TriMinimal Parametrization of Neutrino Mixing Matrix

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Some neutrino Flavor physics

Besides energy and direction, cosmic quanta carry intrinsic information.

For cosmic-rays, it is A and Z;

For photons, it is spin polarization;

For neutrinos, it is flavor:

electron-neutrino(which showers)muon neutrino(whose CC tracks)tau neutrino(which showers below a PeV, tracks above)

Moreover, the flavors mix in a calculable/known way, which means the flavors oscillate in an L/E-dependent way, enabling:

Flavor "Interferometry" over Cosmic baselines !!

IceCube - 2010 completion, 30% 2007





What we know:

$$U = R_{32}(\theta_{32}) U_{\delta}^{\dagger} R_{13}(\theta_{13}) U_{\delta} R_{21}(\theta_{21})$$
(1)
= $\begin{pmatrix} c_{21}c_{13} & s_{21}c_{13} & s_{13}e^{-i\delta} \\ -s_{21}c_{32} - c_{21}s_{32}s_{13}e^{i\delta} & c_{21}c_{32} - s_{21}s_{32}s_{13}e^{i\delta} & s_{32}c_{13} \\ s_{21}s_{32} - c_{21}c_{32}s_{13}e^{i\delta} & -c_{21}s_{32} - s_{21}c_{32}s_{13}e^{i\delta} & c_{32}c_{13} \end{pmatrix}$

where $R_{jk}(\theta_{jk})$ describes a rotation in the *jk*-plane through angle θ_{jk} , $U_{\delta} = \text{diag}(e^{i\delta/2}, 1, e^{-i\delta/2})$, and $s_{jk} = \sin \theta_{jk}$, $c_{jk} = \cos \theta_{jk}$. We have omitted two additional Majorana phases, as they do not affect neutrino oscillations.

flavor fits to data give the following (1σ) and 3σ ranges for the PDG mixing angles [3]:

$$\begin{split} \sin^2\theta_{21} &= 0.32 \ (\pm 0.02) \ ^{+0.08}_{-0.06} \ , \\ \sin^2\theta_{32} &= 0.45 \ (\pm 0.07) \ ^{+0.20}_{-0.13} \ , \\ \sin^2\theta_{13} \ < \ (0.02) \ 0.050 \ . \end{split}$$



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Figure 2.4: Possible neutrino spectra: (a) normal (b) inverted.

The central values of these inferred mixing angles are quite consistent with the TriBiMaximal values [4], given

TriMinimal Parametrization of Neutrino Mixing Pakvasa, Rodejohann, TJW [PRL (2008)]

 $U_{\rm TBM} = R_{32} \left(\frac{\pi}{4}\right) R_{21} \left(\sin^{-1}\frac{1}{\sqrt{3}}\right) = \frac{1}{\sqrt{6}} \left(\begin{array}{ccc} 2 & \sqrt{2} & 0\\ -1 & \sqrt{2} & \sqrt{3}\\ 1 & -\sqrt{2} & \sqrt{3} \end{array}\right)$

The TriMinimal parametrization is given by

$$U_{\text{TMin}} = R_{32} \left(\frac{\pi}{4}\right) U_{\epsilon}(\epsilon_{32}; \epsilon_{13}, \delta; \epsilon_{21}) R_{21} \left(\sin^{-1} \frac{1}{\sqrt{3}}\right)$$
$$= \left(\frac{\sqrt{2} \quad 0 \quad 0}{0 \quad 1 \quad 1}\right) \frac{U_{\epsilon}}{\sqrt{6}} \left(\frac{\sqrt{2} \quad 1 \quad 0}{-1 \quad \sqrt{2} \quad 0}\right),$$
with $U_{\epsilon} = R_{32}(\epsilon_{32}) U_{\delta}^{\dagger} R_{13}(\epsilon_{13}) U_{\delta} R_{21}(\epsilon_{21})$ (8)

chosen to have just the form of the PDG parametrization. And just as in the PDG parametrization, U_{ϵ} is unitary, and therefore so is TriMinimal U_{TMin} . The simplicity of Eq. (8) is a fortuitous result of the fact that it is the middle rotation angle (θ_{13}) in the PDG parametrization that is set identically to zero in the TriBiMaximal scheme. the small ϵ_{jk} are obtained from the (1σ) 3σ ranges of the large oscillation angles in Eq. (6). The allowed ranges are:

- $-0.08 (-0.04) \le \epsilon_{21} \le (0.01) \ 0.07$, (13)
- $-0.18 (-0.10) \le \epsilon_{32} \le (0.04) \ 0.15$, (14)
 - $|\epsilon_{13}| \le (0.14) \ 0.23$, (15)

while the CP-invariant lies in the range $|J_{CP}| \leq$ (0.03) 0.05.

TriMinimal U

$$\begin{split} U_{\rm TMin} &= U_{\rm TBM} - \frac{\epsilon_{21}}{\sqrt{6}} \begin{pmatrix} \sqrt{2} & -2 & 0 \\ \sqrt{2} & 1 & 0 \\ -\sqrt{2} & -1 & 0 \end{pmatrix} - \frac{\epsilon_{32}}{\sqrt{6}} \begin{pmatrix} 0 & 0 & 0 & 0 \\ -1 & \sqrt{2} & -\sqrt{3} \\ -1 & \sqrt{2} & \sqrt{3} \end{pmatrix} - \frac{\epsilon_{13}}{\sqrt{6}} \begin{pmatrix} 0 & 0 & -\sqrt{6} e^{-i\delta} \\ \sqrt{2} e^{i\delta} & e^{i\delta} & 0 \end{pmatrix} \\ &- \frac{\epsilon_{21}^2}{2\sqrt{6}} \begin{pmatrix} 2 & \sqrt{2} & 0 \\ -1 & \sqrt{2} & 0 \\ 1 & -\sqrt{2} & 0 \end{pmatrix} - \frac{\epsilon_{32}^2}{2\sqrt{6}} \begin{pmatrix} 0 & 0 & 0 \\ -1 & \sqrt{2} & \sqrt{3} \\ 1 & -\sqrt{2} & \sqrt{3} \end{pmatrix} - \frac{\epsilon_{13}^2}{2\sqrt{6}} \begin{pmatrix} 2 & \sqrt{2} & 0 \\ 0 & 0 & \sqrt{3} \\ 0 & 0 & \sqrt{3} \end{pmatrix} \\ &- \frac{\epsilon_{21}\epsilon_{32}}{\sqrt{6}} \begin{pmatrix} 0 & 0 & 0 \\ -\sqrt{2} & -1 & 0 \\ -\sqrt{2} & -1 & 0 \end{pmatrix} - \frac{\epsilon_{21}\epsilon_{13}e^{i\delta}}{\sqrt{6}} \begin{pmatrix} 0 & 0 & 0 \\ -1 & \sqrt{2} & 0 \\ -1 & \sqrt{2} & 0 \end{pmatrix} - \frac{\epsilon_{32}\epsilon_{13}e^{i\delta}}{\sqrt{6}} \begin{pmatrix} 0 & 0 & 0 \\ \sqrt{2} & 1 & 0 \\ -\sqrt{2} & -1 & 0 \end{pmatrix} + \mathcal{O}(\epsilon^3) \end{split}$$

- Unitary to order $(eps)^2$
- General, with three (mini-) angles and a phase

Long-Distance QM evolves to Classical Probabilities

of classical probabilities, defined by $(\underline{U})_{\alpha j} \equiv |U_{\alpha j}|^2$. The full matrix of squared elements, through order $O(\epsilon^2)$, is

$$\underline{U}_{\mathrm{TMin}} = \frac{1}{6} \left\{ \begin{pmatrix} 4 & 2 & 0 \\ 1 & 2 & 3 \\ 1 & 2 & 3 \end{pmatrix} - E_1 \begin{pmatrix} 0 & 0 & 0 \\ -1 & 1 & 0 \\ 1 & -1 & 0 \end{pmatrix} - E_2 \begin{pmatrix} 2 & -2 & 0 \\ -1 & 1 & 0 \\ -1 & 1 & 0 \end{pmatrix} - 2\epsilon_{32} \begin{pmatrix} 0 & 0 & 0 \\ 1 & 2 & -3 \\ -1 & -2 & 3 \end{pmatrix} - \epsilon_{13}^2 \begin{pmatrix} 4 & 2 & -6 \\ -2 & -1 & 3 \\ -2 & -1 & 3 \end{pmatrix} \right\}$$

where $E_1 = 2\sqrt{2} \epsilon_{13} \cos \delta + 2 \epsilon_{21} (\epsilon_{13} \cos \delta - 2\sqrt{2} \epsilon_{32})$, and $E_2 = 2\sqrt{2} \epsilon_{21} + \epsilon_{21}^2$. That there are four independent terms in (17) reflects the fact that there are four independent moduli in the neutrino mixing matrix [9].

Some useful results follow immediately from this matrix. For example, in models where neutrinos are unstable, only the lightest neutrino mass-eigenstate arrives at Earth from cosmically-distant sources [10]. Flavor ratios at Earth for the normal mass-hierarchy are $\underline{U}_{e1}: \underline{U}_{\mu 1}: \underline{U}_{\tau 1}$, and for the inverted mass-hierarchy are $\underline{U}_{e3}: \underline{U}_{\mu 3}: \underline{U}_{\tau 3}$. These ratios may be read off directly from the 1st and 3rd columns of Eq. (17). As another example, emanating solar neutrinos are nearly pure ν_2 mass states [11]; consequently, their flavor ratios at Earth are mainly given by the 2nd column of (17).



Figure 2.4: Possible neutrino spectra: (a) normal (b) inverted.

$$\begin{array}{l} \begin{array}{c} \mbox{Decohering the PMNS/tribimaximal}\\ neutrino-mixing matrix \end{array} \\ \hline v_1 & v_2 & v_3 \\ \hline v_1 & v_2 & v_3 \\ \hline v_1 & v_2 & v_3 \\ \hline v_2 & v_3 \\ \hline v_2 & v_4 \\ \hline v_2 & v_1 \\ \hline v_1 & v_2 \\ \hline v_2 & v_1 \\ \hline v_1 & v_2 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_2 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline v_1 & v_1 \\ \hline v_2 & v_1 \\ \hline v_1 & v_1 \\ \hline$$

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Democracy Broken:

1. Galactic β -beam

- 2. Source dynamics (p-p vs. p-gamma) (partial pion-decay chain)
- 3. V decay (15 minutes of fame)
- 4. Vacuum resonance (MaVaNs, LIV vector)
- 5. Pseudo-Dirac ν oscillations

E.g., pure β beam evolves to $\underline{\nu}_{e} \rightarrow 2/3 \underline{\nu}_{1} + 1/3 \underline{\nu}_{2}$ $\rightarrow \underline{\nu}_{e} : \underline{\nu}_{\mu} : \underline{\nu}_{\tau} = 5:2:2$

Mixing (phase-averaged oscillations)



$$P(\nu_{\mu} \rightarrow \nu_{e}) = \left(\sum_{j} U_{\mu j} U_{ej}^{*} e^{-i\phi_{j}}\right) \left(\sum_{k} U_{\mu k} U_{ek}^{*} e^{-i\phi_{k}}\right)^{*}.$$

$$P(\nu_{\mu} \to \nu_{e}) \stackrel{\text{phase average}}{\longrightarrow} \left(\sum_{j} U_{\mu j} U_{ej}^{*} \right) \left(\sum_{k} U_{\mu k} U_{ek}^{*} \right)^{*} \times \delta_{jk} = \sum_{j} |U_{\mu j}|^{2} |U_{ej}|^{2}.$$

Let $(\underline{U})_{\alpha j} \equiv |U_{\alpha j}|^2$. Then,

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) \xrightarrow{\text{phase average}} (\underline{U})(\underline{U})^T$$
.

Neutrino Flavor Ratios for various Astro processes

Process:	Flavors at Source	Flavors at Earth
	$(\nu_e:\nu_\mu:\nu_\tau)$	$(\nu_e:\nu_\mu:\nu_\tau)$
Complete $\pi^{\pm} \to e^{\pm} \stackrel{(-)}{\nu_e} \nu_{\mu} \overline{\nu}_{\mu}$ decay	1:2:0	1:1:1
Incomplete $\pi^{\pm} \to \mu^{\pm} \stackrel{(-)}{\nu_{\mu}}$	0:1:0	4:7:7
β -beam	pure $\overline{\nu}_e$	5:2:2
Virtual Black Hole Spacetime	Any	1:1:1

plus TriMinimal corrections !

(to be determined)

Summary

Neutrinos from the Cosmos:

First non-atmospheric neutrino event is "just around the corner" [like the Higgs and SUSY??]

Next one to ten years will be critical, and, the deities/gods willing, most fruitful !

If the statistics are forthcoming, neutrnio flavor can play a big role.

And the TriMinimal Parametrization keeps it simple.