

# Signals for Compressed SUSY at the LHC

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**Compressed SUSY:** the gluino mass parameter  $M_3$  is taken much smaller than the wino mass parameter  $M_2$  near the GUT scale.

- Ratio of heaviest to lightest superpartner masses is reduced compared to mSUGRA.
- Naturally allows the correct dark matter thermal relic abundance  $0.09 < \Omega_{\text{CDM}} h^2 < 0.13$  by top-squark-mediated LSP annihilation in the early universe:  $\tilde{N}_1 \tilde{N}_1 \rightarrow t \bar{t}$ .
- Ameliorates the SUSY little hierarchy problem (as in Kane-King hep-ph/9810374).

Typically need roughly  $M_3/M_2 \sim 1/3$  at the GUT scale, and:

$$m_t < m_{\tilde{N}_1} < m_t + 100 \text{ GeV}$$

$$m_{\tilde{N}_1} + 25 \text{ GeV} < m_{\tilde{t}_1} < m_{\tilde{N}_1} + 150 \text{ GeV}$$

The LSP ( $\tilde{N}_1$ ) must be heavier than the top quark for natural dark matter from annihilation to tops.

The top squark ( $\tilde{t}_1$ ) must be not too much heavier than the LSP.

The decay  $\tilde{t}_1 \rightarrow t\tilde{N}_1$  is generally forbidden.

In the models to follow,  $\tilde{t}_1 \rightarrow c\tilde{N}_1$ .

To be specific, assume that at the GUT scale:

$$\begin{aligned}M_1 &= m_{1/2}(1 + C_{24}), \\M_2 &= m_{1/2}(1 + 3C_{24}), \\M_3 &= m_{1/2}(1 - 2C_{24}),\end{aligned}$$

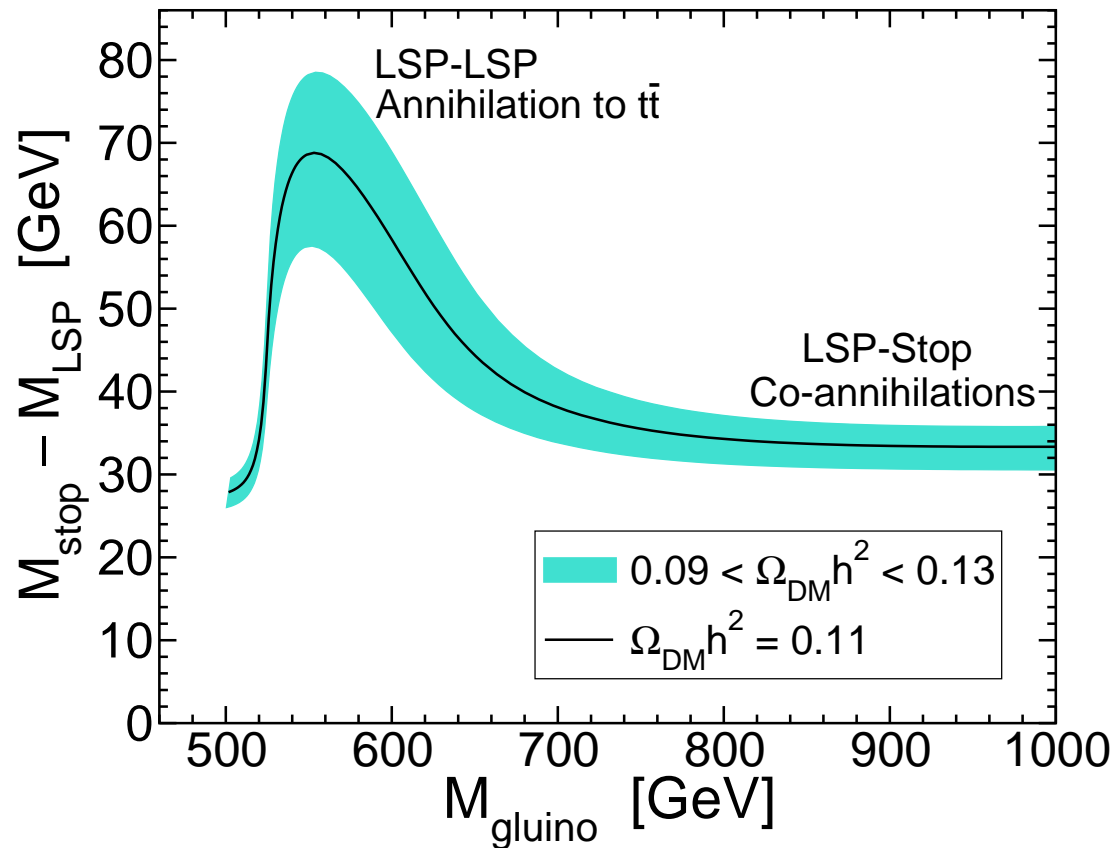
where  $C_{24}$  parameterizes the amount of  $SU(5)$  adjoint **(24)**  $F$ -term SUSY breaking.

$C_{24} = 0$  recovers the usual mSUGRA.

Instead,  $0.15 \lesssim C_{24} \lesssim 0.28$  allows natural top-squark-mediated Dark Matter annihilation.

In this talk, I consider a Model Line with  $C_{24} = 0.21$ ,  $\tan \beta = 10$ ,  $\mu > 0$ , and varying  $M_1$ , with  $m_0$  adjusted to give the right amount of dark matter.

Mass difference  $M_{\tilde{t}_1} - M_{\tilde{N}_1}$  for the dark matter allowed region of the Model Line:



The  $\tilde{N}_1 \tilde{N}_1 \rightarrow t\bar{t}$  “bulge” region has enhanced detection efficiency at the LHC, because the larger mass difference gives harder jets from  $\tilde{t}_1 \rightarrow c\tilde{N}_1$ .

LHC signal:

$$pp \rightarrow \tilde{g}\tilde{g} \rightarrow \begin{cases} t\bar{t}\tilde{t}_1\tilde{t}_1^* \rightarrow t\bar{t}c\bar{c} + \cancel{E}_T & (50\%) \\ tt\tilde{t}_1^*\tilde{t}_1^* \rightarrow tt\bar{c}\bar{c} + \cancel{E}_T & (25\%) \\ \bar{t}\bar{t}\tilde{t}_1\tilde{t}_1 \rightarrow \bar{t}\bar{t}cc + \cancel{E}_T & (25\%) \end{cases}$$

Due to the Majorana gluino, get the distinctive signal:

$$\begin{aligned} \ell^+\ell^+bb + \text{jets} + \cancel{E}_T, \\ \ell^-\ell^-bb + \text{jets} + \cancel{E}_T \end{aligned}$$

Kraml and Raklev studied this for other models in hep-ph/0512284.

Their study differed in having  $\tilde{t}_1$  much lighter, and other squarks much heavier.

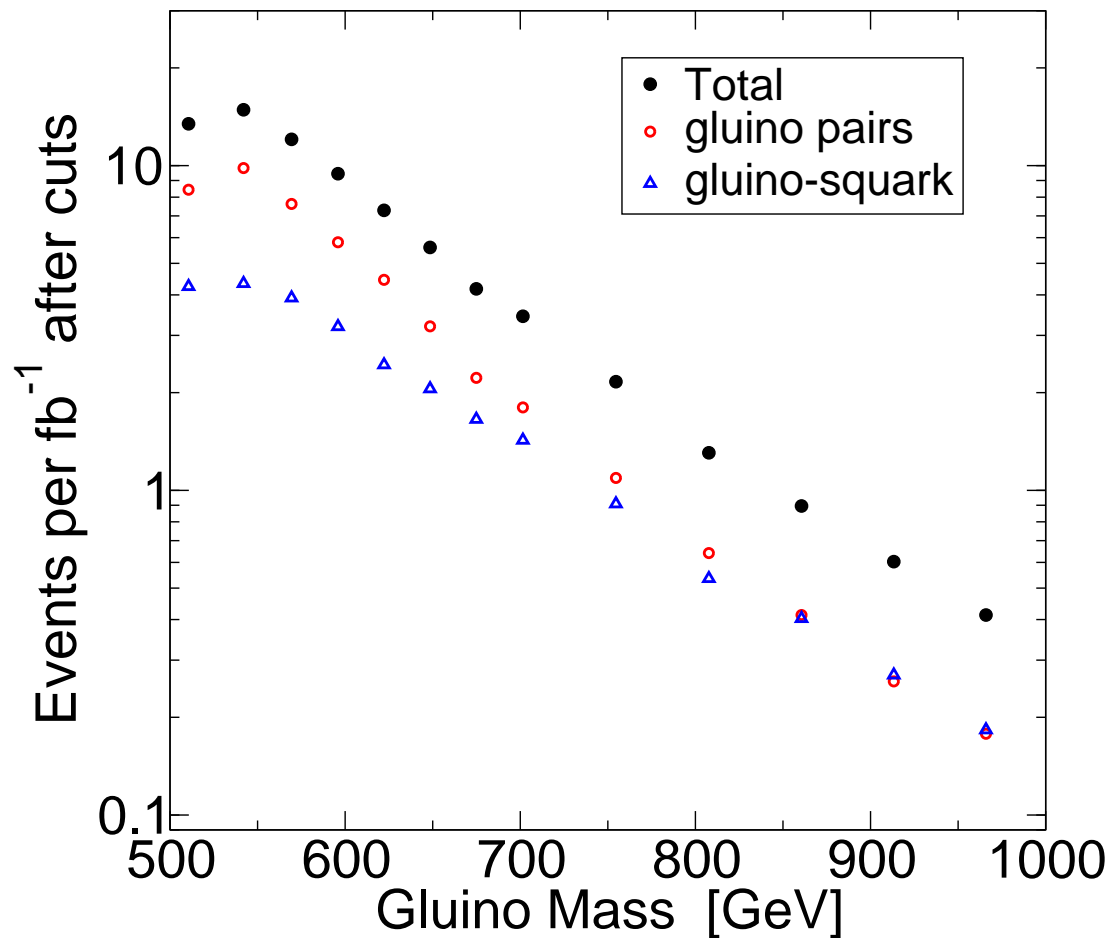
Use MadGraph/MadEvent  $\rightarrow$  Pythia  $\rightarrow$  PGS4 for event generation and detector simulation.

Cuts:

- Exactly 2 Same-Sign isolated leptons with  $p_T > 20$  GeV
- One jet with  $p_T > 100$  GeV, two more with  $p_T > 50$  GeV
- At least two b tags
- Two Same-Sign lepton- $b$  pairs, each consistent with top:  $m(\ell b) < 180$  GeV
- $\cancel{E}_T > 100$  GeV

With these cuts, the Standard Model background is mostly  $t\bar{t}$ , less than 1 event/fb $^{-1}$ . However, depends crucially on fake rates and wrong-sign assignment rates for leptons, which are difficult to anticipate before data taking.

Signal rates, after cuts, for  $bb\ell^+\ell^+\cancel{E}_T$  and  $bb\ell^-\ell^-\cancel{E}_T$ :



Detection prospects will depend on how well Same-Sign lepton backgrounds can be understood, from e.g.  $Z \rightarrow \ell^+\ell^+$ .



## Mass determinations are hard.

- No dilepton mass edges.
- Sleptons, winos, higgsinos nearly decouple from LHC.
- Difficult to disentangle

$$pp \rightarrow \tilde{g}\tilde{g} \rightarrow t\tilde{t}_1^*\tilde{t}_1^*$$

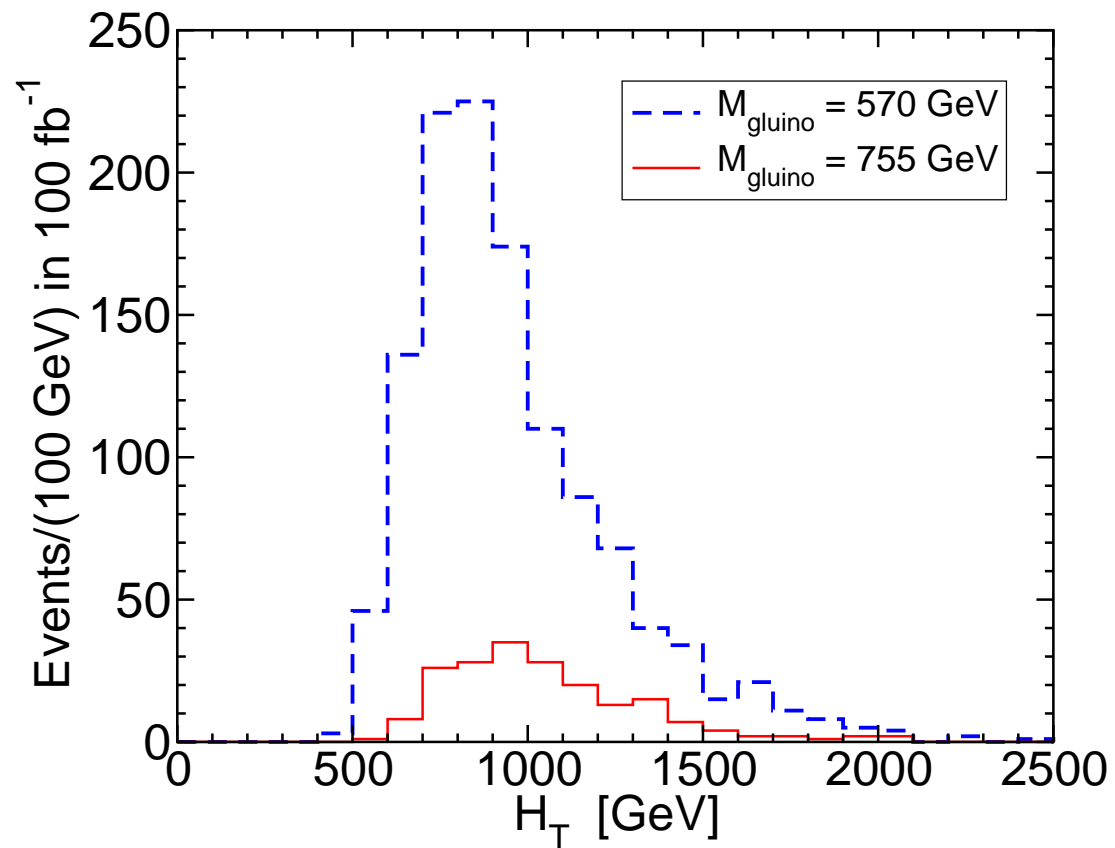
from

$$pp \rightarrow \tilde{g}\tilde{q}_L \rightarrow \tilde{g}\tilde{g}q \rightarrow t\tilde{t}_1^*\tilde{t}_1^*j$$

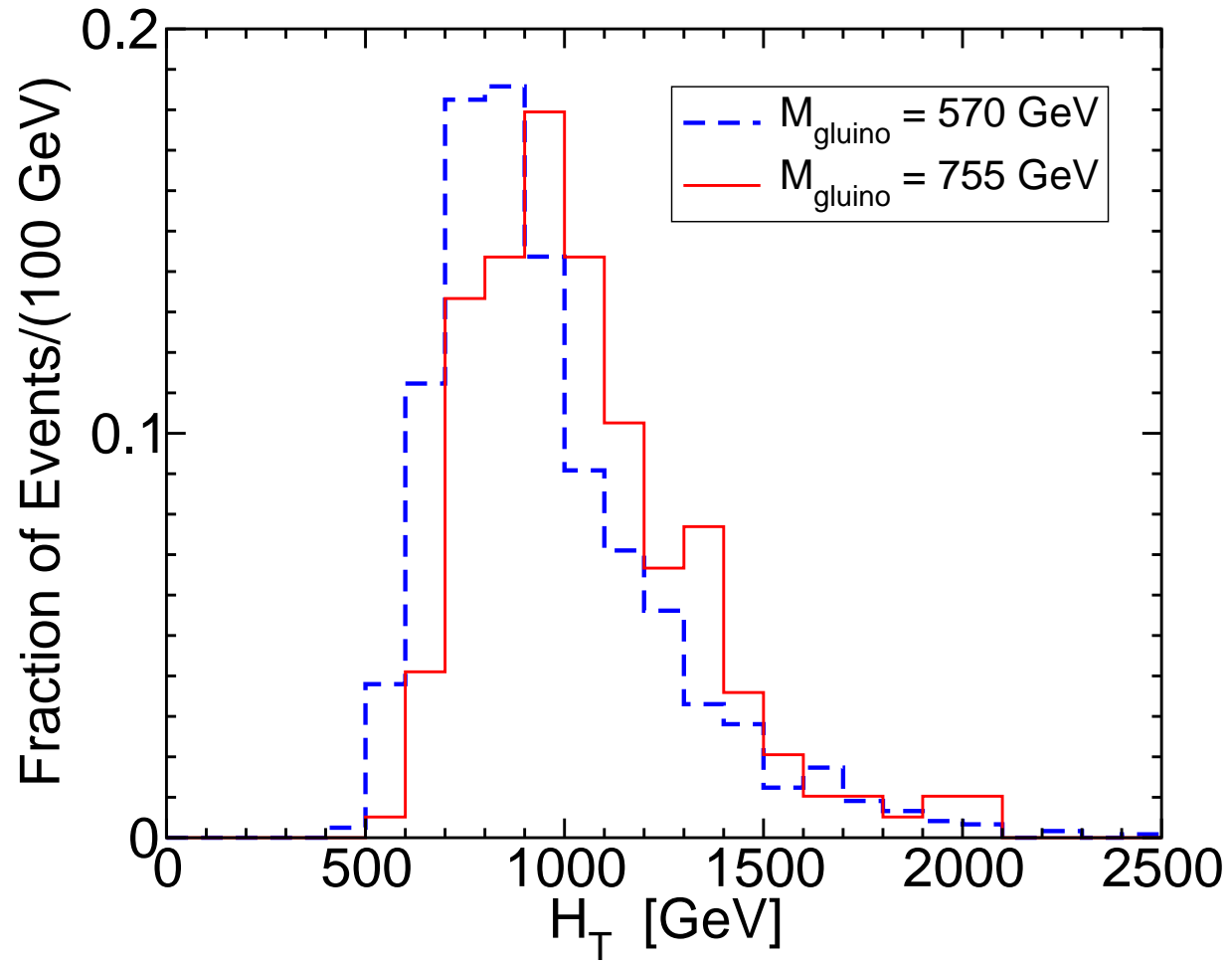
“SUSY is its own background” is especially true in Compressed SUSY.

One attempt:  $H_T = \sum p_T^\ell + \sum p_T^j + \cancel{E}_T$

For  $bb\ell^+\ell^+\cancel{E}_T$  and  $bb\ell^-\ell^-\cancel{E}_T$  events for two points on the Model Line with  $m_{\tilde{g}} = 570$  and 755 GeV:



Rescaled to compare the shapes:



In the classic collider signatures for SUSY,

Invisible LSPs  $\rightarrow$  Missing Energy  $\rightarrow$  No Mass Peaks

Compressed SUSY provides an exception, because of the long lifetime of  $\tilde{t}_1$ :

Stoponium =  $\eta_{\tilde{t}}$  = s-wave  $\tilde{t}_1^* \tilde{t}_1$  bound state

Drees and Nojiri 1994 proposed looking for stoponium in

$$pp \rightarrow \eta_{\tilde{t}} \rightarrow \gamma\gamma$$

Stoponium is very narrow, so the width is effectively that of the detector resolution for diphotons, of order 1% at CMS and ATLAS.

## Stoponium in Compressed SUSY:

- is always stable enough to form
- Binding energy of  $\eta_{\tilde{t}}$  is a few GeV
- $\Gamma_{\eta_{\tilde{t}}}$  is a few MeV
- $M_{\eta_{\tilde{t}}}$  between about 400 and 750 GeV
- $\text{BR}(\eta_{\tilde{t}} \rightarrow gg)$  dominates (but huge background)
- $\text{BR}(\eta_{\tilde{t}} \rightarrow \gamma\gamma) \approx 0.4\%$

The process

$$pp \rightarrow \eta_{\tilde{t}} \rightarrow \gamma\gamma$$

is clearly NOT a discovery mode for supersymmetry.

Importance is that it will give a uniquely precise measurement of the top-squark mass.

Look for a narrow diphoton mass peak against a smoothly falling background.

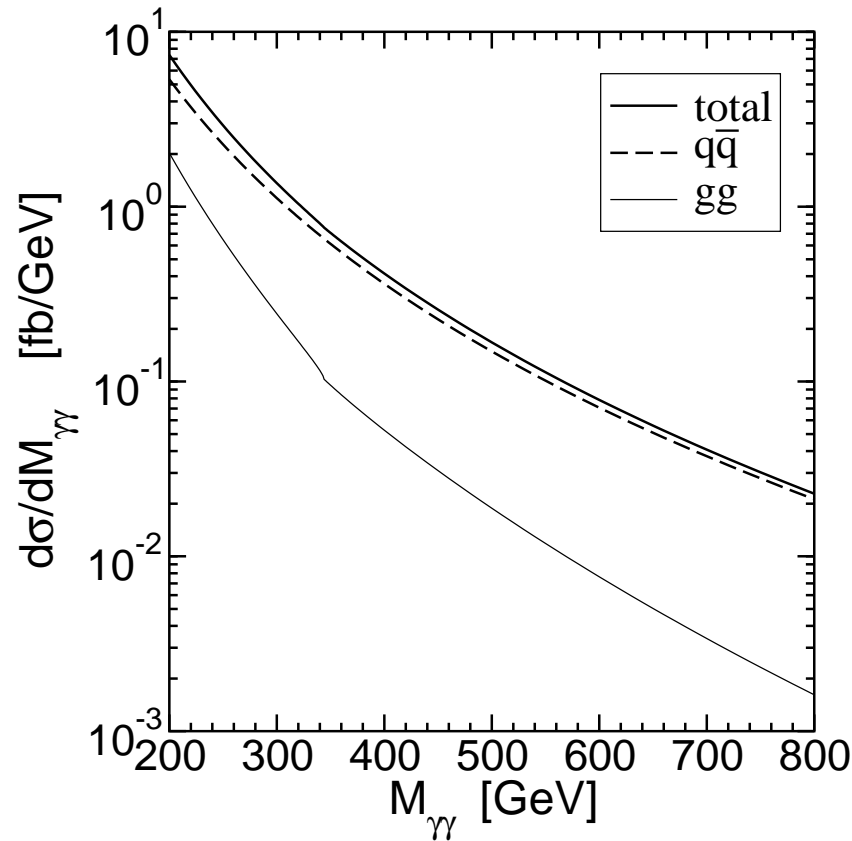
The irreducible physics backgrounds at leading order:

$$q\bar{q} \rightarrow \gamma\gamma \quad (\text{tree-level})$$

$$gg \rightarrow \gamma\gamma \quad (\text{1-loop})$$

Backgrounds from fakes are thought to be smaller.

Diphoton backgrounds at LHC, at leading order,  
after cuts:



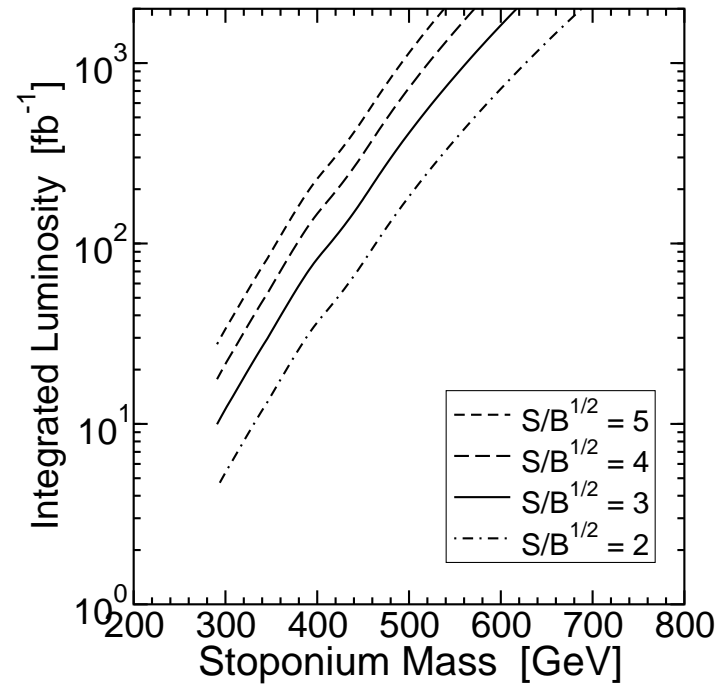
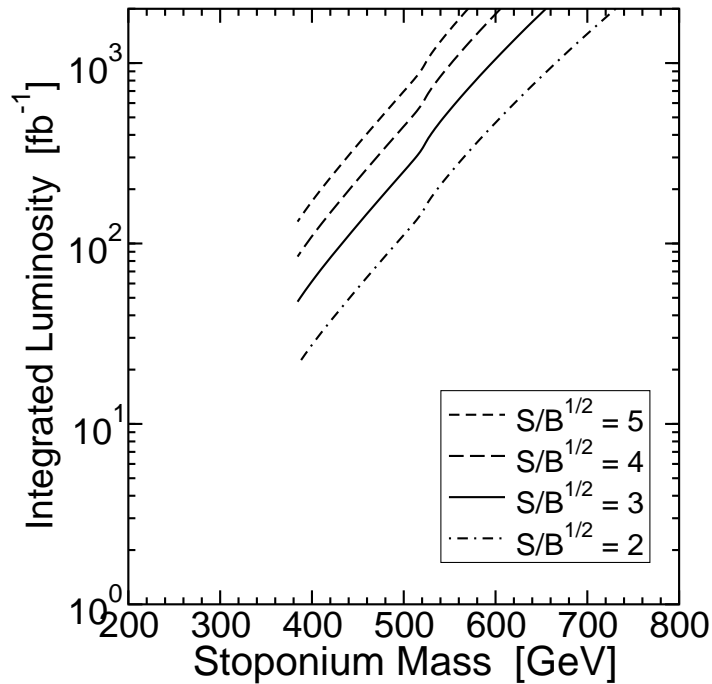
Note: actual background will be obtained from LHC data!



Luminosity needed for expected significances, for two model lines:

$$C_{24} = 0.21, \quad A_0/M_1 = -1$$

$$C_{24} = 0.21, \quad A_0/M_1 = -2$$

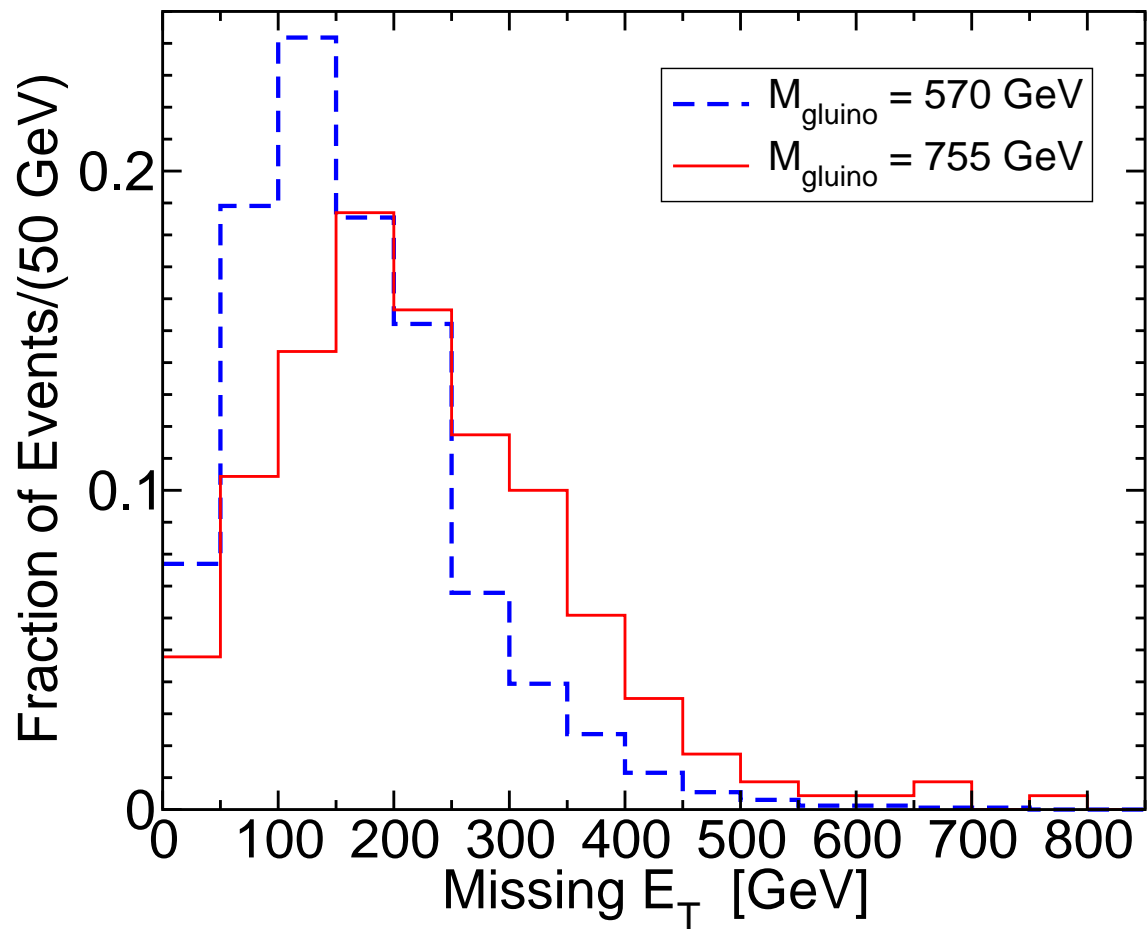


Detectability for  $m_{\eta_{\tilde{t}}} = 500$  GeV will require more than  $100 \text{ fb}^{-1}$ .  
 For smaller  $m_{\eta_{\tilde{t}}}$ , a few  $\times 10 \text{ fb}^{-1}$  might do it (if we're lucky).

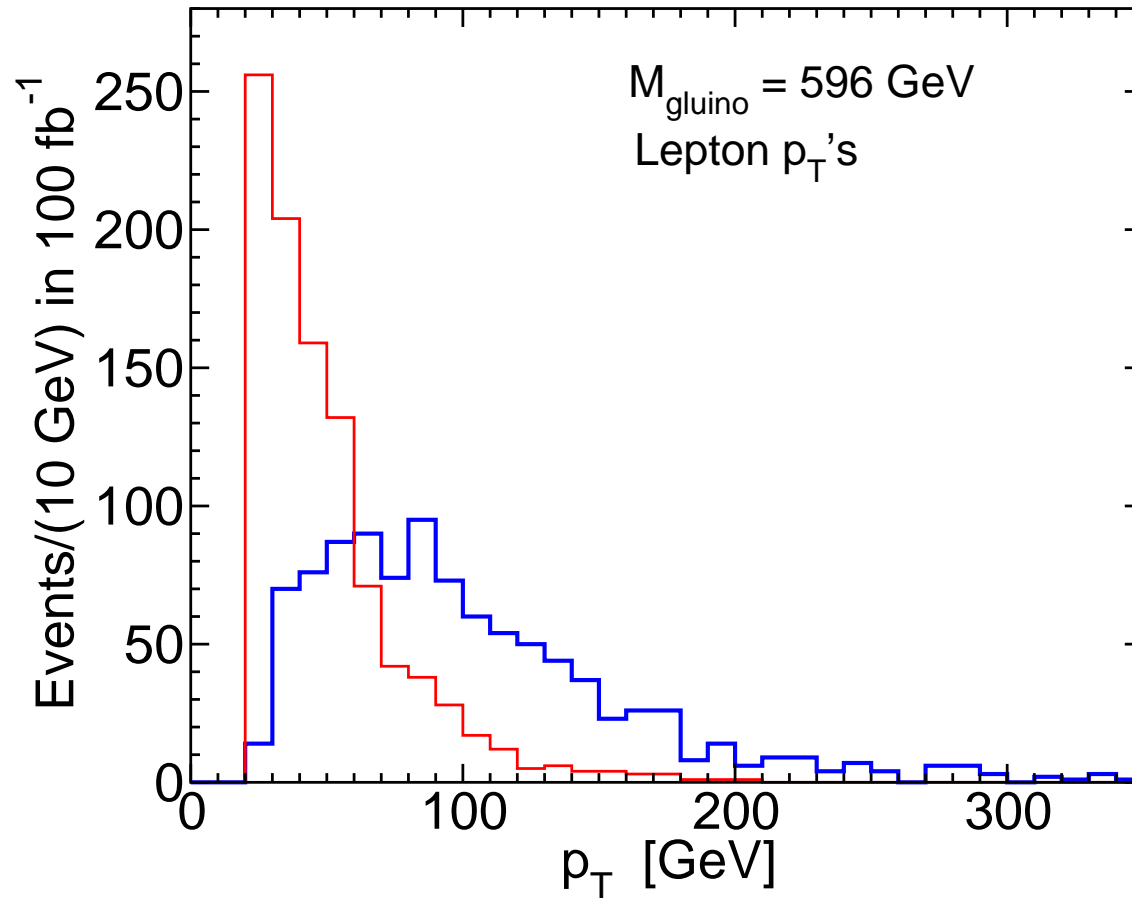
## Summary:

- The talk was its own summary!

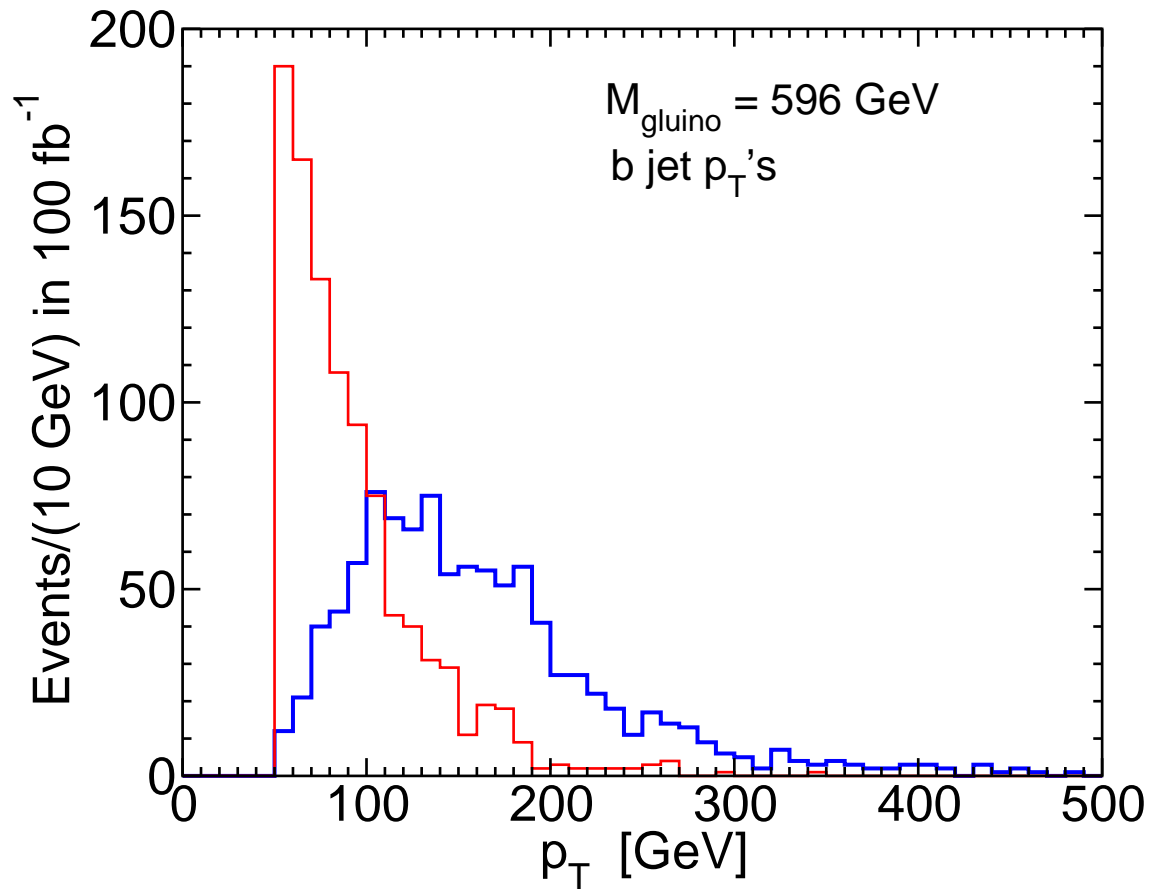
Extra slide:  $\cancel{E}_T$  distribution



# Extra slide: Lepton $p_T$ distributions



Extra slide:  $b$  jet  $p_T$  distributions



# Extra slide: jet $p_T$ distributions

