# Higgs-Radion mixing in RS with Bulk fields

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## Outline

- Introduction
- The radion in RS1
- radion with matter in bulk
- Higgs-radion mixing
- Conclusions

## Introduction

- Warped Extra Dimensions: One compact extra dimension with warped geometry.
- Original setup: Two branes as boundaries and all SM fields on the TeV Brane  $\rightarrow$  RS1.
  - Towers of KK gravitons
  - Radion graviscalar
- More recent setups: Two branes, Higgs field on TeV brane, SM fields in the "bulk".
  - Towers of KK gravitons
  - Towers of KK SM fields
  - Radion graviscalar

- Radion couplings are higgs-like (except to gluons and photons)
- Radion might be the lightest new particle in warped scenarios
- When SM matter in the bulk, KK modes are constrained to be at  $\sim 3$  TeV. The radion could be the only accessible mode from these models.
- Radion can in principle mix with the Higgs

#### The Radion and its interactions

In the RS1 model  $_{[Randall,Sundrum,('98)]}$  the background metric  $g^o_{_{AB}}$  is defined by

$$ds^2 = e^{-2\sigma}\eta_{\mu\nu} dx^{\mu}dx^{\nu} + dy^2$$

with  $\sigma(y) = ky$  and such that a hierarchy is created between the two boundaries at y = 0 and  $y = \pi r_0$  is created.

The linear metric perturbations  $h_{AB}(x, y)$  can be reduced to

$$ds^{2} = \left(e^{-2\sigma}\eta_{\mu\nu} + \left[e^{-2\sigma}h_{\mu\nu}^{TT}(x,y) - \eta_{\mu\nu}r(x)\right]\right)dx^{\mu}dx^{\nu} + \left(1 + 2e^{2\sigma}r(x)\right)dy^{2}$$

(the graviscalar r(x) is massless. A stabilization mechanism providing it with mass is assumed for example[Golberger,Wise('99)])

#### INTERACTIONS

Matter-gravity interactions come from the matter action

$$S_{mat} = \int dx^5 \sqrt{-g} \ \mathcal{L}_{mat}$$

We expand this action in powers of the radion perturbation

$$S_{mat}(r^0) = \int dx^5 \sqrt{-g^{(0)}} \mathcal{L}_{mat}$$

$$S_{int}(r) = -\frac{1}{2} \int dx^5 \sqrt{-g^{(0)}} e^{2\sigma} \left( -T^{\mu}_{\ \mu} + 2T_{55} \right) r(x) \quad [\text{Rizzo}(02), \text{Csaki, Hubisz, Lee}(07)]$$

But the radion r(x) is NOT canonically normalized (canonical kinetic term).

The canonically normalized radion is  $\phi_r(x)\frac{2}{\Lambda_r} = e^{2k\pi r_0}r(x)$ where  $\Lambda_r = \sqrt{6}M_{Pl}e^{-k\pi r_0}$ 

#### $\mathbf{RS1}$ - Matter on the brane

Single radion interaction becomes

$$S_{int}(r) = \frac{1}{\Lambda_r} \int dx^4 T^{\mu}_{\ \mu} \phi_0(x) \quad \Rightarrow \text{Higgs-likecouplings!}$$

gluons 
$$-\frac{\alpha_s}{8\pi} \left[ \sum_i F_{1/2}(\tau_i)/2 - b_3 \right] \frac{\phi_0}{\Lambda_r} G_{\mu\nu} G^{\mu\nu}$$
  
photons  $-\frac{\alpha}{8\pi} \left[ \sum_i e_i^2 N_c^i F_i(\tau_i) - (b_2 + b_Y) \right] \frac{\phi_0}{\Lambda_r} F_{\mu\nu} F^{\mu\nu}$   
massive bosons  $\frac{\phi_0}{\Lambda_r} M_V^2 V^{\alpha} V_{\alpha}$   
fermions  $\frac{\phi_0}{\Lambda_r} m_f \bar{f} f$ 

#### The Radion and Matter in the bulk [Csaki,Hubisz,Lee(07)]

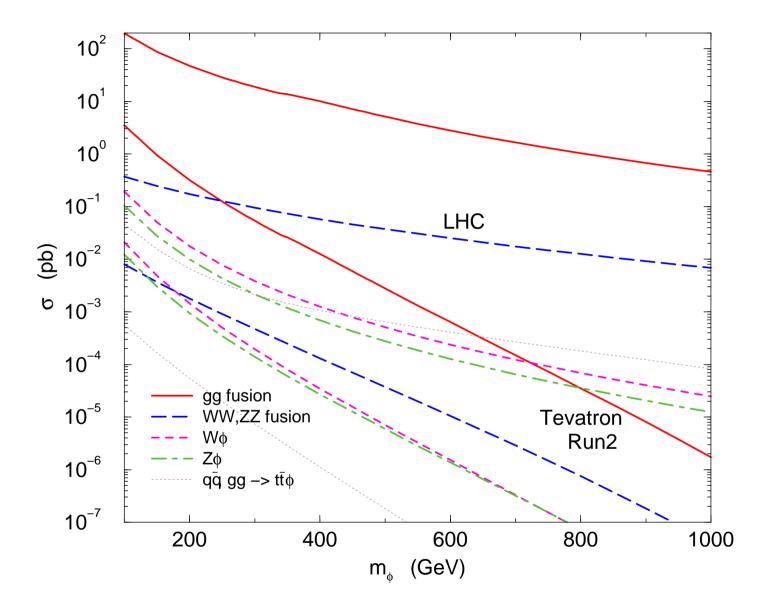
With gauge fields and fermions in the bulk (but Higgs on the TeV brane) we need the new interactions with the radion.

$$S_{int}(r) = -\frac{1}{2} \int dx^5 \sqrt{-g^{(0)}} e^{2\sigma} \left( -T^{\mu}_{\ \mu} + 2T_{55} \right) r(x)$$

- For Massless gauge fields:
  - The  $T_{55}$  term  $\Rightarrow$  tree level coupling *r*-glu-glu and *r*- $\gamma$ - $\gamma$ .
  - Brane localized kinetic terms for gauge fields.
  - Trace anomaly effect
  - Loop contributions (tops and W's)

$$\left[\frac{1-4\pi\alpha(\tau_{UV}^0+\tau_{IR}^0)}{4k\pi r_0}+\frac{\alpha}{8\pi}\left(b-\sum_i\kappa_iF_i(\tau_i)\right)\right]\frac{\phi}{\Lambda_r}F_{\mu\nu}F^{\mu\nu}$$

- Radion interaction with Massive Gauge bosons maintains its main contribution from the boson mass
- Interaction with fermions, although model dependent remains proportional to the mass of the fermion with an  $\mathcal{O}(1)$  coefficient
- Interaction with the higgs is computed as in RS1 since Higgs localized



(from K.Cheung ('00))

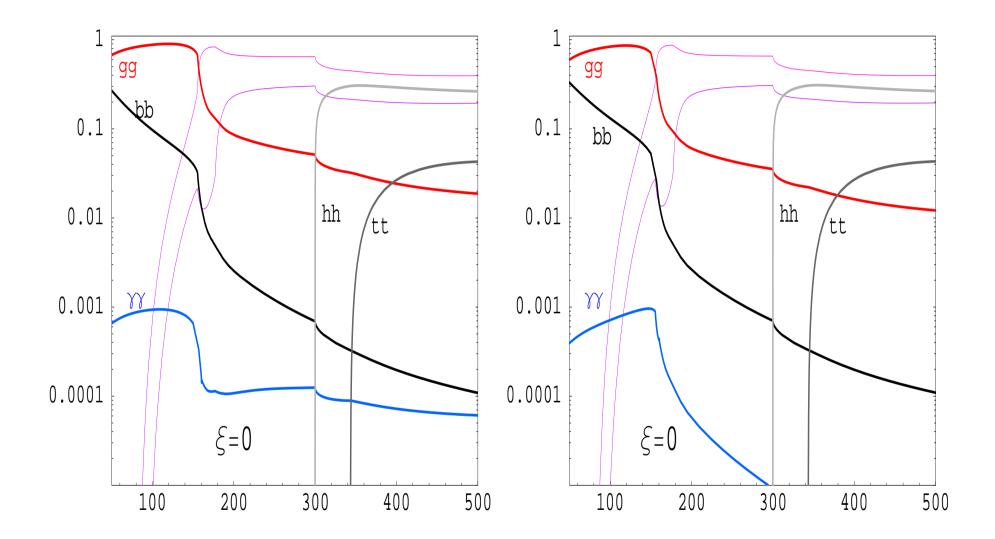


Figure 1: Branchings of the radion vs. its mass  $M_{\phi}$ 

#### **Higgs-radion mixing**

 $[{\rm Giudice, Rattazzi, Wells(00), \ Csaki, Graesser, Kribs(00), \ Han, Kribs, McElrath(01), \ Saki, Graesser, Kribs(00), \ Han, Kribs$ 

Rizzo, Hewett(02), Dominici, Gunion, Grzadkowski, MT(02)], Gunion, MT, Wells(03)]...

We now consider the brane operator:

$$S_{\xi} = \xi \int d^4x \sqrt{g_{ind}} R(g_{ind}) \ H_0^{\dagger} H_0 \,.$$

$$\mathcal{L}_{scalar} = -\frac{1}{2} \left\{ 1 + 6\xi \left( \frac{v_0}{\Lambda_r} \right)^2 \right\} \phi_0 \Box \phi_0 - \frac{1}{2} \phi_0 m_{\phi_0}^2 \phi_0$$
$$-\frac{1}{2} h_0 (\Box + m_{h_0}^2) h_0 - \frac{6\xi v}{\Lambda_r} h_0 \Box \phi_0$$

Radion mass added "by hand".

#### NORMALIZED HIGGS AND RADION PHYSICAL FIELDS

$$h_{0} = \left(\cos\theta - \frac{6\xi\gamma}{Z}\sin\theta\right)h + \left(\sin\theta + \frac{6\xi\gamma}{Z}\cos\theta\right)\phi \equiv dh + c\phi$$

$$\phi_{0} = \left(-\cos\theta\frac{1}{Z}\right)\phi + \left(\sin\theta\frac{1}{Z}\right)h \equiv a\phi + bh$$

with  $\gamma = \frac{v}{\Lambda}$  and Z and  $\theta$  depend on  $m_{\phi}$ ,  $m_h$ ,  $\Lambda$  and  $\xi$ .

 $\Rightarrow$  4 parameters in Higgs-radion sector:  $m_{\phi}$ ,  $m_h$ ,  $\Lambda$  and  $\xi$ 

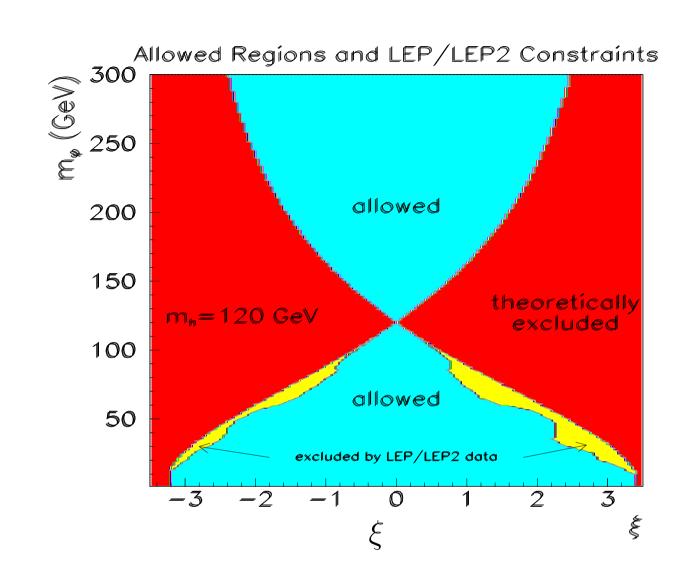


Figure 2: Allowed  $m_{\phi}$  and  $\xi$  for  $m_h = 120 \text{GeV}$  and  $\Lambda = 5 \text{TeV}$ 

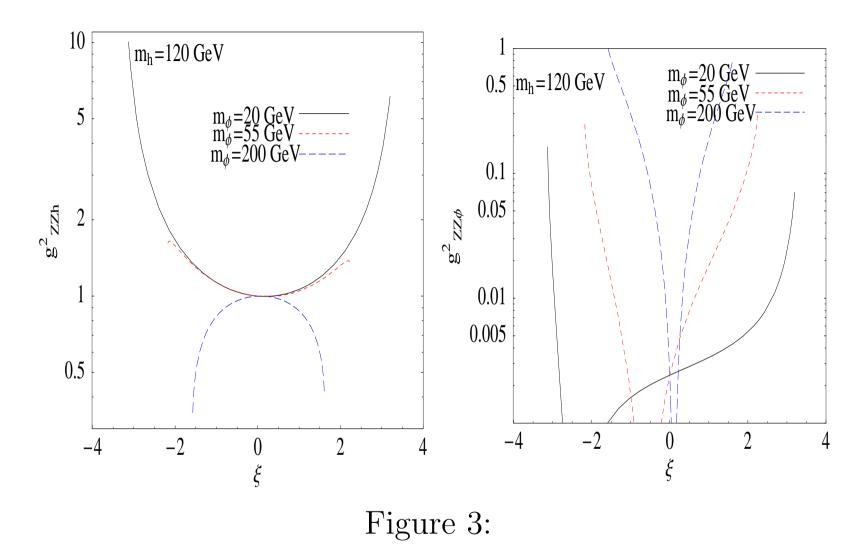
(from Dominici, Gunion, Grzadkowski, M.T. ('02))

#### VV and ff COUPLINGS

$$g_{ZZh} = \frac{g M_z}{c_W} (d + \gamma b) \qquad g_{ZZ\phi} = \frac{g M_z}{c_W} (c + \gamma a)$$
$$g_{f\bar{f}h} = -\frac{g m_f}{2 M_w} (d + \gamma b) \qquad g_{f\bar{f}\phi} = -\frac{g m_f}{2 M_w} (c + \gamma a)$$

Very interesting property of the  $\xi$ -mixing: the different couplings of the physical radion to matter photons, gluons, fermions and massive bosons can vanish at different points in parameter space.

 $\Rightarrow \phi$  can be photon-fobic, gluon-fobic or massive-fobic



(from Dominici, Gunion, Grzadkowski, M.T. ('02))

## Higgs-radion mixing & Matter in the bulk (Preliminary)Bulk MatterRS1-Matter on the brane

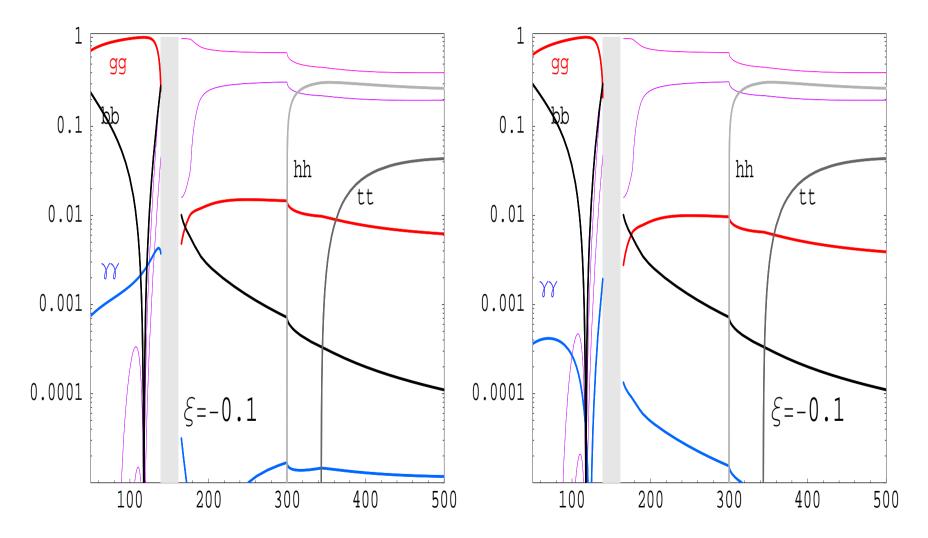


Figure 4: Branchings of the radion vs. its mass  $M_{\phi}$ 

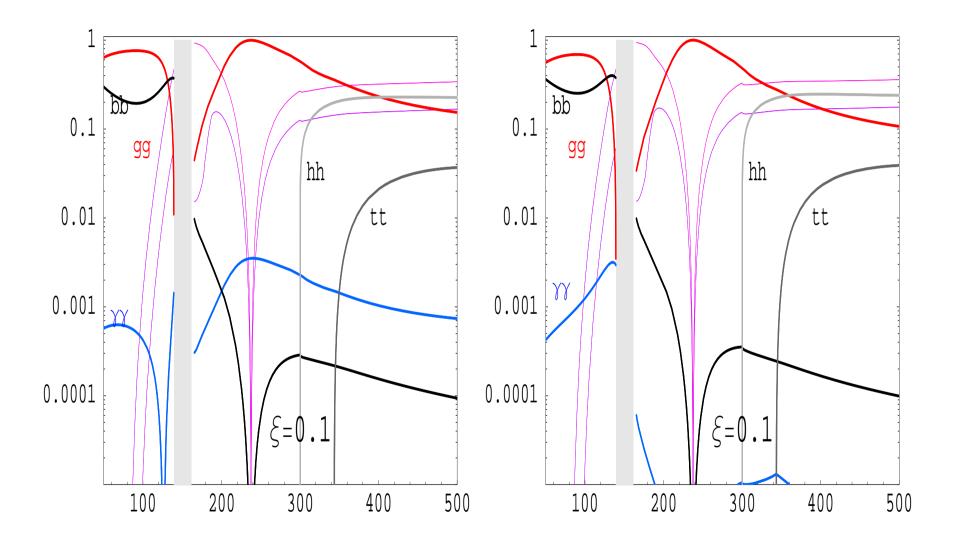


Figure 5: Branchings of the radion vs. radion mass  $M_{\phi}$ 

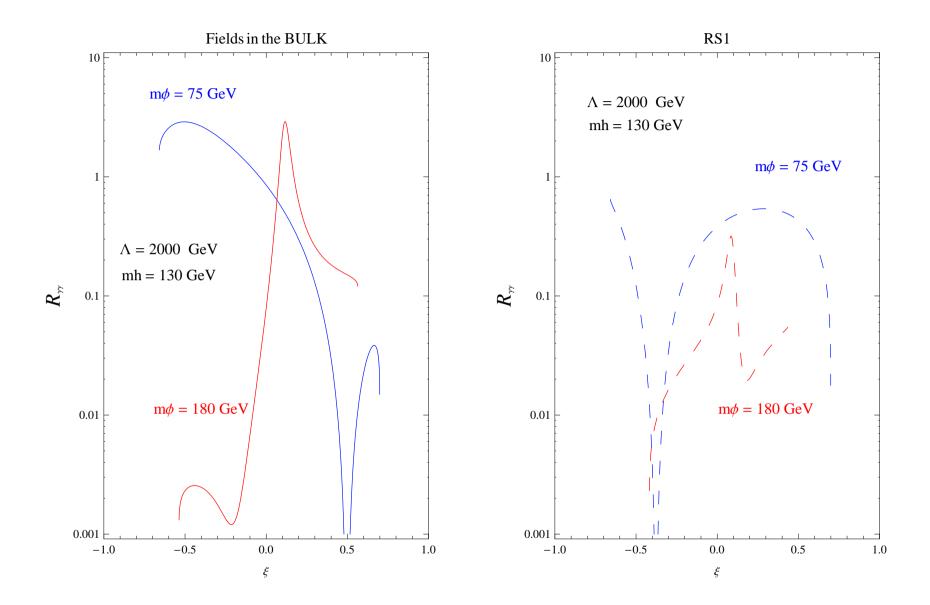
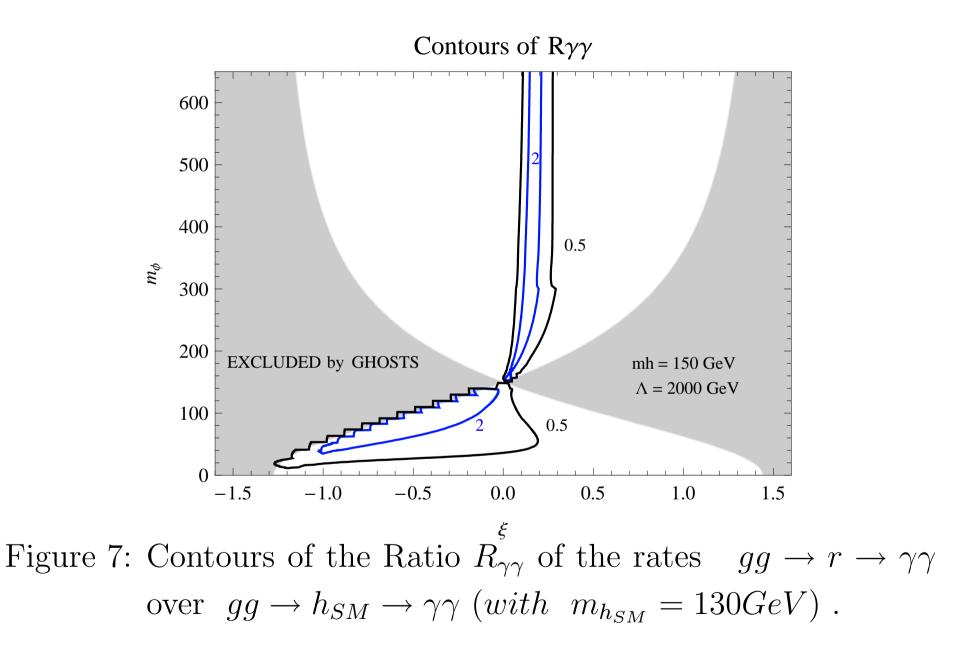


Figure 6: Ratio  $R_{\gamma\gamma}$  of production rates  $gg \to r \to \gamma\gamma$  over  $gg \to h_{SM} \to \gamma\gamma \quad (with \ m_{h_{SM}} = 130 \ GeV)$ .



### Conclusions

- Radion phenomenology is generically very similar to the Higgs search
- $\mathbf{RS1}$  radion is simple and well studied
- Higgs-radion mixing adds interesting properties to both the radion (fobic couplings) and the Higgs (suppressed/enhanced couplings)
- Bulk matter has interesting and unexpected effects in radion pheno, in conjuction with some radion-higgs mixing (preliminary-work in progress).