

# *Eikonal contributions to ultrahigh energy neutrino-nucleon cross sections in LSG models*

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## Motivation

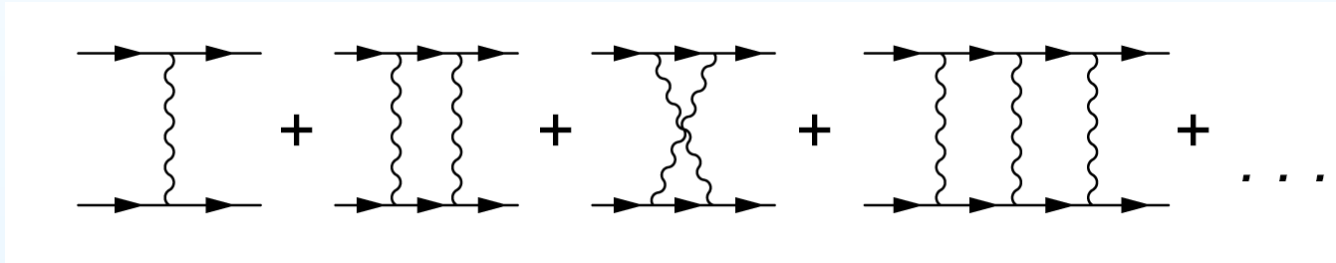
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- There are models that predict the existence of ultra-high energy neutrinos (up to  $10^{20}$  eV).
- The design of neutrino telescopes depends critically on the estimates of neutrino cross sections in the UHE regime.
- The systematic study of the scattering of such high energy neutrinos with baryonic matter might allow us to test new physics effects.
- Calculations of the  $\nu - N$  cross section exist, based on QCD (for example Gandhi, Quigg, Reno and Sarcevic 1998) and new physics effects (for example Jain *et al.* 2002; Hussain and McKay 2004).

# Eikonal approximation

It's a technique for estimating the high energy behavior of a forward scattering amplitude.

In QFT it is obtained by summing the ladder diagrams for boson exchange at all orders.



- $s \gg t$ .
- The deflecting angle is small  $p_1 \sim p'_1, p_2 \sim p'_2$ .
- The approximation doesn't discriminate the spin of the incoming particles which can be considered as scalar fields.

# Eikonal approximation

The tree-level amplitude  $A_{Born}$  is:

$$A_{Born}(t) = g^2 s^r \frac{1}{t - m^2},$$

$r = 1, 2$  rank of boson propagator.

Define:

$$\chi(b_{\perp}) = \frac{1}{2s} \int \frac{d^2 q_{\perp}}{(2\pi)^2} e^{-iq_{\perp} \cdot b_{\perp}} A_{Born}(q_{\perp}^2)$$

And obtain:

$$\begin{aligned} A_{eik} &= -2si \int d^2 b_{\perp} e^{iq_{\perp} \cdot b_{\perp}} \left( \sum_1^{\infty} (i\chi)^n / n! \right) \\ &= -2si \int d^2 b_{\perp} e^{iq_{\perp} \cdot b_{\perp}} (e^{i\chi} - 1). \end{aligned}$$

## Born term divergence in the ADD model

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Even at tree-level  $A_{Born}$  has ultraviolet divergence for  $n \geq 2$

$$A_{Born} = \frac{s^2}{M_D^{n+2}} \int \frac{d^n m}{t - m^2}$$

.

Solution can be a cut-off of the order of  $M_D$  (Giudice, Rattazzi and Wells 1998, 2001. Han, Lykken and Zhang 1999).

This is the upper limit of the effective theory in the hypothesis that the tension of the brane  $M_S$  is infinite .

It can be the finite tension of the brane  $M_S$ , if the standard model fields extend by a length  $\sim 1/M_S$  in the bulk

(Bando, Kugo, Noguchi and Yoshioka 1999).

## Infinite tension case

After dimensional regularization of the Born term, one obtains for the eikonal phase (Giudice, Rattazzi and Wells 2001):

$$\chi = \left[ \frac{(4\pi)^{\frac{n}{2}-1} s \Gamma(\frac{n}{2})}{2M_D^{n+2}} \right] \frac{1}{b^n} \equiv \left( \frac{b_c}{b} \right)^n .$$

Eikonal amplitude:

$$A_{eik} = 4\pi s b_c^2 F_n(b_c q),$$

$$F_n(\eta) = -i \int_0^\infty d\xi \xi J_0(\xi \eta) (e^{i\xi^{-n}} - 1).$$

Cross section:

$$\frac{d^2\sigma}{dx dy} = \sum_{i=\text{flavors}} x f_i(x, xys) s \pi b_c^4(\hat{s}) |F_n(b_c(\hat{s}) \sqrt{xys})|^2$$

## Case of a finite tension $M_S$

If the standard model wave function has a gaussian extra-dimensional extension, one can correct the vertex by

$e^{-\frac{m^2}{2M_S^2}}$  (Lonnblad and Sjudahl 2006). One gets:

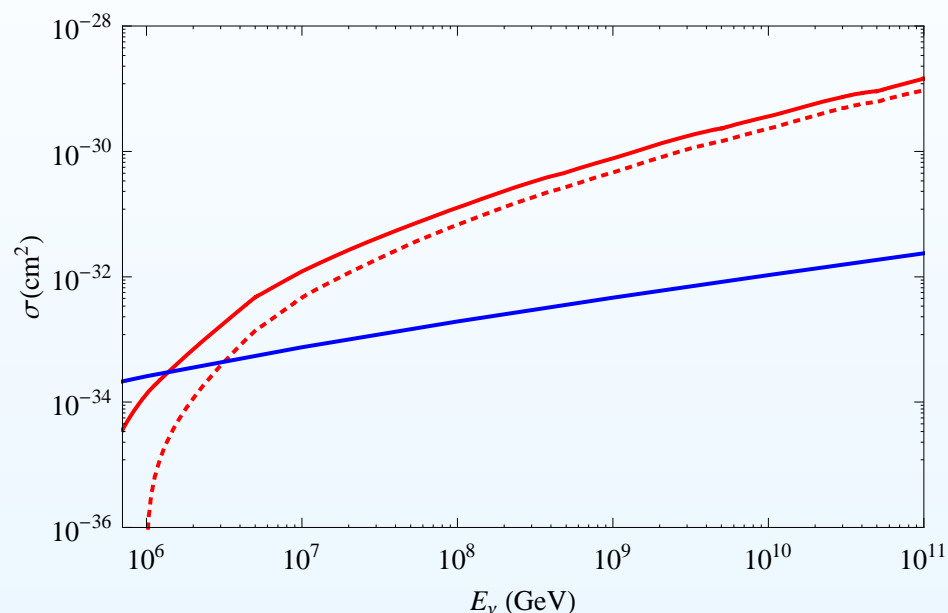
$$A_{Born} = \frac{s^2}{M_D^{n+2}} S_n \int_0^\infty \frac{m^{n-1} dm}{t - m^2} e^{-\frac{m^2}{M_S^2}}.$$

$$A_{eik} = -4\pi i s \int_0^\infty db b J_0(qb) \left[ e^{i\left(\frac{b_c M_S}{2}\right)^n U\left(\frac{n}{2}, 1, \frac{M_S^2 b^2}{4}\right)} - 1 \right].$$

$$\frac{d^2\sigma}{dx dy} = \sum_{i=\text{flavors}} x f_i(x, xys) s \pi |G_n(\sqrt{xys})|^2, \text{ with}$$

$$G_n(z) = -i \int_0^\infty db b J_0(bz) \left[ e^{i\left(\frac{b_c M_S}{2}\right)^n U\left(\frac{n}{2}, 1, \frac{M_S^2 b^2}{4}\right)} - 1 \right].$$

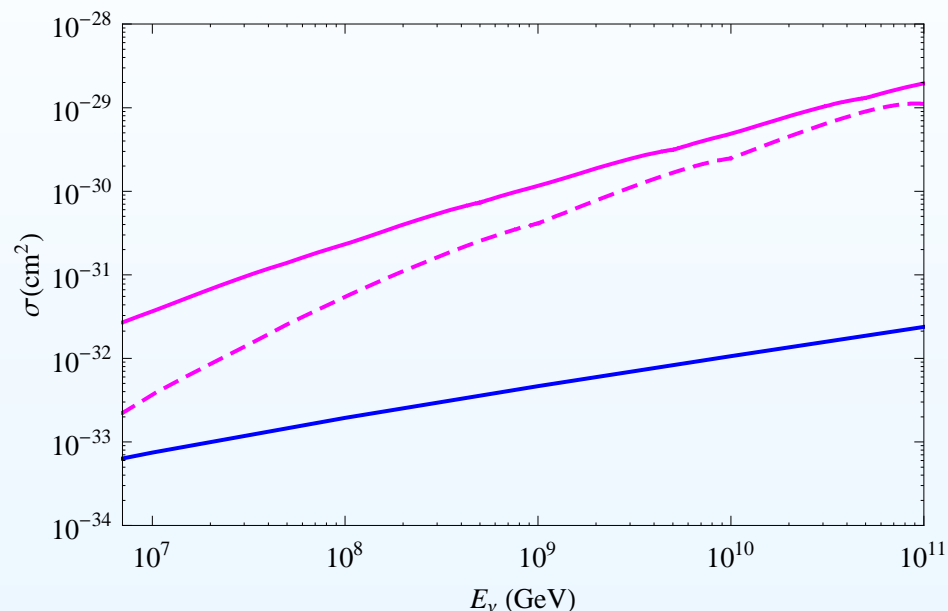
# Cross Sections



- **SOLID RED:** Full eikonal corrected by the Born term,  $M_D = 10^3$  GeV,  $M_S = \infty$ ,  $n = 5$
- **DASHED RED:** Saddle point approximation corrected by the Born term.
- **BLUE:** Standard model neutral current

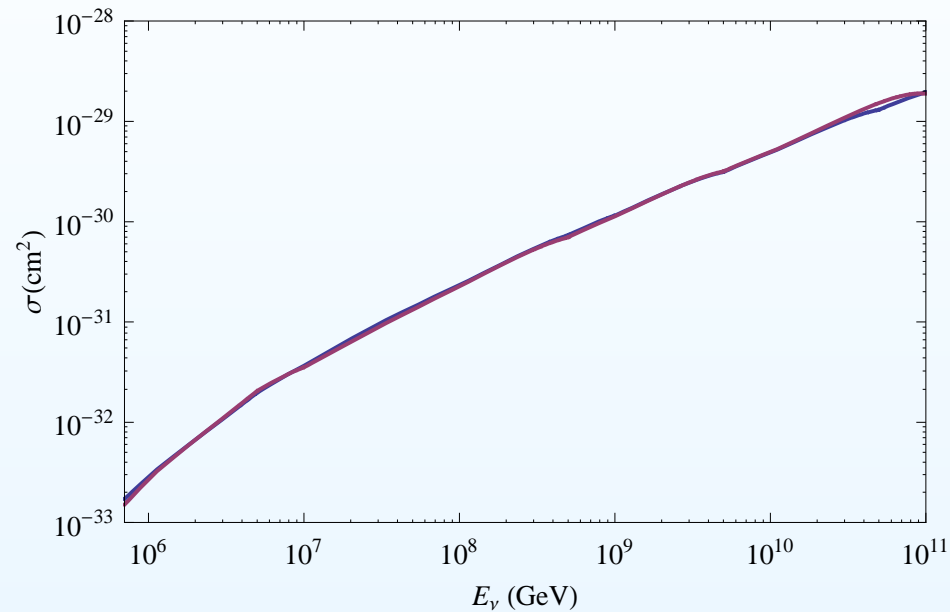


# High energy pure eikonal/ Branes with different tension



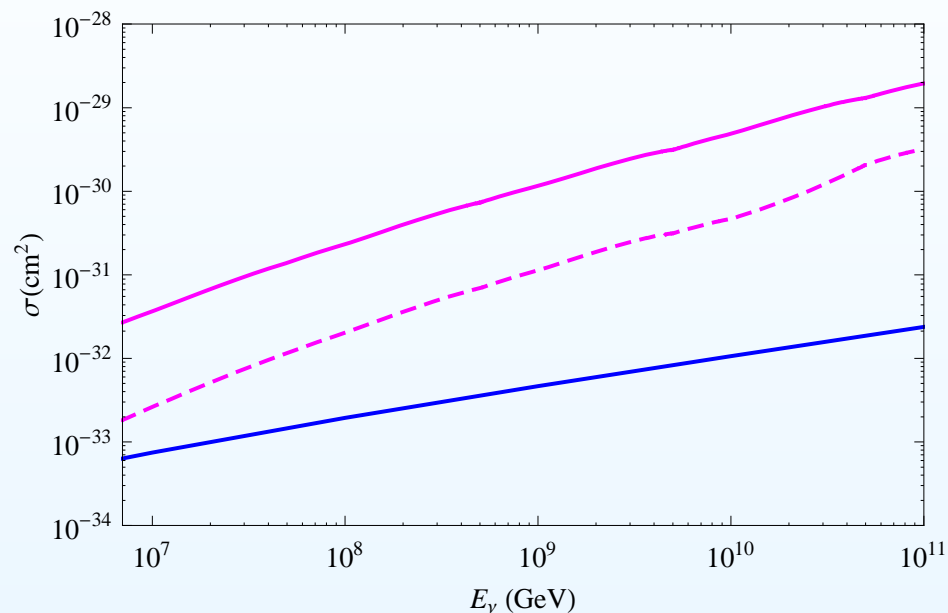
- SOLID MAGENTA:  $M_D = 10^3$  GeV,  $M_S = \infty$ ,  $n = 5$
- DASHED MAGENTA:  $M_D = 10^3$  GeV,  $M_S = 10^3$  GeV,  $n = 5$
- BLUE: Standard model neutral current

## Pure eikonal LSG 1 TeV/ Branes with different tension



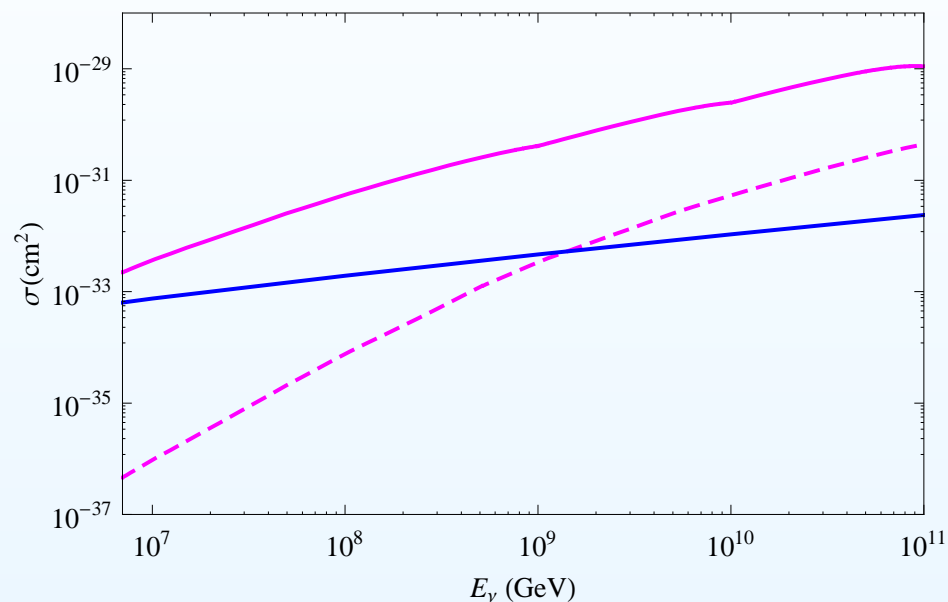
Comparison at  $M_D = 10^3$  GeV between infinite tension and finite  $M_S = 10^4$  GeV tension.  $n = 5$ .

# High energy pure eikonal for different LSG



- SOLID MAGENTA:  $M_D = 10^3$  GeV,  $M_S = \infty$ ,  $n = 5$
- DASHED MAGENTA:  $M_D = 2 \times 10^3$  GeV,  $M_S = \infty$ ,  $n = 5$
- BLUE: Standard model neutral current

## High energy pure eikonal for different LSG



- SOLID MAGENTA:  $M_D = 10^3$  GeV,  $M_S = 10^3$  GeV
- DASHED MAGENTA:  $M_D = 2 \times 10^3$  GeV,  $M_S = 10^3$  GeV
- BLUE: Standard model neutral current

## Discussion

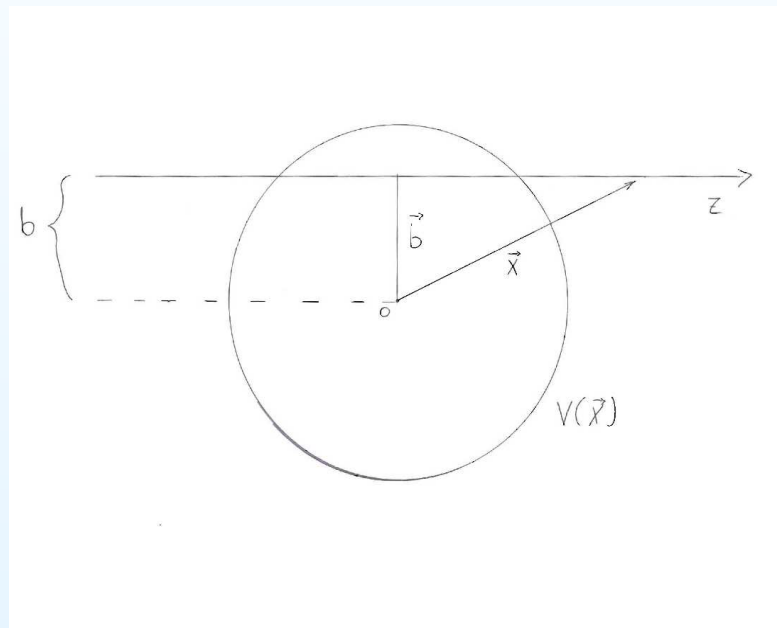
- For the infinite tension case the full eikonal is larger than the saddle point by a factor 2.6 - 1.5
- Three order of magnitude enhancement of the standard model at  $E_\nu = 10^{11}$  GeV
- A "softer" brane suppresses the eikonal cross section by a factor 10 - 2
- Almost perfect overlapping at  $M_S = 10^4$  GeV
- For the thin brane, doubling the LSG suppresses the cross sections by one order of magnitude
- For the thick brane, doubling the LSG lowers the cross section below the standard model at  $E_\nu \leq 10^9$  GeV

## Conclusions

- The brane tension  $M_S$  greatly affects the size of LSG influence on UHE neutrino scattering
- Calculations performed with brane tension of 10 TeV and above are indistinguishable
- The saddle point needs  $qb_c \gg 1$ .  
Even with the maximal  $q$  consistent with the eikonal the saddle point conditions are only marginally satisfied at GZK energies
- Eikonal collisions highly elastic  $\Rightarrow$  consequences for event rate estimates not severe since deposited energy small and corresponding events mixed with highly inelastic processes (black-hole) at energies an order of magnitude lower

## The EA in non-relativistic QM

In non-relativistic QM the interaction is described by the potential  $V(\vec{x})$ . We assume  $E \gg V(\vec{x})$ , so that the particle moves in a straight line.



The wave function is given by:

## The EA in non-relativistic QM

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$$\psi(\vec{x}) \sim e^{iS(\vec{x})}$$

The Hamilton-Jacobi equation is (if  $E \gg V$ ):

$$\frac{(\nabla S)^2}{2m} + V = E = \frac{p^2}{2m}$$

Solve along straight path  $z$ :

$$S \sim pz - \frac{m}{p} \int_{-\infty}^z V(\sqrt{b^2 + z'^2}) dz'$$

So that:

$$\psi(\vec{b} + z\hat{z}) \sim \frac{1}{(2\pi)^{3/2}} e^{ipz} e^{\frac{-im}{p} \int_{-\infty}^z V(\sqrt{b^2 + z'^2}) dz'}$$



## The EA in non-relativistic QM

Remember the Scattering Amplitude:

$$f(\vec{p}, \vec{p}') = -(2\pi)^2 m \langle \vec{p}' | V | \vec{p} \rangle$$

With the previous approx it becomes:

$$f(\vec{p}, \vec{p}') = -\frac{m}{2\pi} \int d^3 x' e^{-i\vec{p}' \cdot \vec{x}'} V(\sqrt{b^2 + z'^2}) e^{i\vec{p} \cdot \vec{x}'} \\ \times e^{\frac{-im}{p} \int_{-\infty}^{z'} V(\sqrt{b^2 + z''^2}) dz''}$$

This is valid for  $E \rightarrow \infty$  with  $\vec{q} = \vec{p}' - \vec{p}$  kept finite, so that  $q_3 \sim 0$ .

## The EA in non-relativistic QM

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We can carry out the integration in  $z$  to obtain:

$$\begin{aligned} f(\vec{q}) &= -\frac{m}{2\pi} \int d^2b e^{-i\vec{q}\cdot\vec{b}} \int dz V(z) e^{\frac{-im}{p} \int_{-\infty}^z V(z') dz'} \\ &= -\frac{ip}{2\pi} \int d^2b e^{-i\vec{q}\cdot\vec{b}} [e^{i\chi(b)} - 1], \end{aligned}$$

where

$$\chi(b) = -\frac{m}{p} \int_{-\infty}^{\infty} V(\sqrt{b^2 + z^2}) dz$$

is the eikonal phase.