

Regularization-Independent Inflaton Spectrum

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With Leonard Parker

Supported by Lynde and Harry Bradley
Foundation and Wisconsin Space Grant

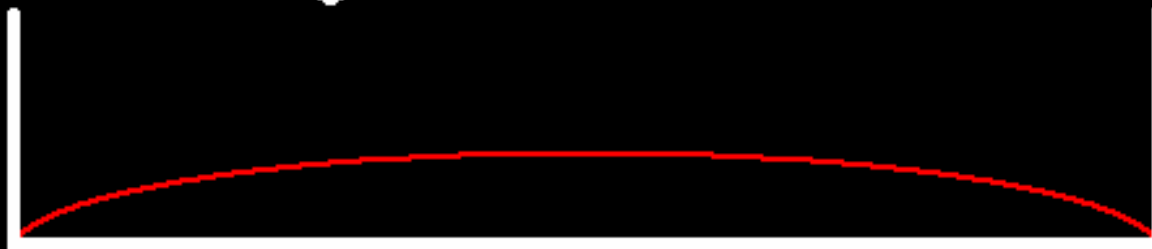
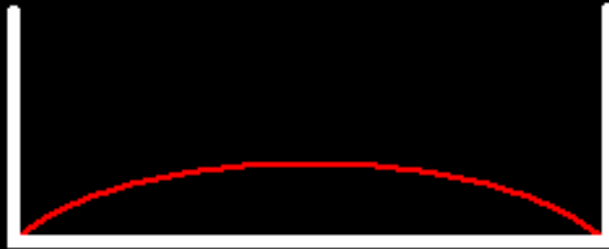
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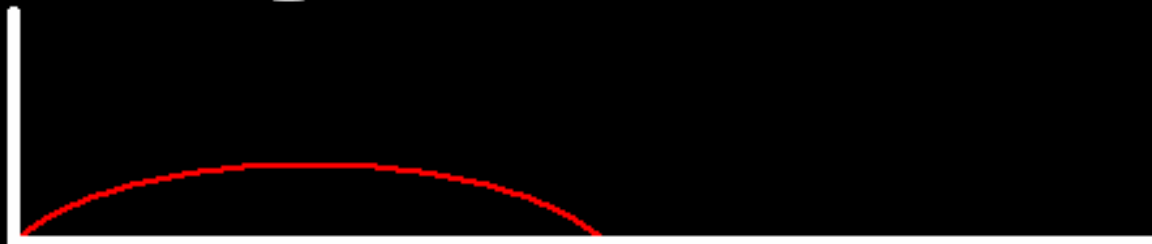


PARTICLE IN A BOX

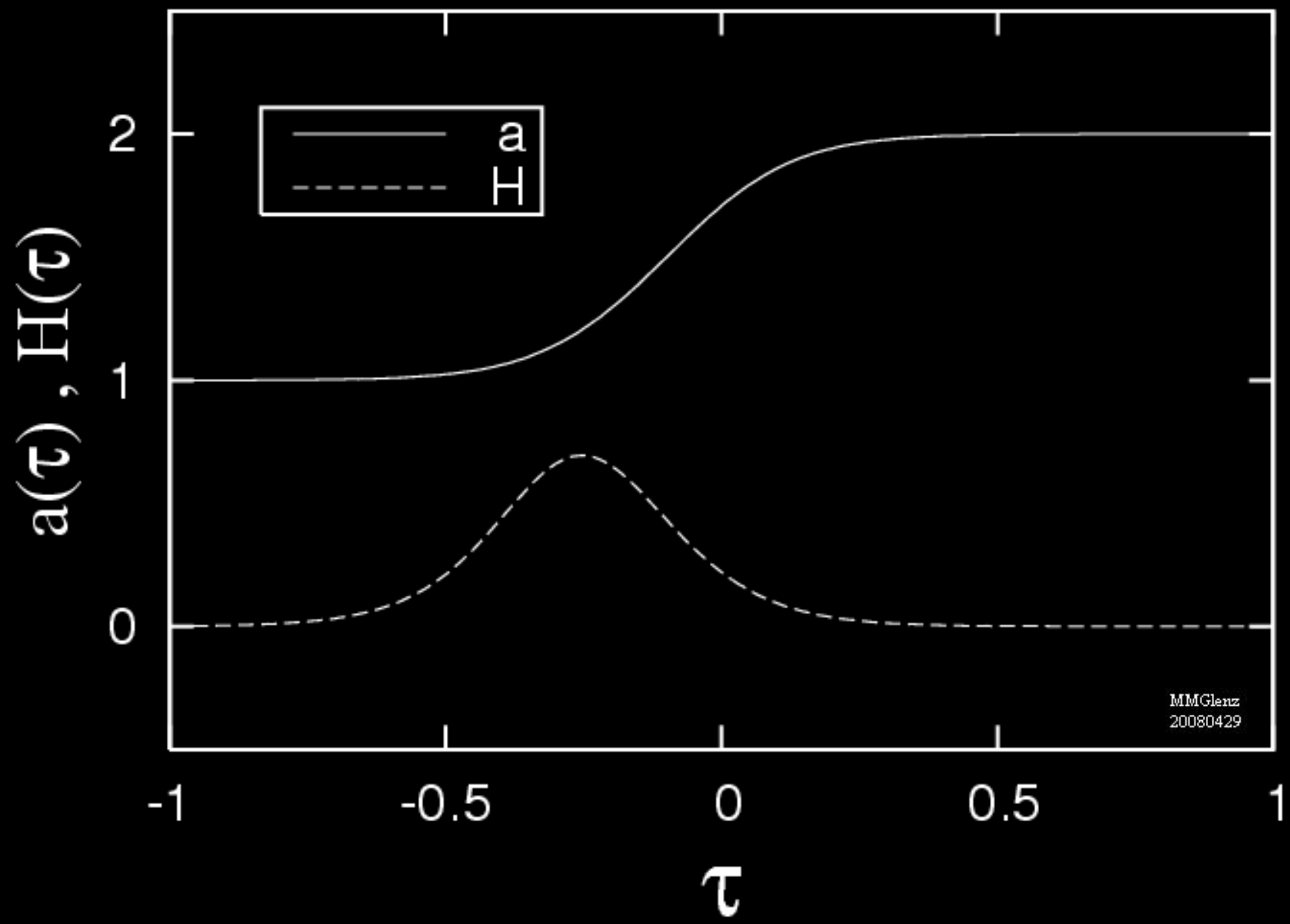


SLOW

OR



FAST



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EARLY- AND LATE-TIME
VACUUMS ARE RELATED
BY A BUGOLIUBOV
TRANSFORMATION.

$$a_{\vec{k}} = \alpha_k A_{\vec{k}} + \beta_k^* A_{-\vec{k}}^\dagger$$

$$\langle N_{\vec{k}} \rangle_{t \rightarrow \infty} = \langle 0 | a_{\vec{k}}^\dagger a_{\vec{k}} | 0 \rangle = |\beta_k|^2$$

DEFINITIONS:

$$H = \frac{da/dt}{a}$$

$$q_2 = \frac{k}{a_{final} H_{inflation}}$$

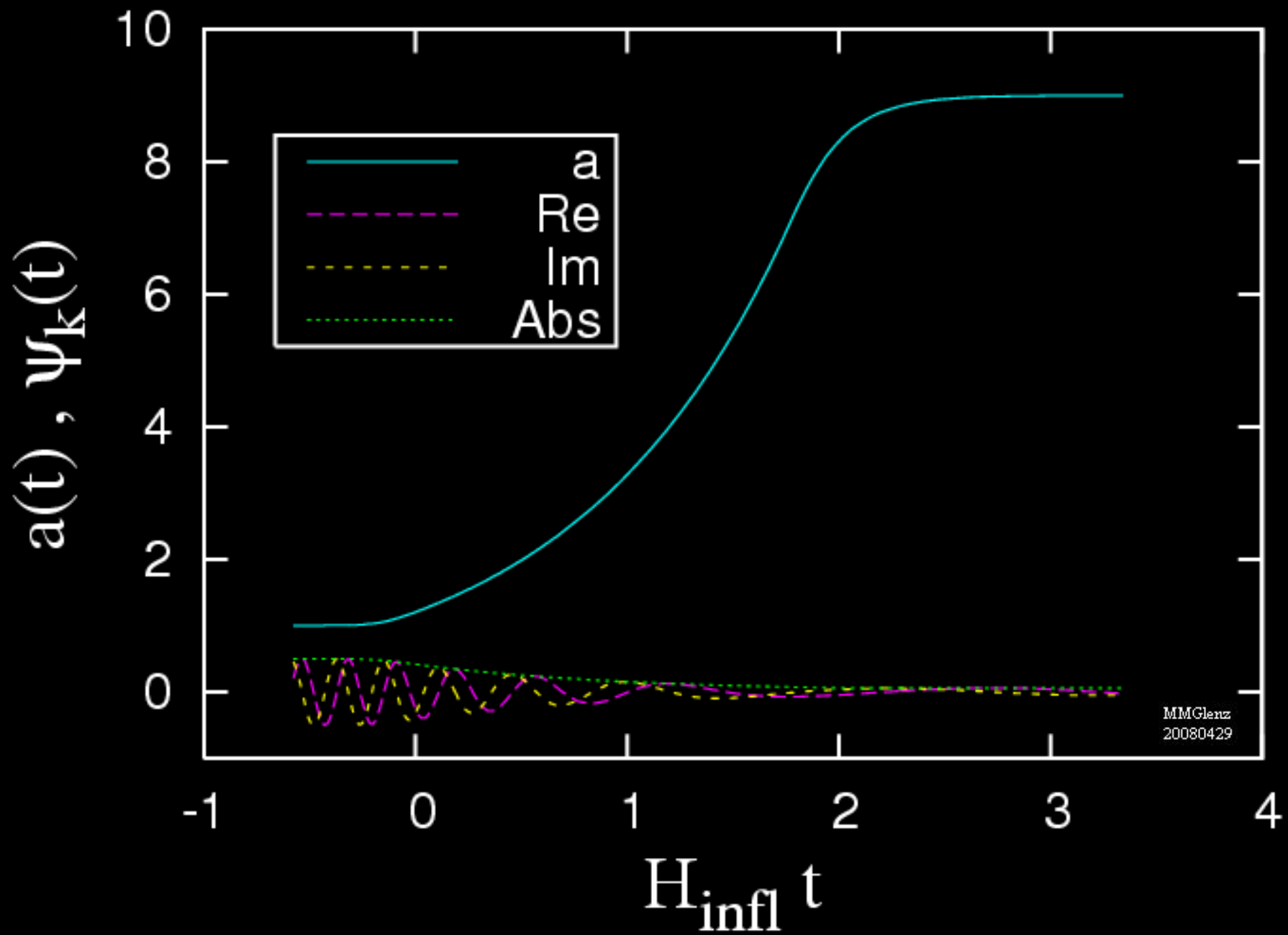
$$m_H = \frac{m}{H_{inflation}}$$

$$\text{e-folds} = \int H dt$$

WE WANT TO KNOW
THE LATE-TIME
PARTICLE PRODUCTION

$$\phi(\vec{x}, t) = \phi(t) + \delta\phi(\vec{x}, t)$$

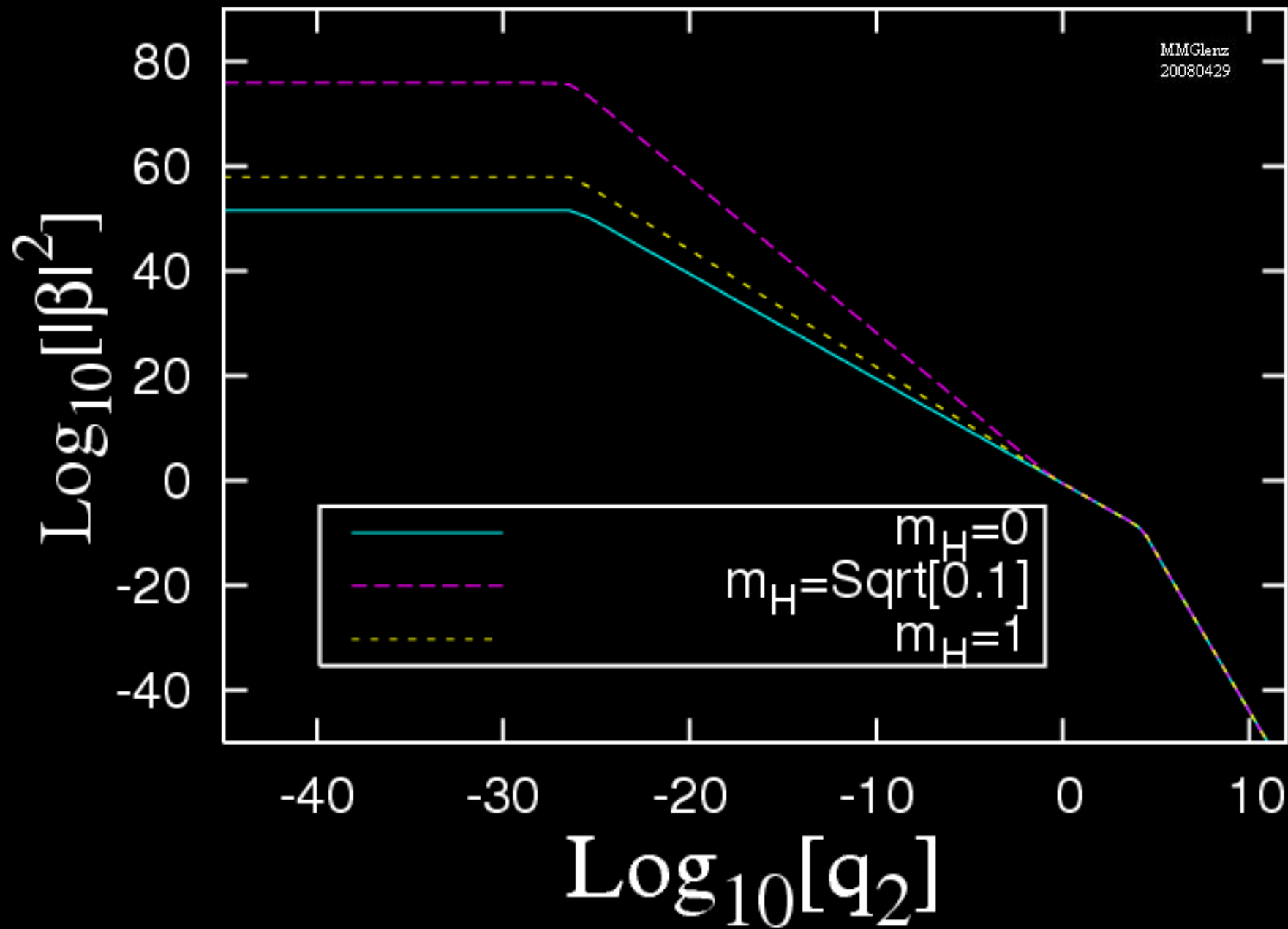
$$\delta\phi = (v)^{-1/2} \sum_{\vec{k}} [a_{\vec{k}} \psi_k(t) e^{i\vec{k} \cdot \vec{x}} + H.c.]$$



EVOLUTION EQUATION

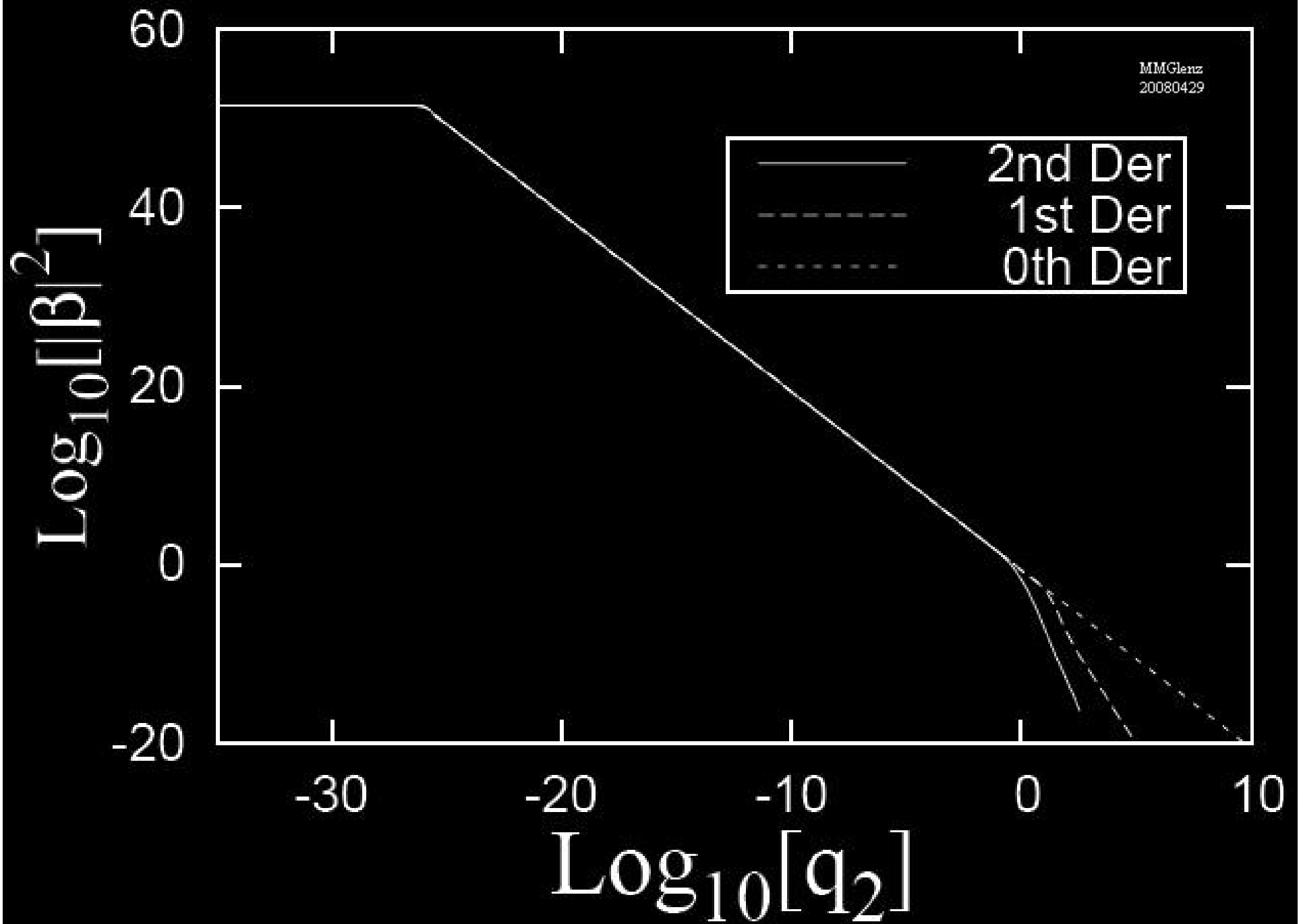
$$\partial_t^2 \delta\phi + 3H\partial_t \delta\phi - a^{-2}(t) \sum_{i=1}^3 \partial_i^2 \delta\phi + m(\phi^{(0)})^2 \delta\phi = 0$$

$$m(\phi^{(0)})^2 = \frac{d^2 V}{d(\phi^{(0)})^2}$$



$$n_s = 4 - \sqrt{9 - 4m_H^2}$$

$\delta_H: V = \frac{1}{2}m^2\phi^2$	$H = 10^{12}\text{GeV}$	$H = 10^{14}\text{GeV}$	$H = 10^{16}\text{GeV}$
$m_H = \frac{1}{10000}$	2.263×10^{-4}	2.263×10^{-2}	2.263×10^0
$m_H = \frac{1}{100}$	2.263×10^{-6}	2.263×10^{-4}	2.263×10^{-2}
$m_H = \frac{1}{10}$	2.258×10^{-7}	2.258×10^{-5}	2.258×10^{-3}
$m_H = \frac{1}{4}$	8.917×10^{-8}	8.917×10^{-6}	8.917×10^{-4}
$m_H = 1$	1.903×10^{-8}	1.903×10^{-6}	1.903×10^{-4}



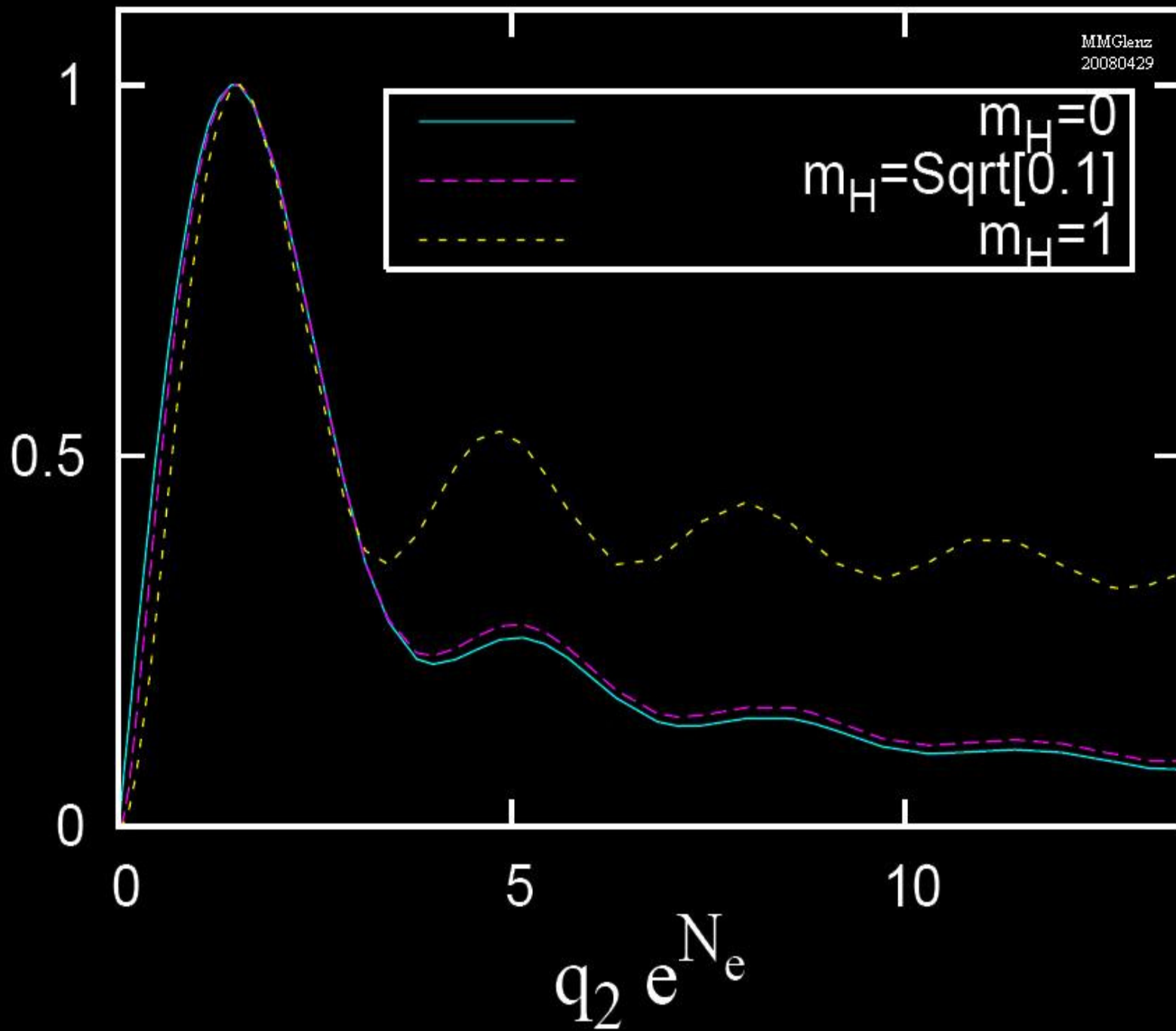
$$\langle E \rangle \simeq \frac{H_{infl}^4 e^{2Y}}{16\pi^2}$$

$$Y \sim \ln \left(\frac{a_{2f}}{a_{2f} - a_{1f}} \right)$$

$$\langle |\delta\phi^2| \rangle = \frac{1}{2(a_{2f}L)^3} \sum_k \left[\frac{1 + 2|\beta_k|^2}{\sqrt{(k/a_{2f})^2 + m^2}} \right]$$

$$Z \equiv \frac{q_2 |\beta_{q_2}|^2 H_{infl}^2}{2\pi^2 \sqrt{1 + \frac{m_H^2}{q_2^2}}}$$

Normalized Z/H^2



Thank You

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