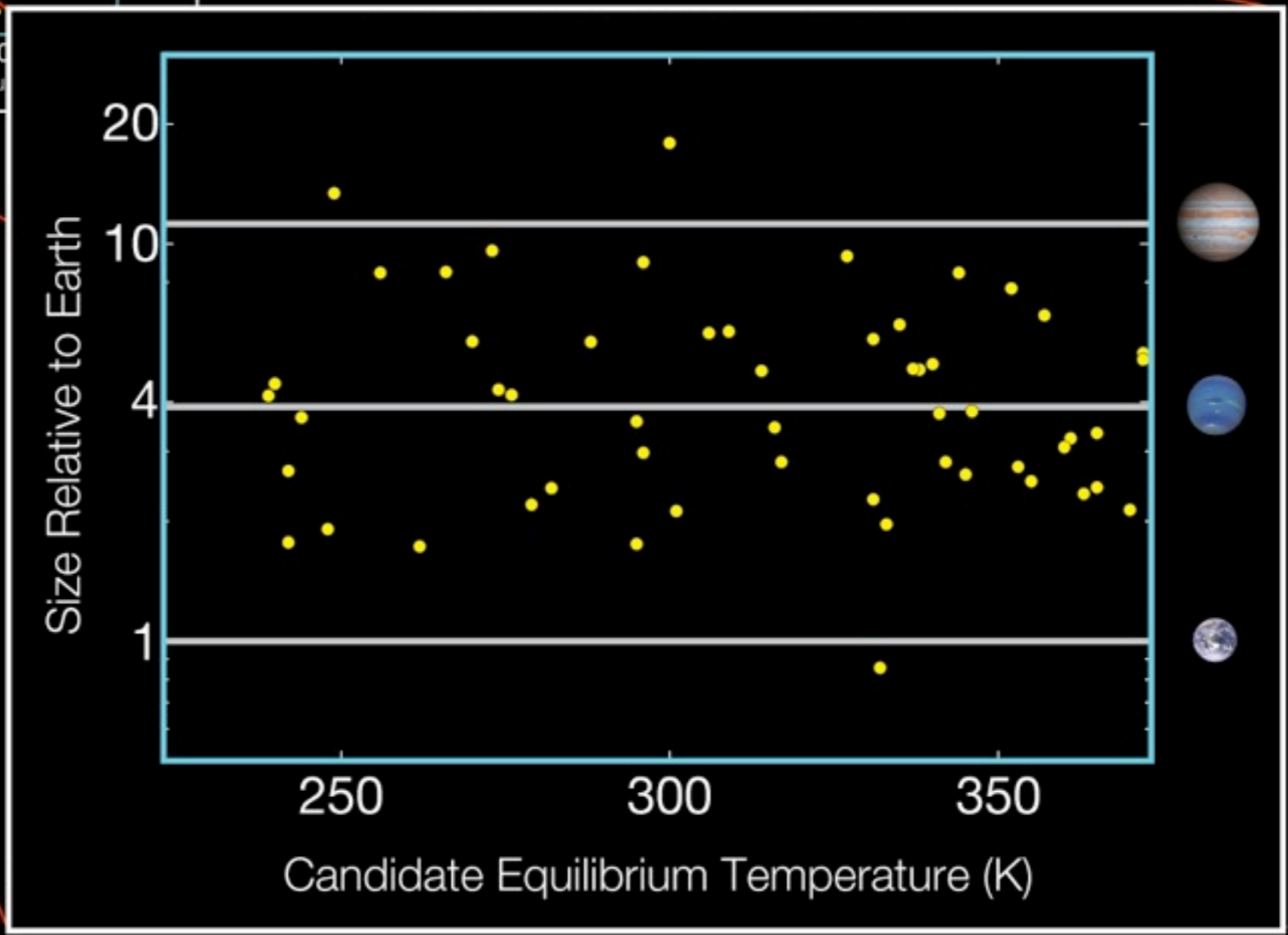
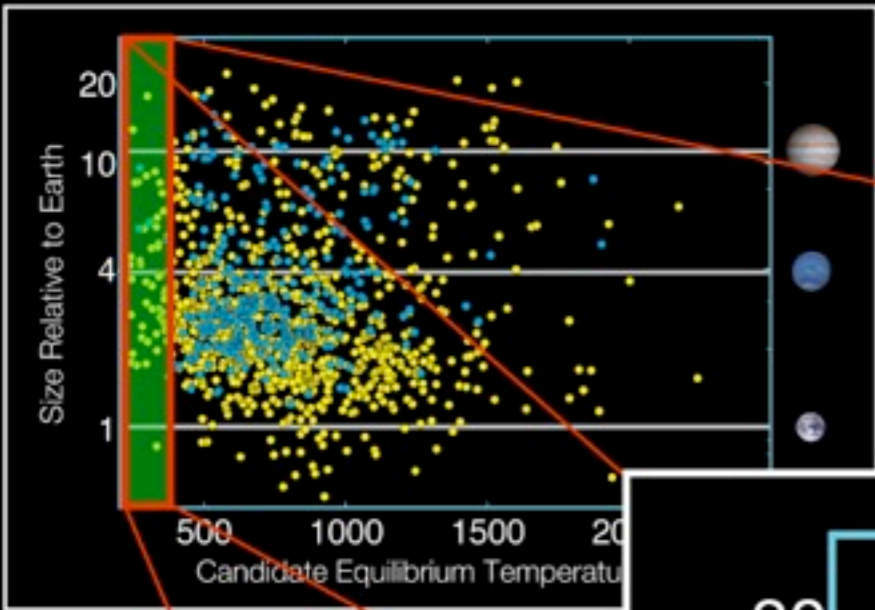
Borucki et al. 2011, arXiv: 1102.0541

Kepler Candidates as of February 1, 2011



Borucki et al. 2011, arXiv: 1102.0541

Kepler Planet Candidates In the Habitable Zone



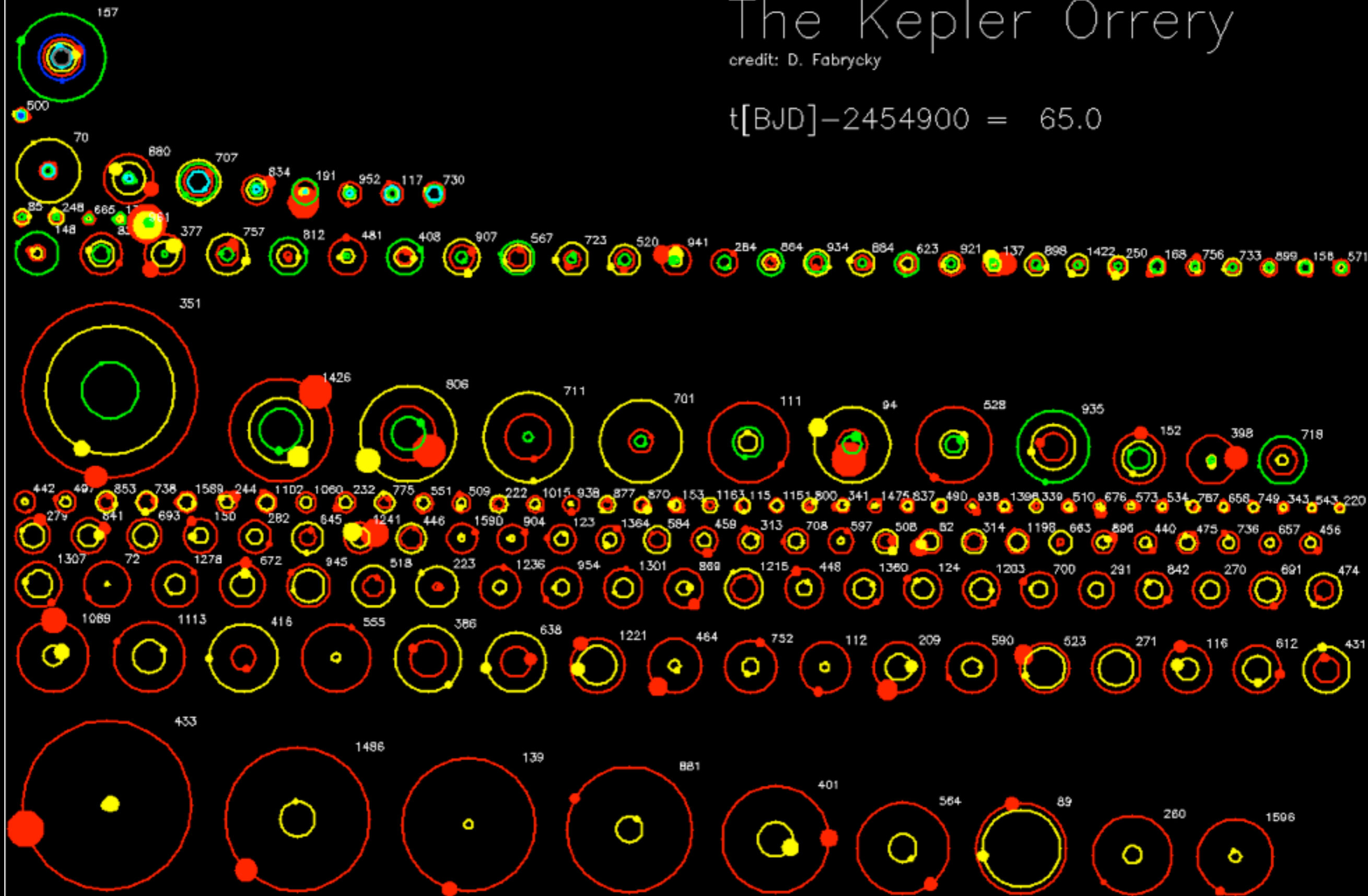
170 Stars (17%) with Multiple Transit Signatures

170 Stars (17%) with Multiple Transit Signatures

The Kepler Orrery

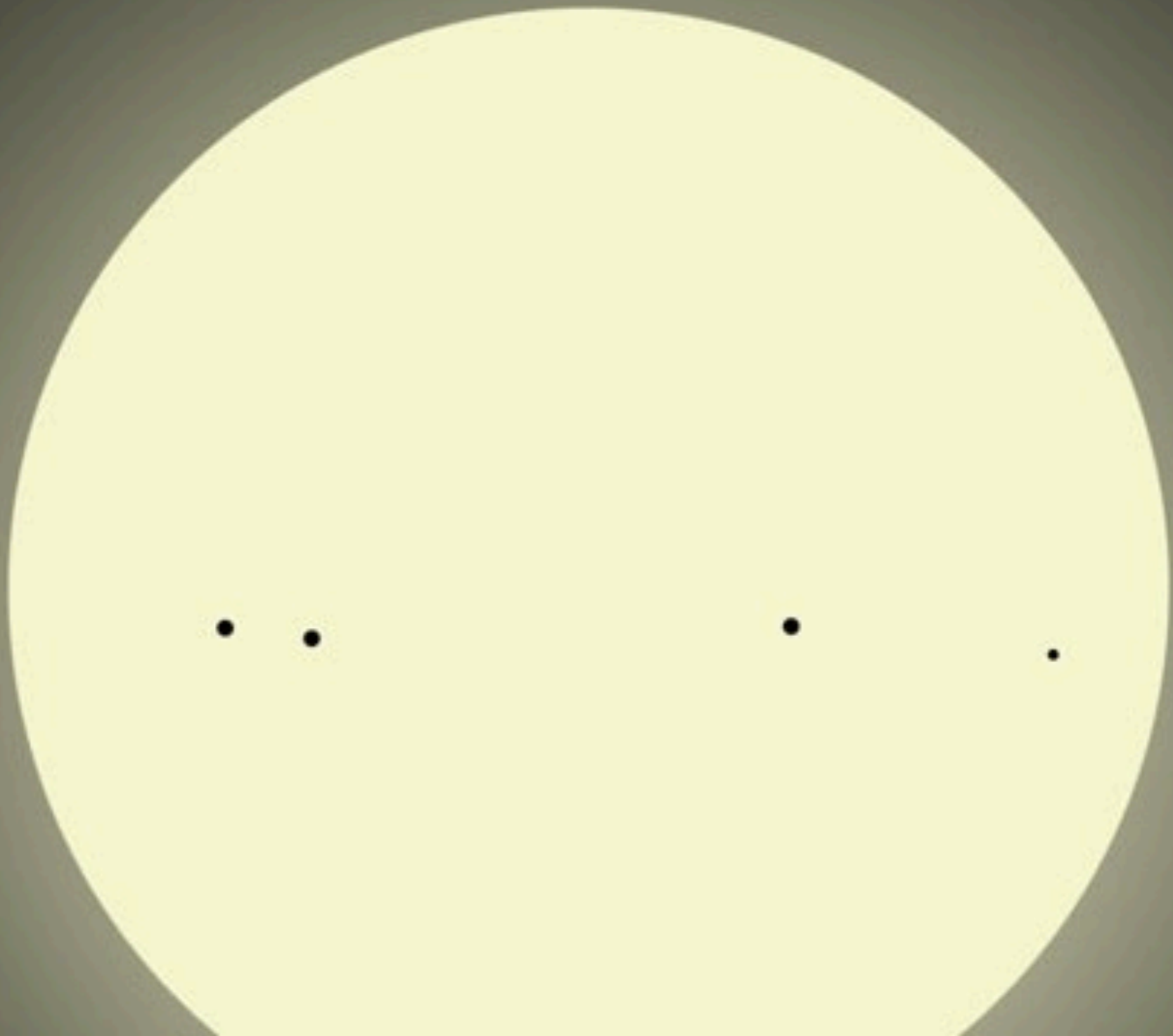
credit: D. Fabrycky

$t[\text{BJD}] - 2454900 = 65.0$



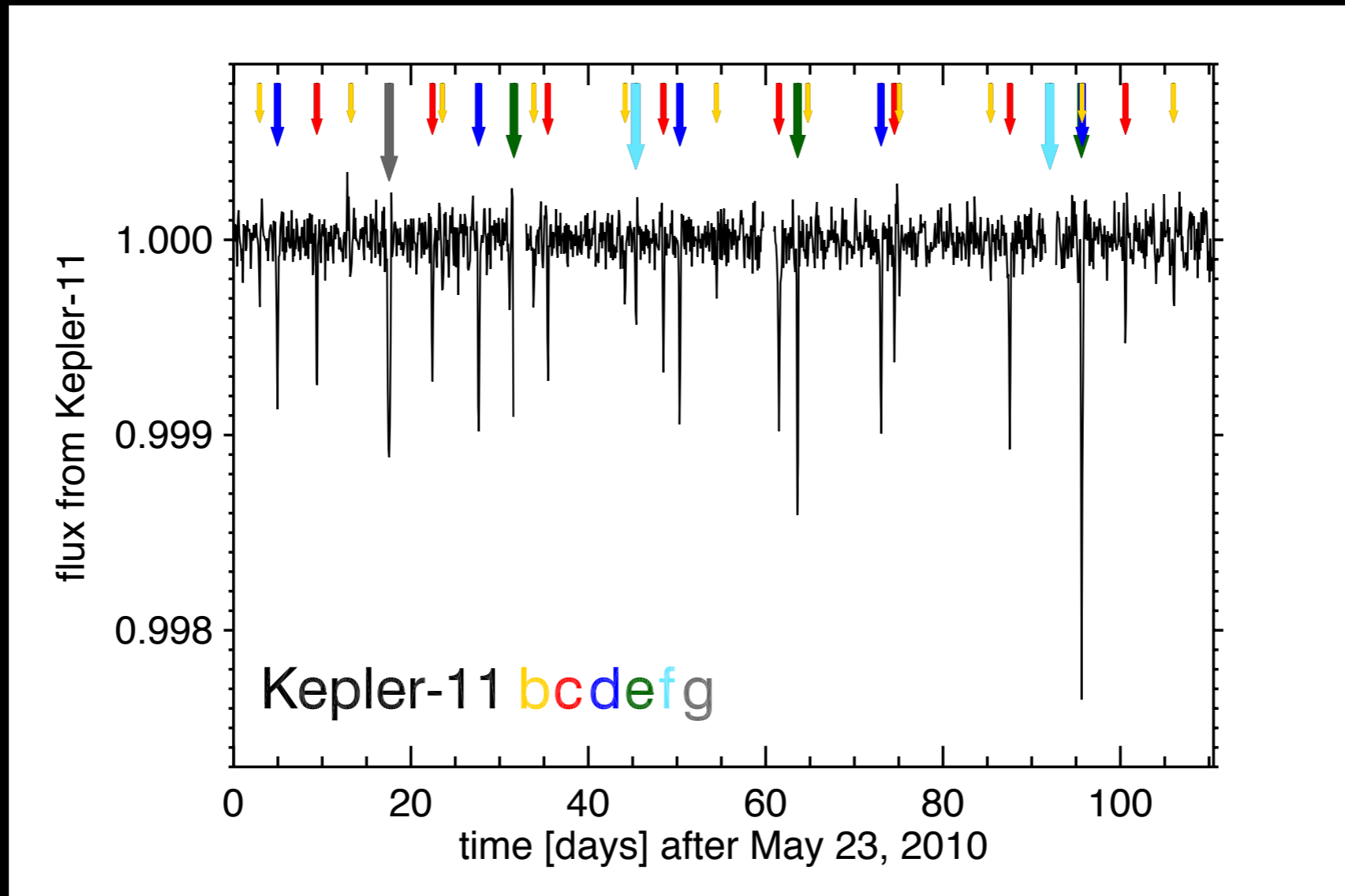
170 Stars (17%) with Multiple Transit Signatures





Lissauer et al. 2011, Nature, 470, 53

Kepler-11: Light Curve



Periods: 10.3, 13.02, 22.69, 32.00, 46.69, 118.38 days

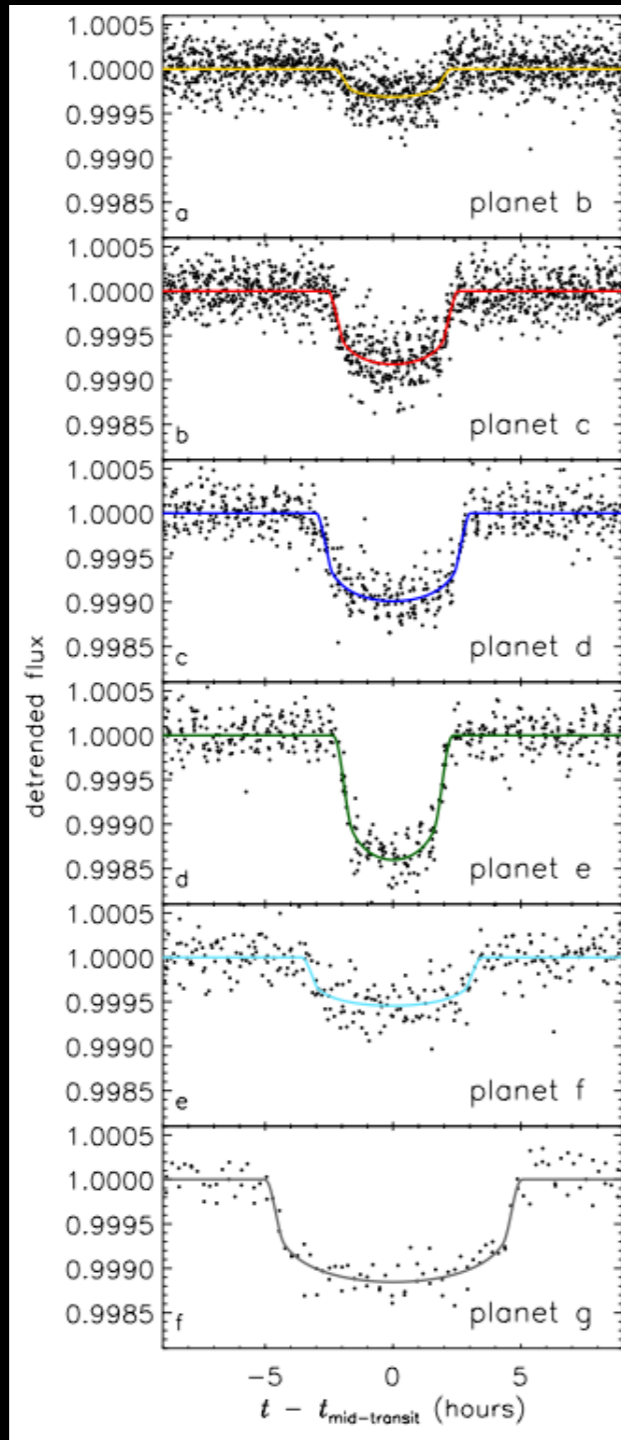
Radii: 1.97, 3.15, 3.43, 4.52, 2.61, 3.66

Masses: 4.3, 13.5, 6.1, 8.4, 2.3, ?

Lissauer et al. 2011, Nature, 470, 53

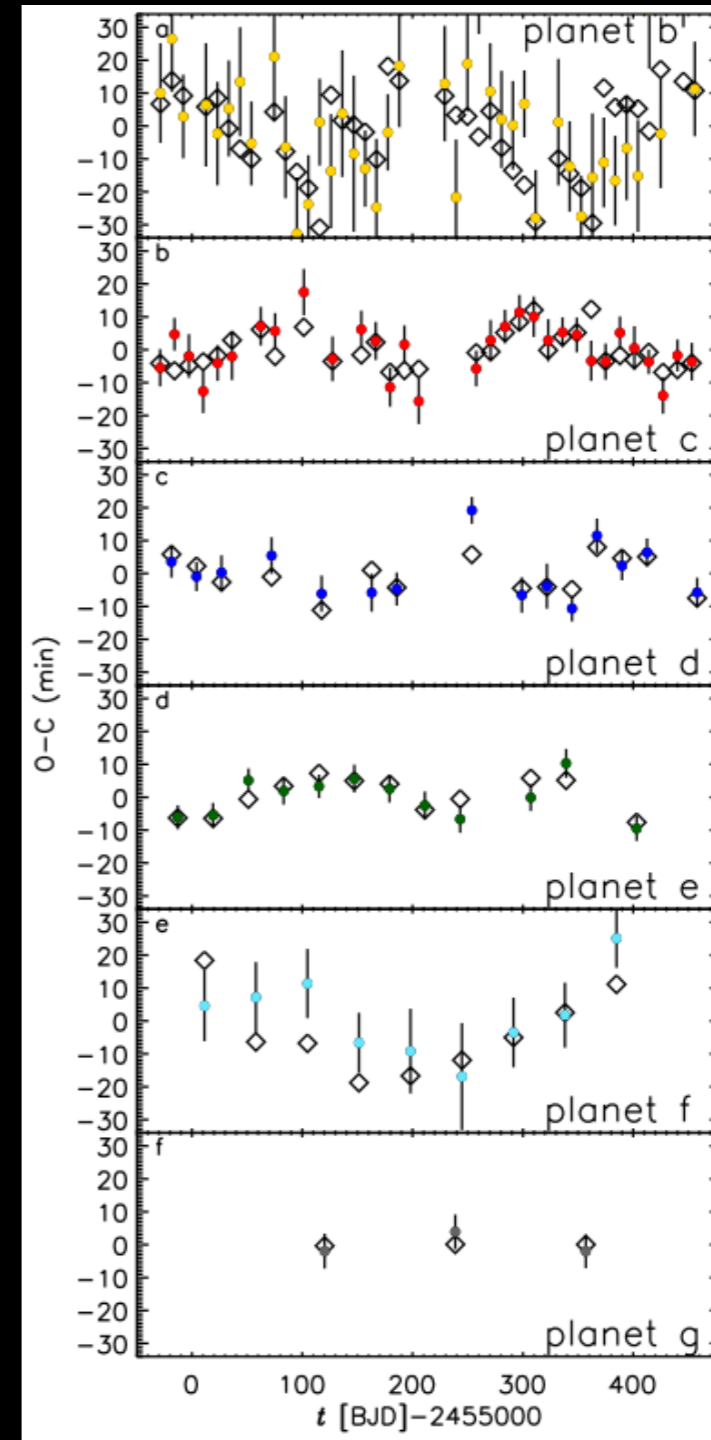
American Physical Society, May 1, 2011, Anaheim, CA

Transit Timing Variations



← Transits

TTV's →



Lissauer et al. 2011, Nature, 470, 53

American Physical Society, May 1, 2011, Anaheim, CA





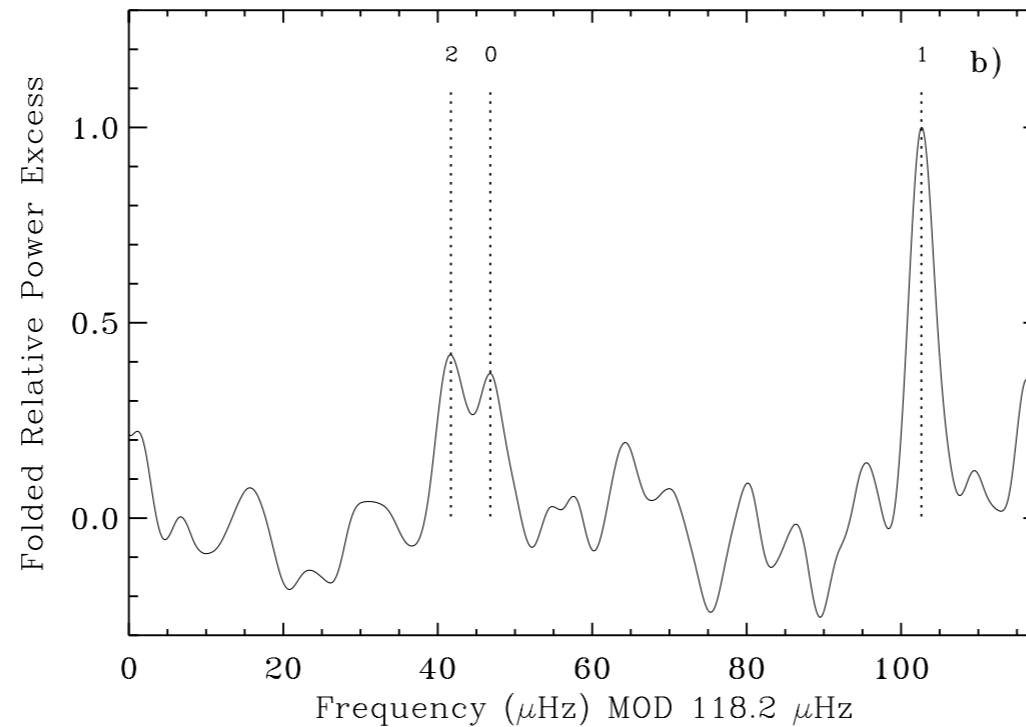
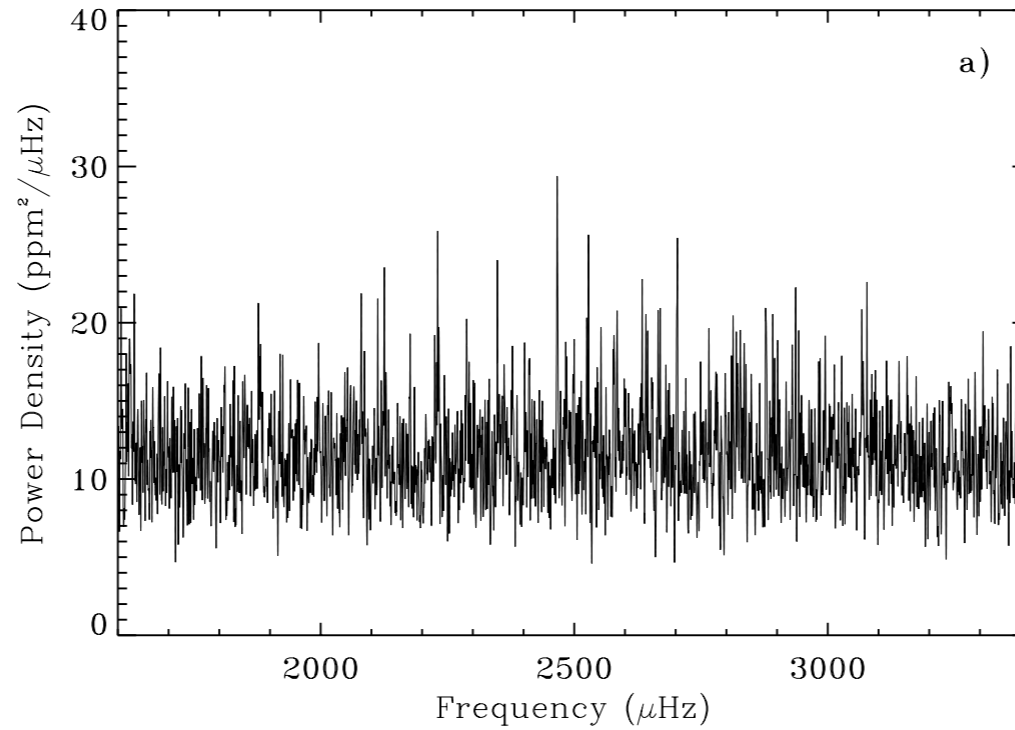
Stellar Properties

Kepler-10

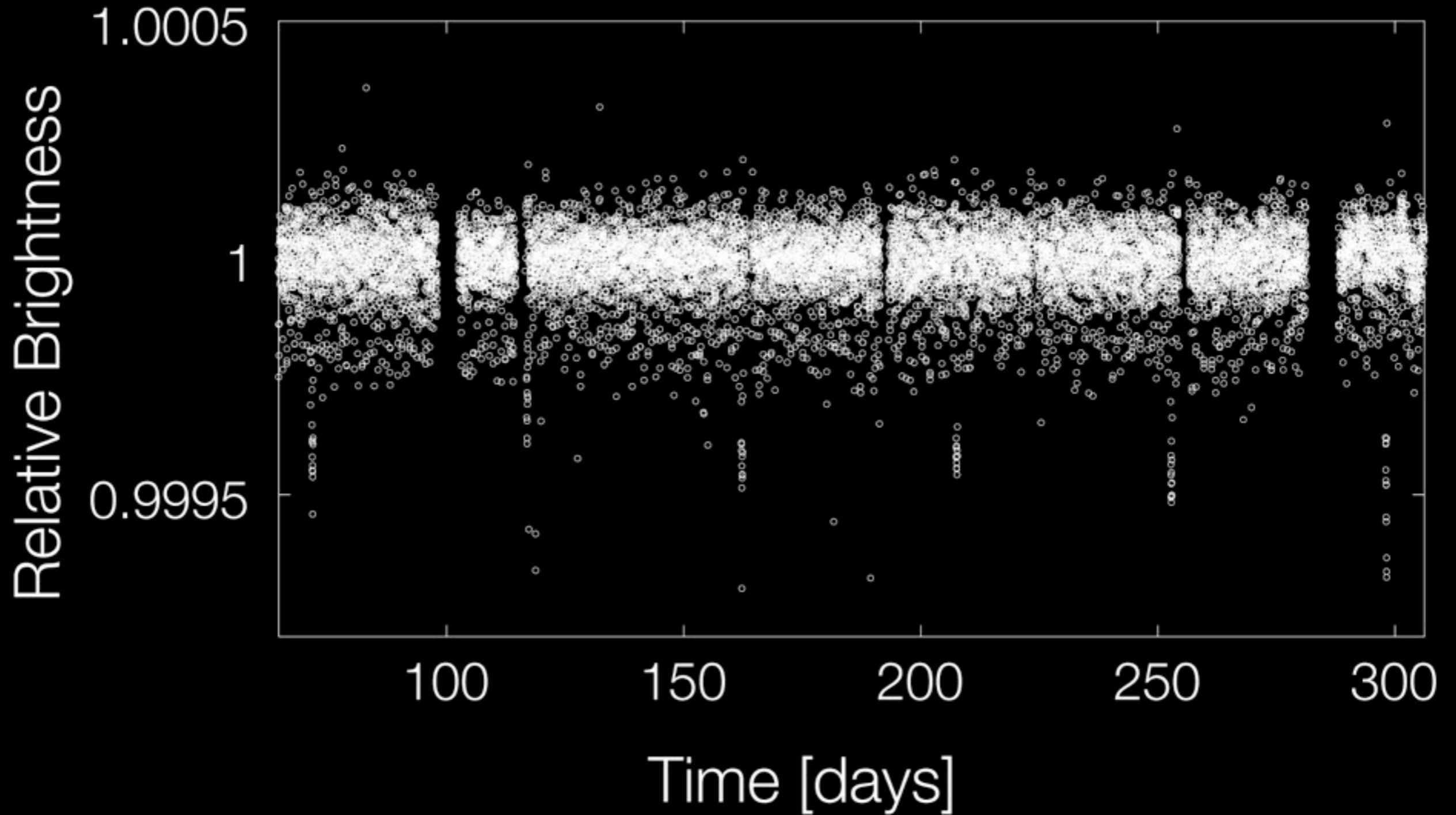
- G4V
- Mass = 0.90 \odot
- Radius = 1.06 R_{\odot}
- Age > 8 Gyr
- Distance = 560 Light-years



Asteroseismology



Kepler-10b Light Curve

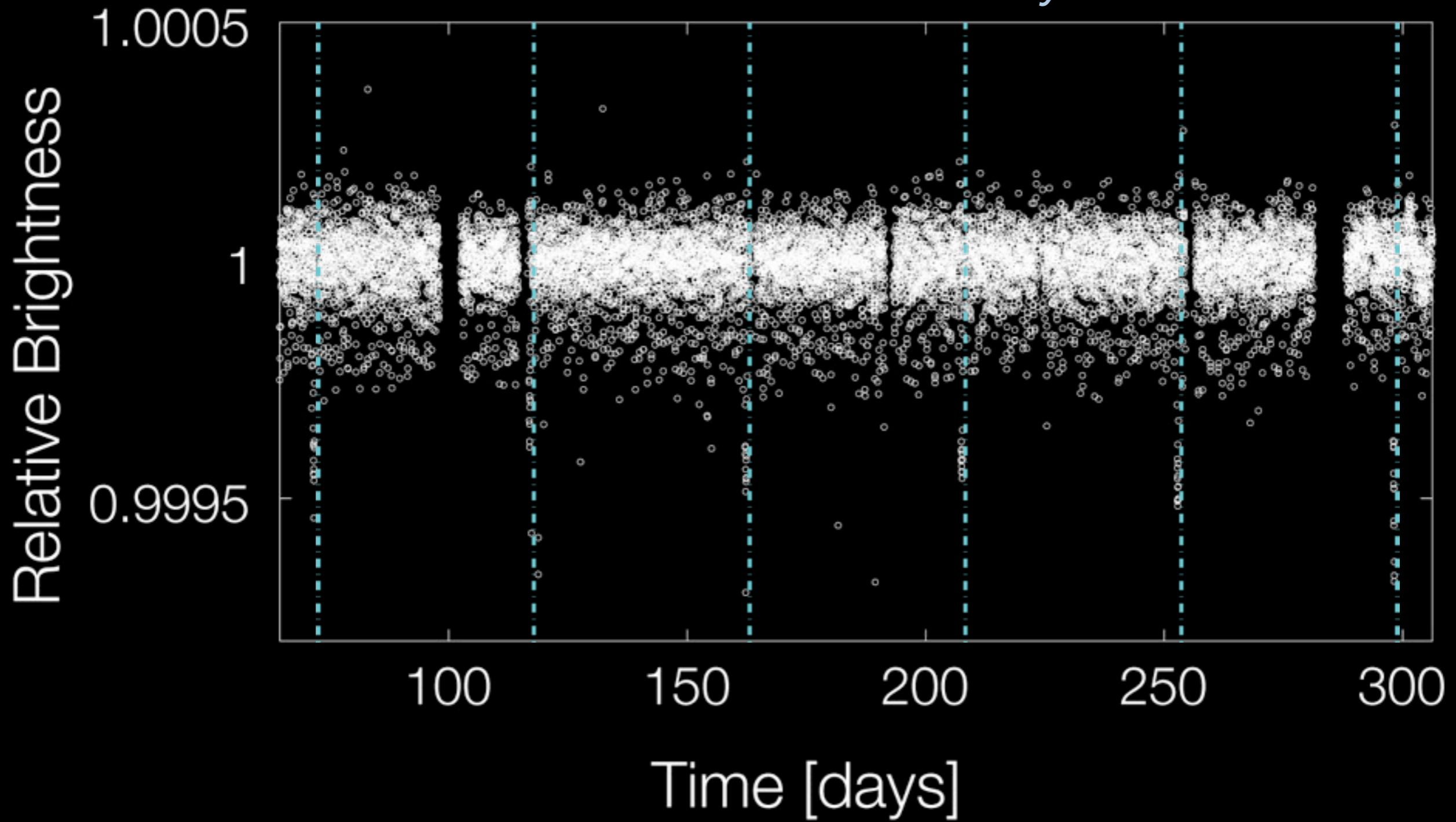


American Physical Society, May 1, 2011, Anaheim, CA

Kepler-10b Light Curve



Period = 45.29 days

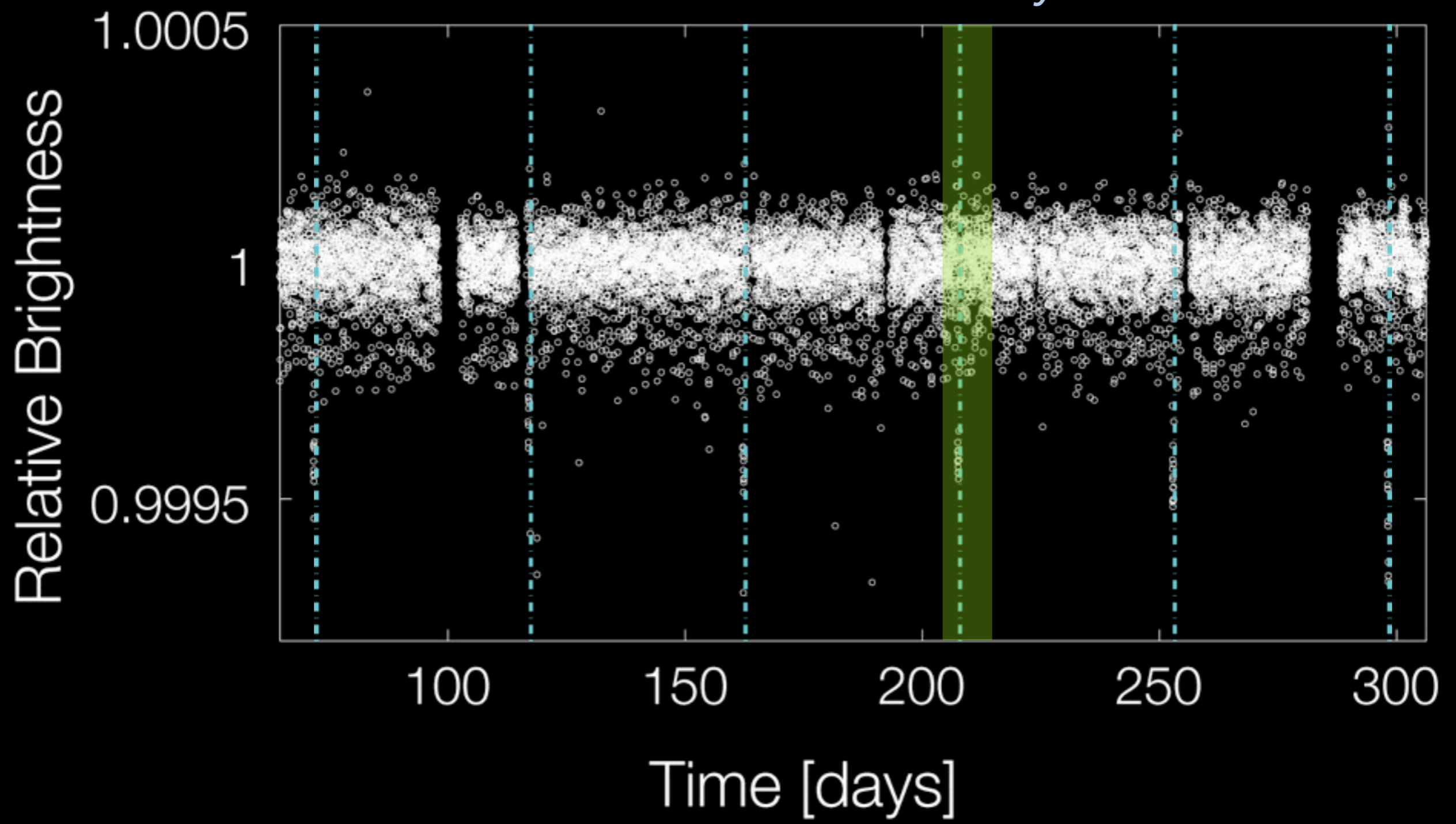


American Physical Society, May 1, 2011, Anaheim, CA

Kepler-10b Light Curve



Period = 45.29 days

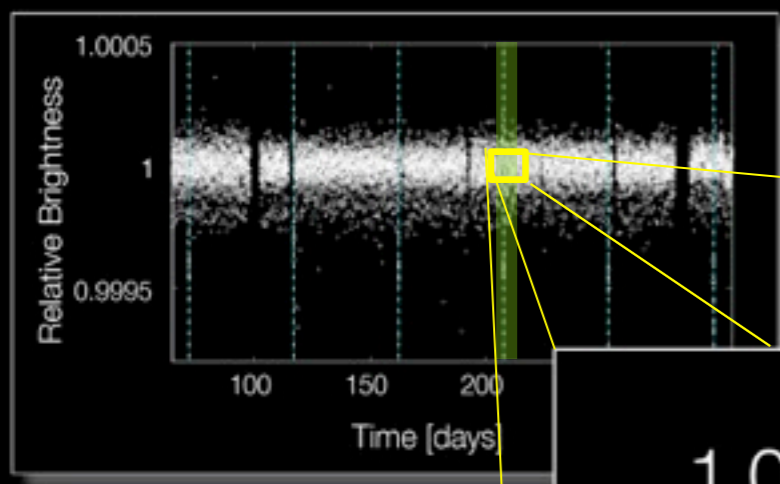


American Physical Society, May 1, 2011, Anaheim, CA

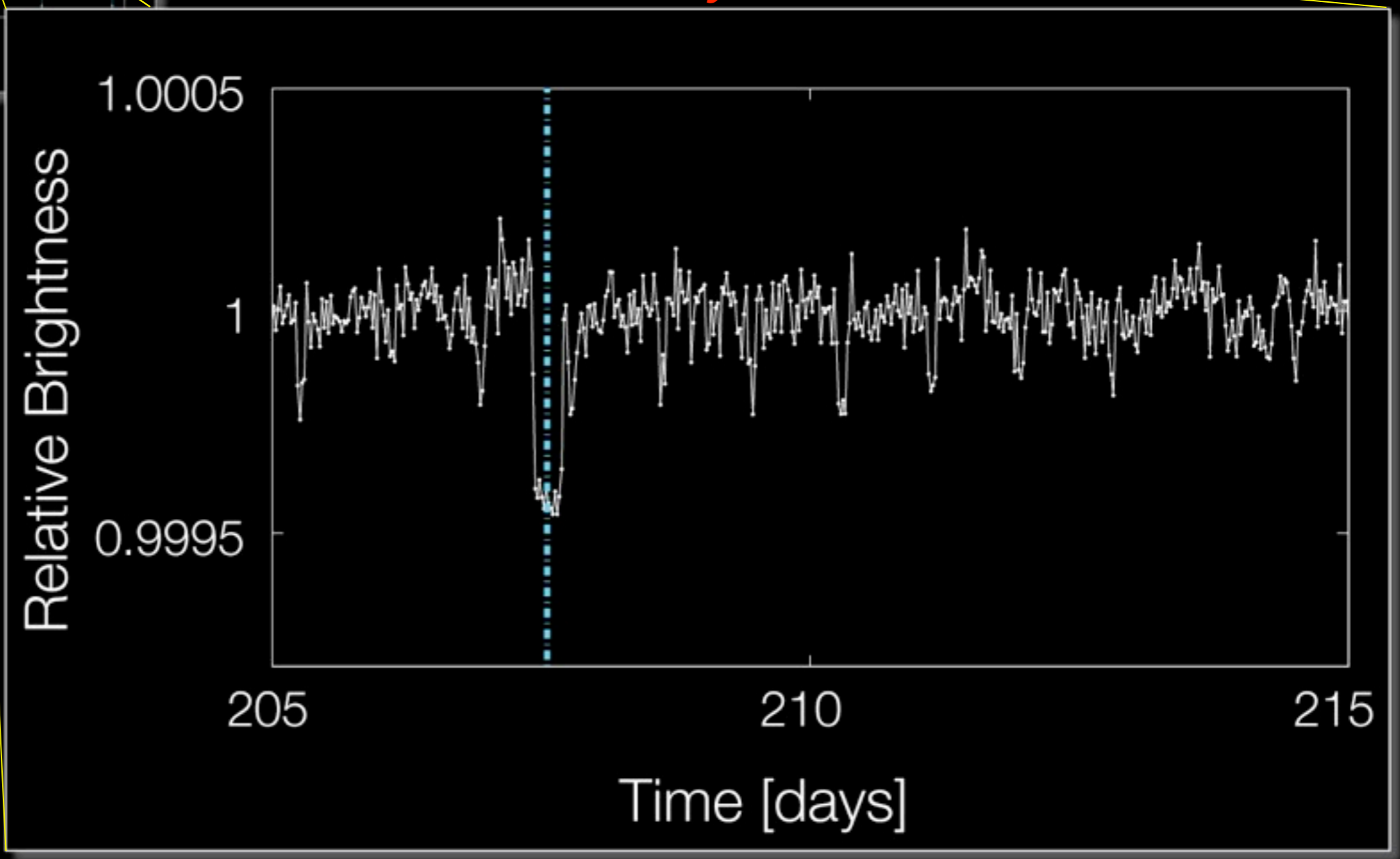


Kepler-10b Light Curve

Period = 45.29 days



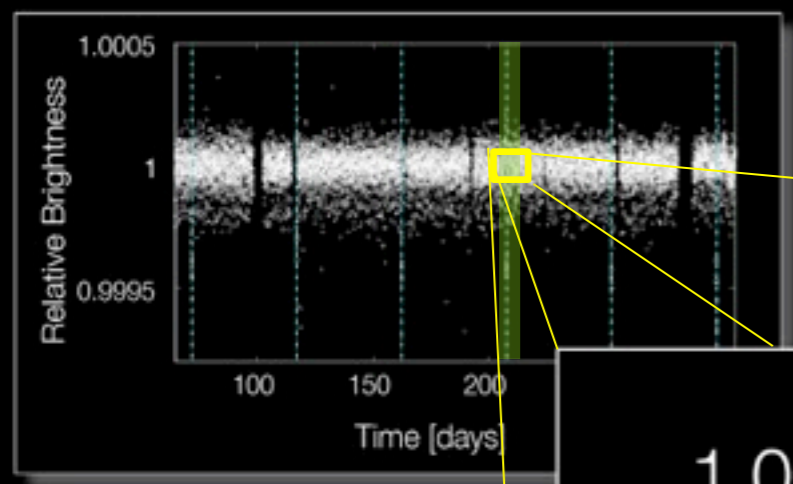
Period = .84 days



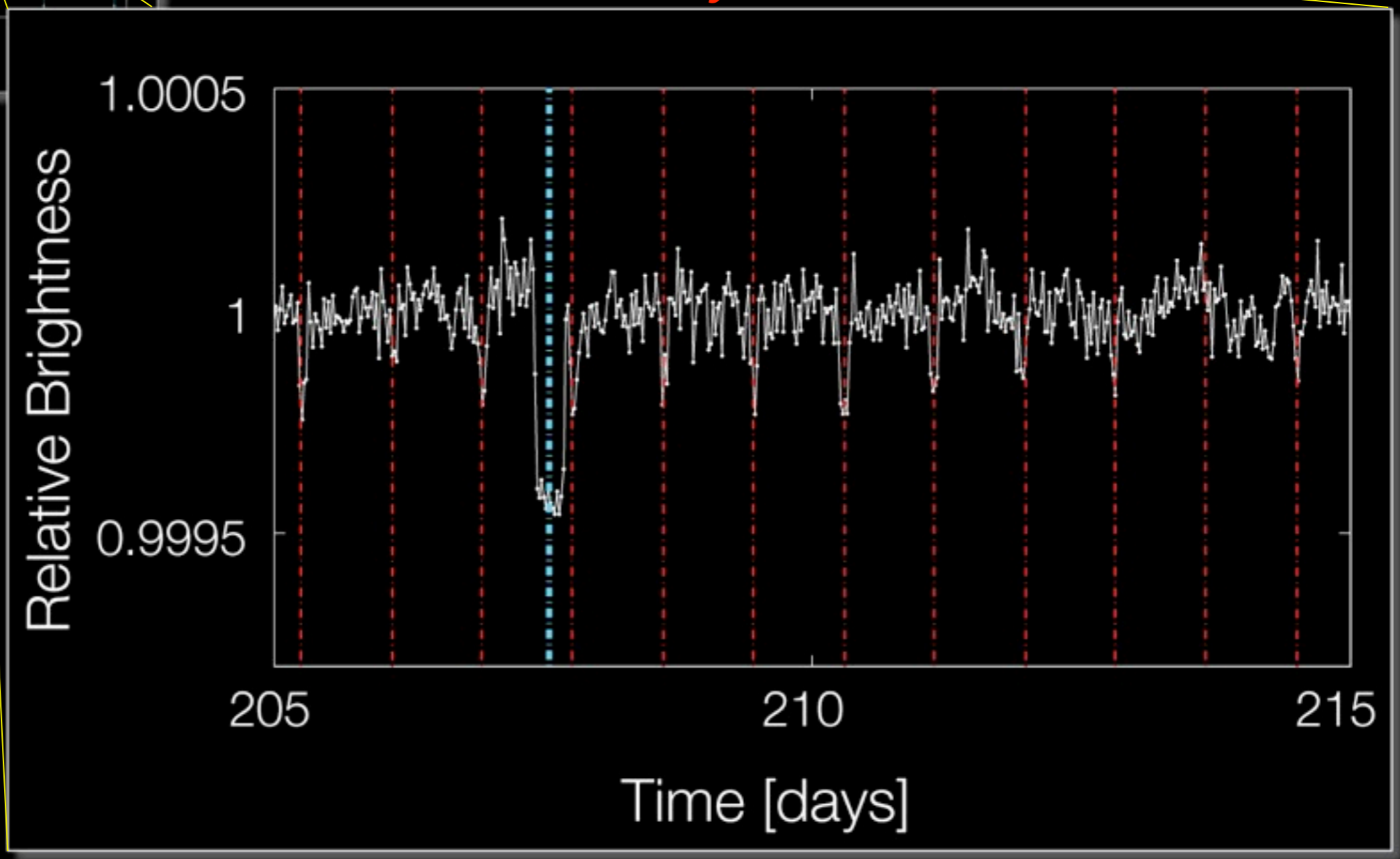
American Physical Society, May 1, 2011, Anaheim, CA

Kepler-10b Light Curve

Period = 45.29 days

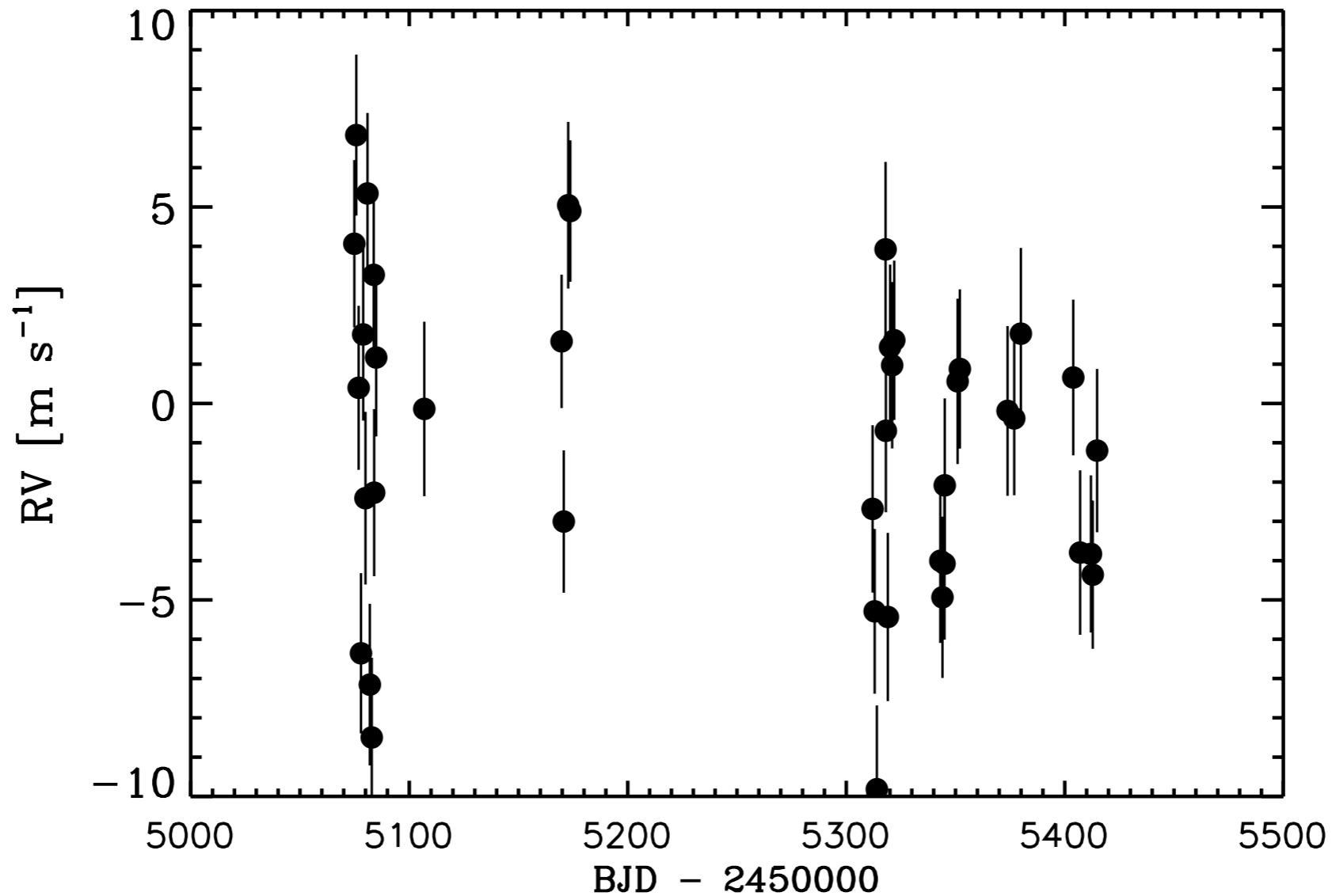


Period = .84 days

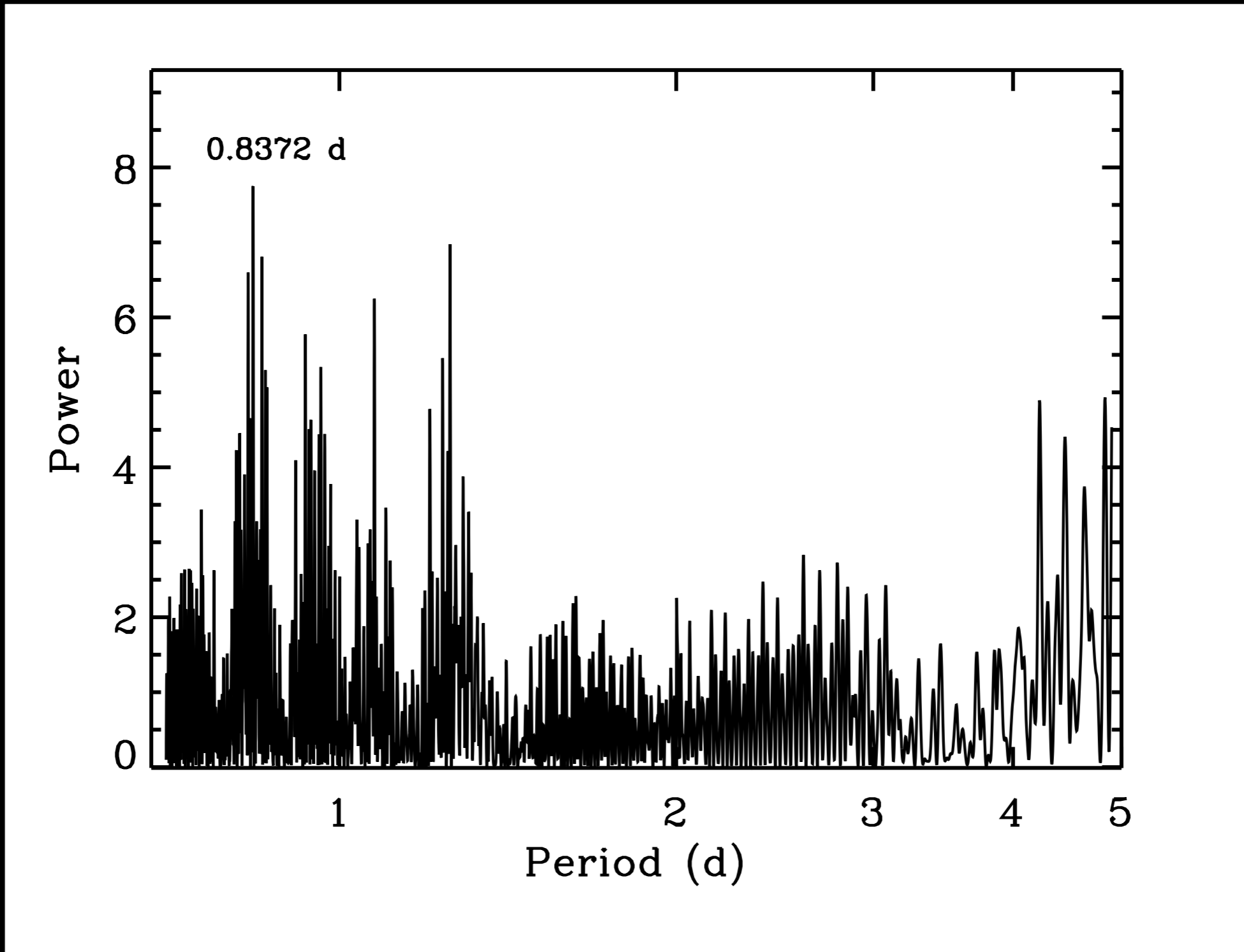


American Physical Society, May 1, 2011, Anaheim, CA

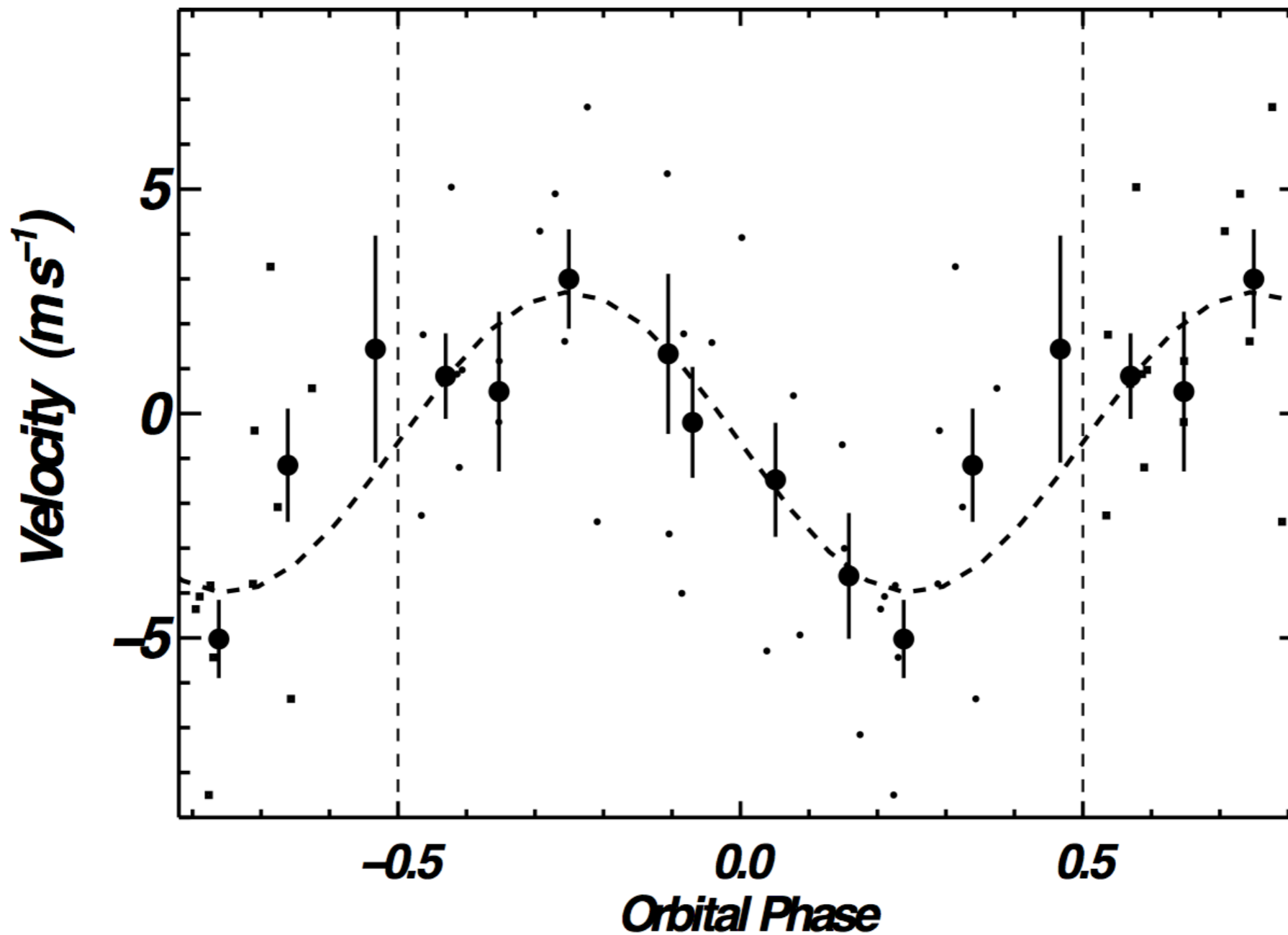
High Precision Doppler



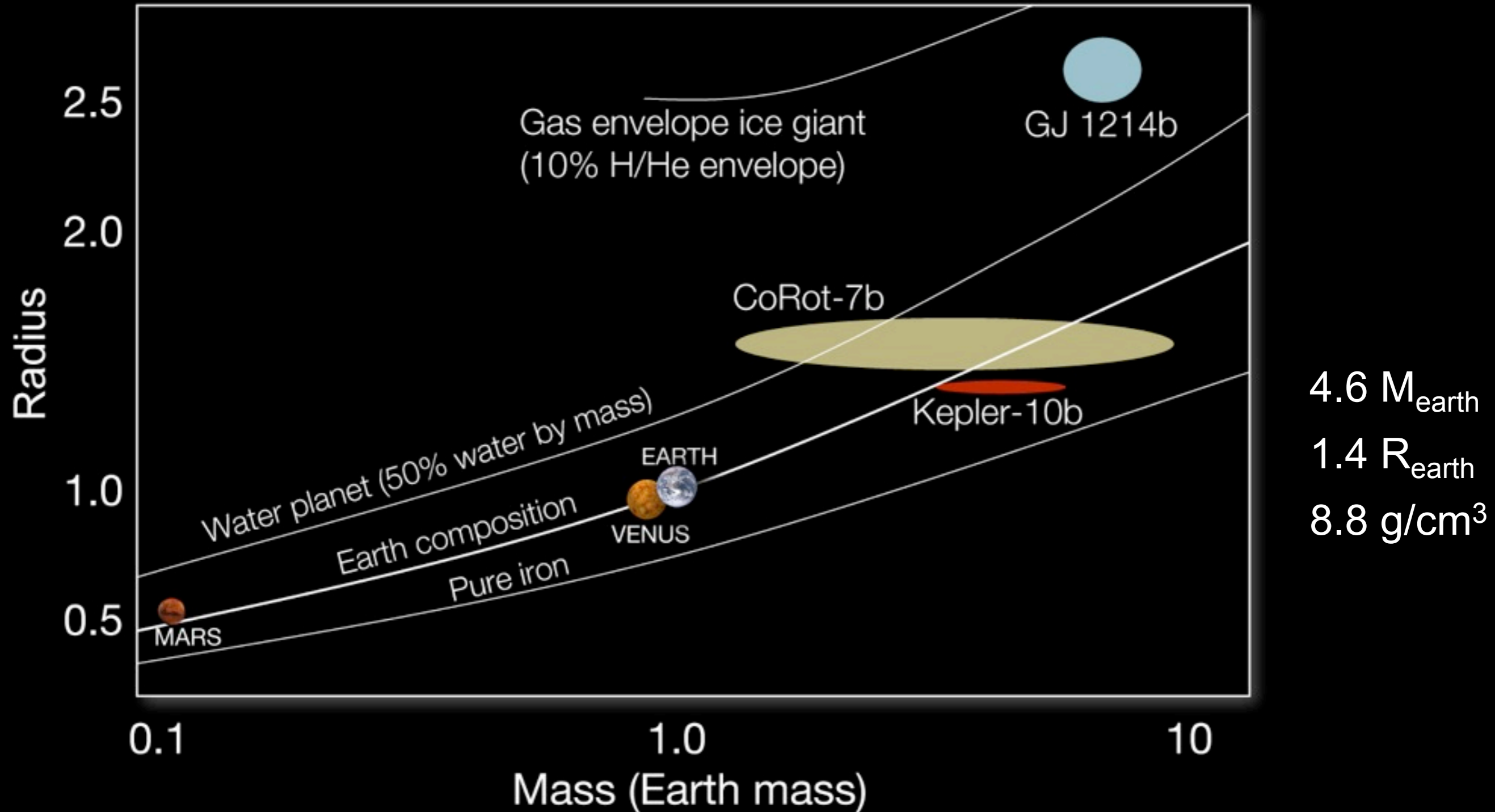
Doppler: Periodogram



Doppler: Model

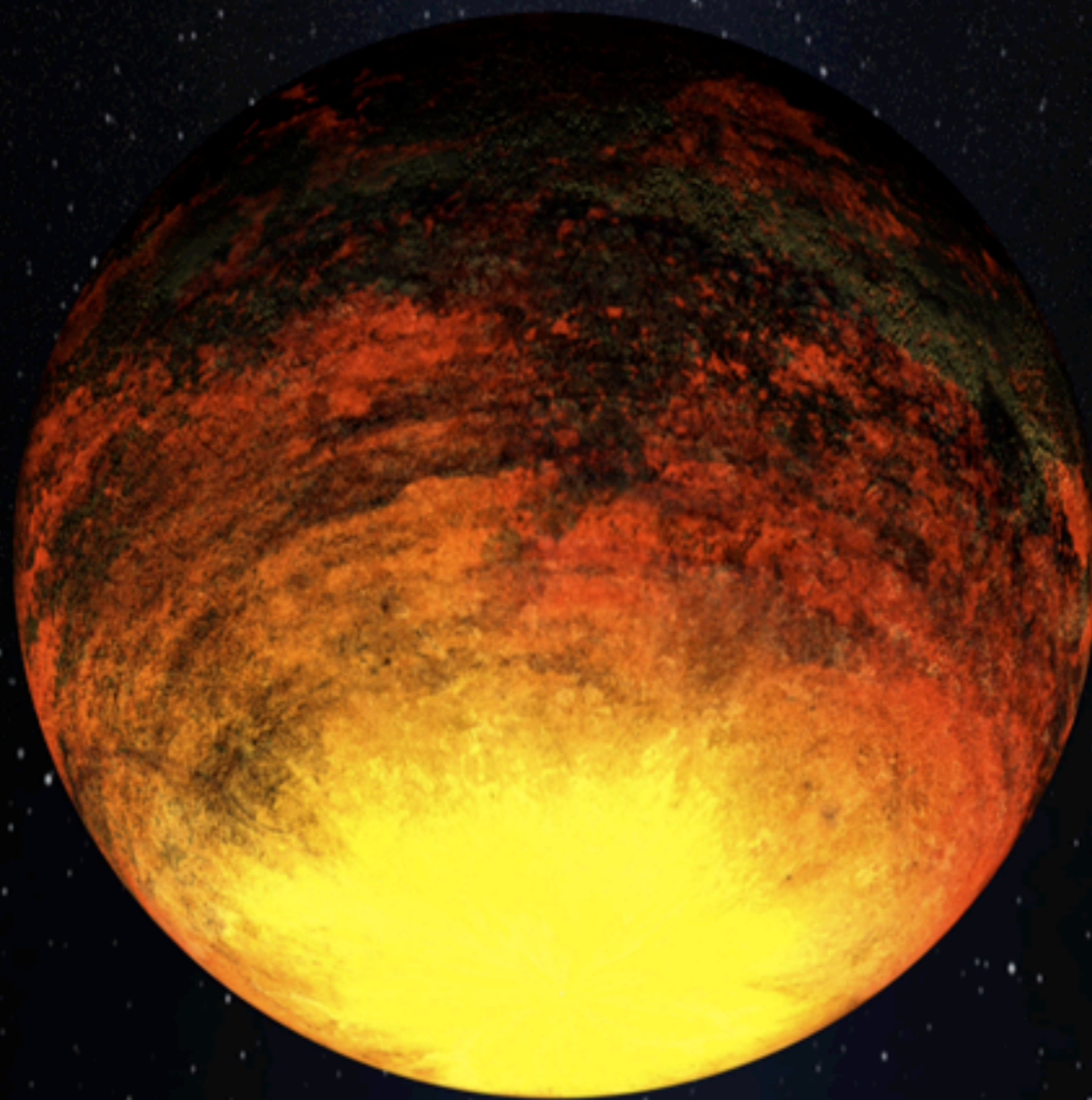


Kepler-10b: Composition

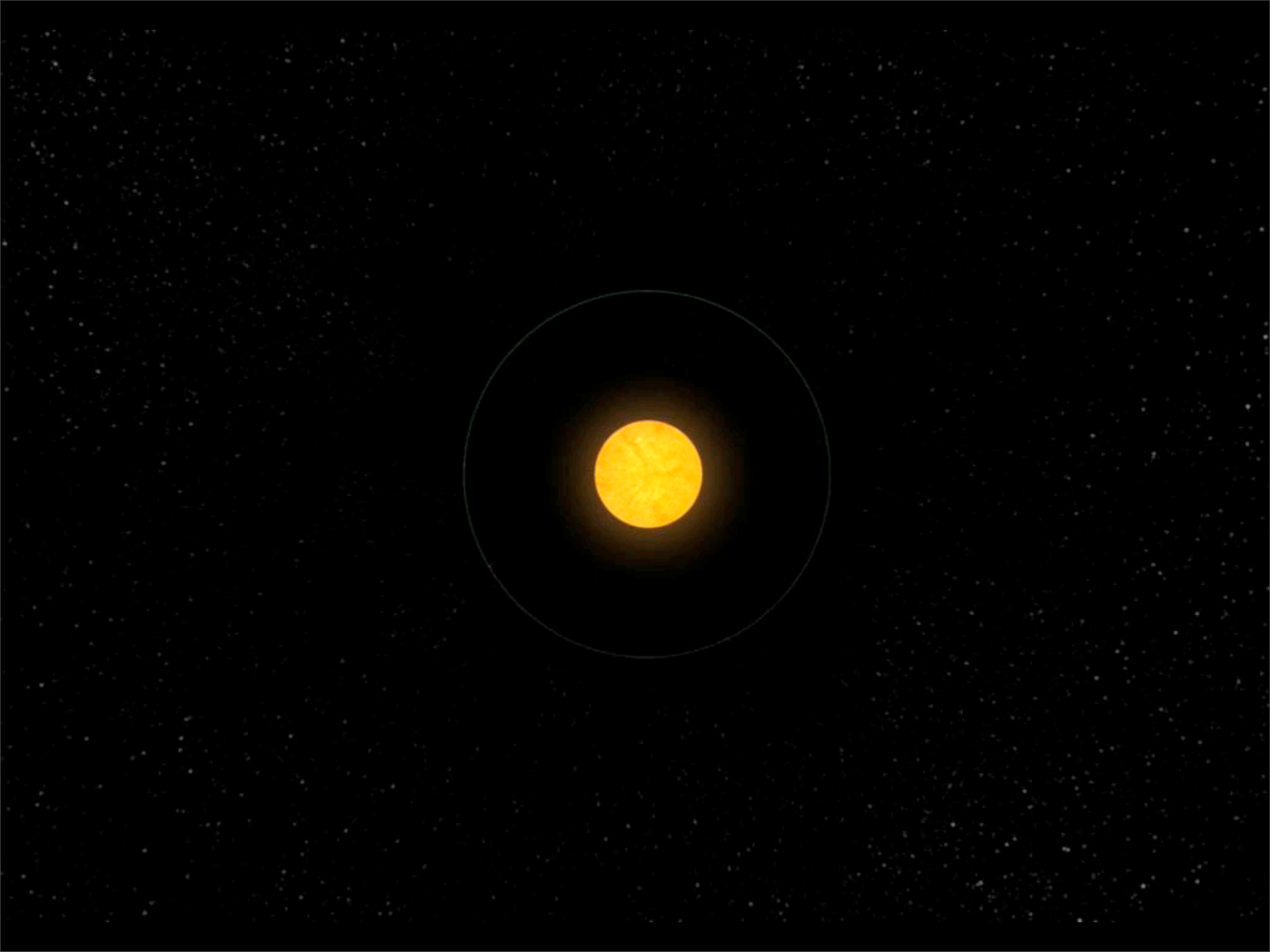




Kepler's First Rocky Planet



Kepler-10b
Batalha et al 2011, ApJ, 729, 27

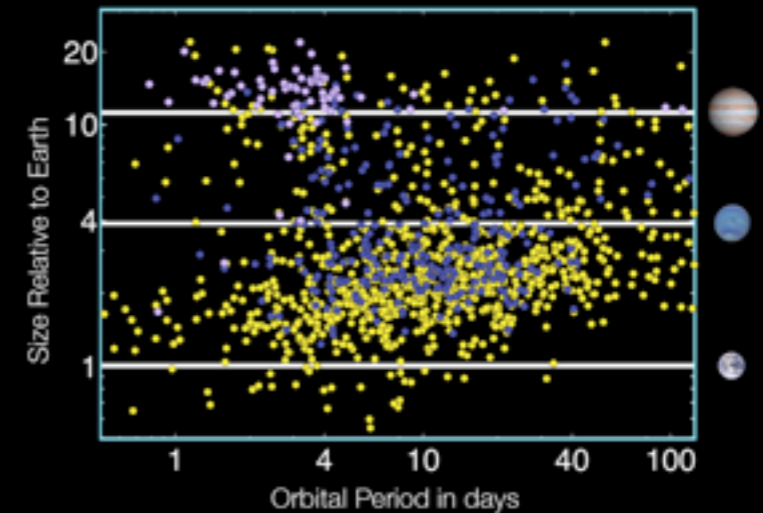


Occurrence Rates



Correcting for:

- geometric probability
- sensitivity biases



Borucki et al 2011, arXiv:1102.0541:

10.5% for Earth-size candidates, 6.5% for super-Earth size candidates, 17.5% for Neptune-size candidates, 2.5% for Jupiter-size candidates.

False positive rate unknown

Pipeline completeness unknown

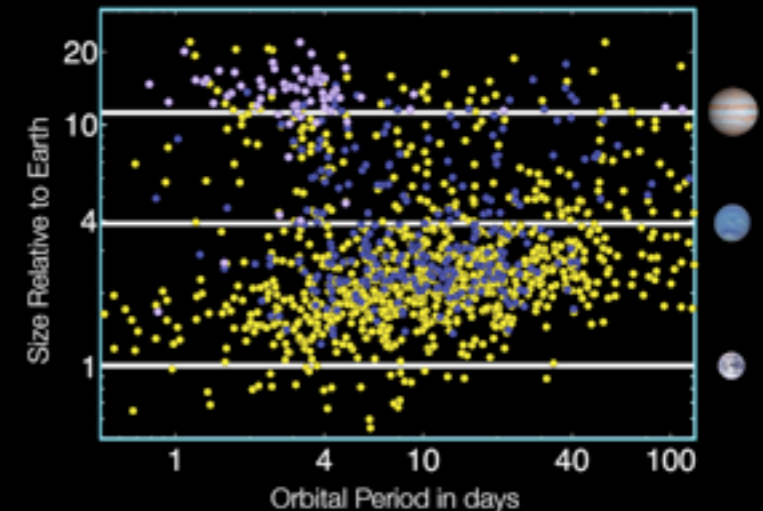


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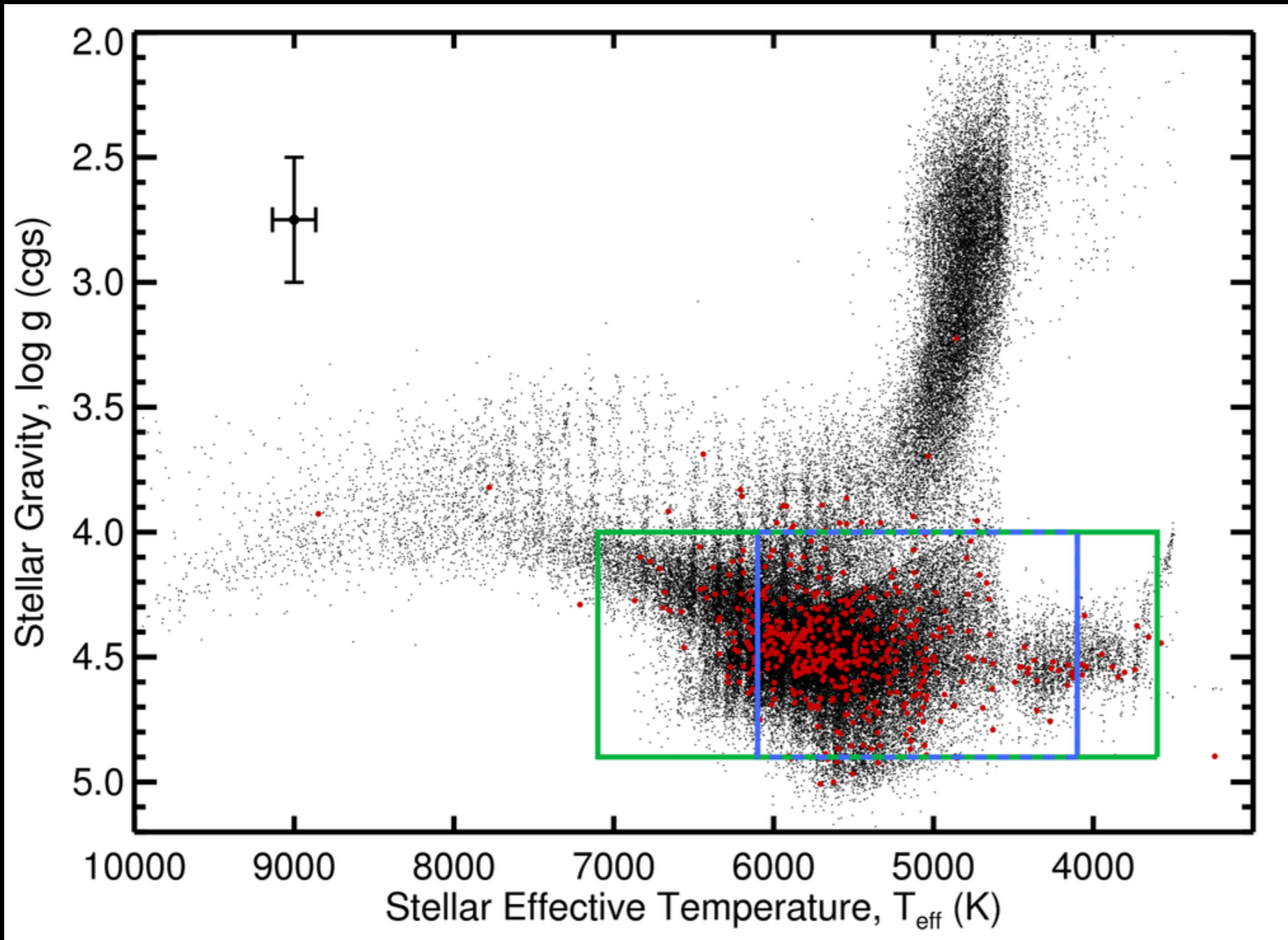
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Occurrence Rates



Kepler
A Search for Terrestrial Planets

Subsample of stars: bright GK Main Sequence



$T_{\text{eff}}: 4100-6100 \text{ K}$
 $\log g: 4.0-4.9$
 $K_p: < 15$

Subsample of candidates:

$P < 50 \text{ days}$
 $R_p > 2 R_e$
 $SN > 10$

Howard et al. 2011, arXiv: 1103.2541

American Physical Society, May 1, 2011, Anaheim, CA

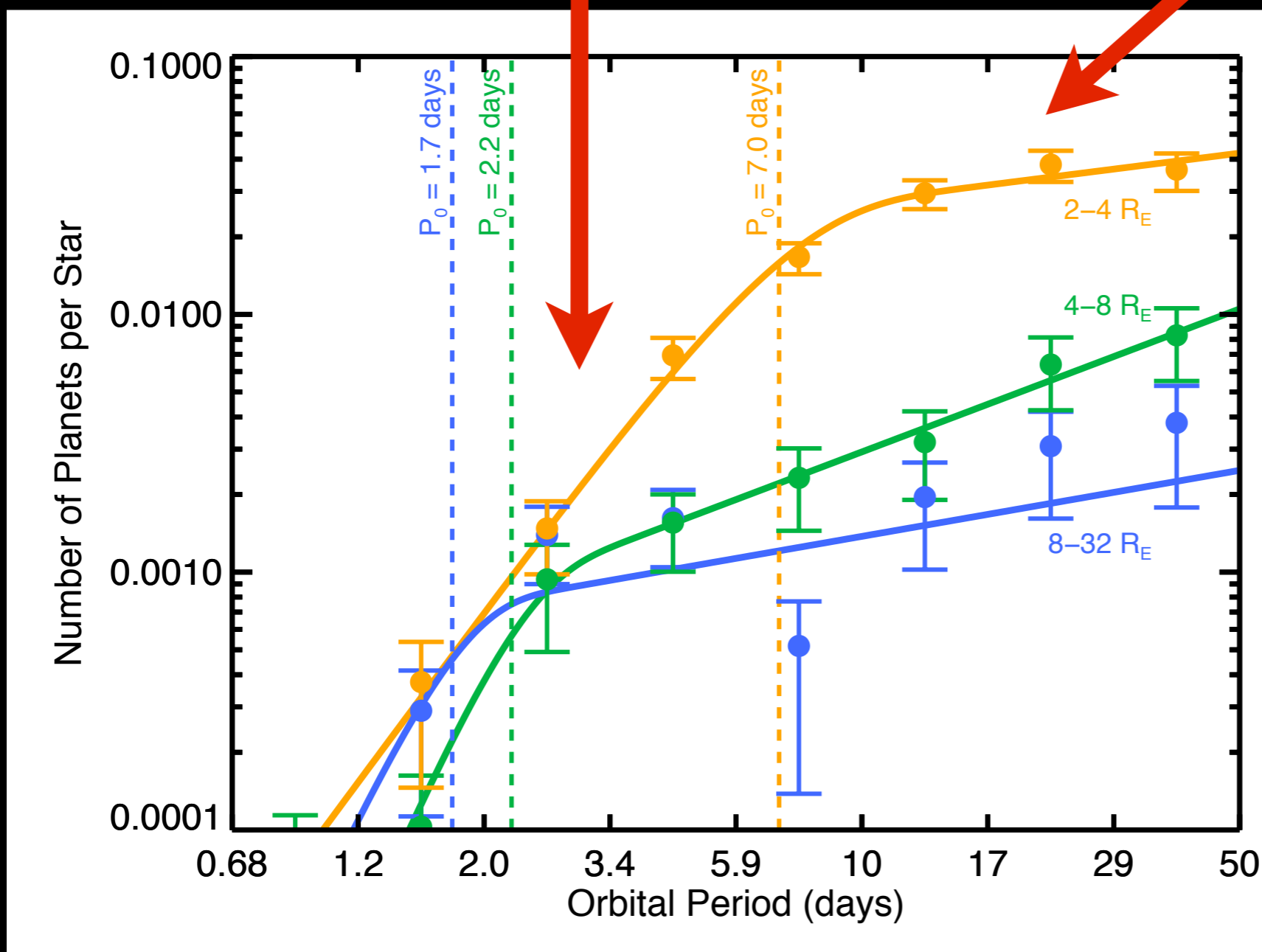


Occurrence Rate vs. Period



Exponential Cutoff

Power Law



Power law with Exponential Cutoff:
 $dN/d\log P = kP^\beta (1 - \exp(-(P/P_0)^\gamma))$

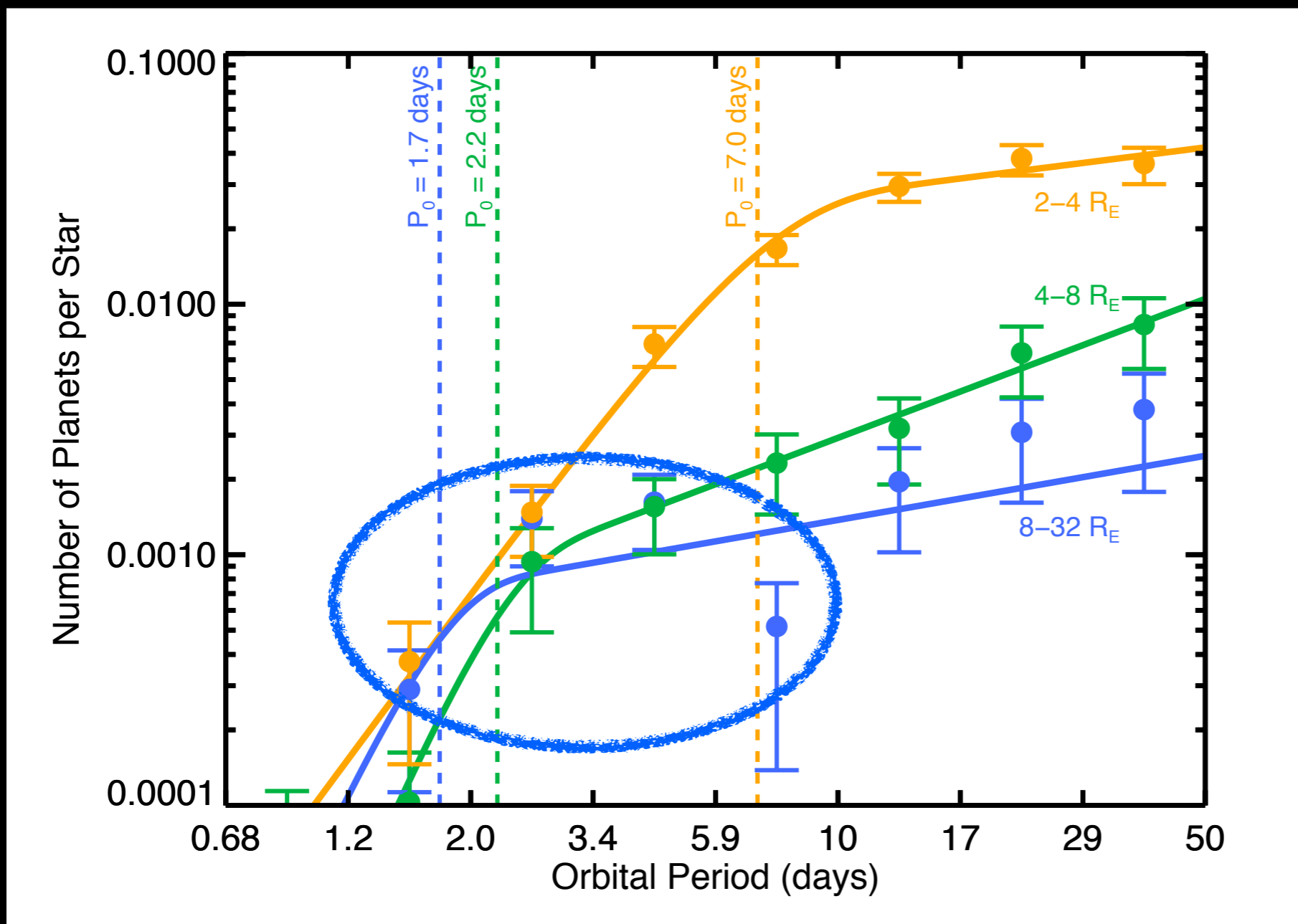
Total occurrence rates qualitatively similar to RV results.

Howard et al. 2011, arXiv: 1103.2541

American Physical Society, May 1, 2011, Anaheim, CA



Occurrence Rate vs. Period



Hot-Jupiters
($P < 10$ days):

Kepler: 0.004
RV: 0.012

3x's lower in
Kepler sample

Howard et al. 2011, arXiv: 1103.2541

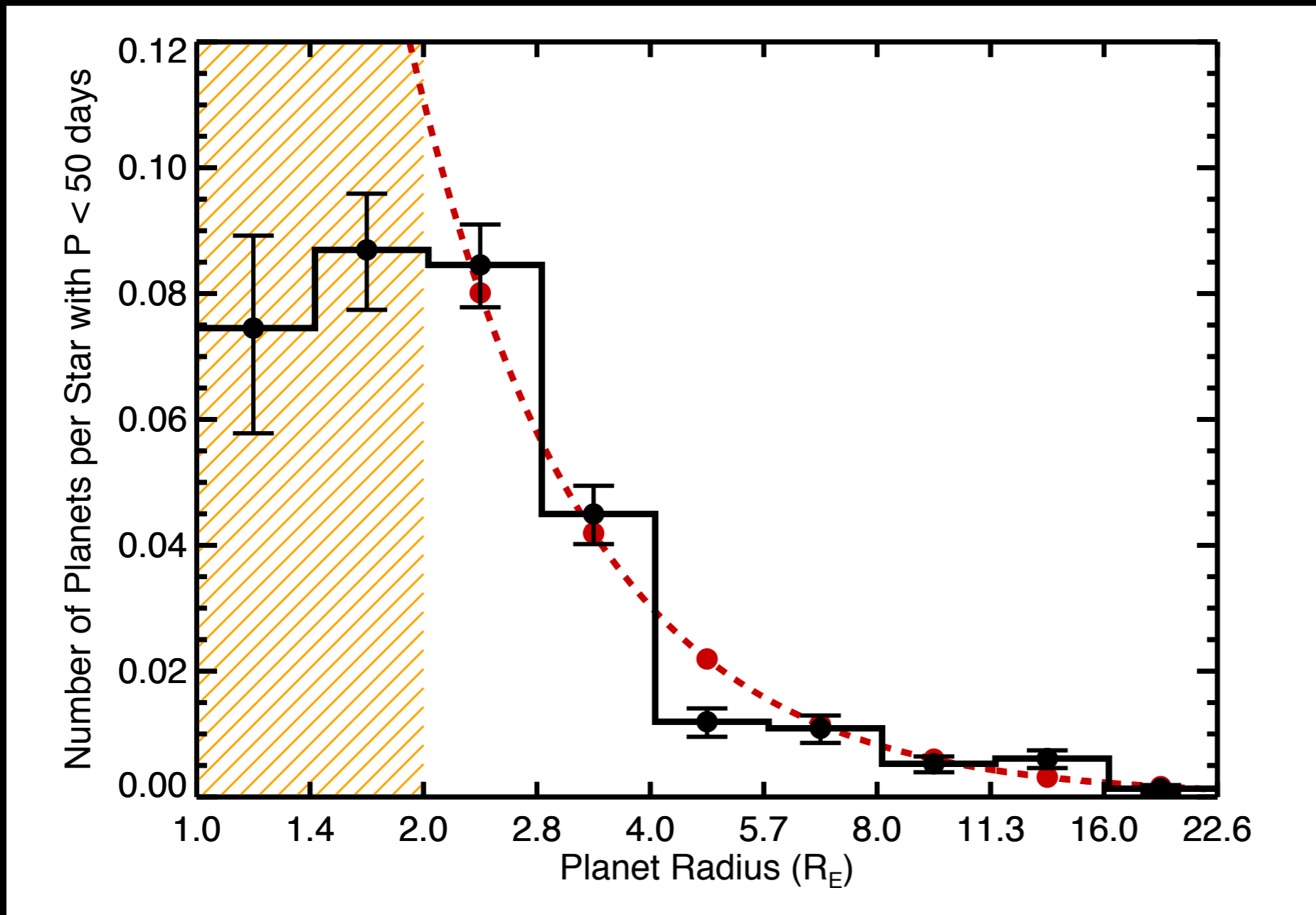
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Occurrence Rate vs. Radius



Power law increase in occurrence rate for smaller planets



$$dN/d\log R = kR^\alpha$$

$$k = 2.9 \pm 0.5$$

$$\alpha = -1.92 \pm 0.11$$

Howard et al. 2011, arXiv: 1103.2541

American Physical Society, May 1, 2011, Anaheim, CA



Statistics suggest that small planets are going to be common



produced by Dana Berry

"One day, from the shores of a new world, we'll gaze at the sea that took us there. And its waves will be stars."

Rui Borges, Sagan Day Essay Contest
<http://kepler.nasa.gov/education/sagan>

